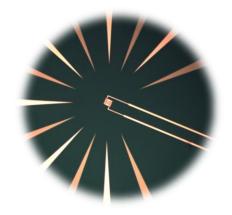
# Expanding the search for light dark matter with cryogenic sensors

A direct measurement with a TES arXiv:2506.18982

José Alejandro Rubiera Gimeno for the ALPS TES team





QS4Physics kickoff, 20.11.2025

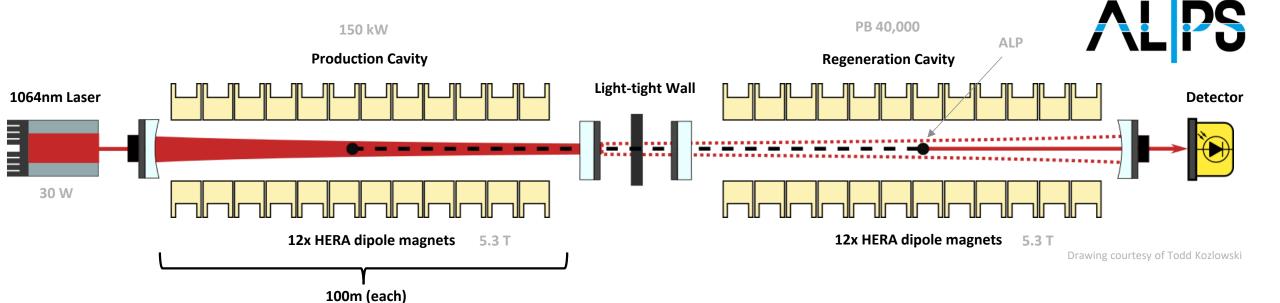






## Original challenge: The ALPS II experiment

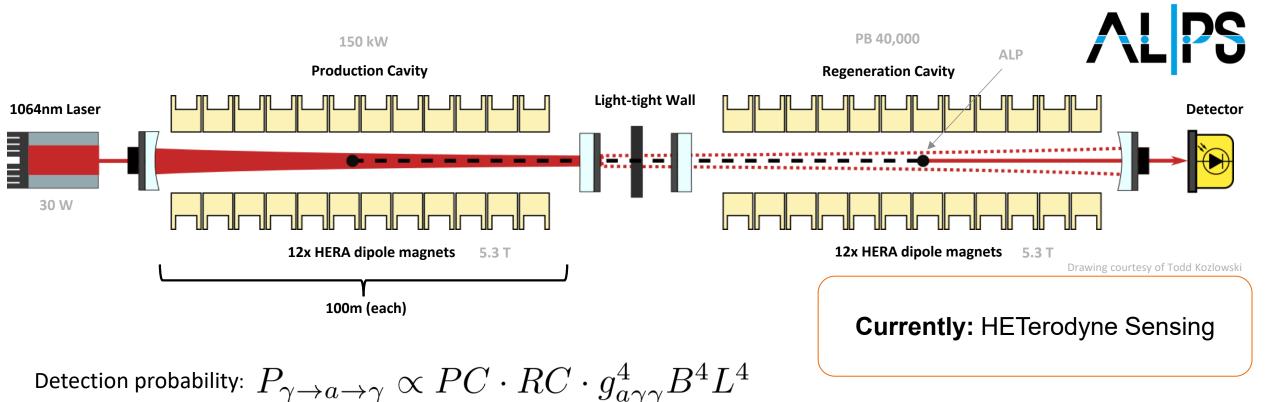
A light-shining-through-walls experiment



Detection probability: 
$$P_{\gamma \to a \to \gamma} \propto PC \cdot RC \cdot g_{a\gamma\gamma}^4 B^4 L^4$$

## Original challenge: The ALPS II experiment

A light-shining-through-walls experiment

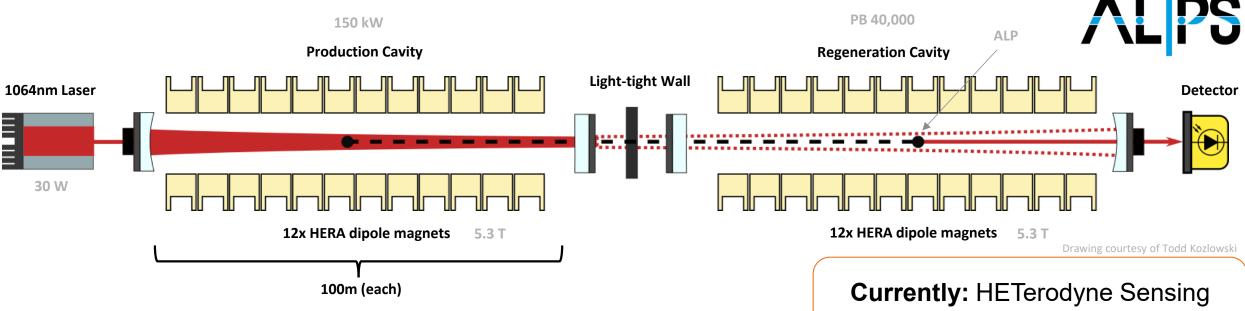


Expected rate of low energy ( $\sim 1.16\,\mathrm{eV}$ ) photons:

$$\dot{N}_{\gamma} \approx 2.8 \cdot 10^{-5} \frac{\gamma}{\mathrm{s}} \approx 1 \frac{\gamma}{\mathrm{day}}$$
 (for  $g_{a\gamma\gamma} = 2 \cdot 10^{-11} \, \mathrm{GeV}^{-1}$ )

## Original challenge: The ALPS II experiment

A light-shining-through-walls experiment



Detection probability: 
$$P_{\gamma \to a \to \gamma} \propto PC \cdot RC \cdot g_{a\gamma\gamma}^4 B^4 L^4$$

Expected rate of low energy ( $\sim 1.16\,\mathrm{eV}$ ) photons:

$$\dot{N}_{\gamma} \approx 2.8 \cdot 10^{-5} \frac{\gamma}{\mathrm{s}} \approx 1 \frac{\gamma}{\mathrm{day}}$$
 (for  $g_{a\gamma\gamma} = 2 \cdot 10^{-11} \, \mathrm{GeV}^{-1}$ )

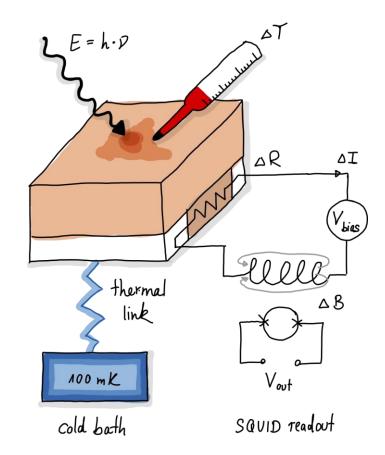
Future Option: **TES** 

#### Single-photon detection requirements for ALPS II:

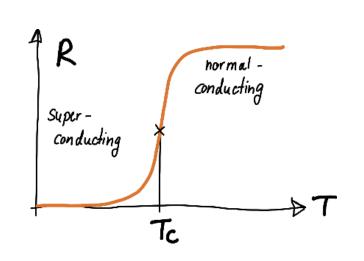
- Low energy photon detection
- Low background (< 1 photon/day)
- High detection efficiency

