

# 1S-3S CW spectroscopy on deuterium atoms

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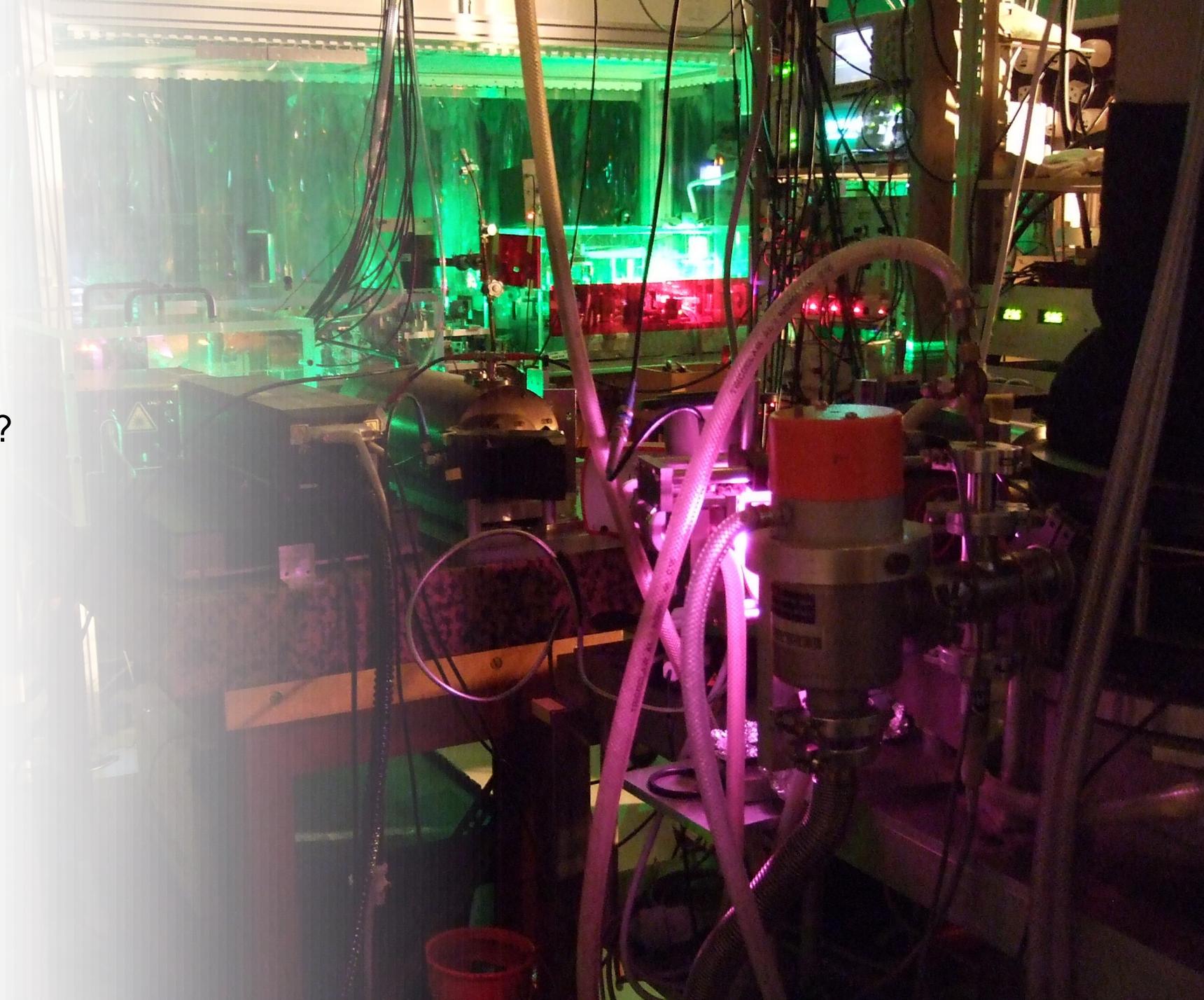


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PREN - June 2023

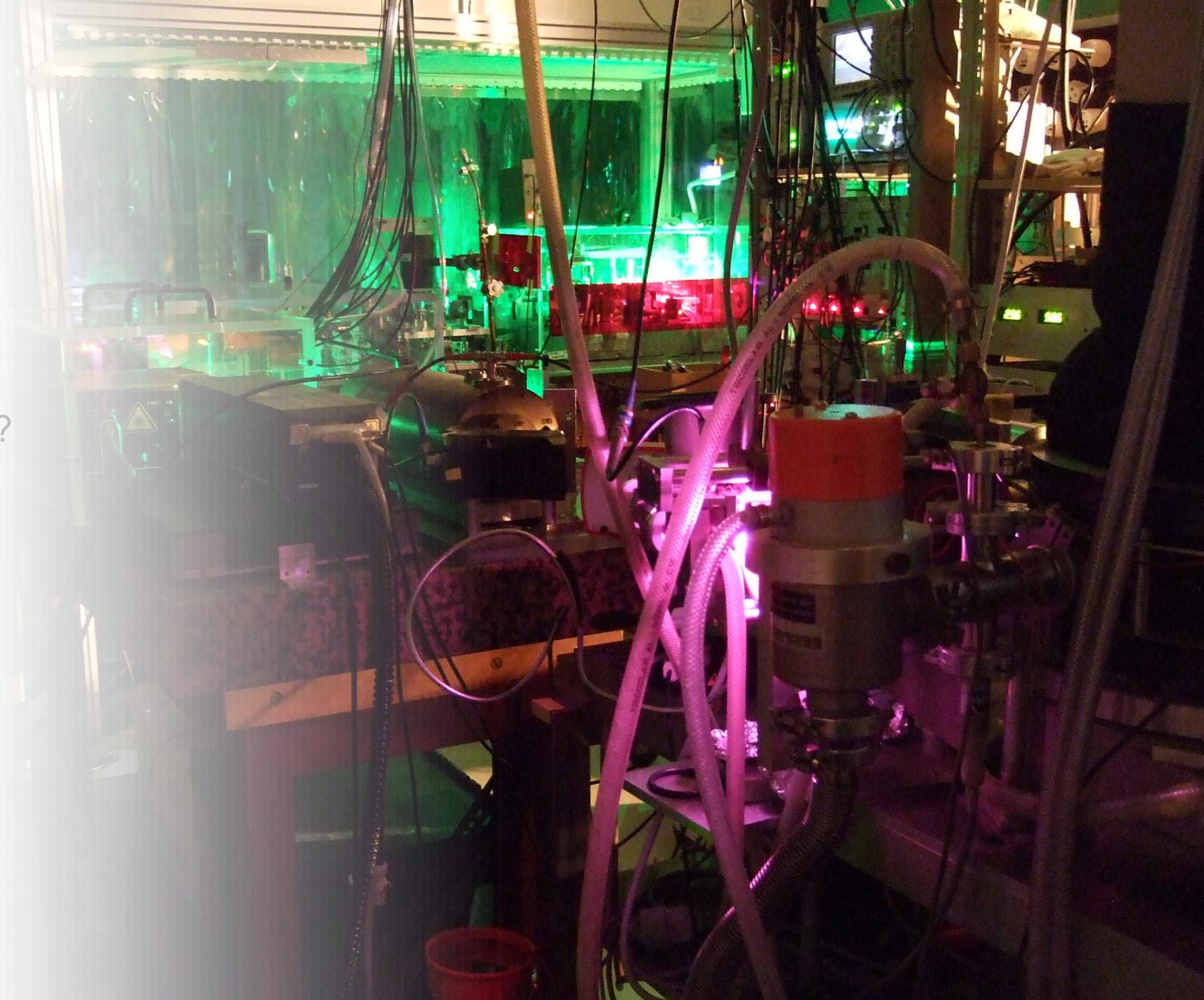
# Overview

- The experiment
- Dealing with systematics
- A new systematics effect ?
- Preliminary result ?

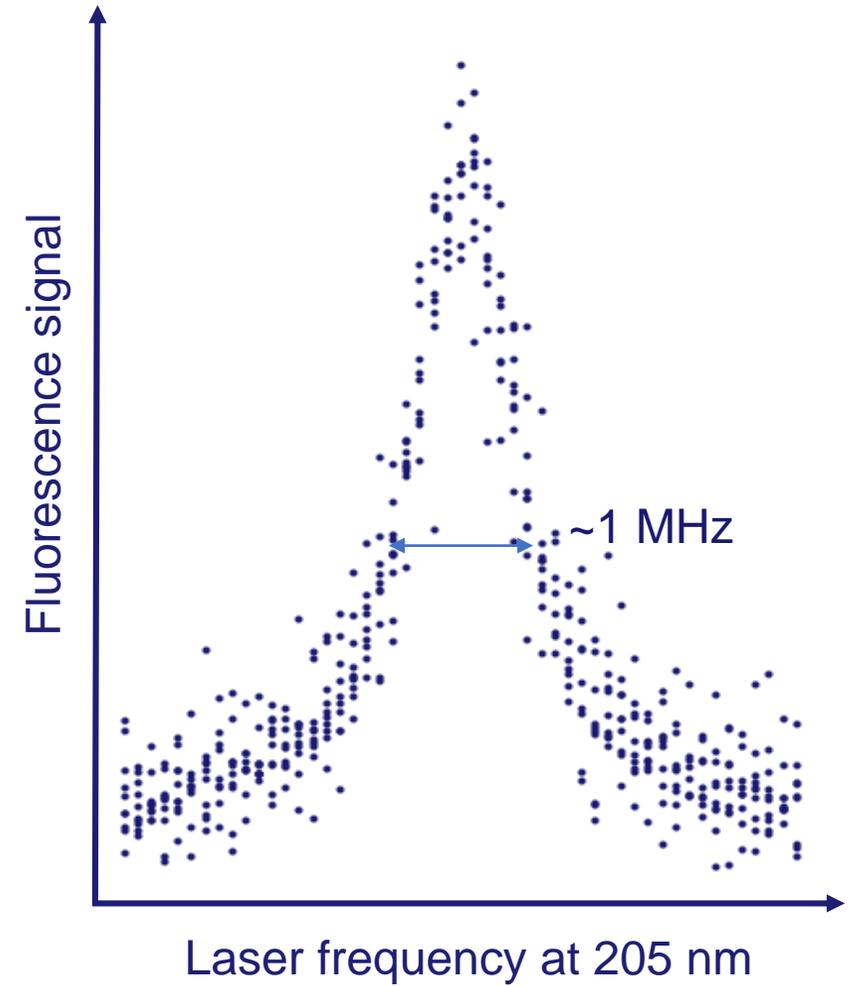
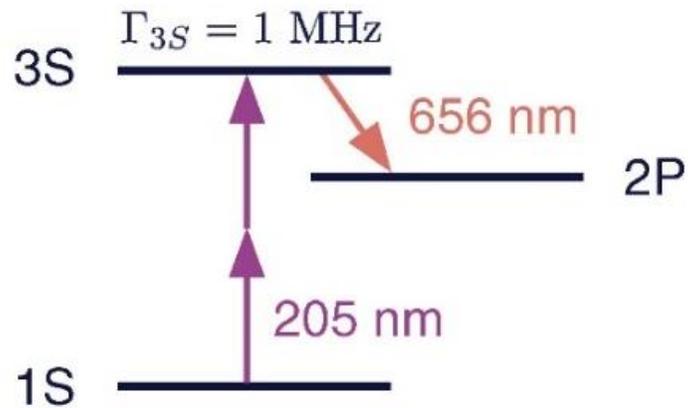
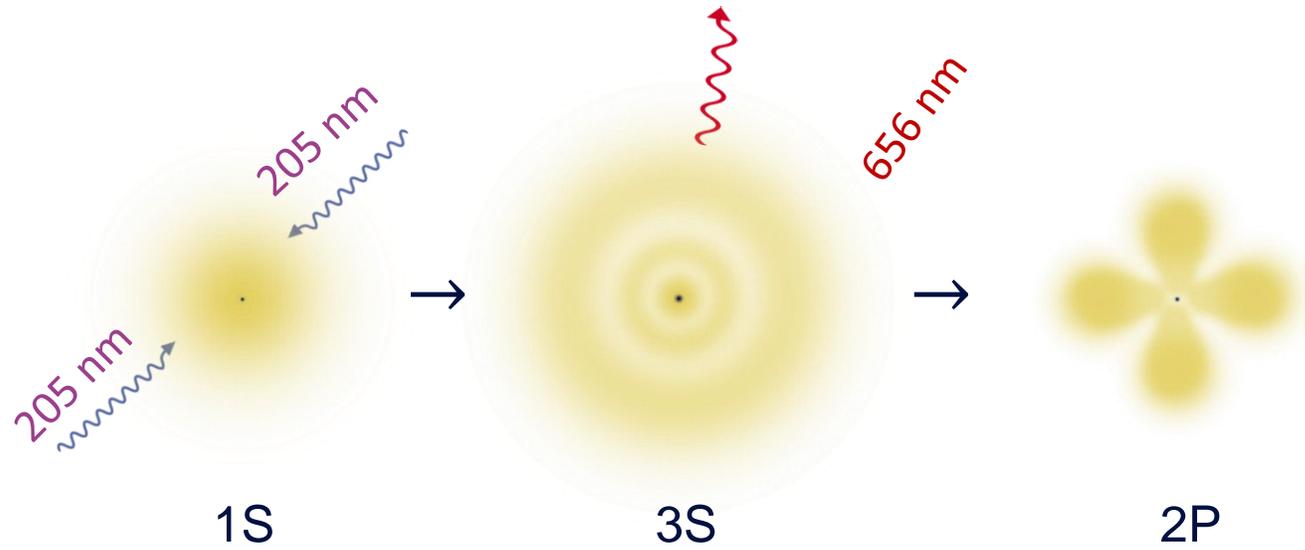


# Overview

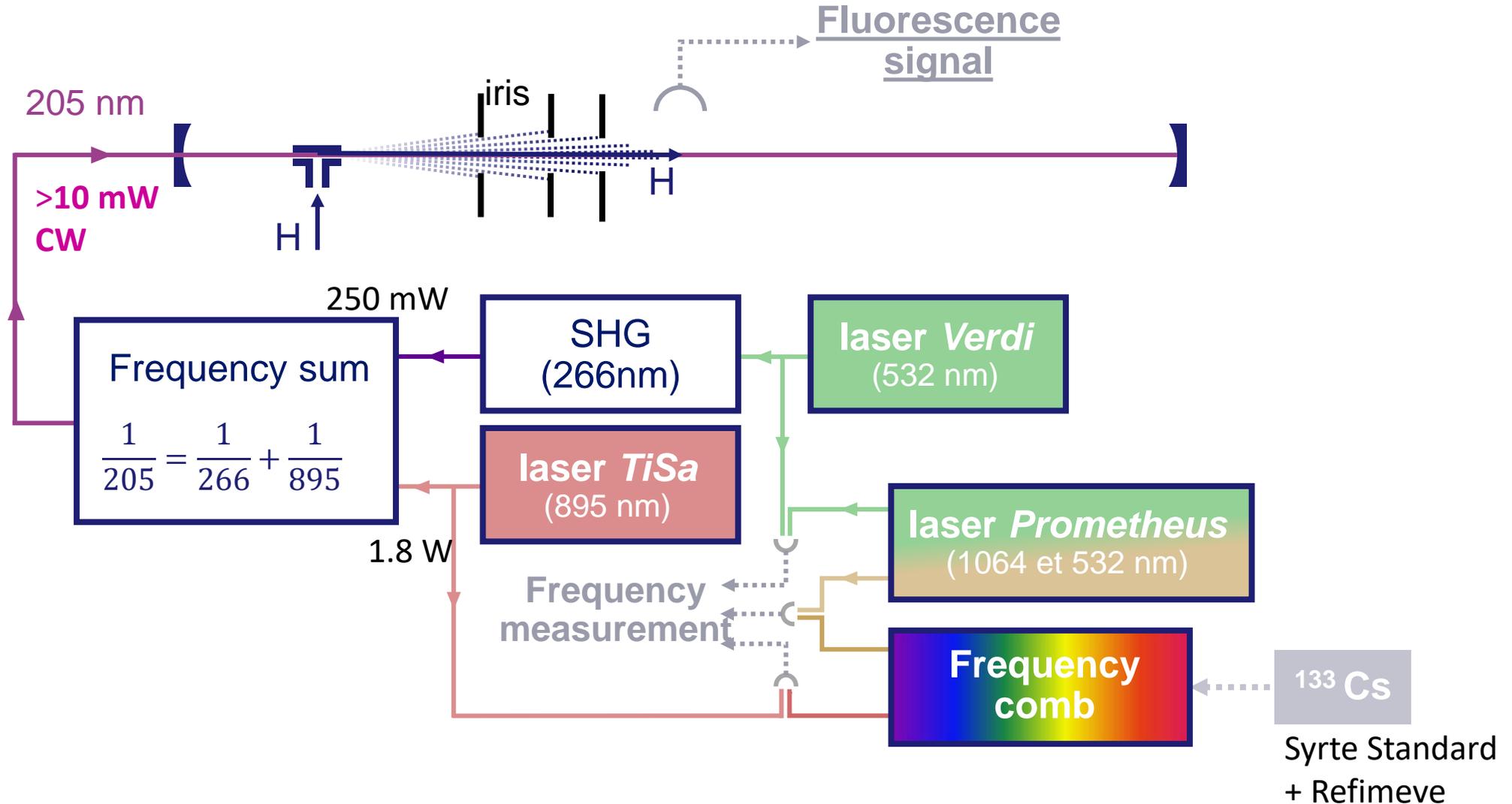
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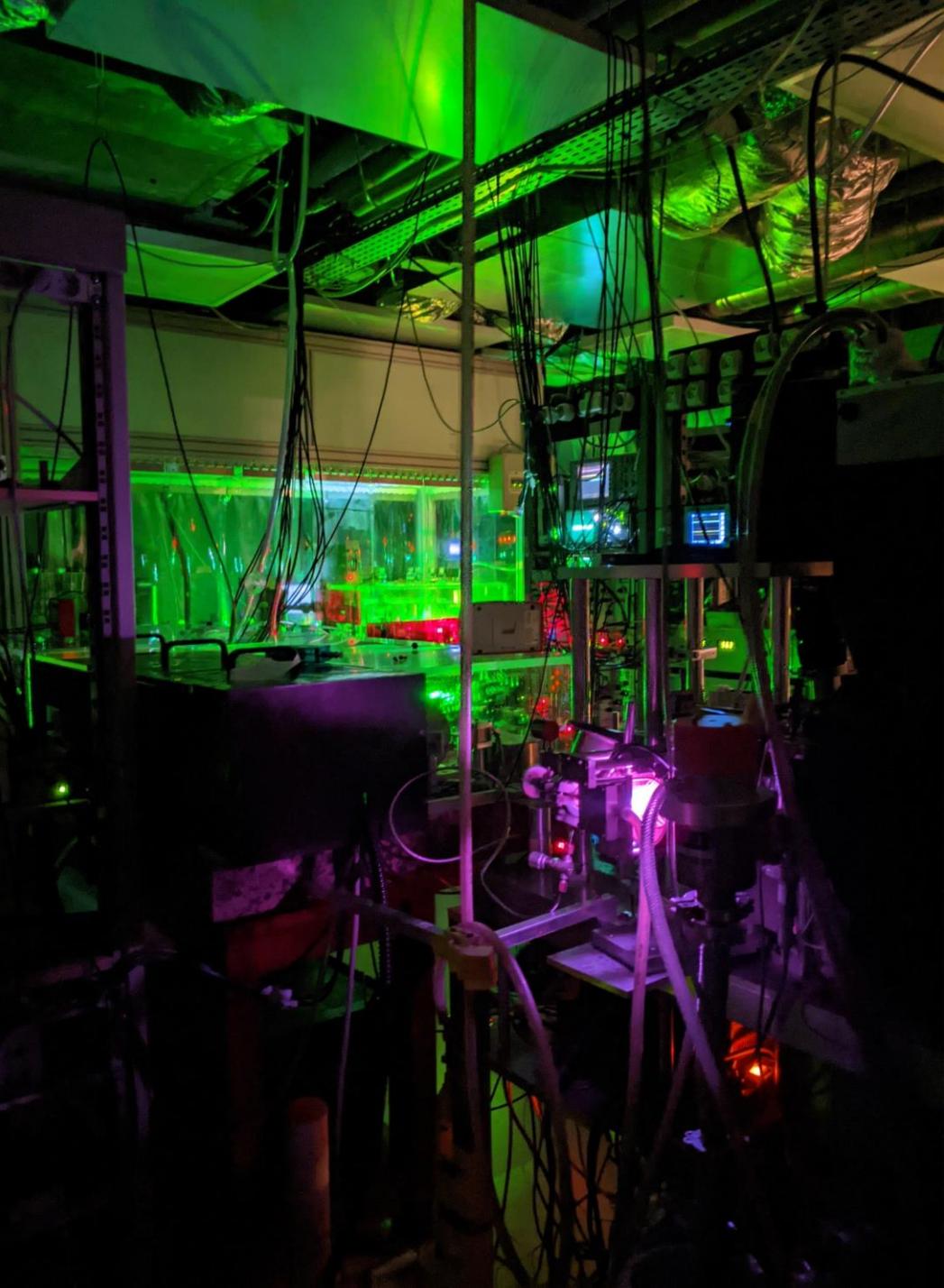