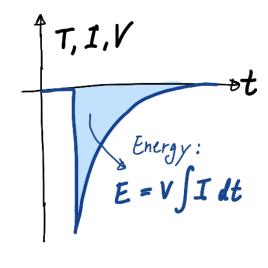
## **Reminder: Transition Edge Sensor**

#### **Working Principle**

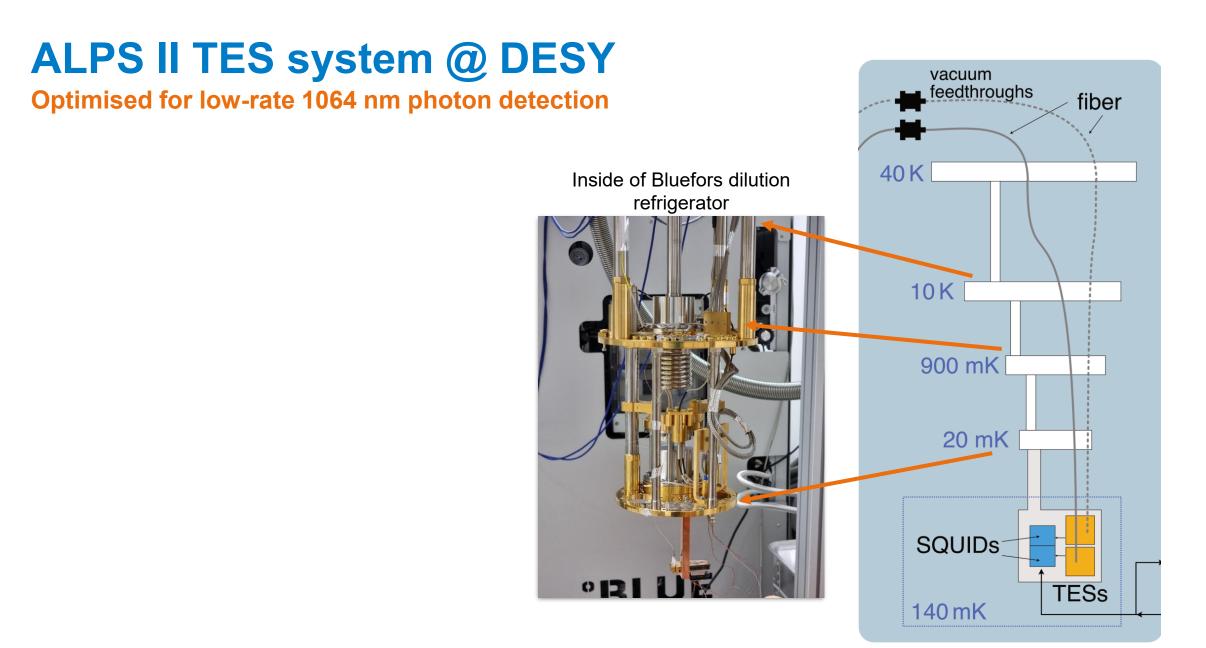


\*Courtesy of Katharina-Sophie Isleif





- Cryogenic microcalorimeter operated in superconducting transition region
- Connected to a colder thermal bath
- Working point controlled by a voltage bias circuit
- Change in resistance produced by energy deposition



Shah, R., Isleif, KS., Januschek, F. et al. Characterising a Single-Photon Detector for ALPS II. J Low Temp Phys (2022). https://doi.org/10.1007/s10909-022-02720-0

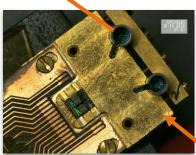
## **ALPS II TES system @ DESY**

**Optimised for low-rate 1064 nm photon detection** 

Area = 25  $\mu$ m x 25  $\mu$ m Thickness = 20 nm 0.2 ng of Tungsten



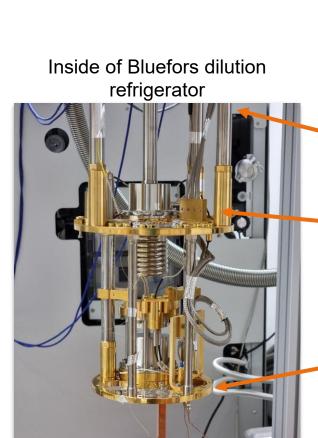
127 nm SiNx 2 nm a-Si 20 nm W  $2\,\mathrm{nm}$  a-Si  $132\,\mathrm{nm}\,\,\mathrm{SiNx}$ 80 nm Ag Si substrate

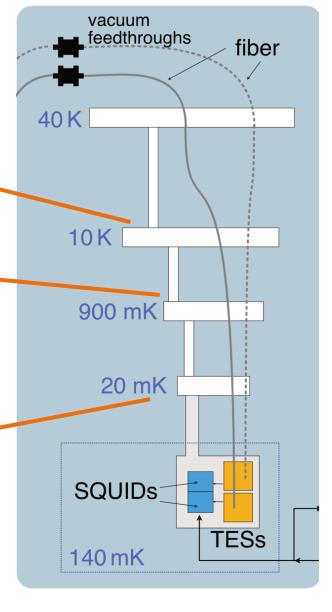


Sensors provided by: NIST

Packaging and SQUIDs provided by:







## **ALPS II TES system @ DESY**

Optimised for low-rate 1064 nm photon detection

## Single-photon detection requirements for ALPS II:

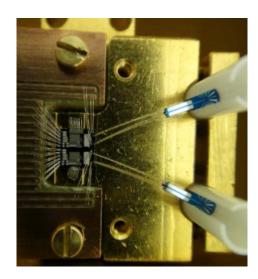
- Low energy photon detection
- Low background (< 1 photon/day)</li>
- High detection efficiency



#### Highly sensitive Transition Edge Sensor

- Low noise
- High efficiency:  $\epsilon > 80 \%$
- Good energy resolution: FWHM = 0.14 eV
- Ultra-low background rate at 1064nm:  $b = 6.9 \cdot 10^{-6} cps$

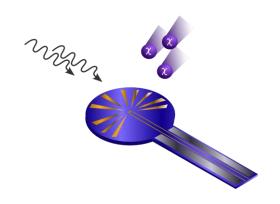


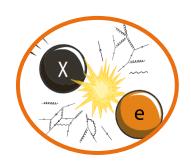




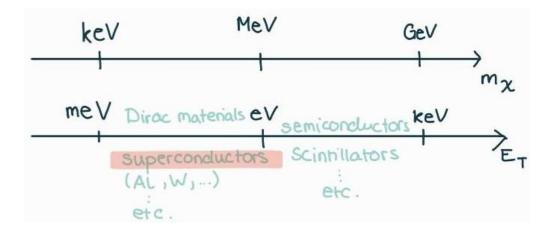
Shah, R., Isleif, KS., Januschek, F. et al. Characterising a Single-Photon Detector for ALPS II. J Low Temp Phys (2022). <a href="https://doi.org/10.1007/s10909-022-02720-0">https://doi.org/10.1007/s10909-022-02720-0</a>
J. A. Rubiera Gimeno et al., "A TES system for ALPS II - Status and Prospects", PoS EPS-HEP2023 (2023) 567, doi:10.22323/1.449.0567