# 1S–3S Hydrogen CW spectroscopy

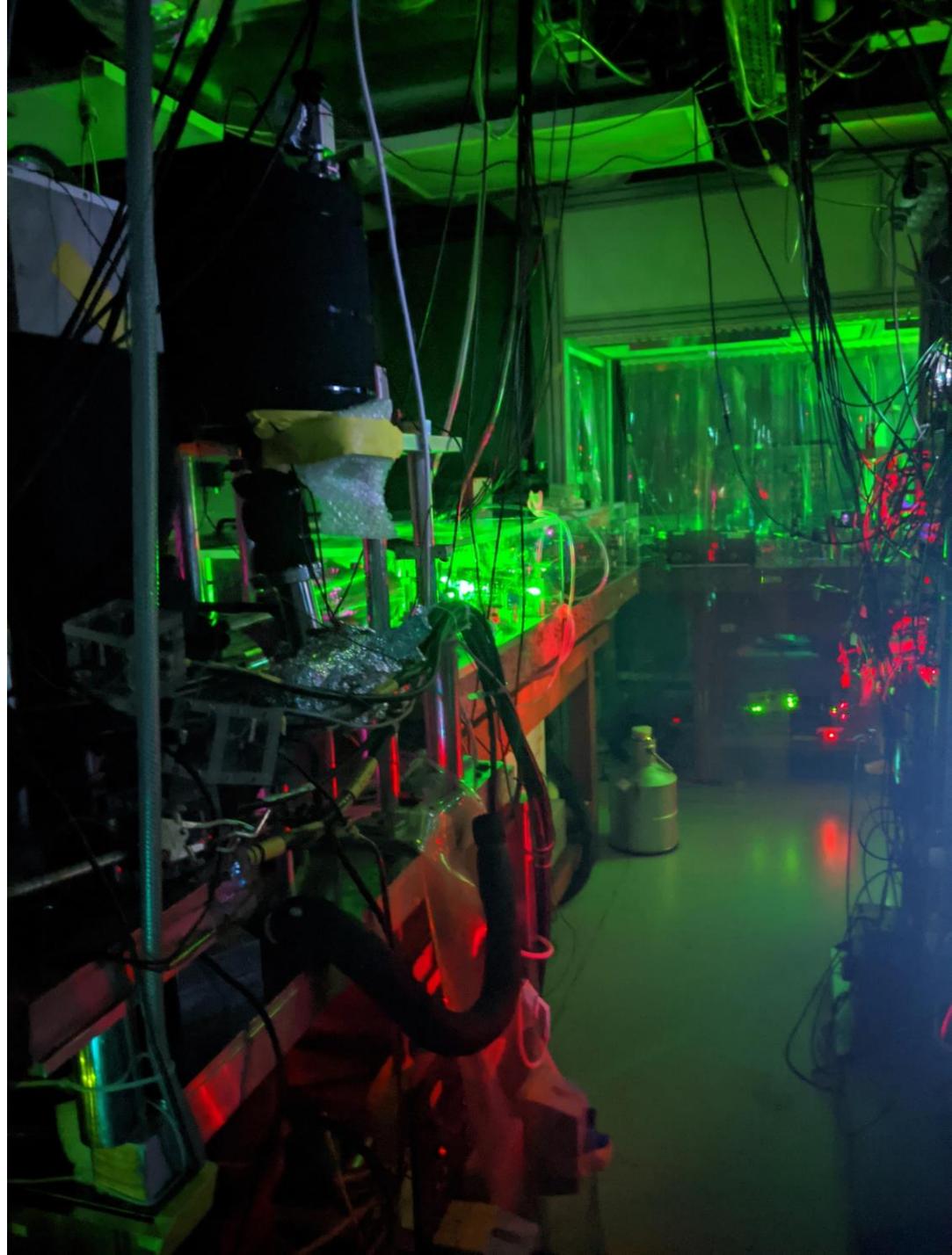


# Overview of the experiment





## Overview of the experiment



# Spectroscopy 1S-3S on Deuterium

Some of the results obtained during the *PhD thesis S. Thomas*  
- dec. 2021

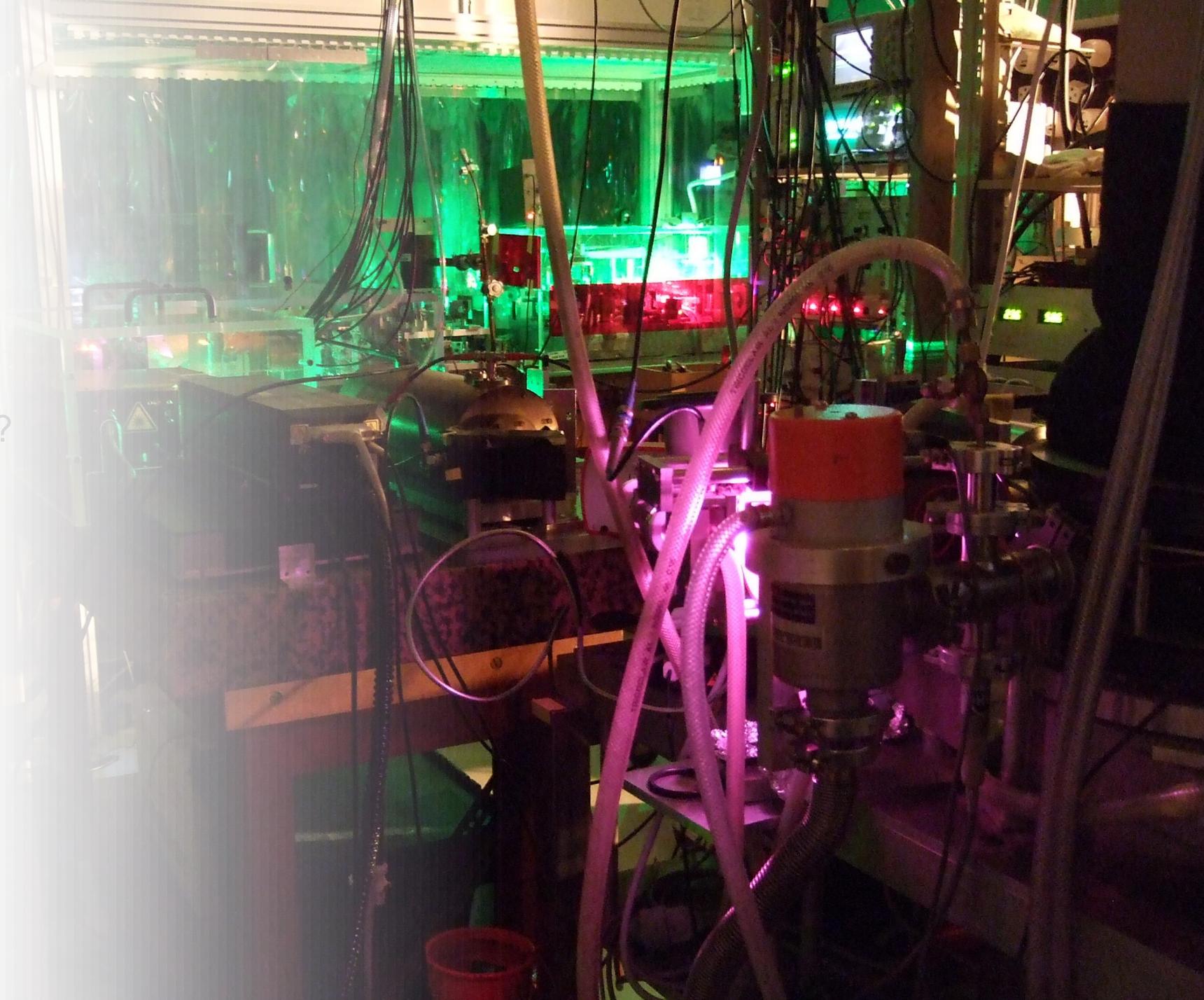
*Based on the campaign measurement on Deuterium atoms:*

**From 20 October to 17 December 2020**

- 9434 spectra
- 3 values of pressure
- 4 values of magnetic field (x 2 direction) (***new B field at “low” 20G***)
- measurement of the laser intensity (AC. Stark shift estimation)

# Overview

- The experiment
- **Dealing with systematics**
- A new systematics effect ?
- Preliminary result ?



# Dealing with systematics ...

1S - 3S spectroscopy on electronic Deuterium:

A few  $10^{-12}$  relative uncertainty targeted

**One of the main works:** identifying systematics effects and try to compensate or characterize them !

# Pressure shift

Origin: Collisions with rest gas in the chamber

=> broadening ~2 kHz << negligible (natural linewidth 1 MHz)

=> non negligible shift of the center of the line

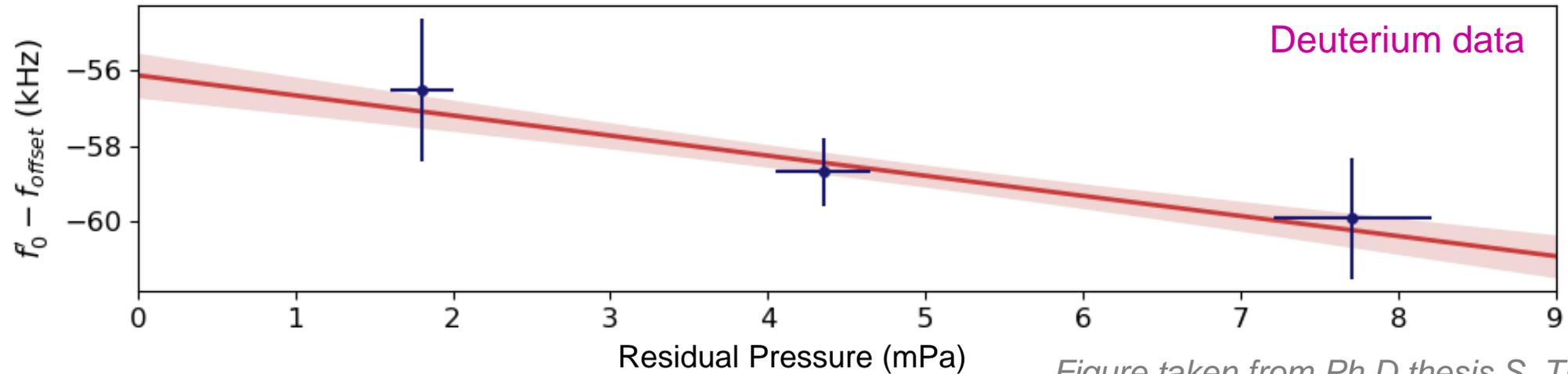


Figure taken from Ph.D thesis S. Thomas (2021)

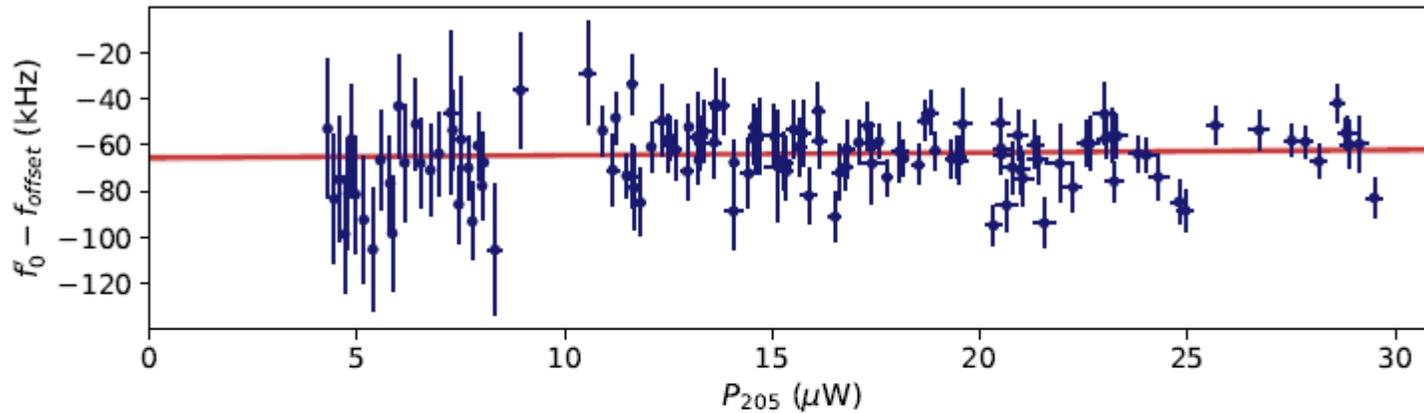
$$f'_0 - f_{offset} = -0,527(344) \times P_f$$

kHz/mPa

Pressure inside the vacuum chamber (mPa)

# Light shift (AC Stark shift)

- Proportional to the laser intensity inside the interaction chamber

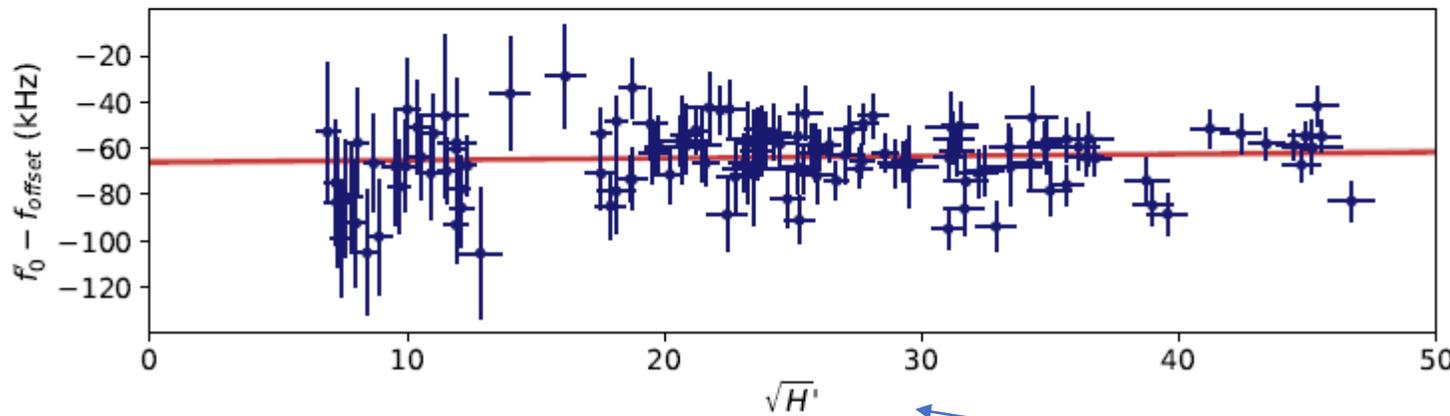


Deuterium data

$$f'_0 - f_{offset} = 0,115(176) \times P_{205}$$

kHz/ $\mu$ W

Transmitted 205nm light ( $\mu$ W)  
(through the exit mirror of the Fabry  
Perot cavity & pickup window)

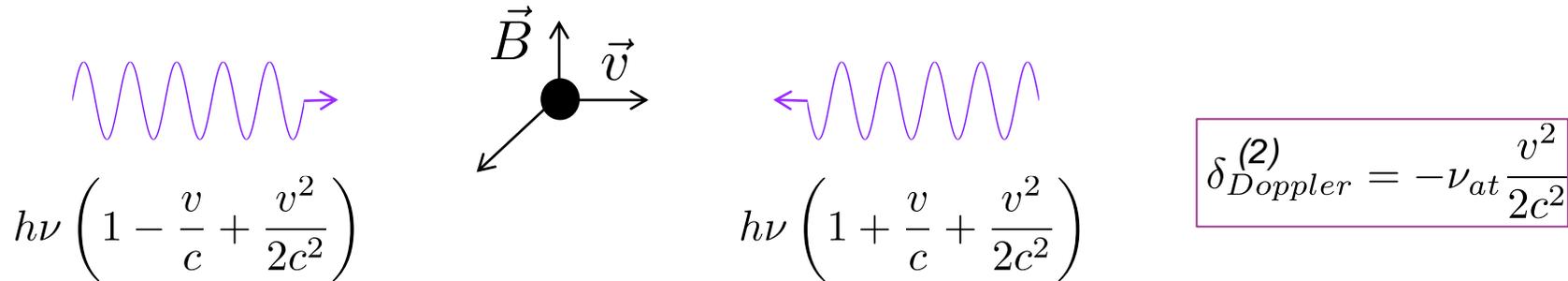


$$f'_0 - f_{offset} = 0,088(112) \times \sqrt{H'}$$

kHz/ $\sqrt{\text{counts/s}}$

Height of the signal  
(counts/s)

# 2<sup>nd</sup> Order Doppler effect frequency shift



$$h\nu \left( 1 - \frac{v}{c} + \frac{v^2}{2c^2} \right)$$

$$h\nu' \left( 1 + \frac{v}{c} + \frac{v'^2}{2c^2} \right)$$

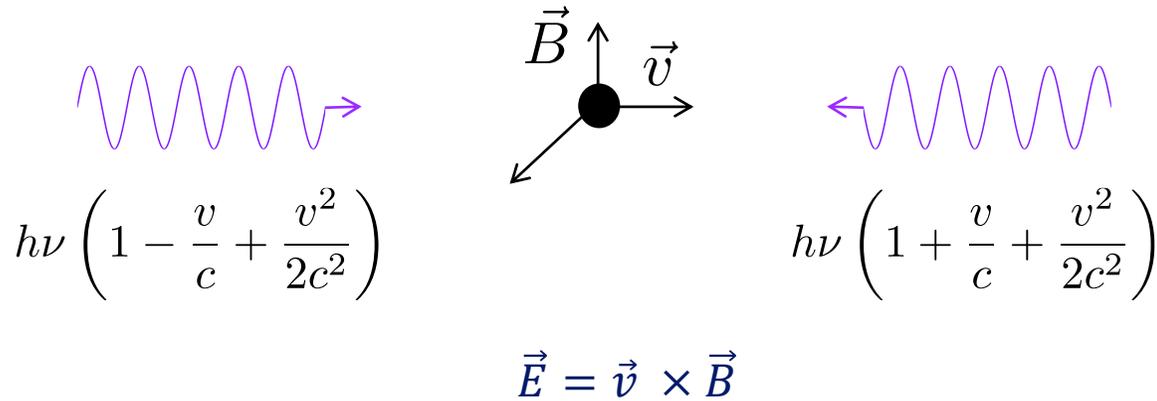
$$\delta_{Doppler}^{(2)} = -\nu_{at} \frac{v^2}{2c^2}$$

Need to determine the **velocity distribution** of the H beam ...

And no 1-photon transition easily achievable ( 121nm laser) for  
1<sup>st</sup> order Doppler broadening measurement

# 2<sup>nd</sup> Order Doppler effect

Compensating it ?



$$\delta_{Doppler}^{(2)} = -\nu_{at} \frac{v^2}{2c^2}$$

**Idea principle:** inducing a Motional Stark effect

$$\delta_{Stark} = \frac{E^2}{\Delta\nu_{SP}} = \frac{v^2 B^2}{\Delta\nu_{SP}}$$

To compensate 2<sup>nd</sup> order Doppler

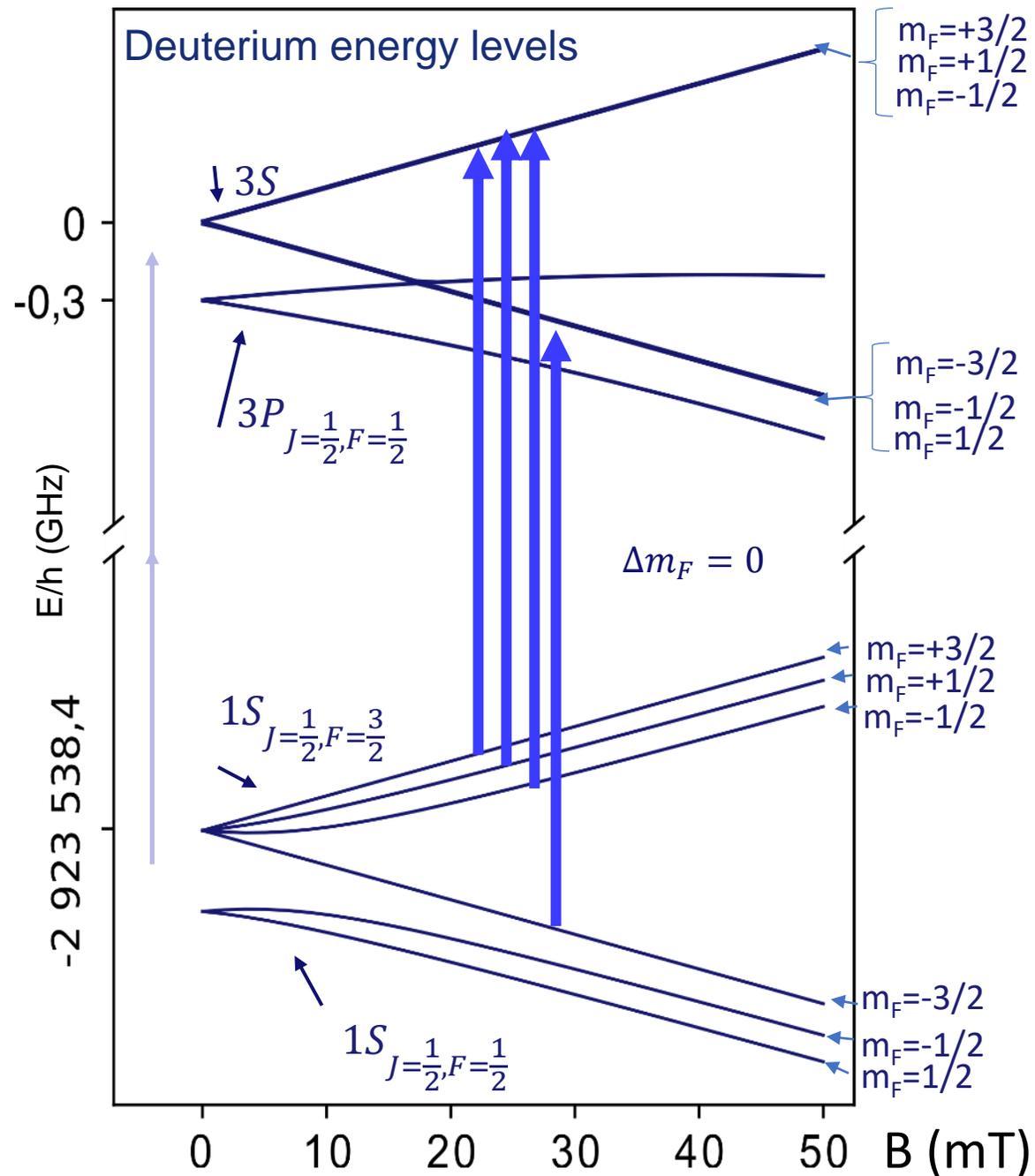
# 2<sup>nd</sup> Order Doppler effect

*Avoiding it ?*

Let's zoom in hyperfine structure

2- $\gamma$  transition Selection rules

$$\Delta m_F = 0$$



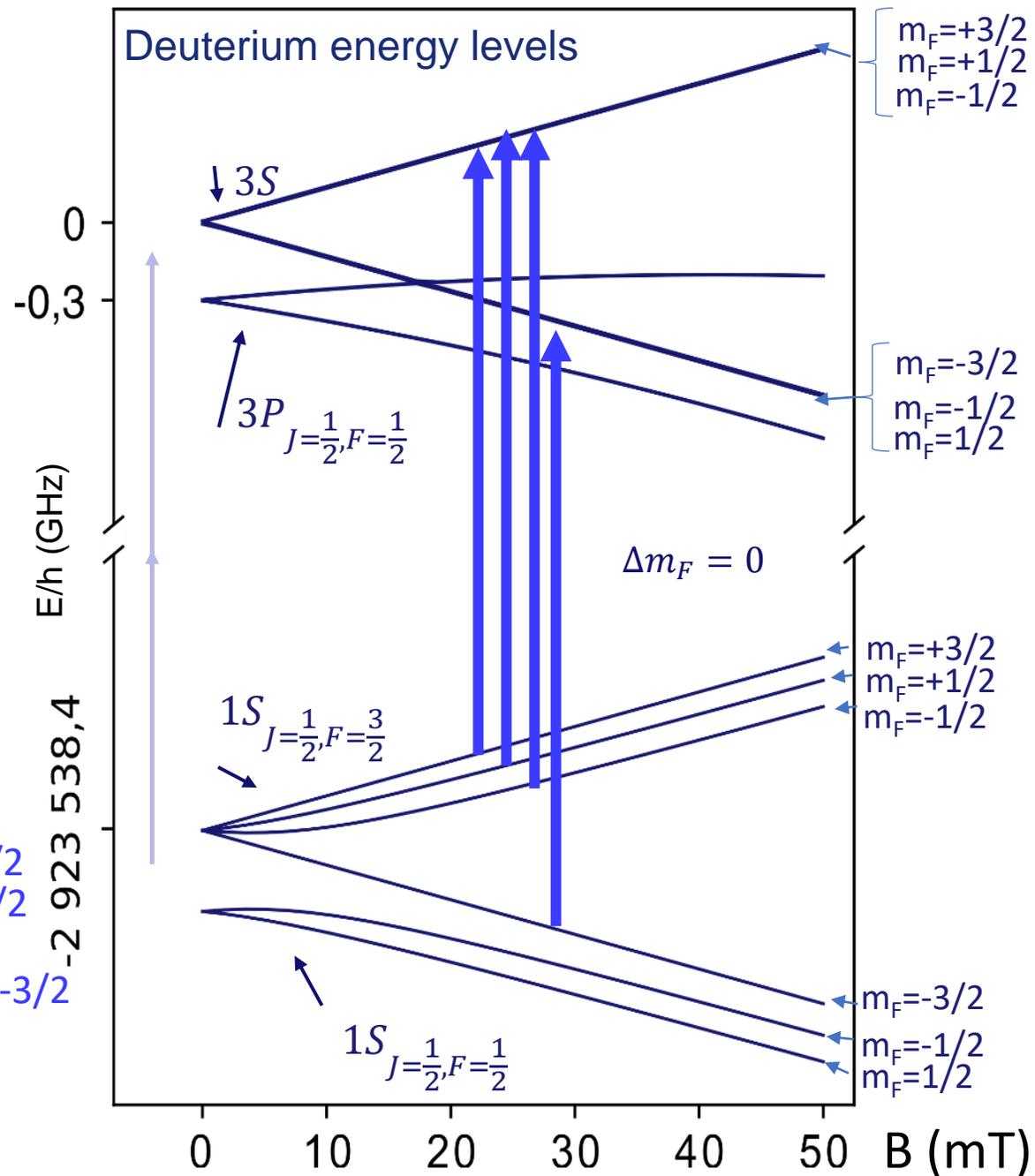
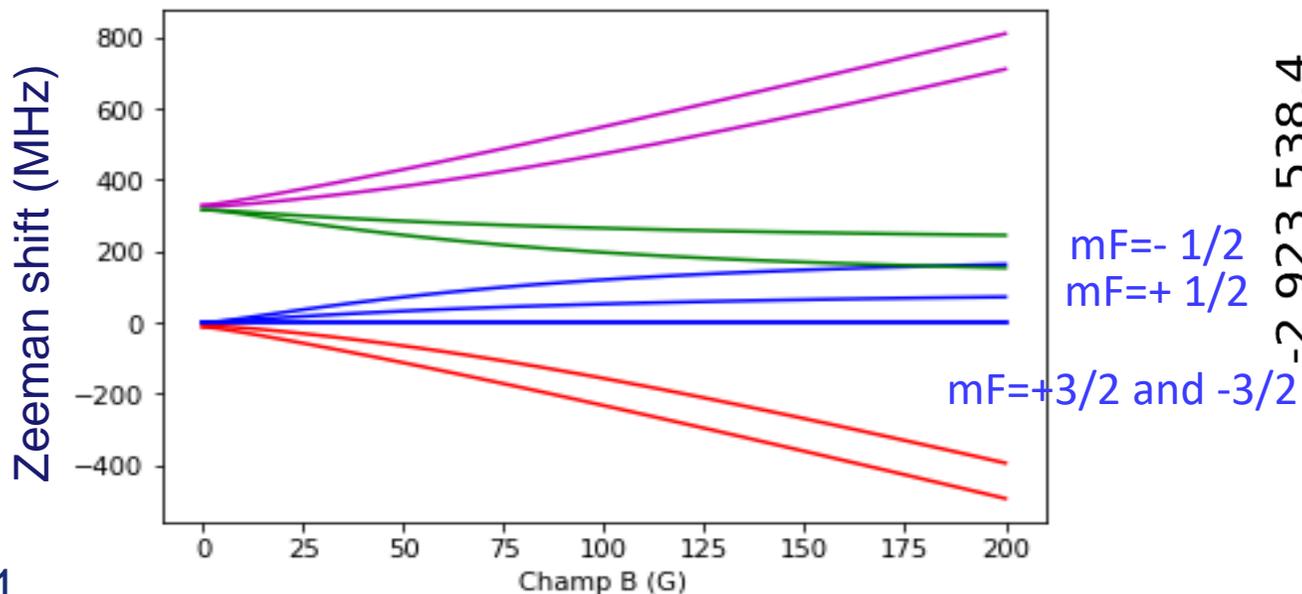
# 2<sup>nd</sup> Order Doppler effect

## Avoiding it ?

Let's zoom in hyperfine structure

Transitions  $1S\ F=3/2 \leftrightarrow 3S\ F=3/2$

Position of the 1S-3S transitions lines ( $\Delta m_F = 0$ )



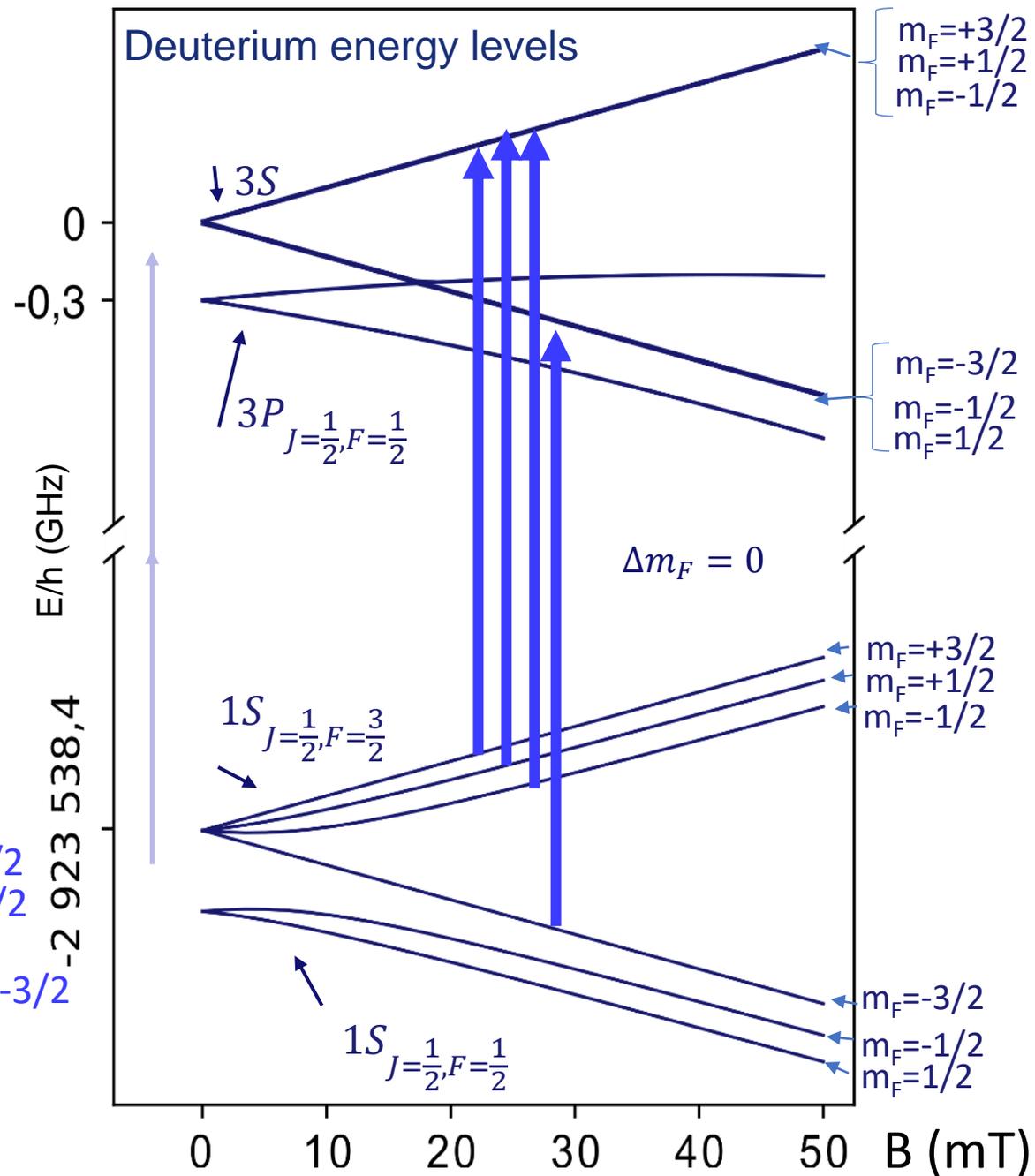
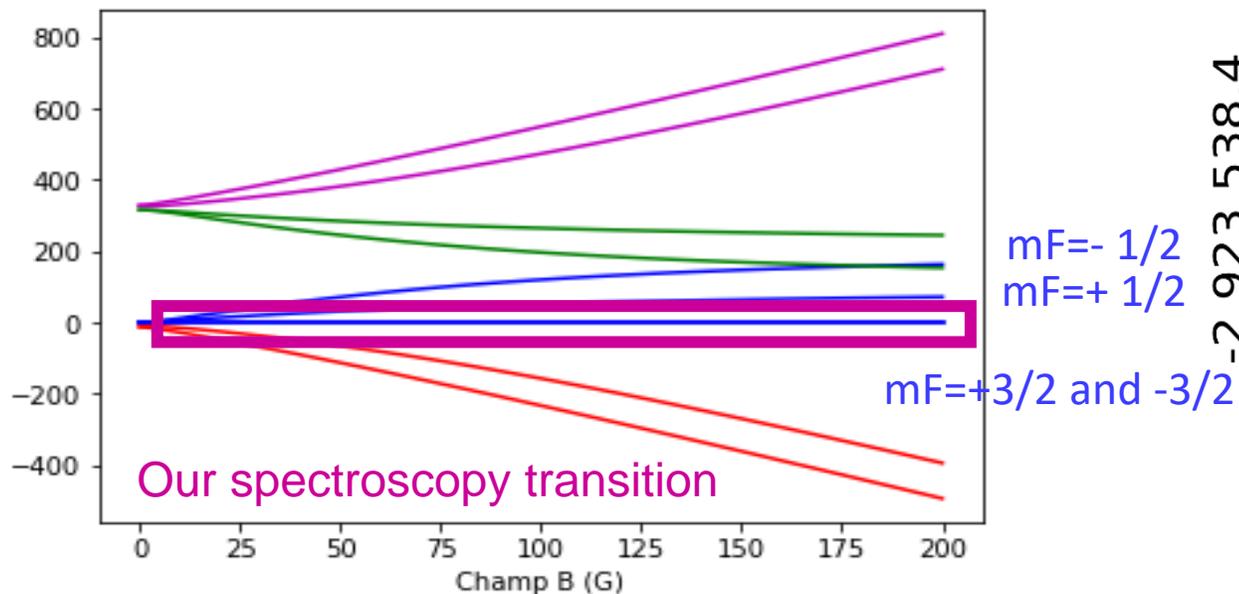
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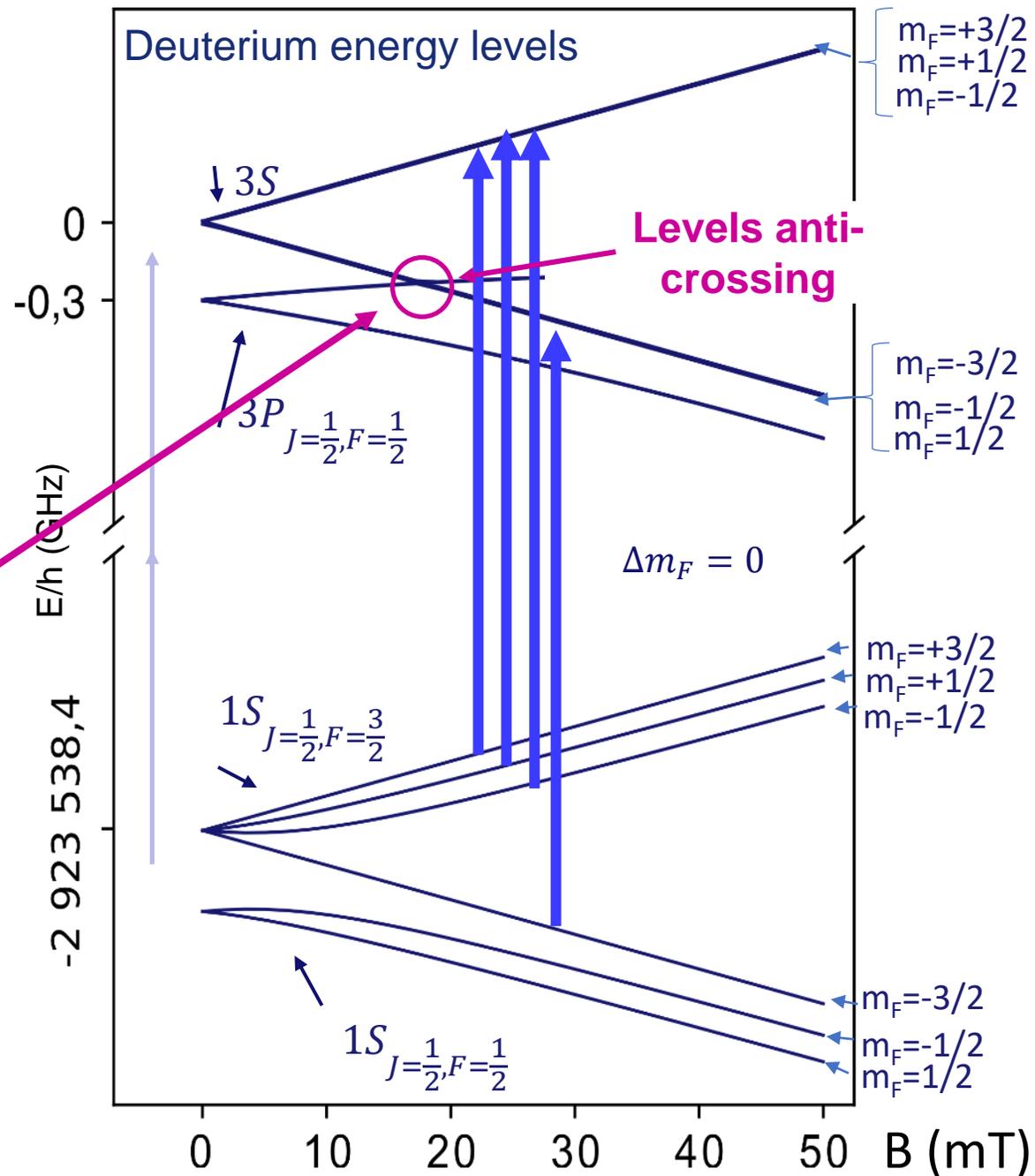
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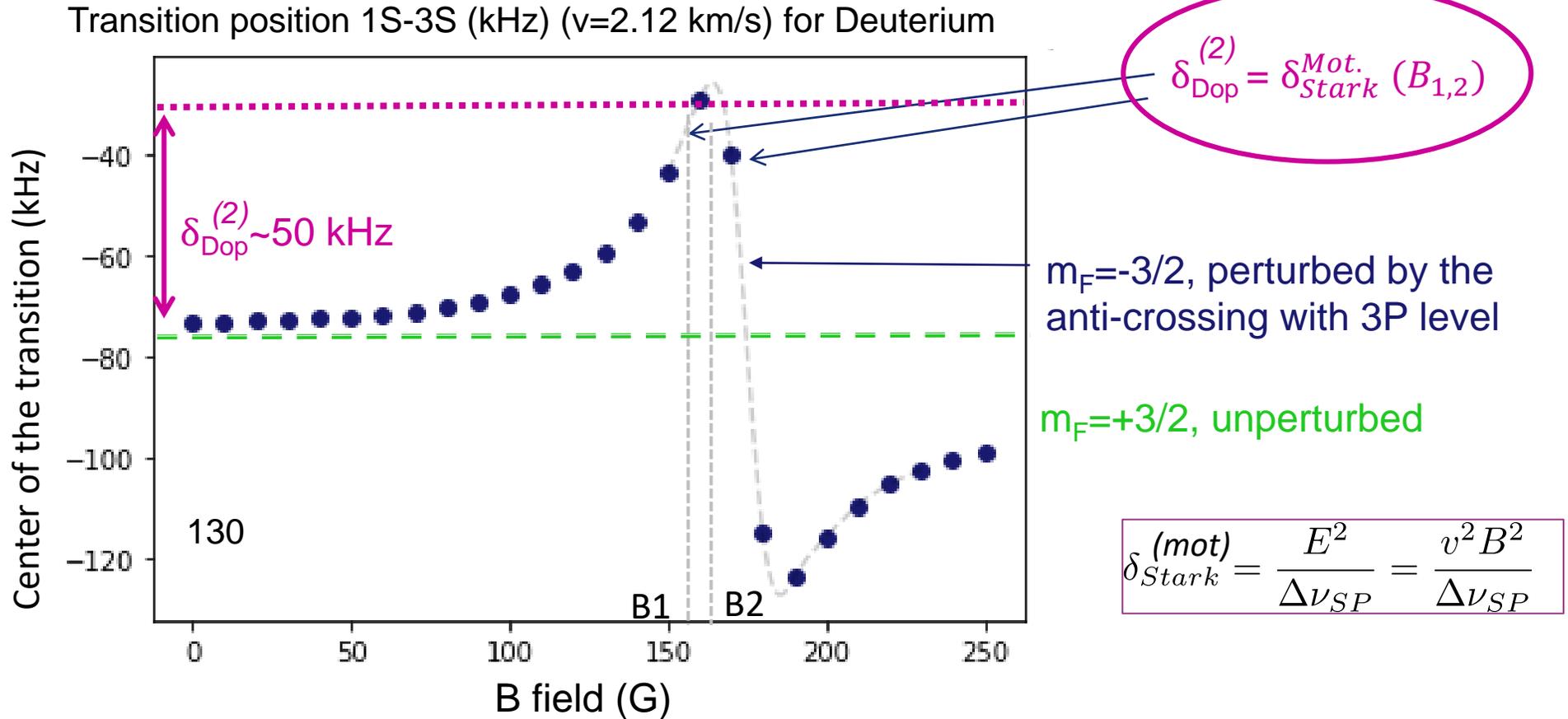
Transitions  $1S\ F=3/2 \leftrightarrow 3S\ F=3/2$