### Can we use this for a dark matter search?





$$E_{\rm T_{\rm max}} = E_{\rm kin} \sim m_{\chi} v^2 \sim 10^{-6} m_{\chi}$$

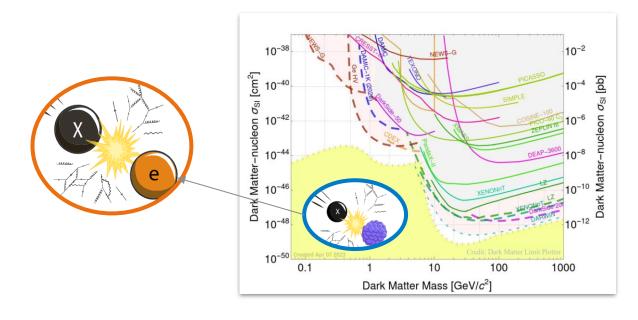


## TES can enable low-threshold energy-resolved light DM searches

**TES for DM searches** 

#### Why TES for Dark Matter?

- Direct DM detection via electron scattering opens access to sub-MeV DM
- TESs offer sub-eV thresholds
- Prior work with Superconducting Nanowire Single Photon Detector (SNSPD) able to set new limits with 4 dark counts in 180 h with 0.73 eV energy threshold Hochberg, Y. et al. arXiv:2110.01586 (2021)
- lacked energy resolution → TES can provide spectral information



PRL 116, 011301 (2016)

PHYSICAL REVIEW LETTERS

week ending 8 JANUARY 201

3

#### Superconducting Detectors for Superlight Dark Matter

Yonit Hochberg, <sup>1</sup> Yue Zhao, <sup>2</sup> and Kathryn M. Zurek <sup>1</sup>

<sup>1</sup>Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA and Berkeley Center for Theoretical Physics, University of California, Berkeley, California 94720, USA

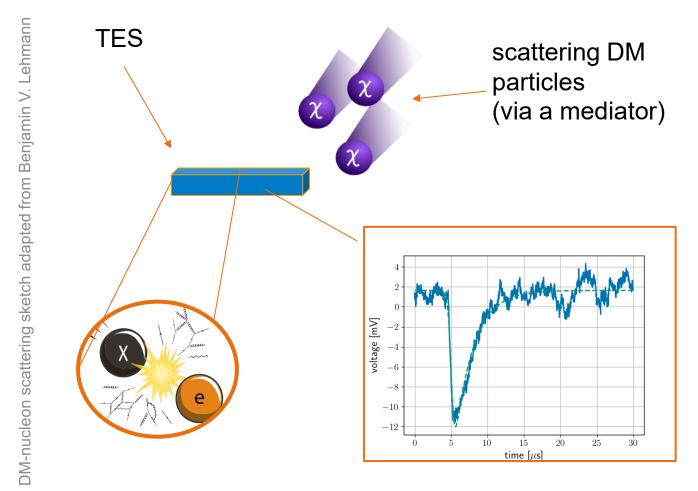
<sup>2</sup>Department of Physics, Stanford Institute for Theoretical Physics, Stanford University, Stanford, California 94305, USA (Received 8 June 2015; revised manuscript received 21 October 2015; published 7 January 2016)

We propose and study a new class of superconducting detectors that are sensitive to  $\mathcal{O}(\text{meV})$  electron recoils from dark matter–electron scattering. Such devices could detect dark matter as light as the warm dark-matter limit,  $m_X \gtrsim 1$  keV. We compute the rate of dark-matter scattering off of free electrons in a (superconducting) metal, including the relevant Pauli blocking factors. We demonstrate that classes of dark matter consistent with terrestrial and cosmological or astrophysical constraints could be detected by such detectors with a moderate size exposure.

DOI: 10.1103/PhysRevLett.116.011301

## **Proof-of-principle: TES for Direct Dark Matter Searches**

TES as simultaneous target & sensor



- First attempt to use TES as target & sensor for DM
- Direct DM detection via electron-scattering
- Should yield similar energy-dependent response as photon-interactions
- Enables small-scale high-sensitivity DM search
- Recently proposed as well in

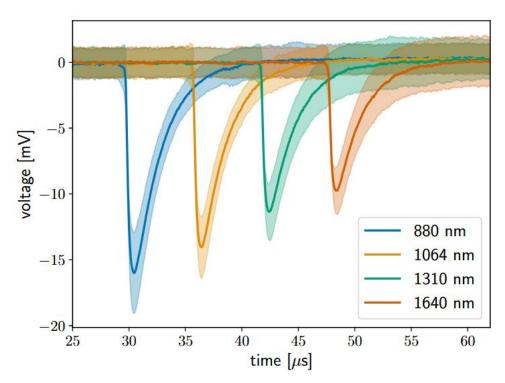
Muping Chen et al.: "Light Dark Matter Detection with Sub-eV Transition-Edge Sensors", <a href="mailto:arXiv:2506.10070">arXiv:2506.10070</a>

#### **Challenges:**

- low mass (0.2 ng)
- small area (25 x 25 μm)
- limited knowledge about broadband response

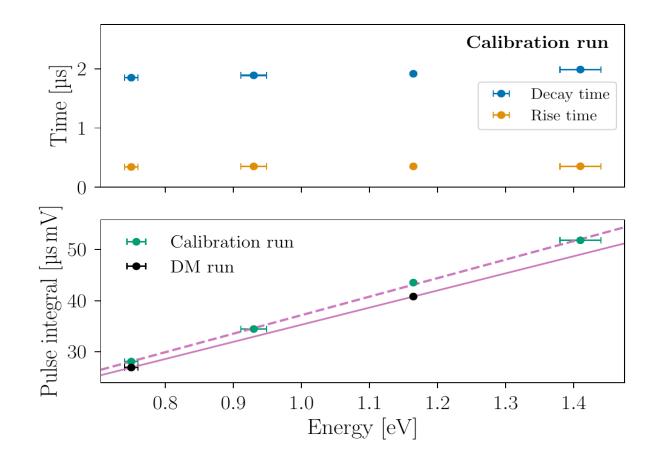
## Can we use our TES for other energies?

#### Going to lower energies



#### Confirmed assumptions:

- constant rise and decay time response
- linear pulse integral/height response for different energies.