Motional Stark:  
mixing S – P states = anti-crossing

$$\delta_{Stark} = \frac{E^2}{\Delta\nu_{SP}} = \frac{v^2 B^2}{\Delta\nu_{SP}}$$

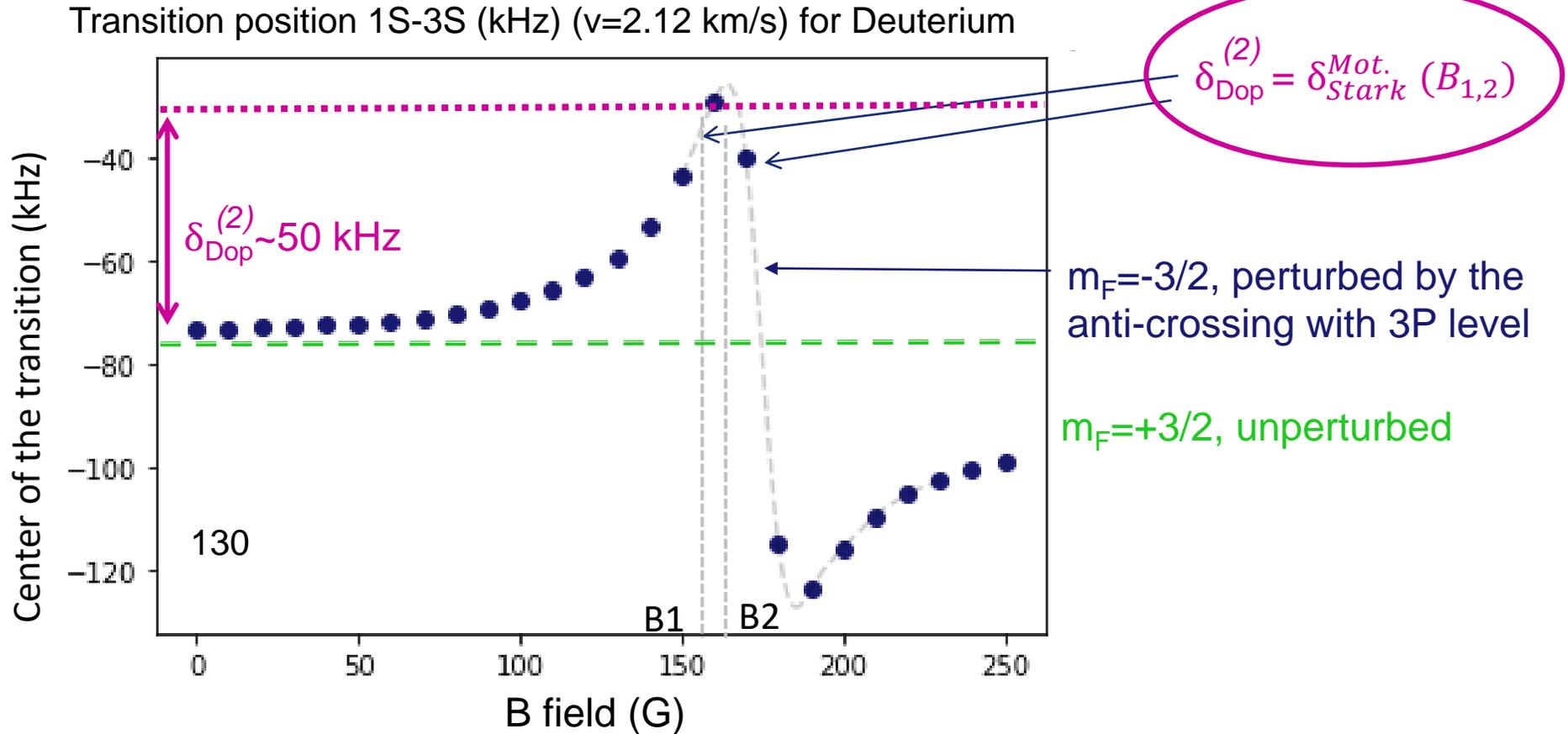


# Avoiding 2<sup>nd</sup> order Doppler effect



- Total compensation of 2<sup>sd</sup> Ord. Doppler for  $B = B1$  and  $B2$  for  $m_F = -3/2$  transition

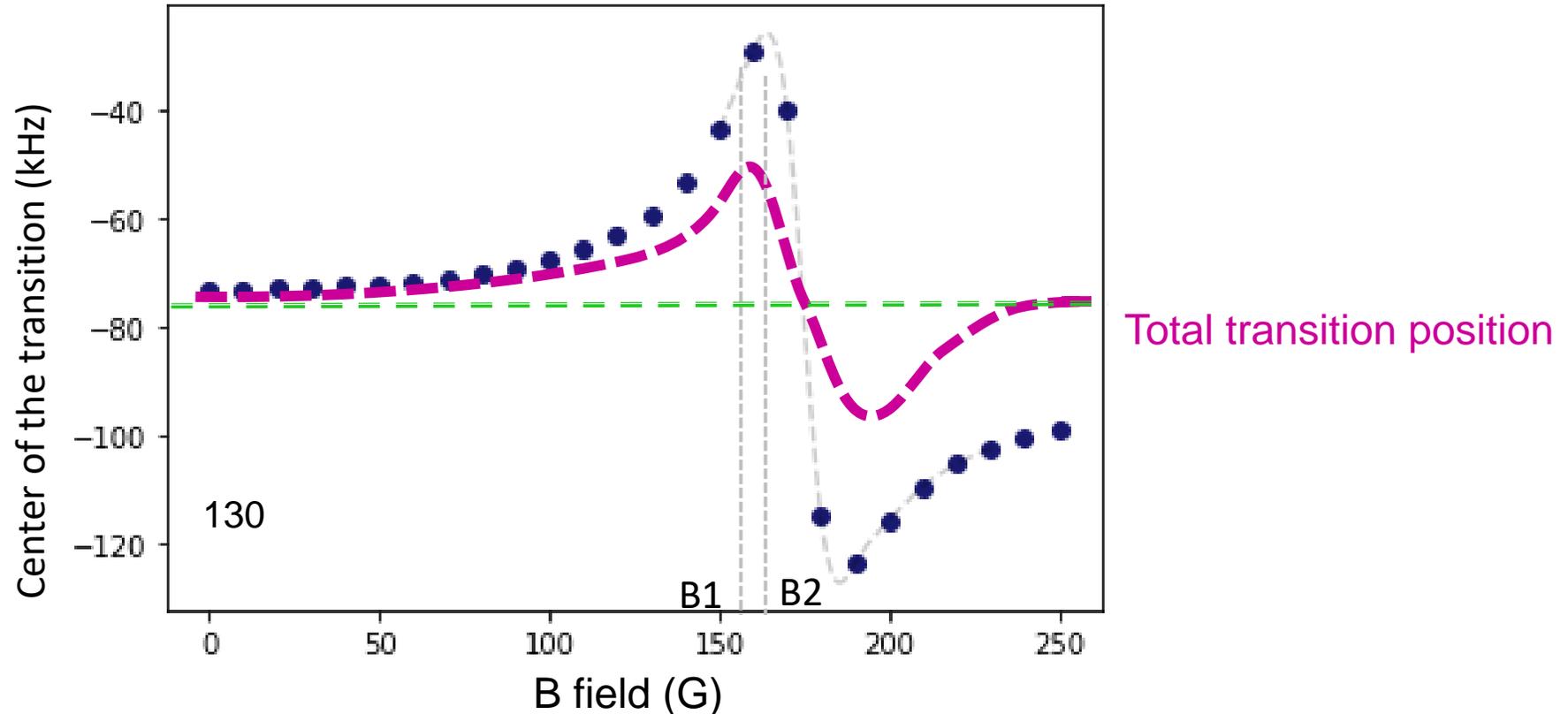
# Avoiding 2<sup>nd</sup> order Doppler effect



- Total compensation of 2<sup>sd</sup> Ord. Doppler for  $B = B1$  and  $B2$  for  $m_F = -3/2$  transition
- For  $v = 2.3$  km/s: the transitions  $m_F = -3/2$  and  $m_F = 3/2$  are split by  $\sim 50$  kHz for  $B = B1$  and  $B2$ .
- The 3S natural bandwidth  $\sim 1$  MHz  $\Rightarrow$  **Both lines are excited**

# Avoiding 2<sup>nd</sup> order Doppler effect

Transition position 1S-3S (kHz) ( $v=2.12$  km/s) for Deuterium



The two lines cannot be resolved  $\Rightarrow$  **partial compensation only** due to Motional Stark

$\Rightarrow$  **Determination of the velocity distribution** by fitting the dispersion curve with several data points for different B (« pink » profile)

# Examples of fits to extract the best velocity distribution

Our theoretical fluorescence lines: take into account:

- Zeeman, Motional Stark, 2<sup>sd</sup> Doppler effects
- Integrated over all the velocity distribution.

Fits give the “zero field” frequency (*corrected from the above shifts*)

The theoretical lines depend on:

- B field
- Velocity distribution of the H beam (2 parameters to model it)

Protocol: it our data sets for each recorded B field, with various velocity distribution parameters.

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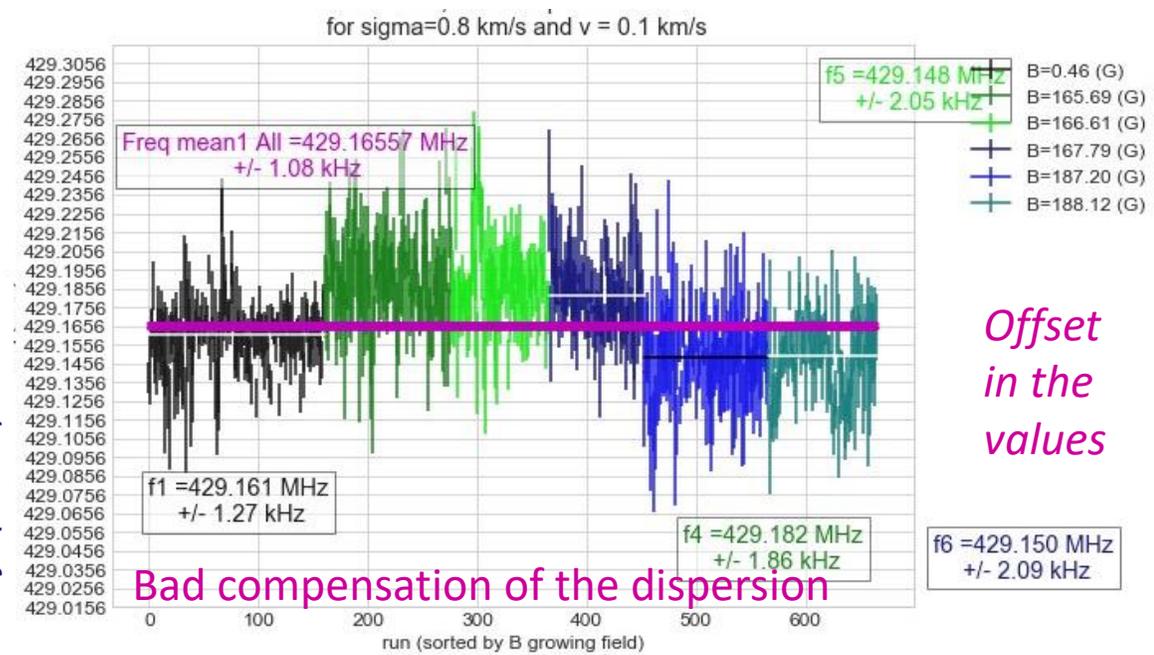
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Extracted transition frequency (MHz)



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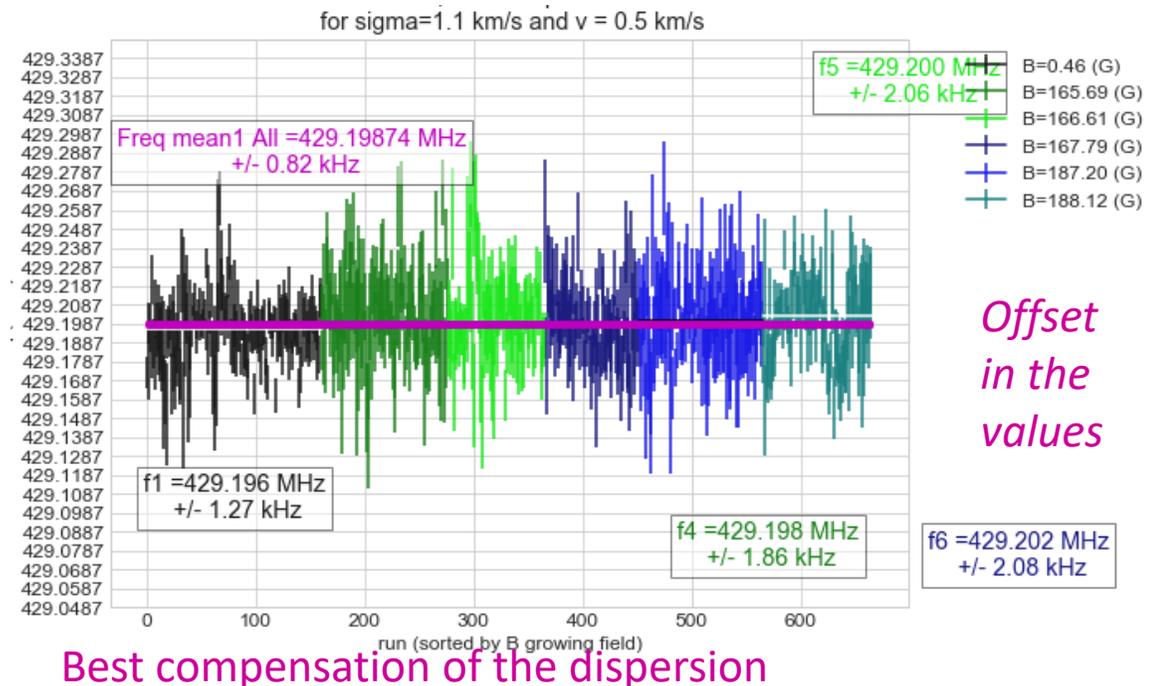
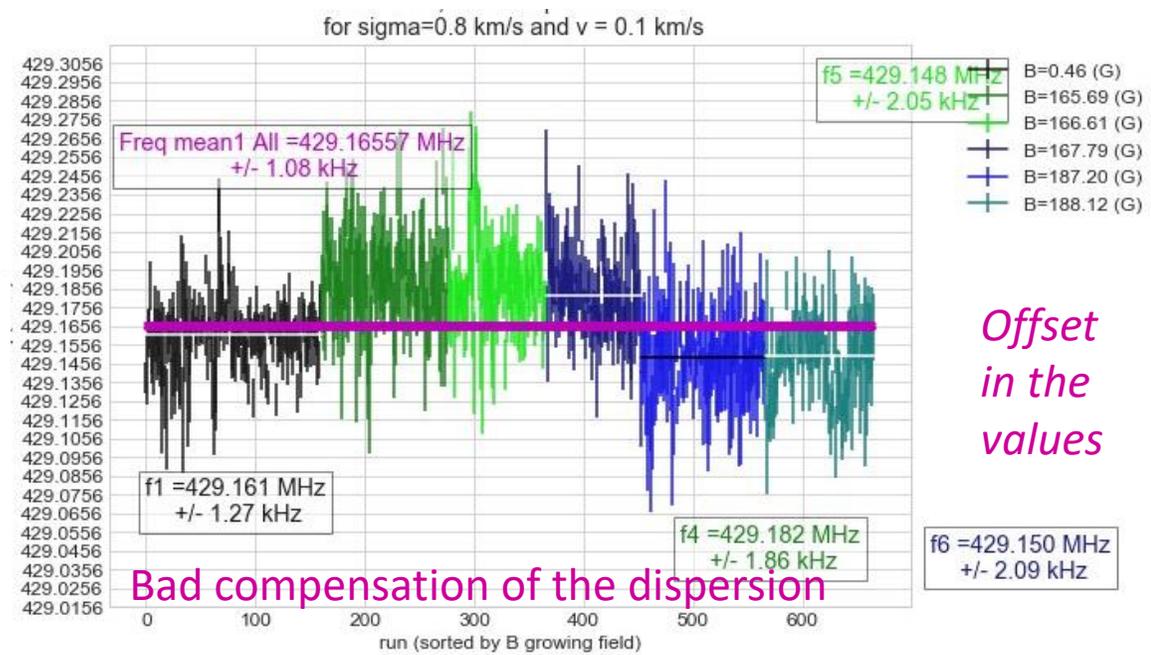
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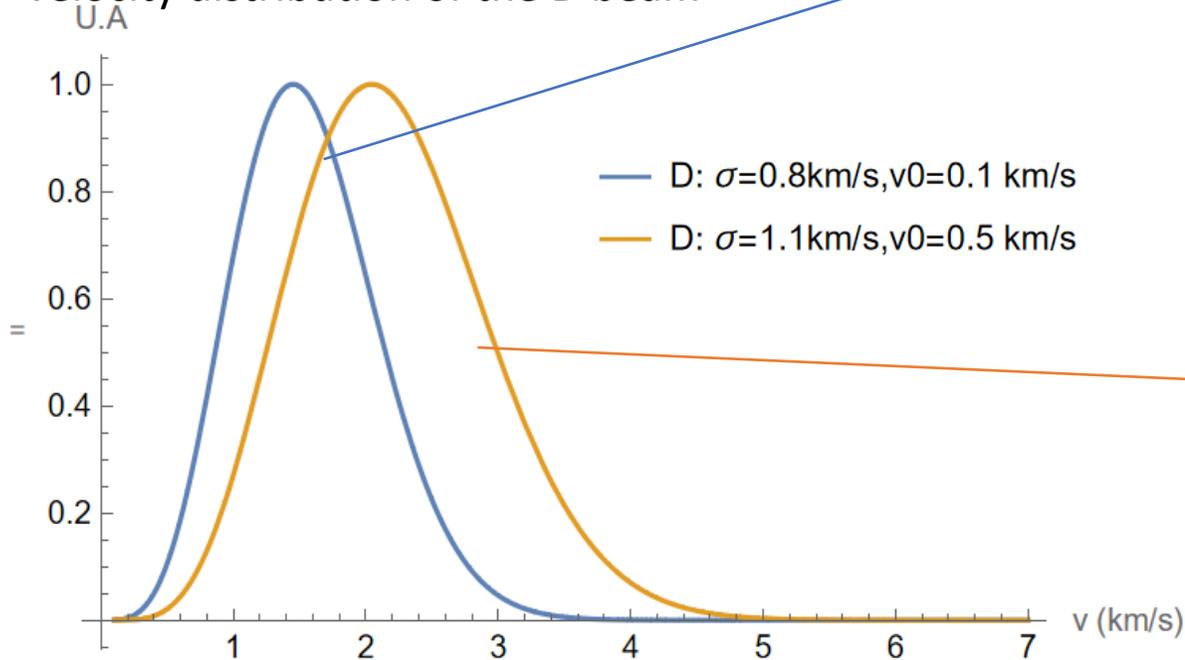
Protocol: it our data sets for each recorded B field, with various velocity distribution parameters.

Extracted transition frequency (MHz)



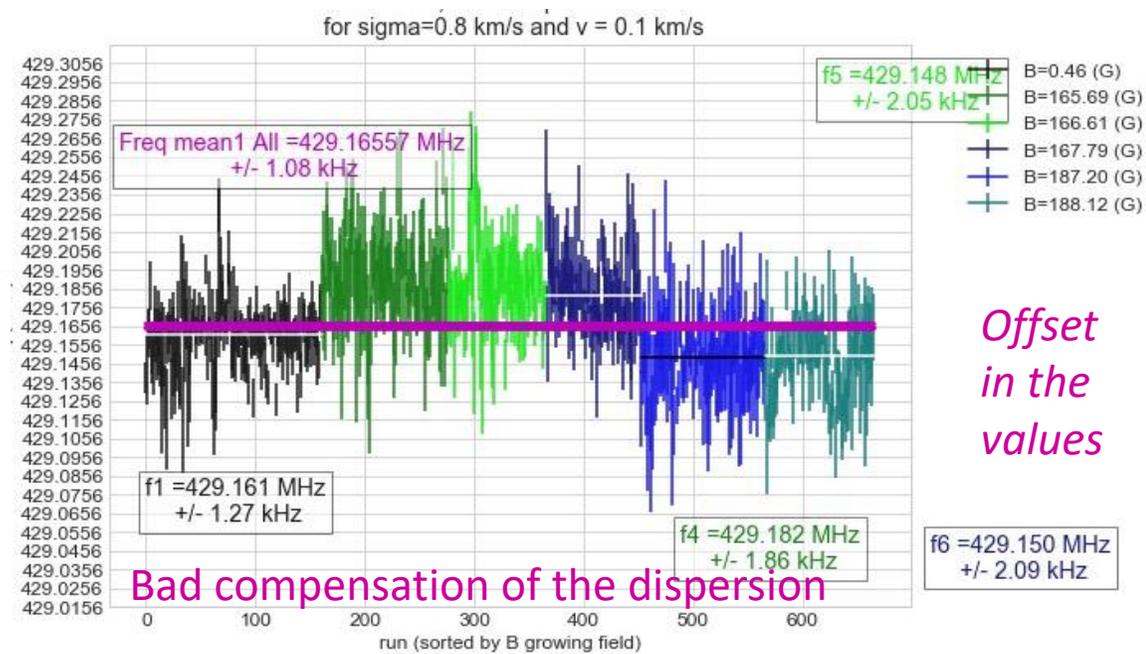
# Examples of fits to extract the best velocity distribution

Velocity distribution of the D beam



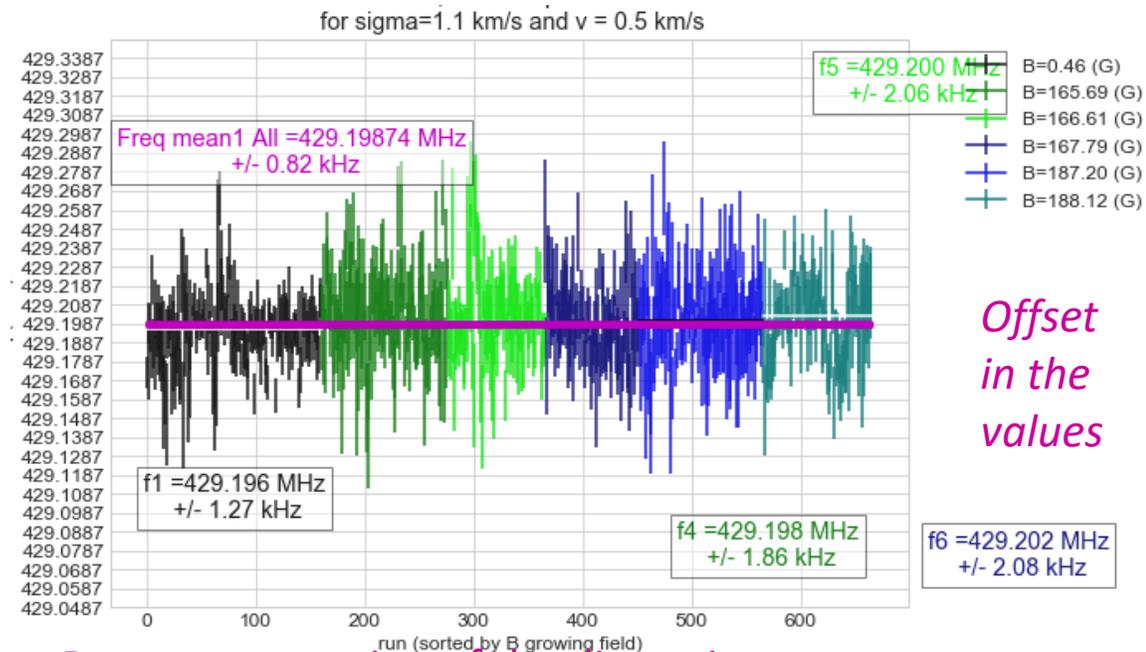
$$f(v, \sigma, v_0) \propto v f_M(v, \sigma) P[\Psi(z), \text{Kn}] \exp(-v_0/v)$$

Extracted transition frequency (MHz)



*Offset in the values*

*Bad compensation of the dispersion*

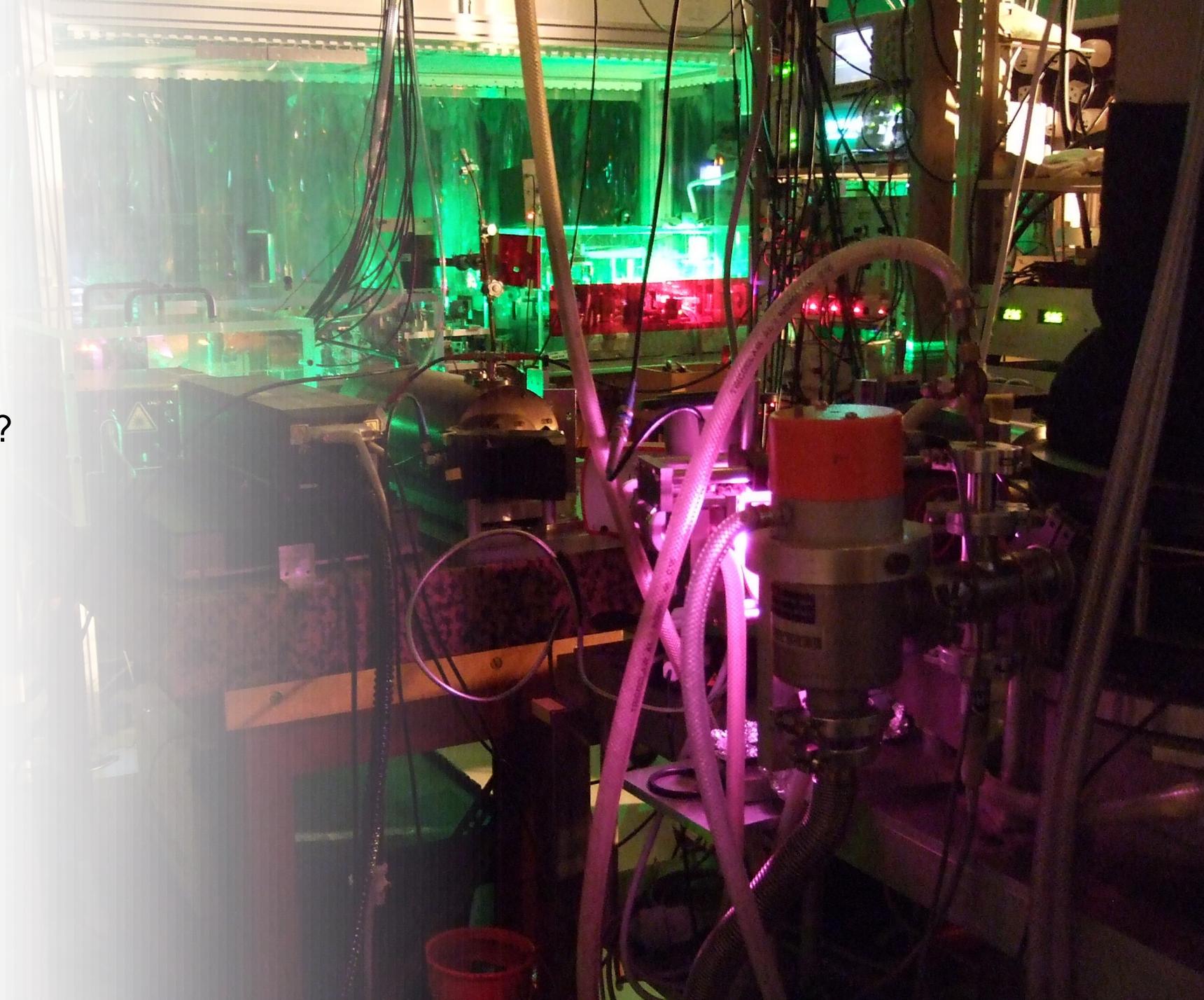


*Offset in the values*

*Best compensation of the dispersion*

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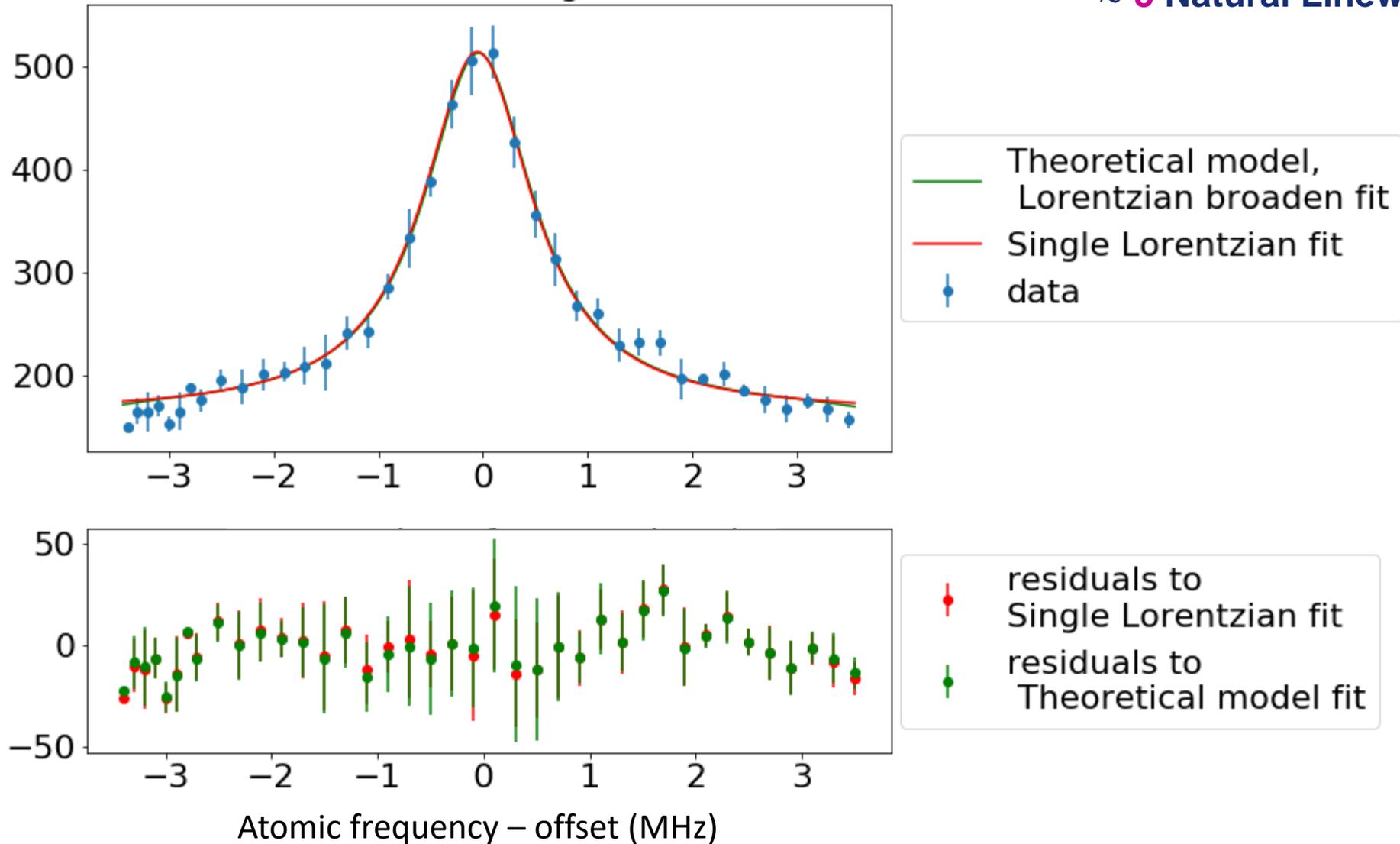
- The experiment
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# For the campaign, we recorded data over [-3,3] MHz range

Fit on the center data  $B = 0.46$  G  
central range

Frequency Center range scan  
 $\approx 5$  Natural Linewidth



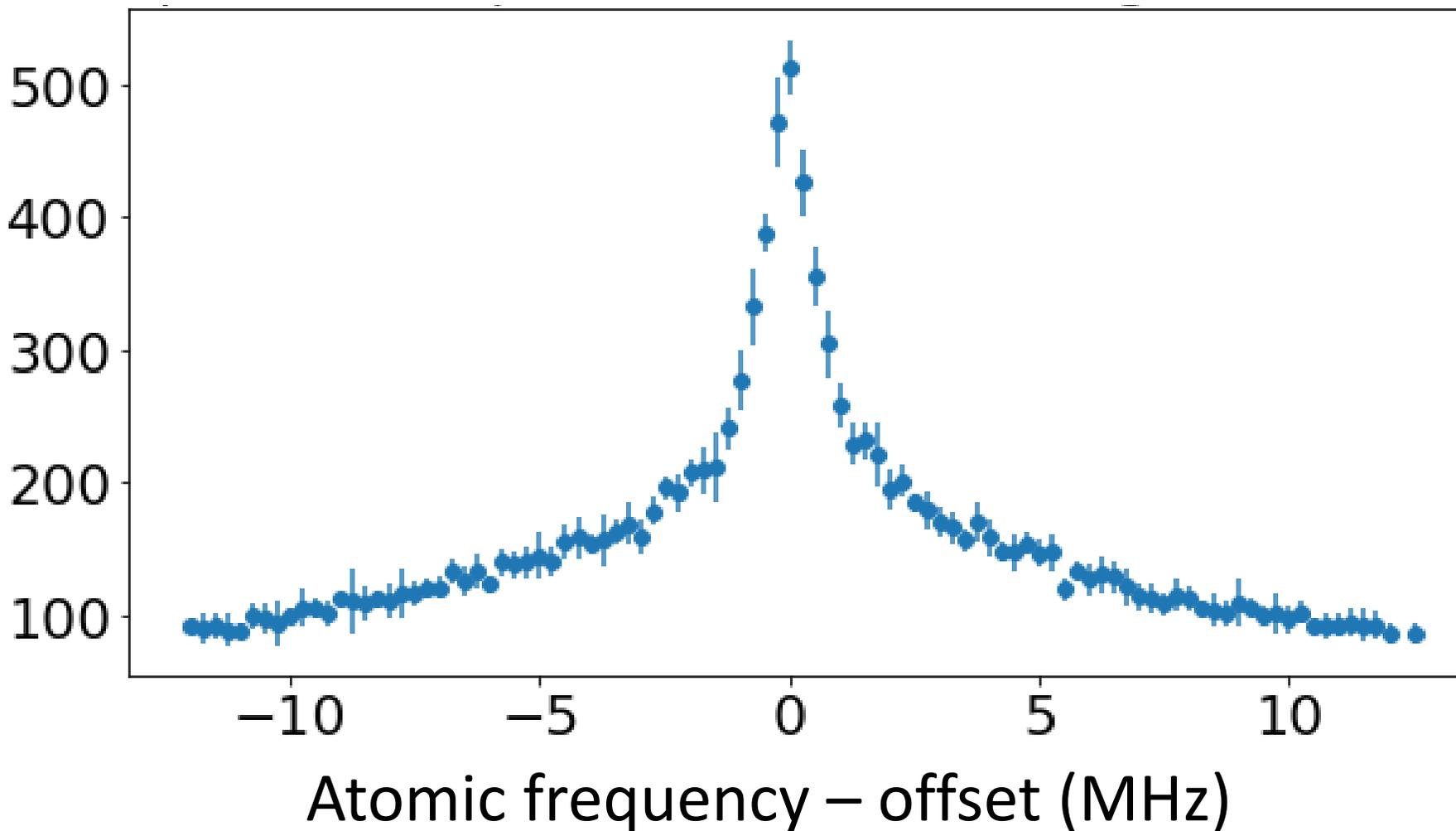
# Improvements of the experiment during the last years

- Recent Work to increase the frequency stability and tunability of the laser
  - **Before: usual** scan  $\pm 2.5 - 3$  MHz scan ( $\approx 5$  Natural Linewidth) –  $\sim 6$  min for 1 run (= 10 scans)
  - possible up to  $\pm 5$  MHz ( $\approx 10$  Natural Linewidth) –  $\sim 11$  min
  - **New system:**  $\pm 10 - 30$  MHz scan possible ( $\approx 20 - 60$  Natural Linewidth)  $\sim 20$  min  $\Rightarrow$  A few “wide” scans recorded

# With a wider laser frequency scan:

Deuterium spectrum  
14/12/2020, run 38  
B=0.46 G

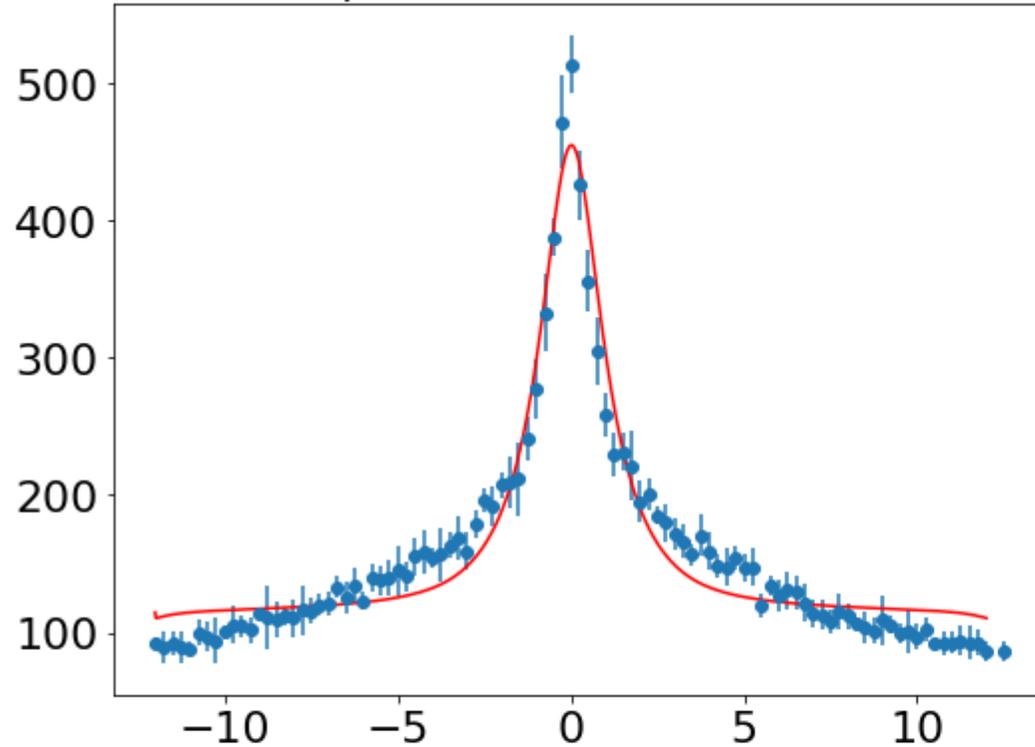
Wide spectrum, recorded for B=0.46 G



Averaged  
1 run  
(10 scans)

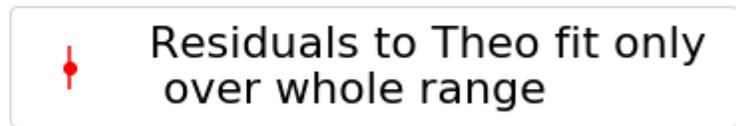
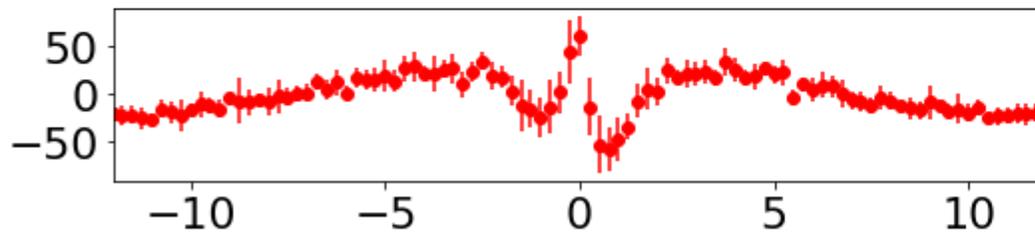
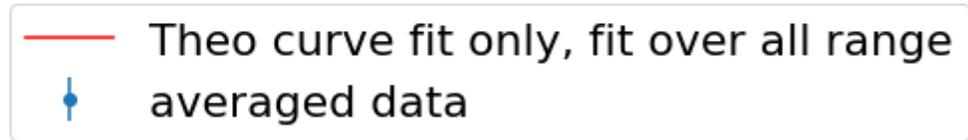
# Which fit model ?