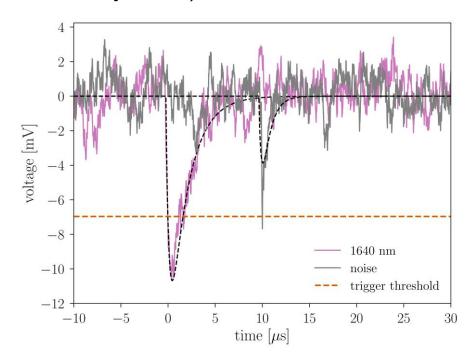


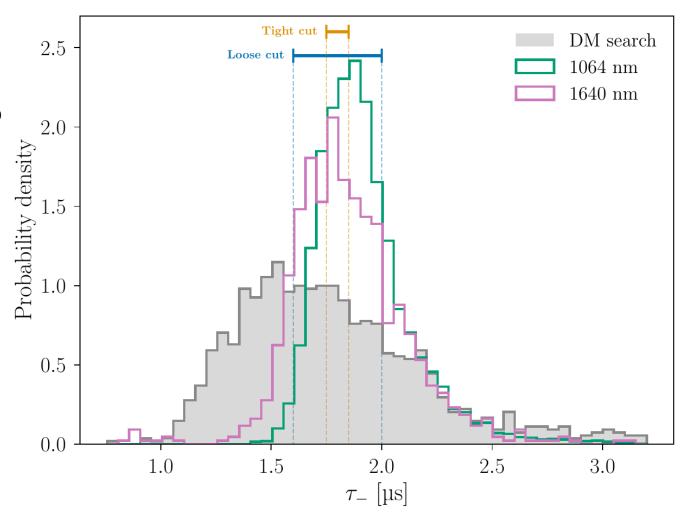
→ isolating photon-like signals based on calibration and simulation results of broad-band TES response is possible

#### **Event selection**

#### **Evaluating a 489 h dedicated DM search run**

- 489 h DM search measurement
- Cuts based on calibration samples
- Looking for photon-like shapes (in particular also rise, decay times)

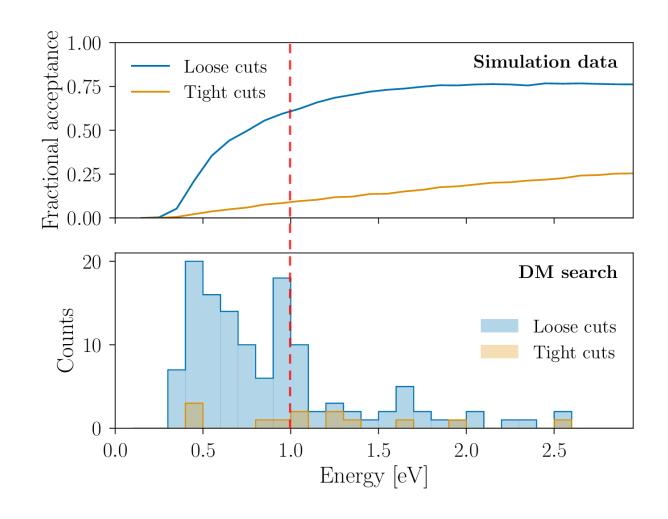




## **Observed event spectrum**

#### Remaining events after two different cut schemes

- Two different cut schemes with different acceptance (based on calibration & simulations) for different energies
- Sensitivity also below 1 eV
- 126 events (loose) and 13 events (tight) remaining after selection
- Resulting spectrum compared with expected DM model
- Used to set exclusion limits on DM-electron scattering cross section with 95% C.L. for ~MeV light DM



## TES sensitivity for DM-electron scattering

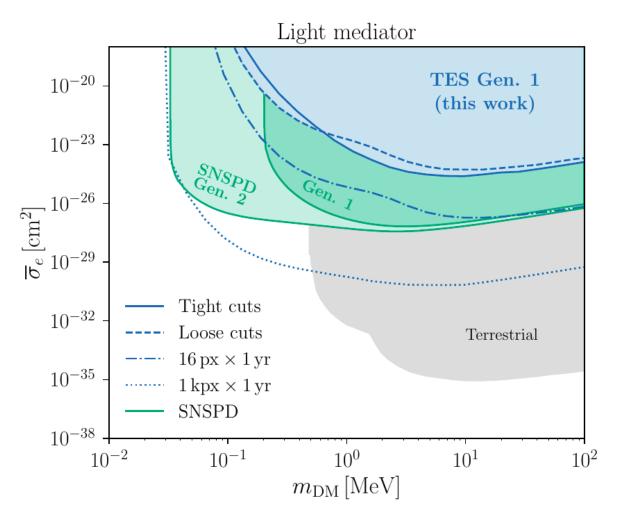
**Compared to similar searches** 

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow

- Results of the 0.2 ng dark matter search: exclusion limits on DM-electron scattering cross section with 95% C.L. for ~MeV light DM
- Interesting in particular for the low energy region
- Here: Compared to 1st and 2nd generation SNSPD DM search

Important note: Overburden effects need to be considered above approximately the order of 10<sup>-24</sup> cm<sup>2</sup> depending on the model and might relax the constraints.

→ Proof-of-concept



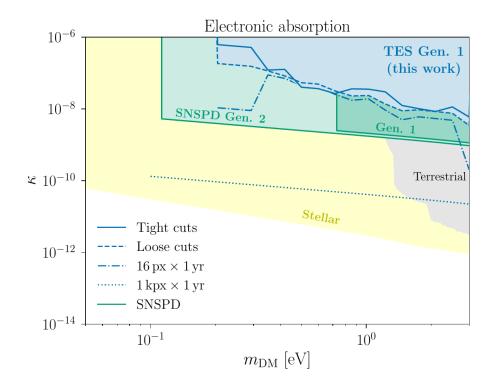
### **Additional Dark Matter interactions**

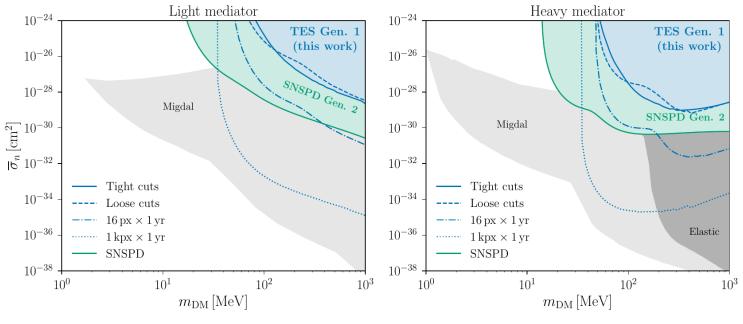
#### **Electronic absorption and DM-nucleon scattering**

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow

- Results can be used to set constraints on dark photon absorption
- and DM-nucleon scattering as well

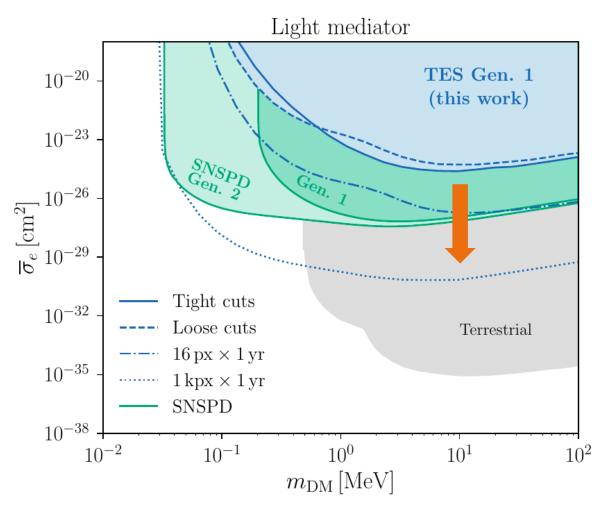
S.M.Griffin et al., "Dark Matter-Electron Detectors for Dark Matter-Nucleon Interactions", arXiv:2412.16283v1





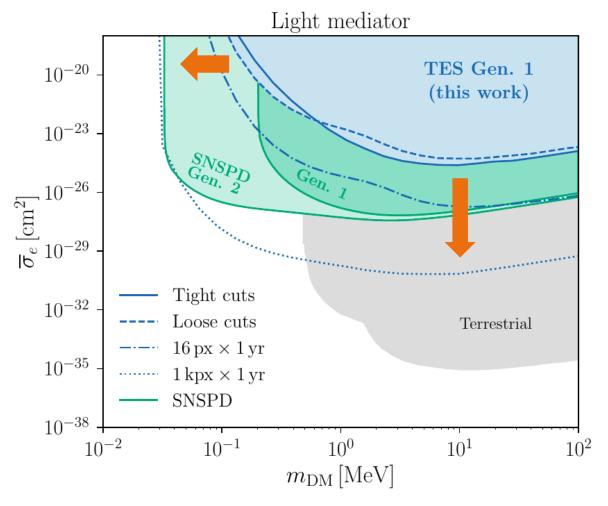
- Scaling up the system (going for lower cross sections)
- Complex TES readout for larger systems
- New possibilities with Kinetic Inductance Current Sensors

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow



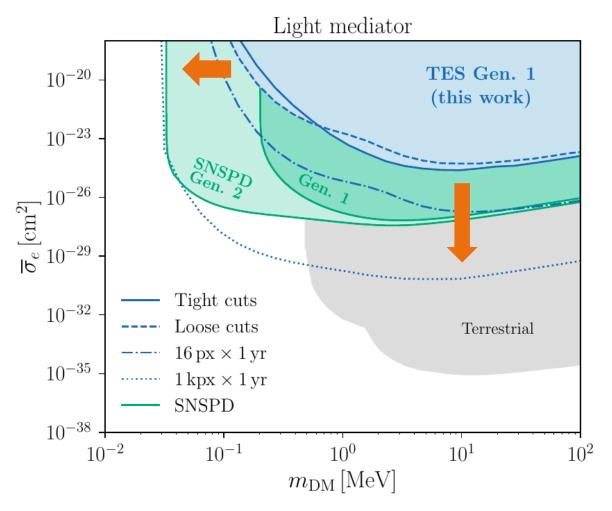
- Scaling up the system (going for lower cross sections)
- Complex TES readout for larger systems
- New possibilities with Kinetic Inductance Current Sensors
- Reduction of TES system noise (going for lower masses):
  - Improves energy resolution
  - Allows lower energy threshold
- Usually lower Tc or smaller sensor

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow



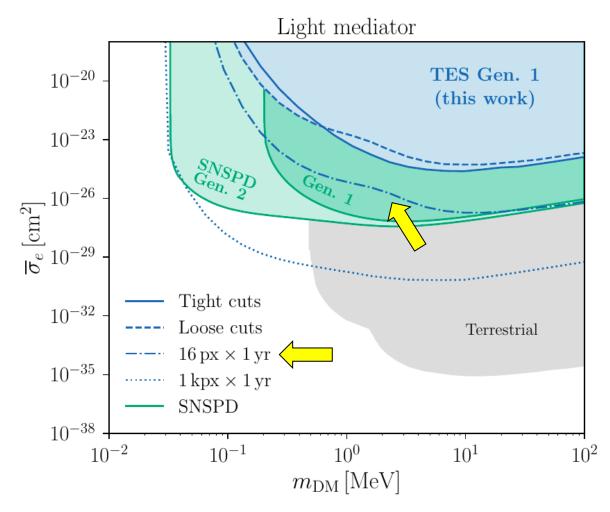
- Scaling up the system (going for lower cross sections)
- Complex TES readout for larger systems
- New possibilities with Kinetic Inductance Current Sensors
- Reduction of TES system noise (going for lower masses):
  - Improves energy resolution
  - Allows lower energy threshold
- Usually lower Tc or smaller sensor
- Reduction of photon-like background

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow



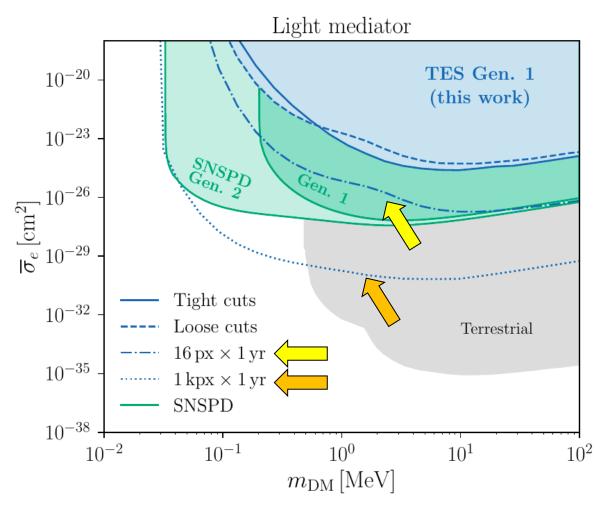
- Scaling up the system (going for lower cross sections)
- Complex TES readout for larger systems
- New possibilities with Kinetic Inductance Current Sensors
- Reduction of TES system noise (going for lower masses):
  - Improves energy resolution
  - Allows lower energy threshold
- Usually lower Tc or smaller sensor
- Reduction of photon-like background

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow



- Scaling up the system (going for lower cross sections)
- Complex TES readout for larger systems
- New possibilities with Kinetic Inductance Current Sensors
- Reduction of TES system noise (going for lower masses):
  - Improves energy resolution
  - Allows lower energy threshold
- Usually lower Tc or smaller sensor
- Reduction of photon-like background

C. Schwemmbauer et al., "First direct search for light dark matter interactions in a transition-edge sensor", <u>arXiv:2506.18982</u> → journal submission to follow

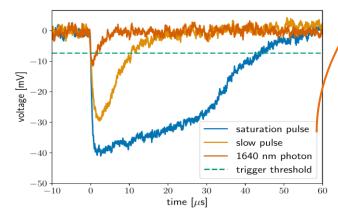


SNSPD Gen. 2 results based on: L. Baudis et al., "A New Bite Into Dark Matter with the SNSPD-Based QROCODILE Experiment", <u>arXiv:2412.16279</u>

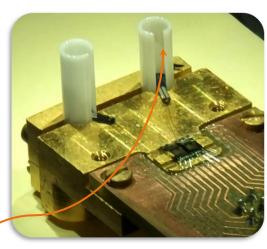
## **Optimised Sensors**

#### **Reducing background further**

- The sensor of this study was optimised for 1064 nm photon detection
  - → other sensors?
- Background reduction?
  - Most (non baseline) triggers from interactions in Si substrate



#### ALPS TES module

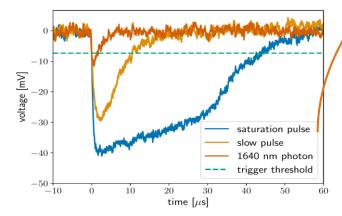


Background simulations with GEANT4+COMSOL confirmed dominant contribution of radioactive decays from Zirconia fiber sleeves