Comparison models



$$F(\nu_L) = F_{theo}(\nu_L, \nu_0, \sigma_0)$$

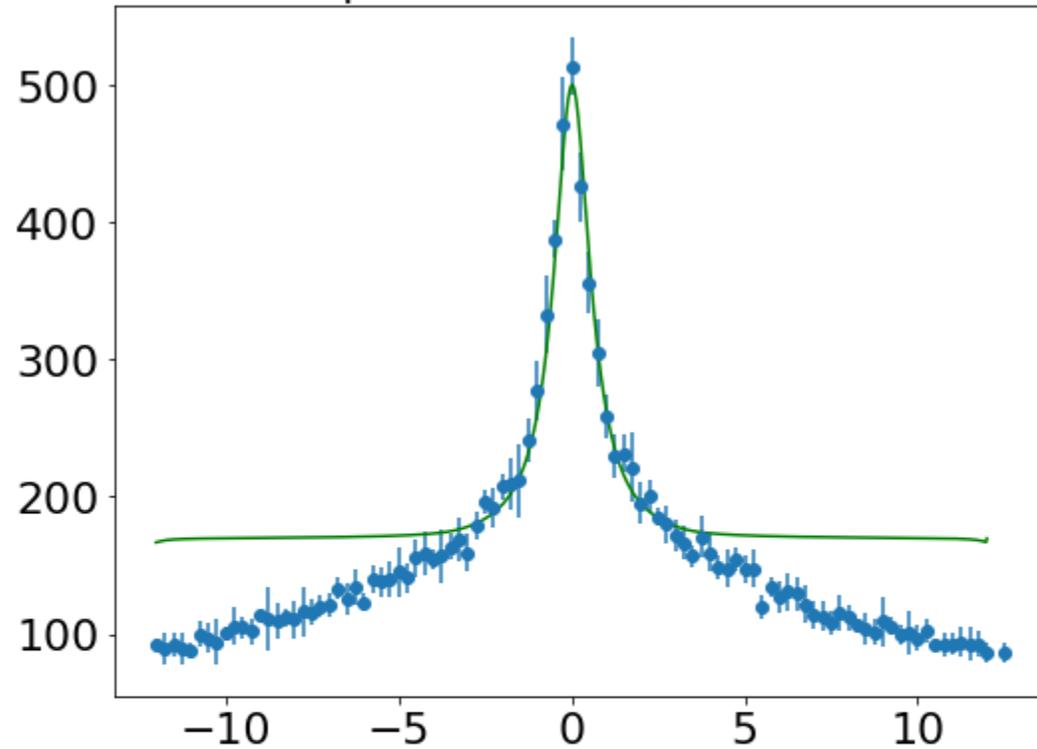
Fitting range : all data



Atomic frequency – offset (MHz)

# Which fit model ?

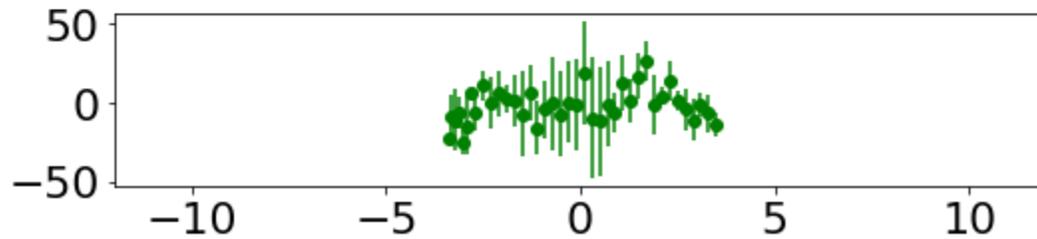
Comparison models



$$F(\nu_L) = F_{theo}(\nu_L, \nu_0, \sigma_0)$$

Fitting range : central data only

— Theo curve fit only, fit over central range  
• averaged data



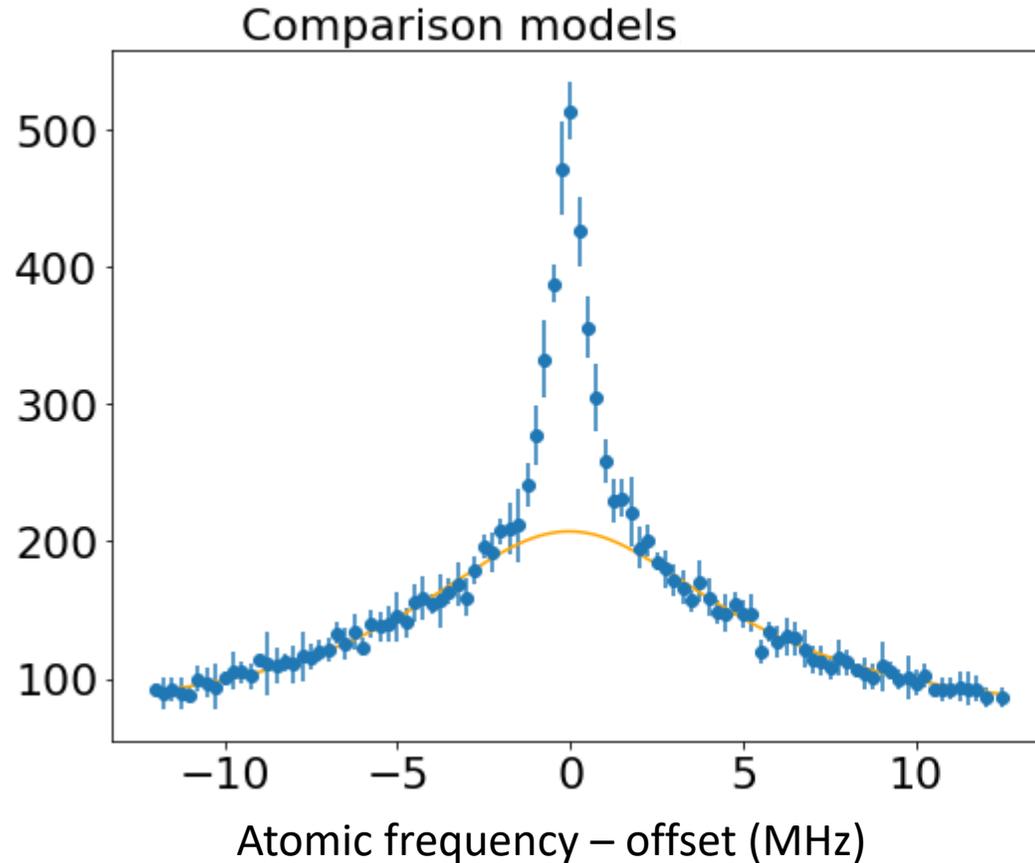
• Residuals to Theo fit only over central range

Atomic frequency - offset (MHz)

# Which fit model ?

$$F(\nu_L) = Lor(\nu_L, \nu_{Bump}, \Gamma_{Bump})$$

Fitting range : edge data only



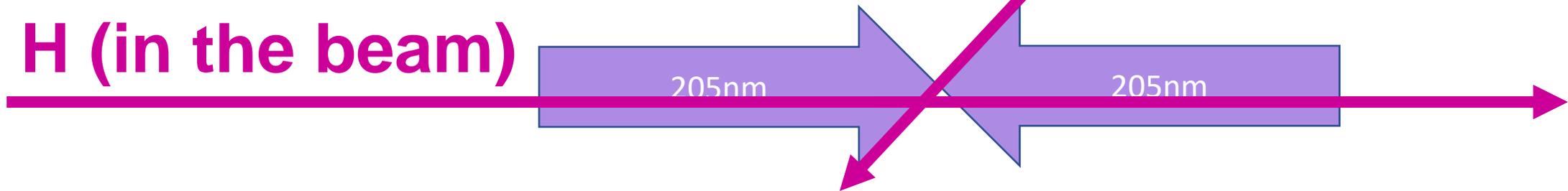
Presence of a “Pedestal fluorescence signal” = “Bump” signal

# Which fit model ?

Origin of this pedestal ?

Presence of residual H gas in the chamber.

H (residual gas in the chamber)



⇒ Contribution to a **broaden** fluorescence signal due to **short interaction time**

New model to fit:

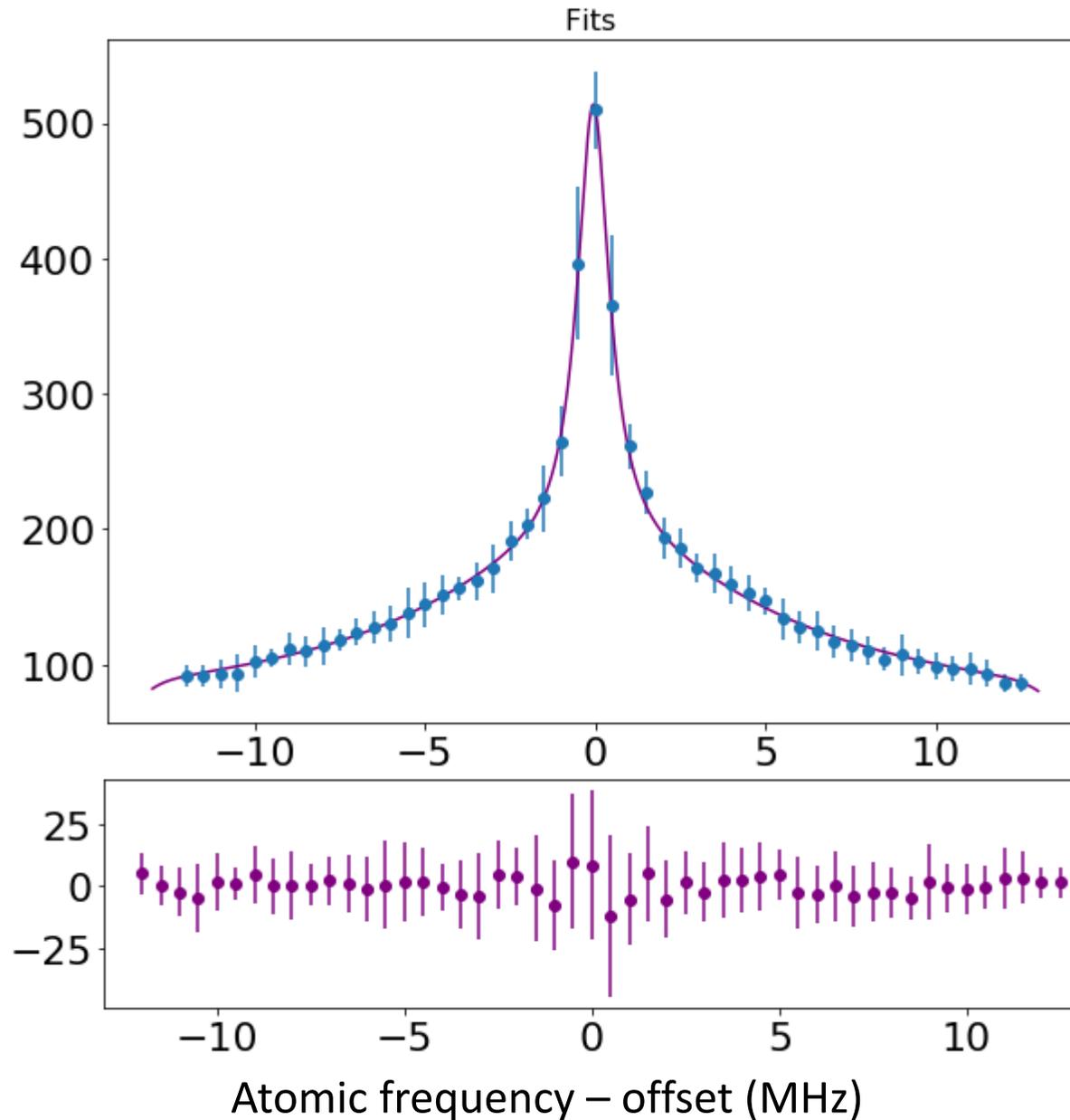
$$F(\nu_L) = F_{theo}(\nu_L, \nu_0, \sigma_0) + F_{theo}(\nu_L, \nu_B, \sigma_B) + B$$

# Which fit model ?

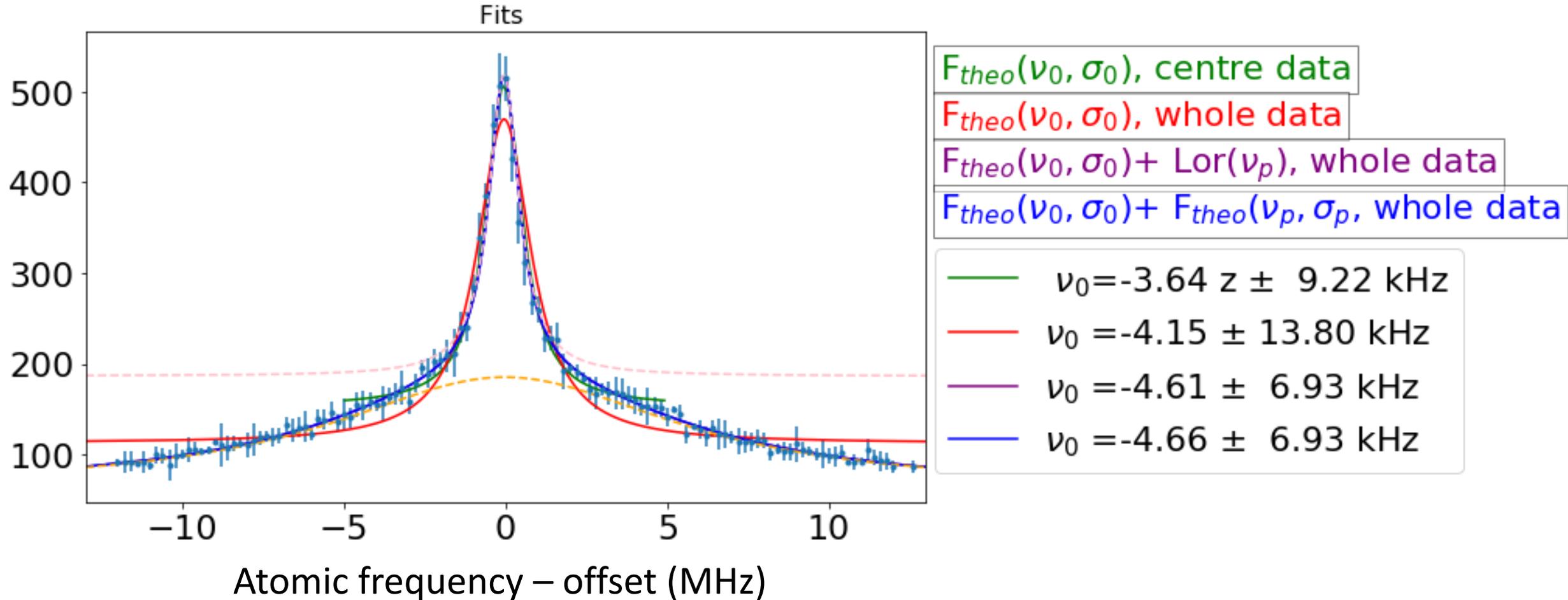
$$F(\nu_L) = F_{theo}(\nu_L, \nu_0, \sigma_0) + F_{theo}(\nu_L, \nu_B, \sigma_B)$$

Fitting range : all data

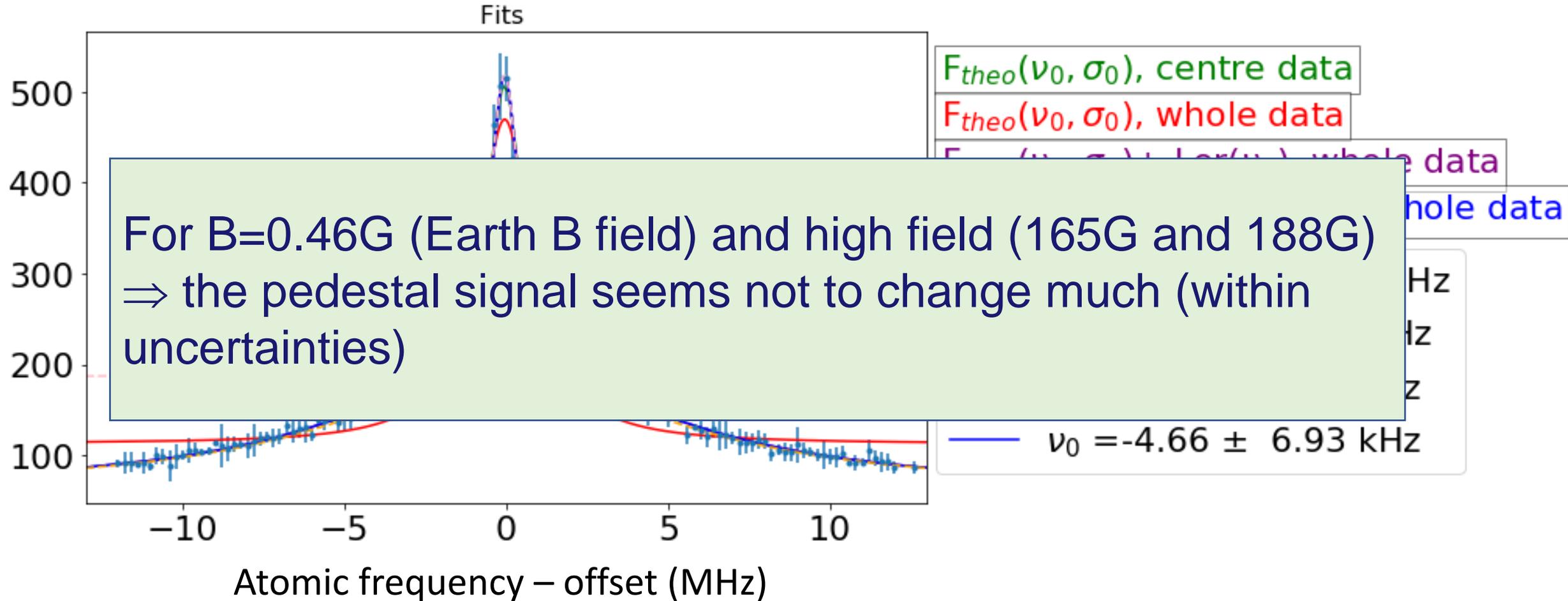
 Theo curve + Extra broaden Theo curve for Bump, whole data