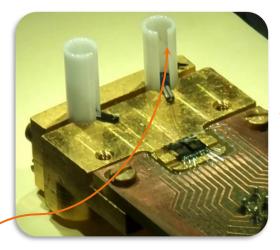
## **Optimised Sensors**

#### **Reducing background further**

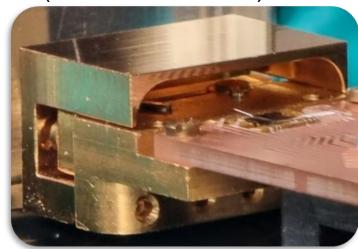
- The sensor of this study was optimised for 1064 nm photon detection
  - → other sensors?
- Background reduction?
  - Most (non baseline) triggers from interactions in Si substrate



#### ALPS TES module



no fiber sleeves (not needed for DM)

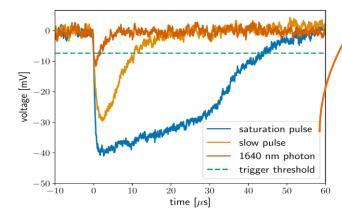


Background simulations with GEANT4+COMSOL confirmed dominant contribution of radioactive decays from Zirconia fiber sleeves

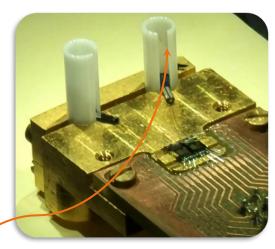
## **Optimised Sensors**

#### **Reducing background further**

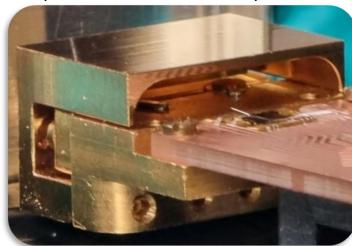
- The sensor of this study was optimised for 1064 nm photon detection
  - → other sensors?
- Background reduction?
  - Most (non baseline) triggers from interactions in Si substrate



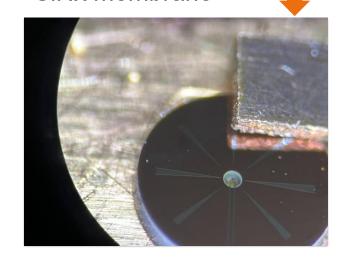
#### ALPS TES module



no fiber sleeves (not needed for DM)



TES chip placed on SiNx membrane



Background simulations with GEANT4+COMSOL confirmed dominant contribution of radioactive decays from Zirconia fiber sleeves

## **Status of the project**

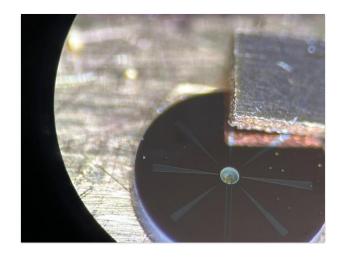
- Currently testing new membrane TES optimized for low background
- Preliminary results look promising:
  - Better energy resolution

Previous sensor: FWHM = 0.14 eV

Membrane sensor: FWHM = 0.11 eV



TES chip placed on SiNx membrane



## **Status of the project**

- Currently testing new membrane TES optimized for low background
- Preliminary results look promising:
  - Better energy resolution

Previous sensor: FWHM = 0.14 eV

Membrane sensor: FWHM = 0.11 eV

Lower raw background rate

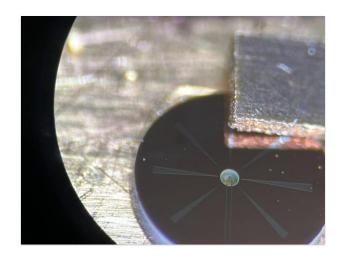
Rate at a pulse equivalent to 1064nm photons

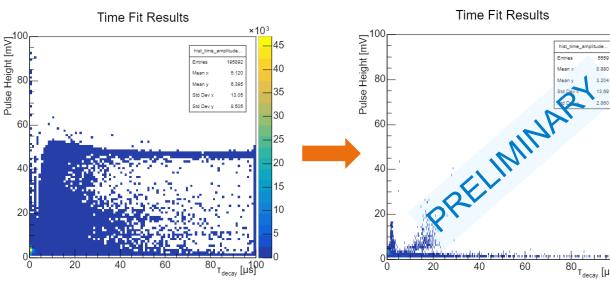
Previous sensor:  $b = 1.04 \cdot 10^{-2} cps$ 

Membrane sensor:  $b = 1.30 \cdot 10^{-4} cps$ 



## TES chip placed on SiNx membrane

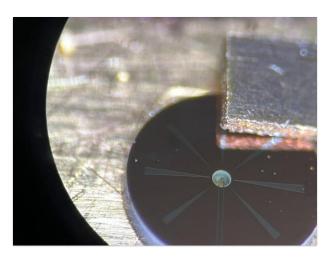




## **Next steps**

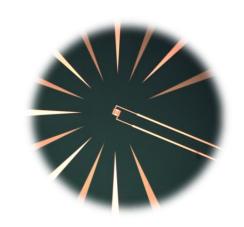
- Dedicated DM run with new membrane TES optimized for low background
- Continue development on more robust and precise background simulations
- Feasibility study of larger system together with other QS4Physics partners, in particular KIT

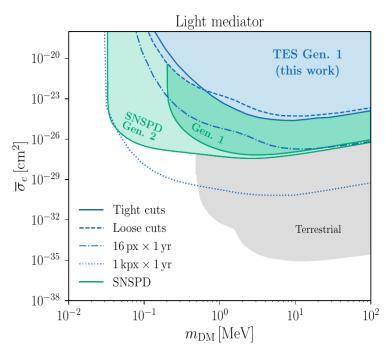
TES chip placed on SiNx membrane



## **Summary**

- First direct search for light dark matter interactions in a TES (use of the TES as target and sensor simultaneously)
  - → Successful proof-of-concept study
  - Here: 0.2 ng detector in a system originally for the ALPS II experiment was used → Small setup, not optimized for a DDM search
- Improving the limits:
  - Even better understanding/reduction of background could increase the sensitivity considerably
  - DM run with membrane TES sensor
  - Larger systems with optimized sensors
- Within the QS4Physics project we aim to evaluate the feasibility and efficacy of a larger system



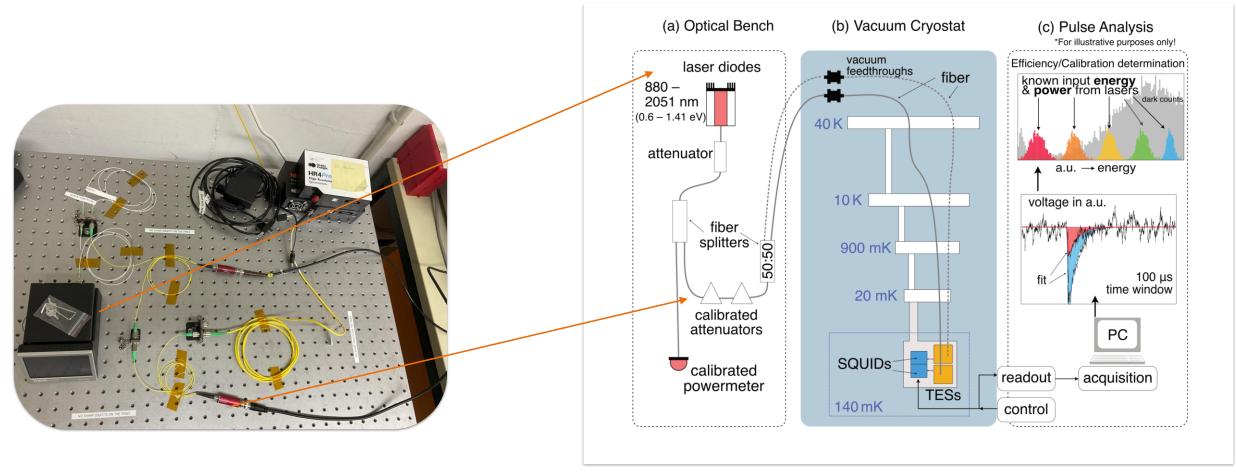


arXiv:2506.18982

## Backup

## Can we use our TES for other energies?

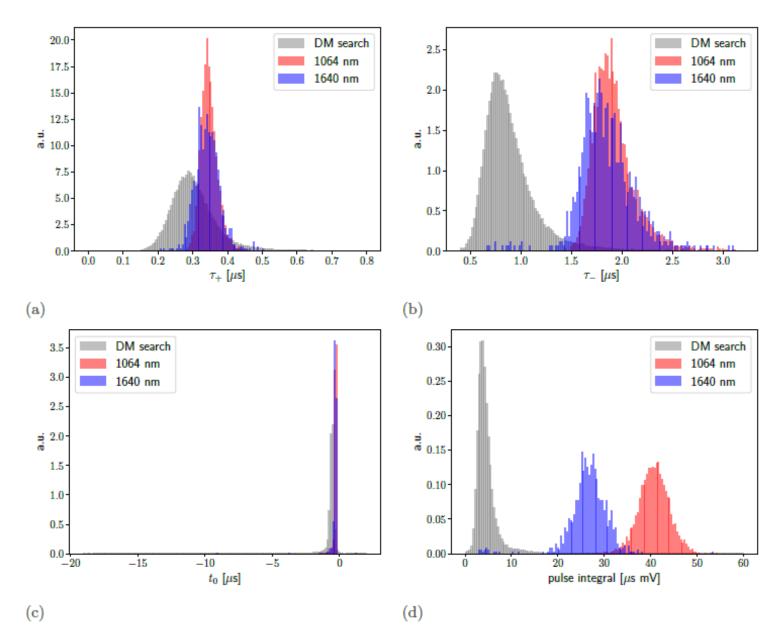
**Calibration Setup** 



#### **Event selection: Cuts**

#### Loose and tight cuts

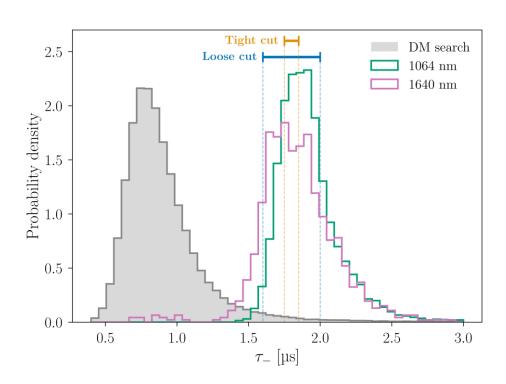
	Parameter	Loose	Tight
Cuts	$\tau_+$ [µs]	(0.3, 0.4)	(0.325, 0.375)
	$ au$ [ $\mu \mathrm{s}$ ]	(1.6, 2.0)	(1.75, 1.85)
	$ au [\mu  ext{s}] \ \chi^2_{ ext{red}}$	(0.9, 1.1)	(0.95, 1.05)
	$t_0$ [ $\mu s$ ]	(-0.5, 0.5)	(-0.5, 0.5)
	$\mathcal{R}_{ ext{dev}}$	(0.73, 1.10)	(0.73, 1.10)
Results	Event rate [Hz] Survival [%]	$7.2 \times 10^{-5}$ $0.107$	$7.4 \times 10^{-6}$ $0.011$

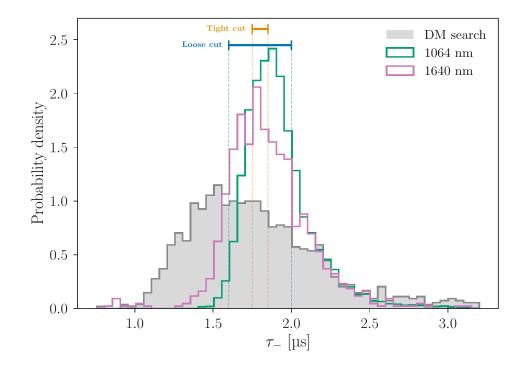


## **Event selection: Cuts**

#### **Loose and tight cuts**

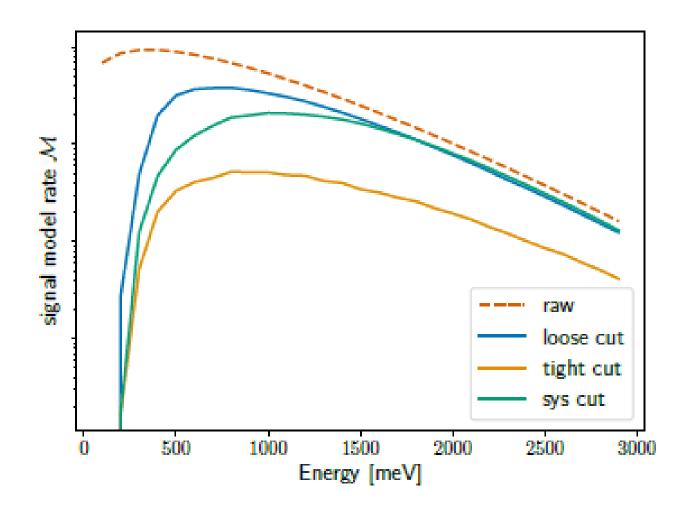
	Parameter	Loose	$\mathbf{Tight}$
Cuts	$\tau_+$ [µs]	(0.3, 0.4)	(0.325, 0.375)
	$ au$ [ $\mu \mathrm{s}$ ]	(1.6, 2.0)	(1.75, 1.85)
	$ au^{}[\mathrm{\mu s}] \ \chi^2_{\mathrm{red}}$	(0.9, 1.1)	(0.95, 1.05)
	$t_0 \; [\mu \mathrm{s}]$	(-0.5, 0.5)	(-0.5, 0.5)
	$\mathcal{R}_{ ext{dev}}$	(0.73, 1.10)	(0.73, 1.10)
Results	Event rate [Hz] Survival [%]	$7.2 \times 10^{-5}$ $0.107$	$7.4 \times 10^{-6}$ $0.011$





## **Predictions**

#### **Dark matter model dependent**



## **Energy resolution**

#### PhD Thesis José Alejandro Rubiera Gimeno (DESY, now @ HSU)

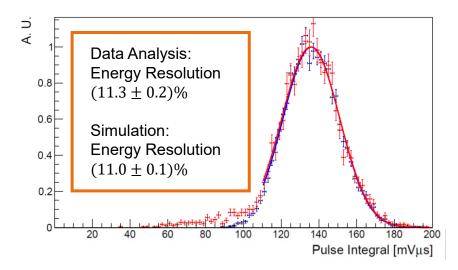
 Energy resolution of the TES can be understood and reproduced from the noise