# What does it change for the fitted centre frequency?

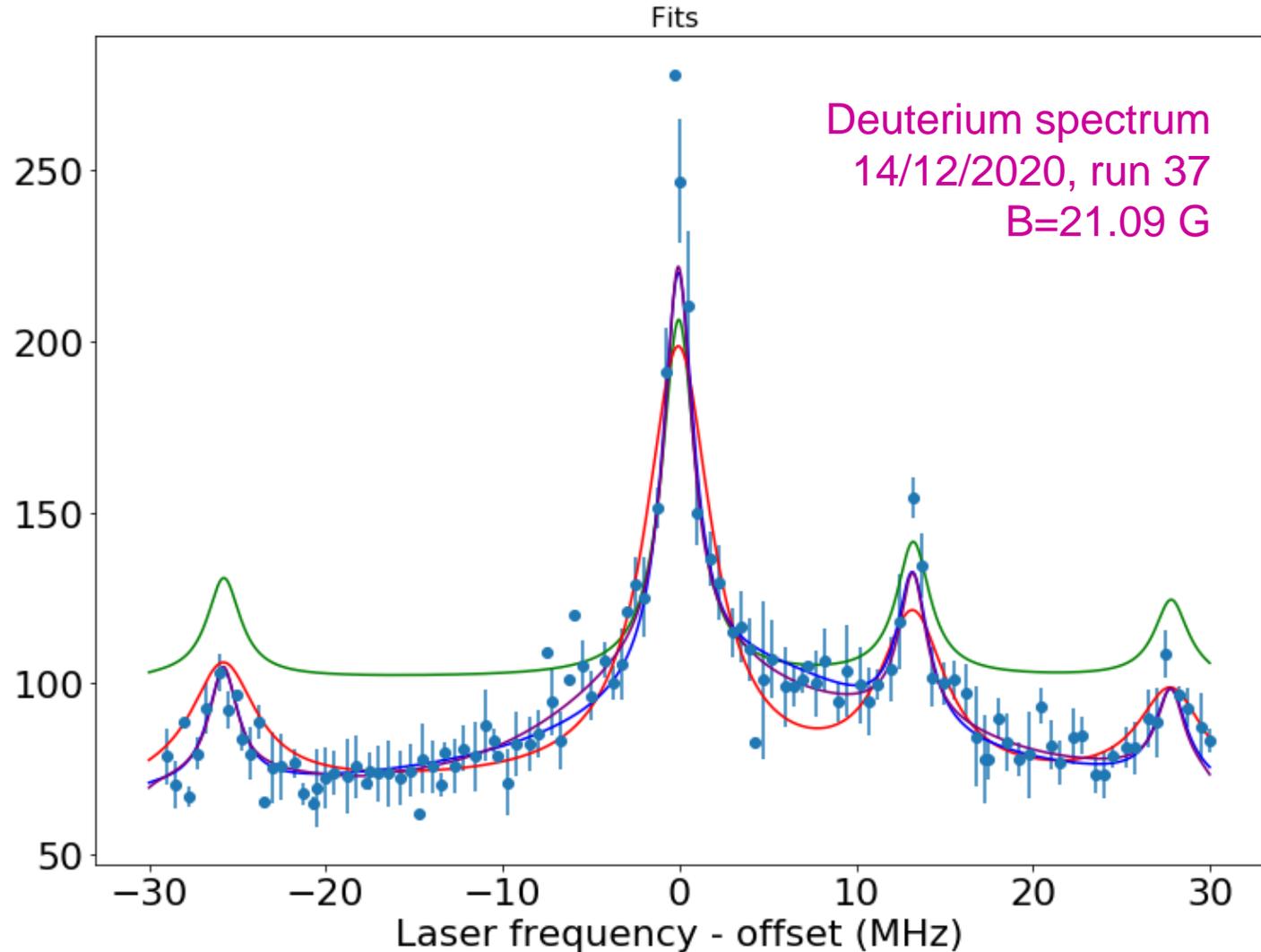
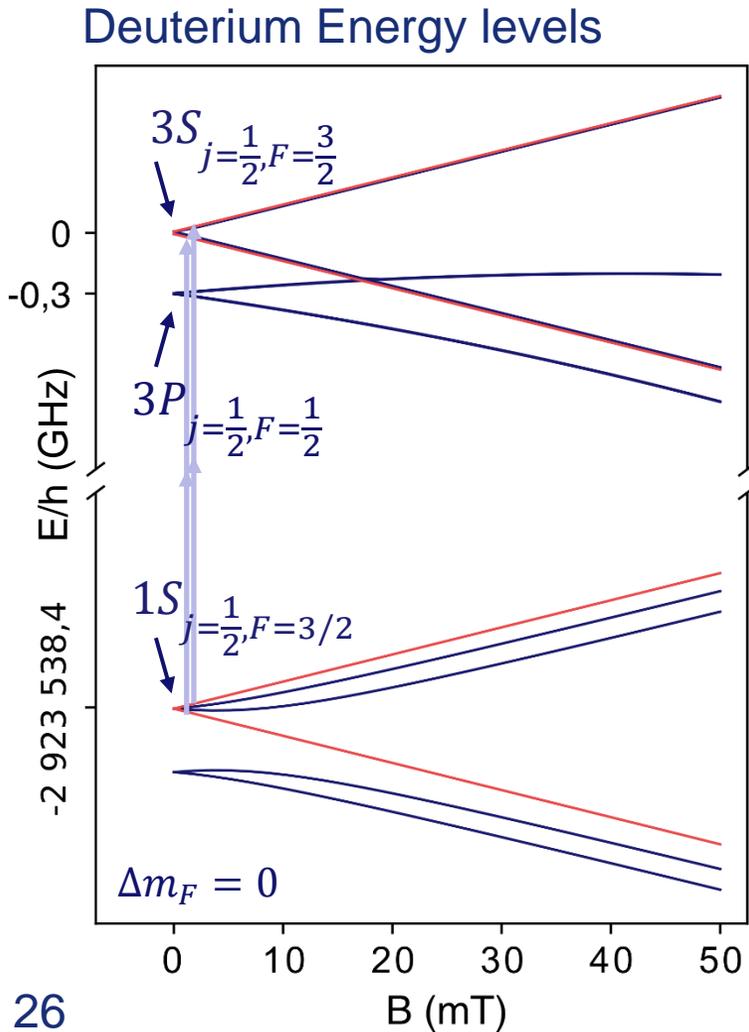


# What does it change for the fitted centre frequency?

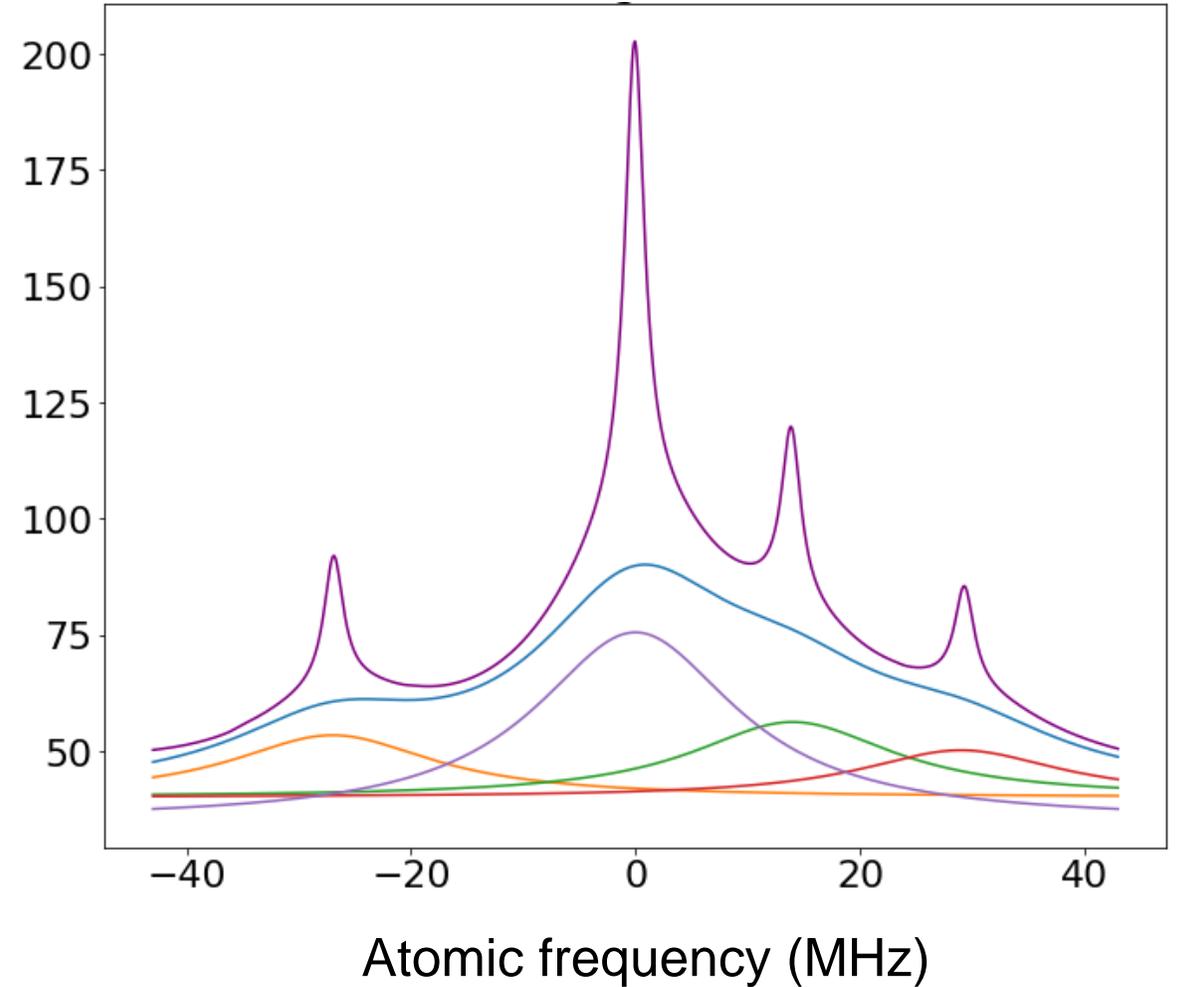
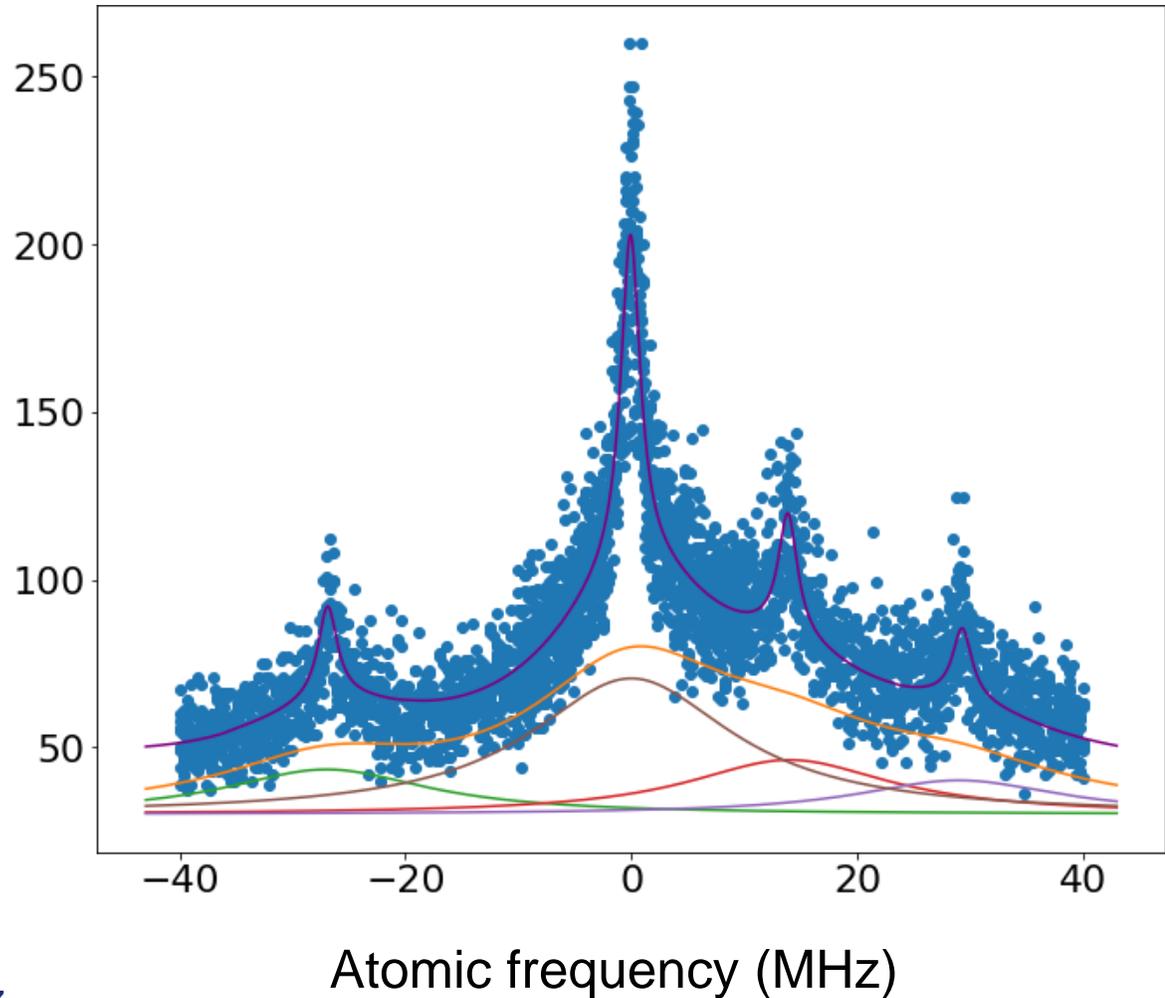


# Issue with the new “low” field data ...

Story more complex at weak B field:  $B \sim 20$  G

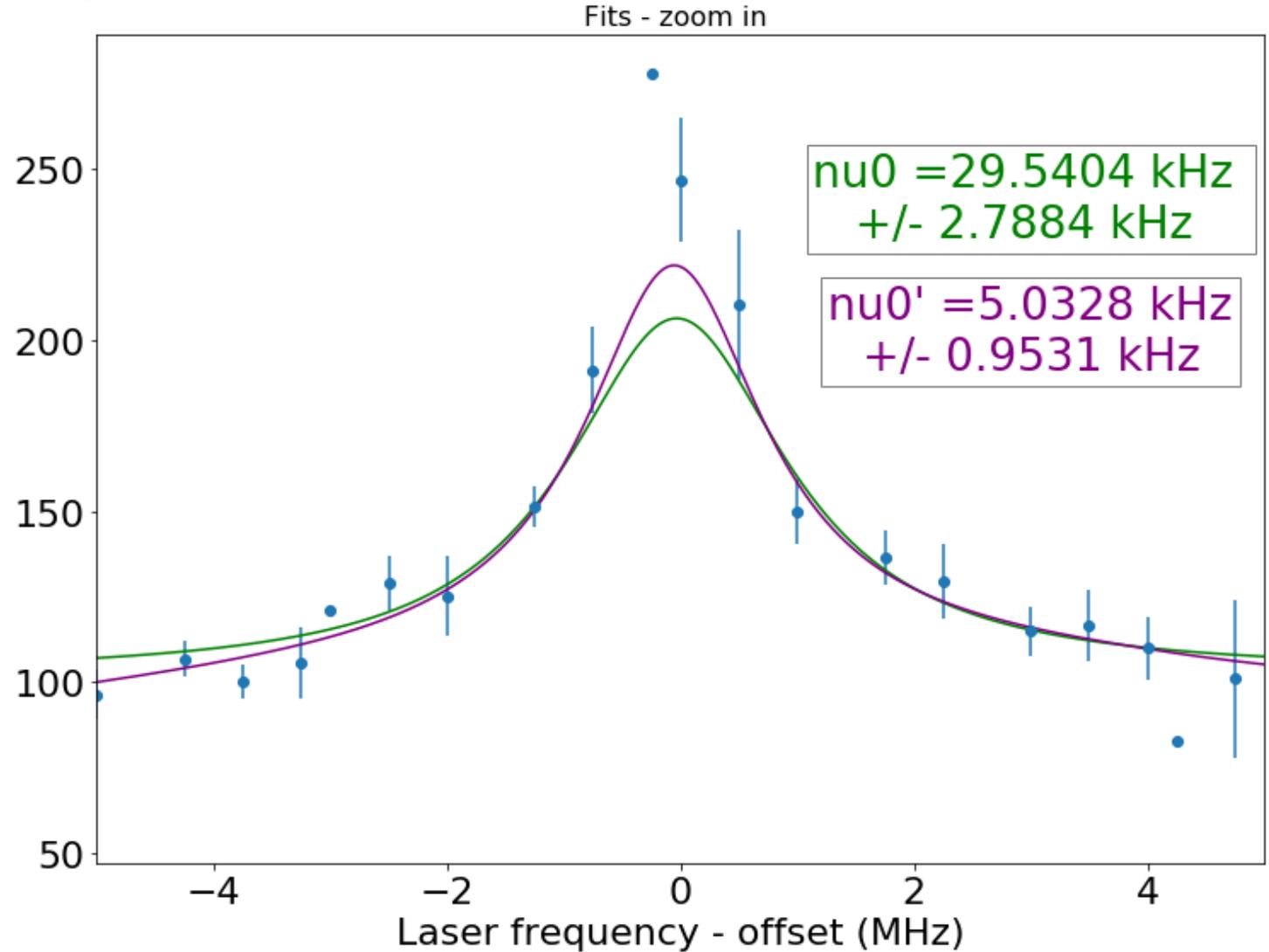
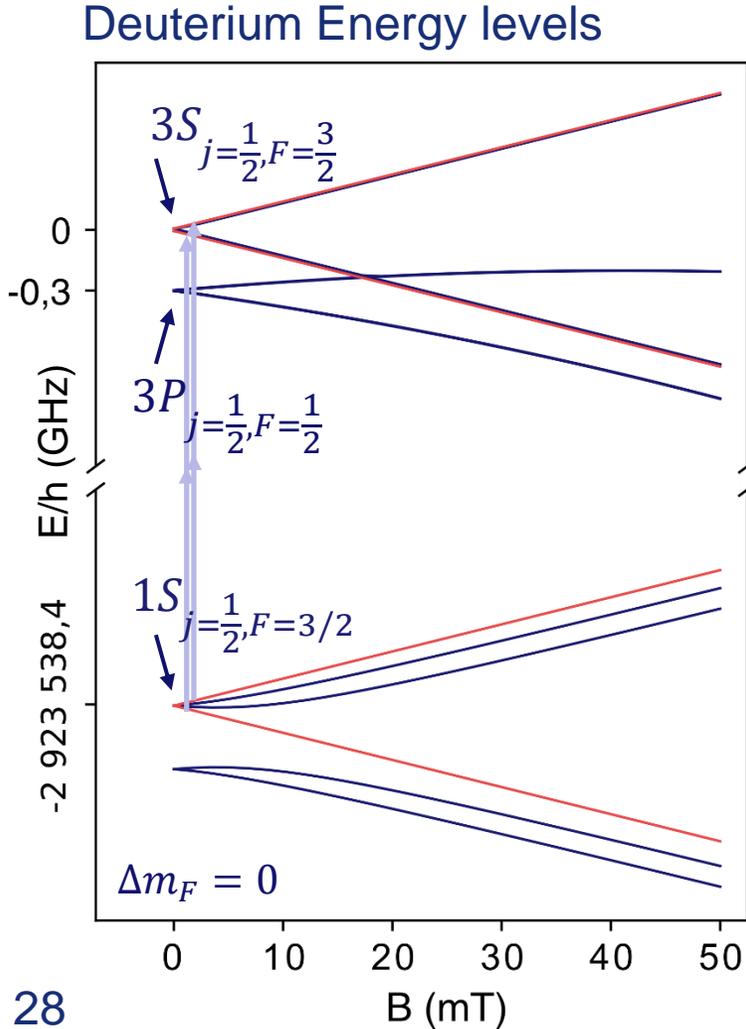


# Asymmetry of the pedestal due to neighbouring lines at $B \sim 20$ G



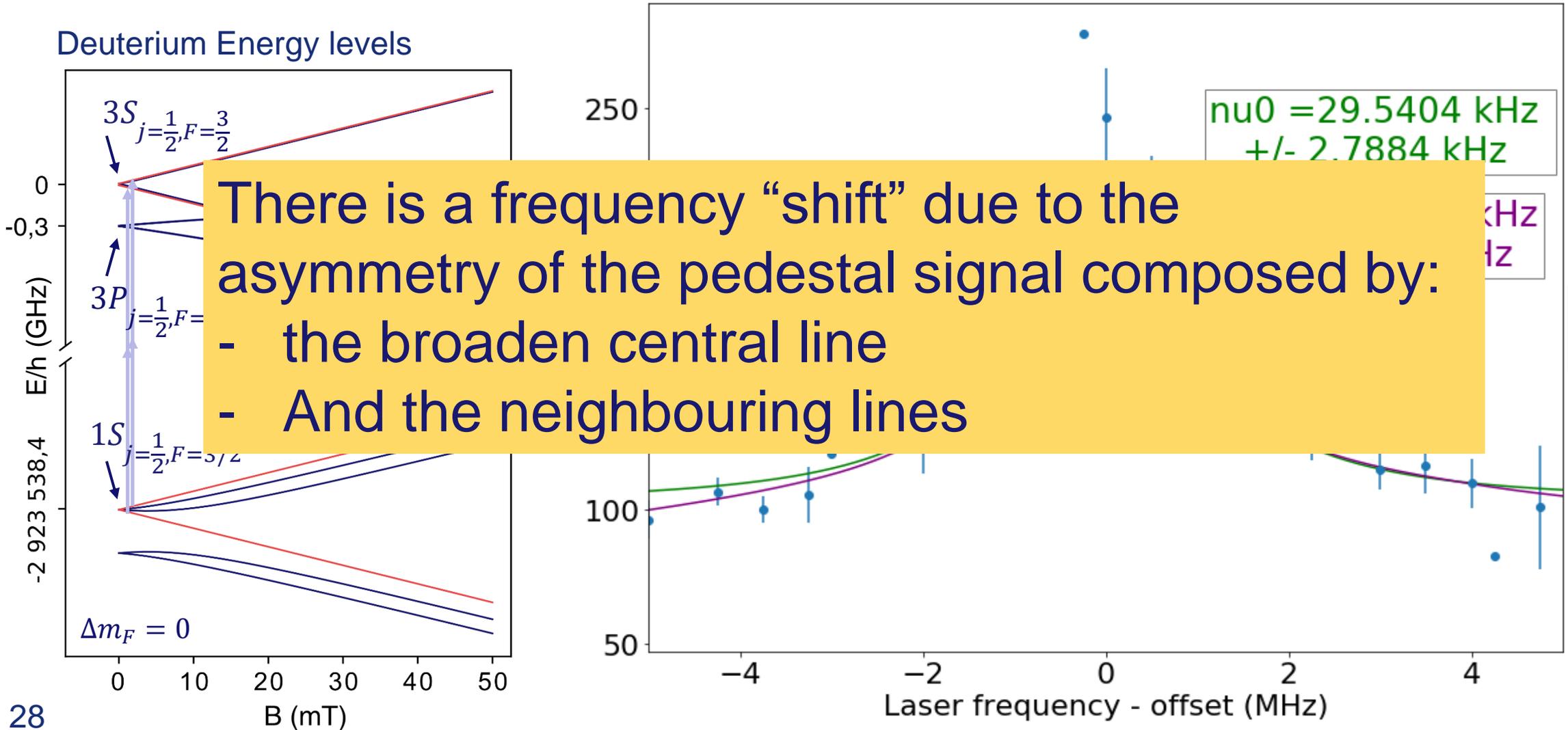
# Issue with the new “low” field data ...