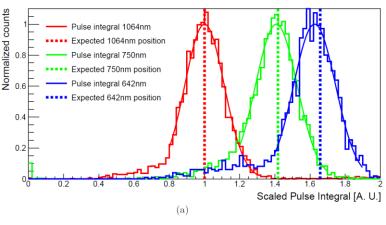
J. A. Rubiera Gimeno et al., "The TES detector of the ALPS II experiment", Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment 1046 (January 2023) 167588, doi:10.1016/j.nima.2022.167588

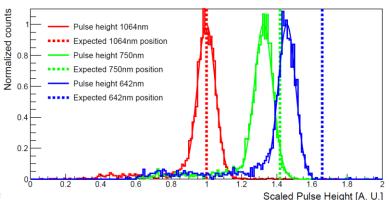
- With improved analysis, we can reach an energy resolution of about 5%
  - Faster fitting in the frequency domain
  - Access to physical properties of the sensor
  - Improvement in energy resolution by a factor of 2

J. A. Rubiera Gimeno et al., "A TES system for ALPS II - Status and Prospects", PoS EPS-HEP2023 (2023) 567, doi:10.22323/1.449.0567

Pulse height, FFT (1064 nm) 
$$\frac{\sigma}{\mu}$$
 100% = (5.31 ± 0.06)%







#### Intrinsic dark counts

PhD Thesis Rikhav Shah (Mainz, now UHH)

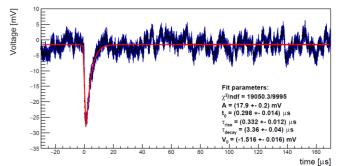


With data analysis rate of events in the order of 10<sup>-2</sup> Hz could be reduced to a dark count rate of 6.9 · 10<sup>-6</sup> Hz over 20 days for 1064 nm photons with an acceptance greater than 90%

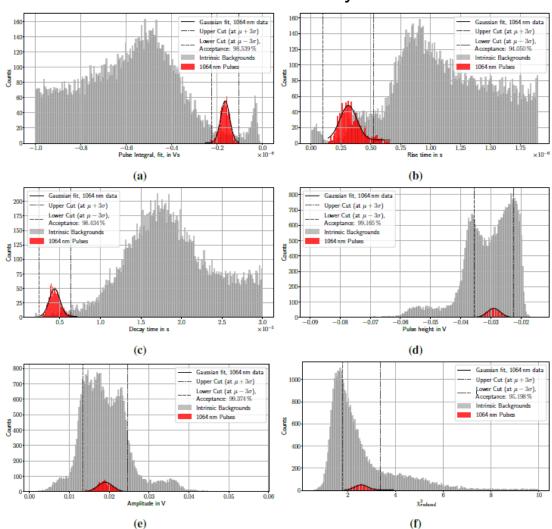
Shah, R., Isleif, KS., Januschek, F. *et al.* Characterising a Single-Photon Detector for ALPS II. *J Low Temp Phys* (2022). https://doi.org/10.1007/s10909-022-02720-0

 Currently investigating improvements with Machine Learning Algorithms

Meyer, M., Isleif, KS, Januschek, F et al. A First Application of Machine and Deep Learning for Background Rejection in the ALPS II TES Detector https://doi.org/10.1002/andp.202200545



#### Cut-based-analysis



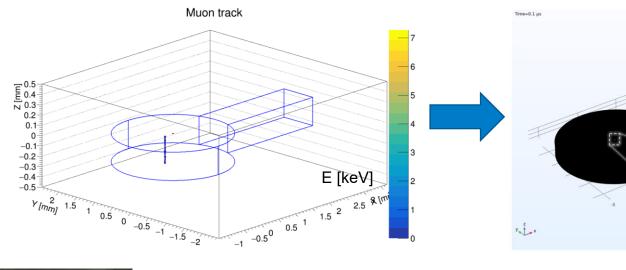
## Simulations of noise, background and signals

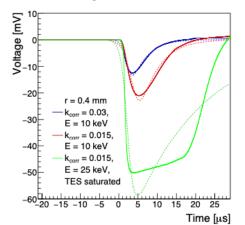
PhD Thesis José Alejandro Rubiera Gimeno (DESY, now HSU)

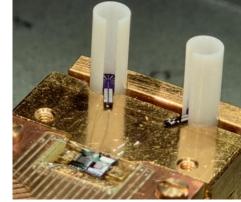
- Simulation of noise for understanding and improving energy resolution as well as reduction of noise counts
- Simulation of TES signal→ understanding of the system, improving selection criteria
- Simulation of background → background reduction

Simulation of energy deposition of radiation (muon, gamma) in Geant4

Simulation of TES physics and transport of heat in silicon substrate using COMSOL Multiphysics



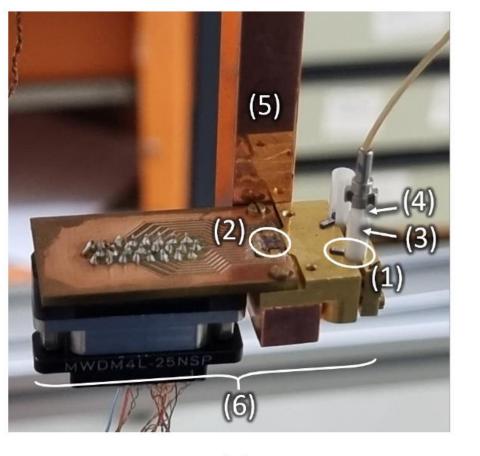


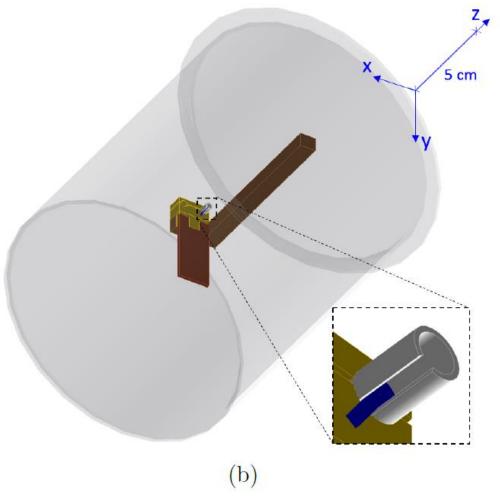


Radioactivity of the Zirconia from the fiber sleeve can explain the majority of the remaining "intrinsic" background!

Black Body Radiation simulation framework for simulating spectrum detected by the TES with different suppression methods

## **Geant4 simulation geometry**





(a)

## Simulation of intrinsic background

Zirconia as main background contributor

- Shape describe different types of events
- Rate of events passing the threshold are compatible with data

Counts Pulse Height [mV] 300 Simulation 250 Saturation pulses 200 50 150 40 100 30 Slow pulses 50 Below trigger threshold 10

Rates from Geant4 + COMSOL simulation

Zirconia: rate  $\in [0.33 ; 2.1] \cdot 10^{-2} \text{ cps}$ 

Muon: rate =  $8 \cdot 10^{-5}$  cps

Rate from data:  $2.1 \cdot 10^{-2}$  cps