Story more complex at weak B field: like 20 G



# Issue with the “low” field data

Story more complex at weak B field:  $B \sim 20$  G



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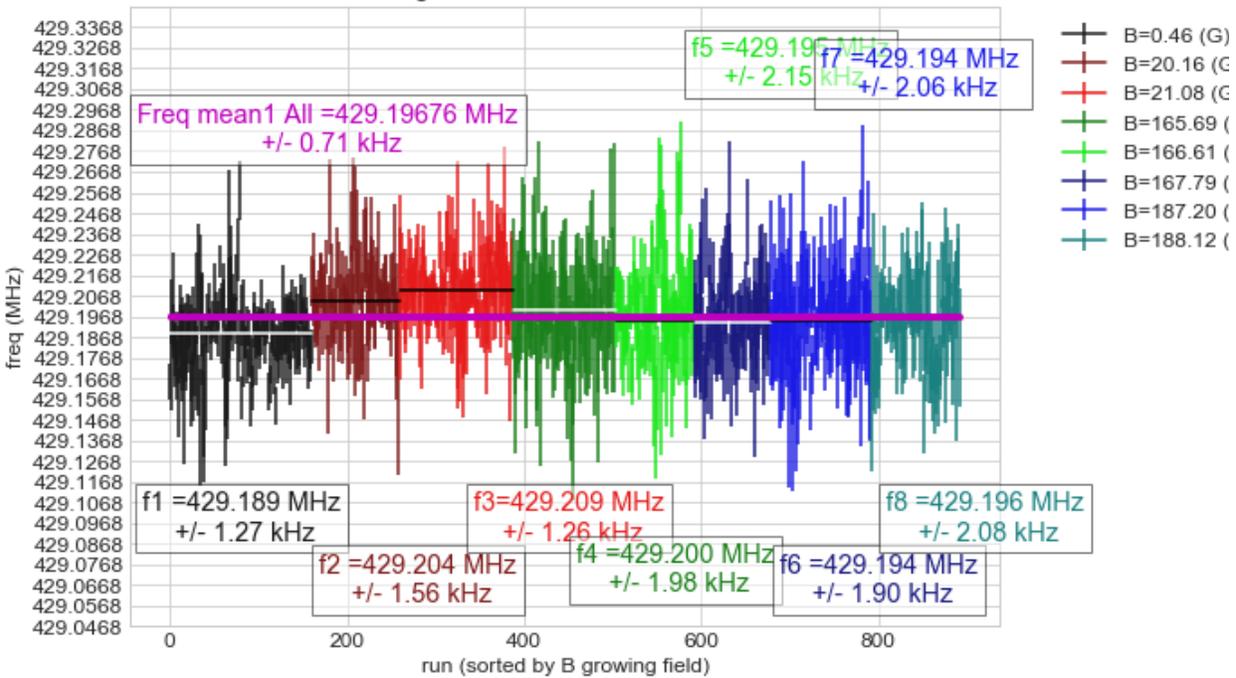


# Data analysis

- From 20 October to 17 December 2020 – B field recorded:  
-0.46 , **+20.16**, **-21.09**, +165.99, -166.91,-167.80, +187.20,  
-188.12 G
- Data recorded over [-3,3] MHz atomic frequency range  $\Rightarrow$  impossible to well determine the pedestal parameters in such a short frequency range (centre, bandwidth, amplitude) for the fits.
- **Issues in the fits for the B=20.16 and B=-21.09 G** because of the neighbouring lines (asymmetrical pedestal signal leads to a wrong positioning of the centre line)

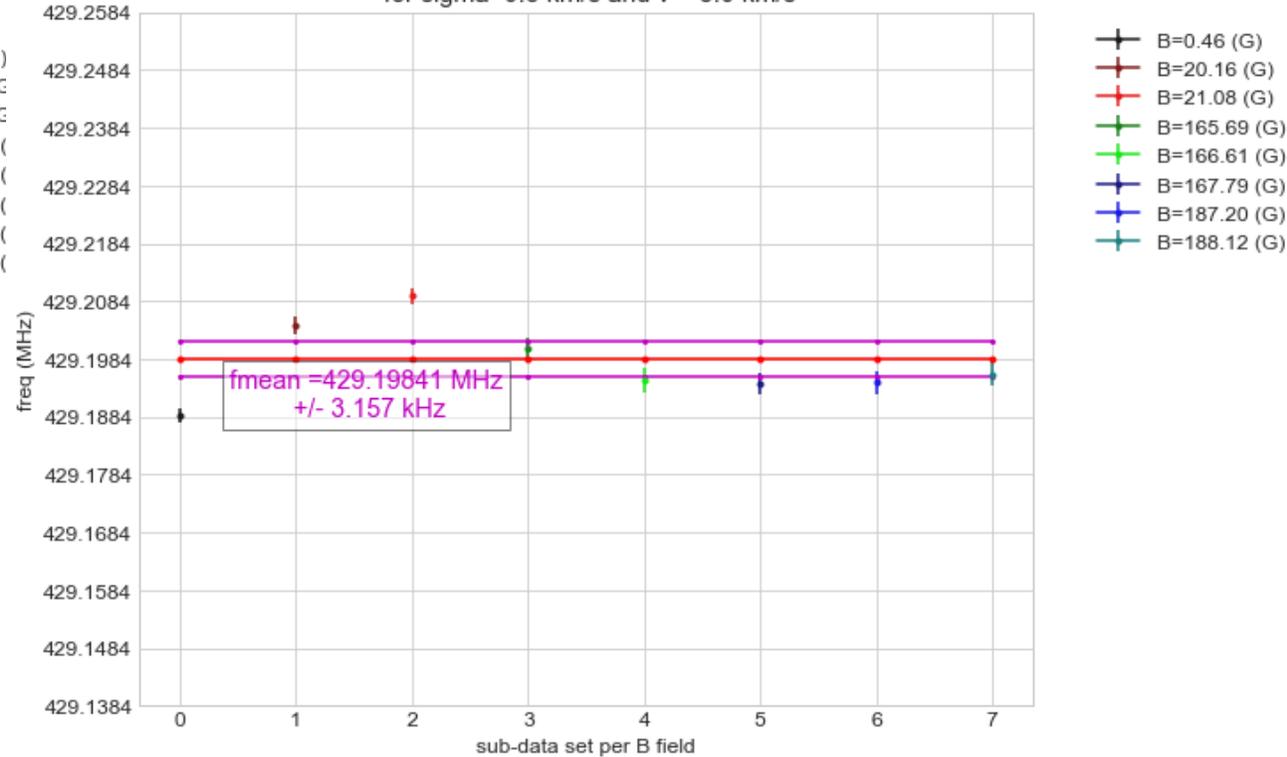
# Example of the issue with the B=20G data

all B: fitted freq for each run  
for sigma=0.8 km/s and v = 5.0 km/s



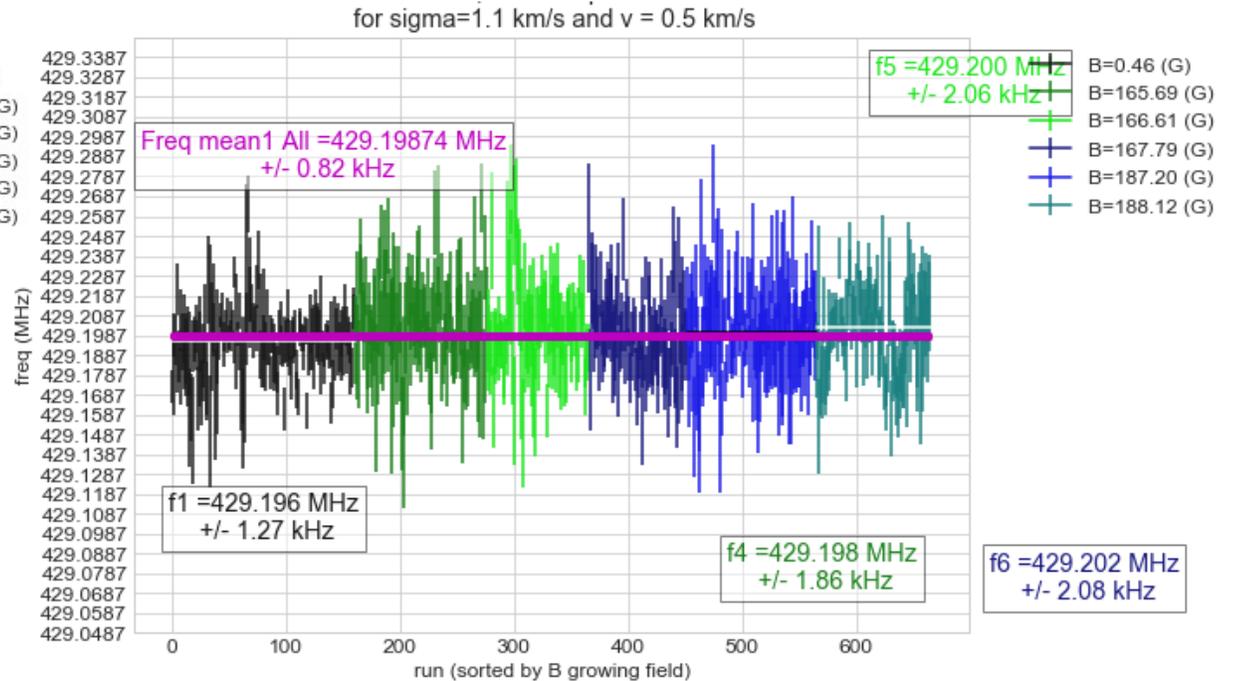
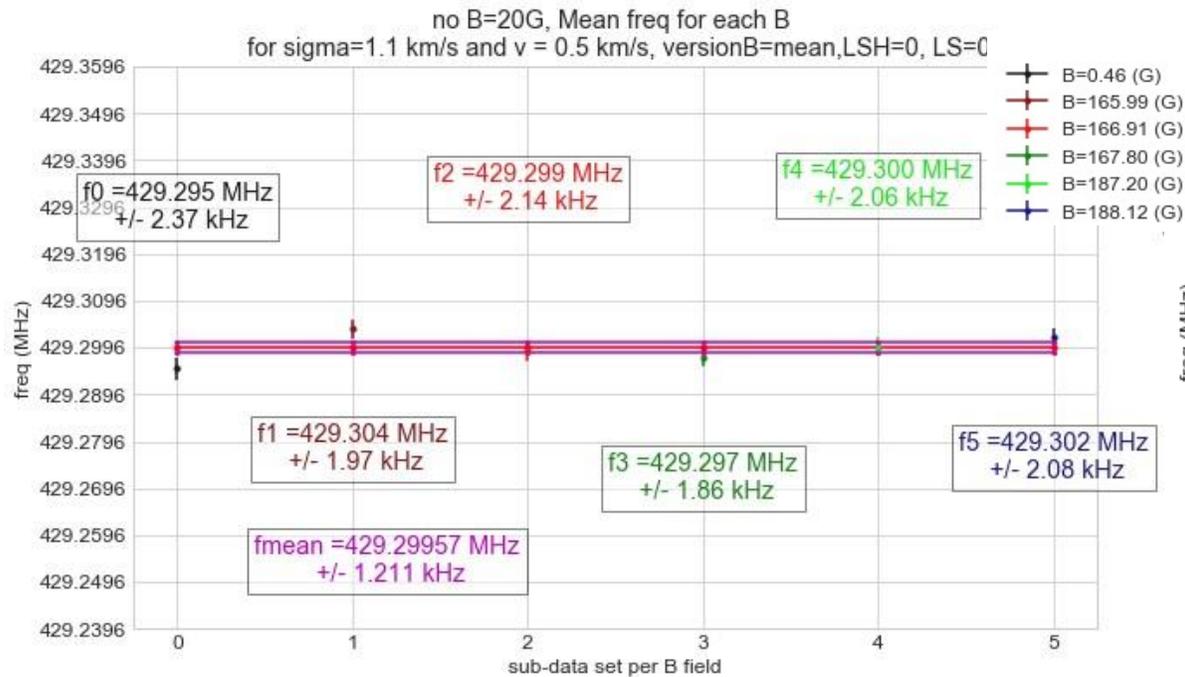
Best velocity distribution (sigma=0.8 km/s, v0=5 km/s)

all B: Mean freq for each B  
for sigma=0.8 km/s and v = 5.0 km/s



Even for the “best parameters” that minimize  $\chi^2$  of the determination of velocity distribution: not a “good” correction of dispersion type profile of the data separated by B fields

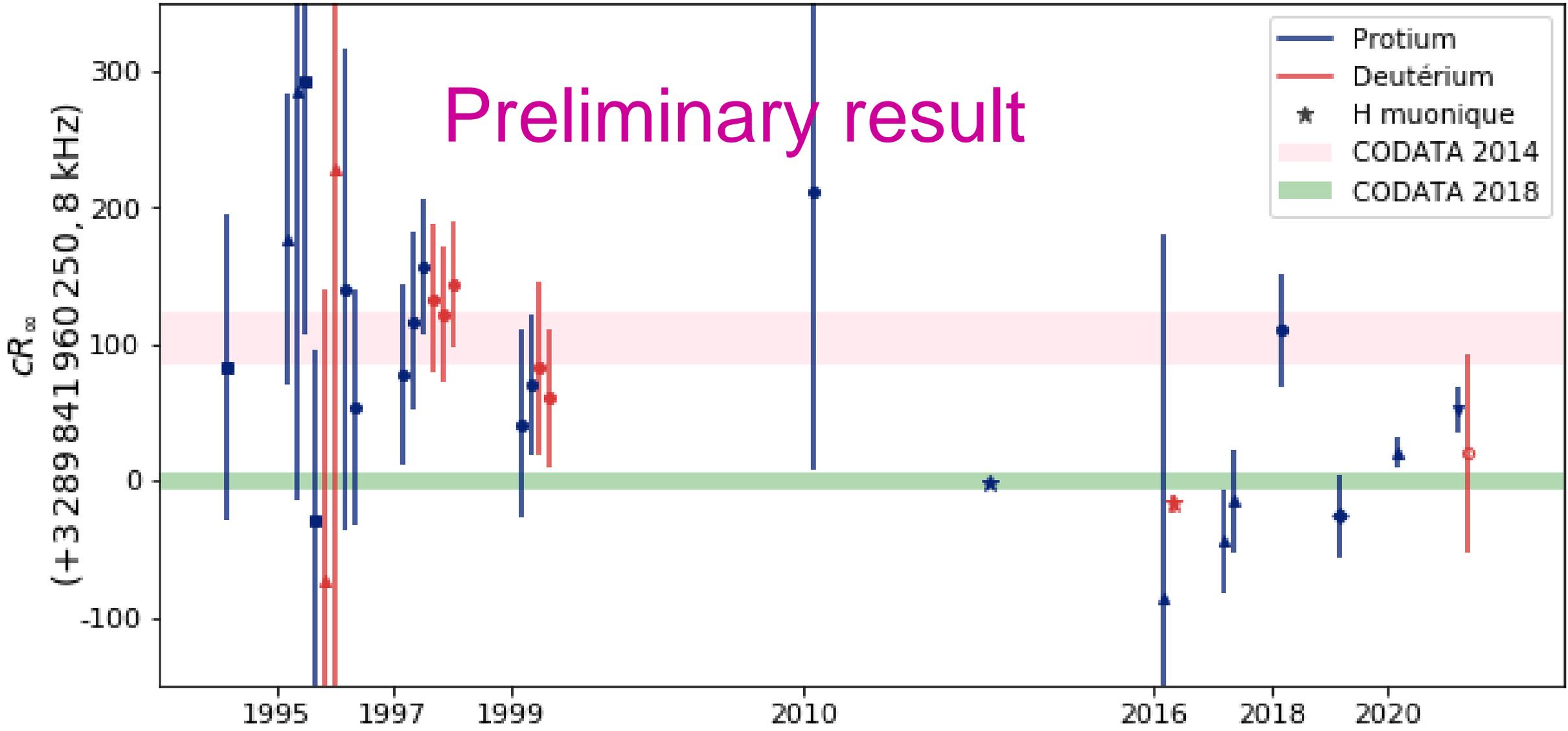
# Whereas, excluding the B=20 G data

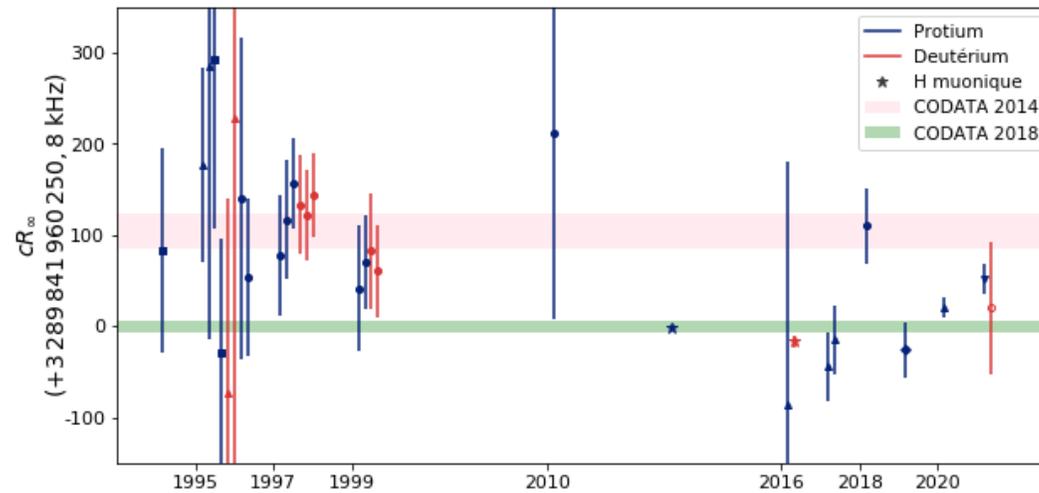
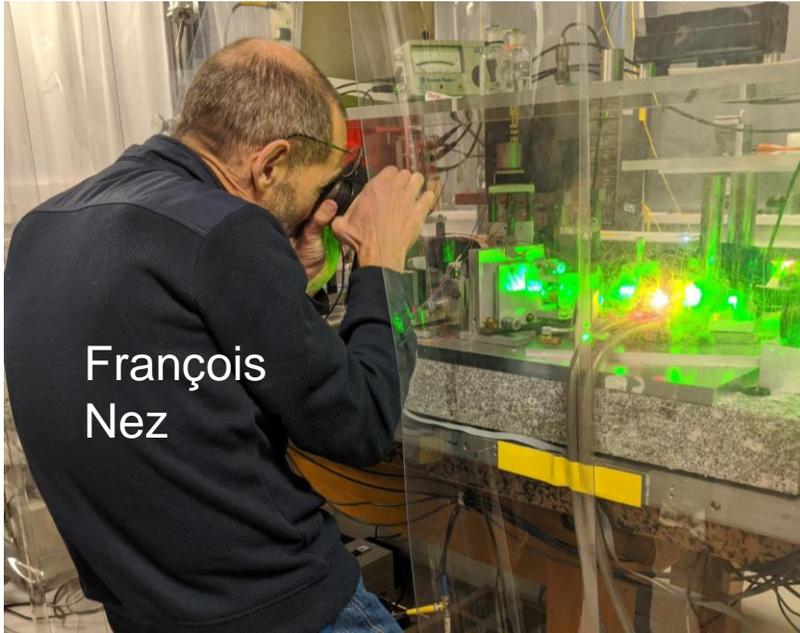


| contributions                                                    | In frequencies                | comments                                                                                                                                                                                            |
|------------------------------------------------------------------|-------------------------------|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| $v_{\text{fit}} (\delta v_{\text{fit}})$<br>1S-3S F=3/2 -> F=3/2 | 2 923 538 429 299.7 (1.5) kHz | Take into account the<br>- the SOD, the Zeeman effect, and the motional Stark shifts<br>- Statistical uncertainty<br>- Uncertainty of the $v_{\text{mean}}(B=0.46\text{G})$ due to 20mG uncertainty |
| $\Delta v_{\text{pressure shift}} (\delta v_{\text{PS}})$        | +2.6 (0.9) kHz                |                                                                                                                                                                                                     |
| $\Delta v_{\text{light shift}} (\delta v_{\text{LS}})$           | -2.5 (3.0) kHz                |                                                                                                                                                                                                     |
| $\Delta v_{\text{quantum interference}} (\delta v_{\text{QI}})$  | +0.6 (0.2) kHz                | H. Fleurbaey, et al <a href="#">Phys. Rev. A 95, 052503 (2017)</a> .                                                                                                                                |
| $\Delta v_{\text{correctionSyrte}}$                              | -0.171( <0.010) kHz           | From Syrte (maser drift in time)                                                                                                                                                                    |
| $\Delta v_{\text{Bump}} (\delta_{\text{Bump}})$                  | 0.0 (3.0) kHz                 | To be investigated                                                                                                                                                                                  |
| Tot<br>1S-3S F=3/2 -> F=3/2                                      | 2 923 538 429 300.1 (4.5) kHz | Not corrected of the HFS                                                                                                                                                                            |

Preliminary result

Preliminary result

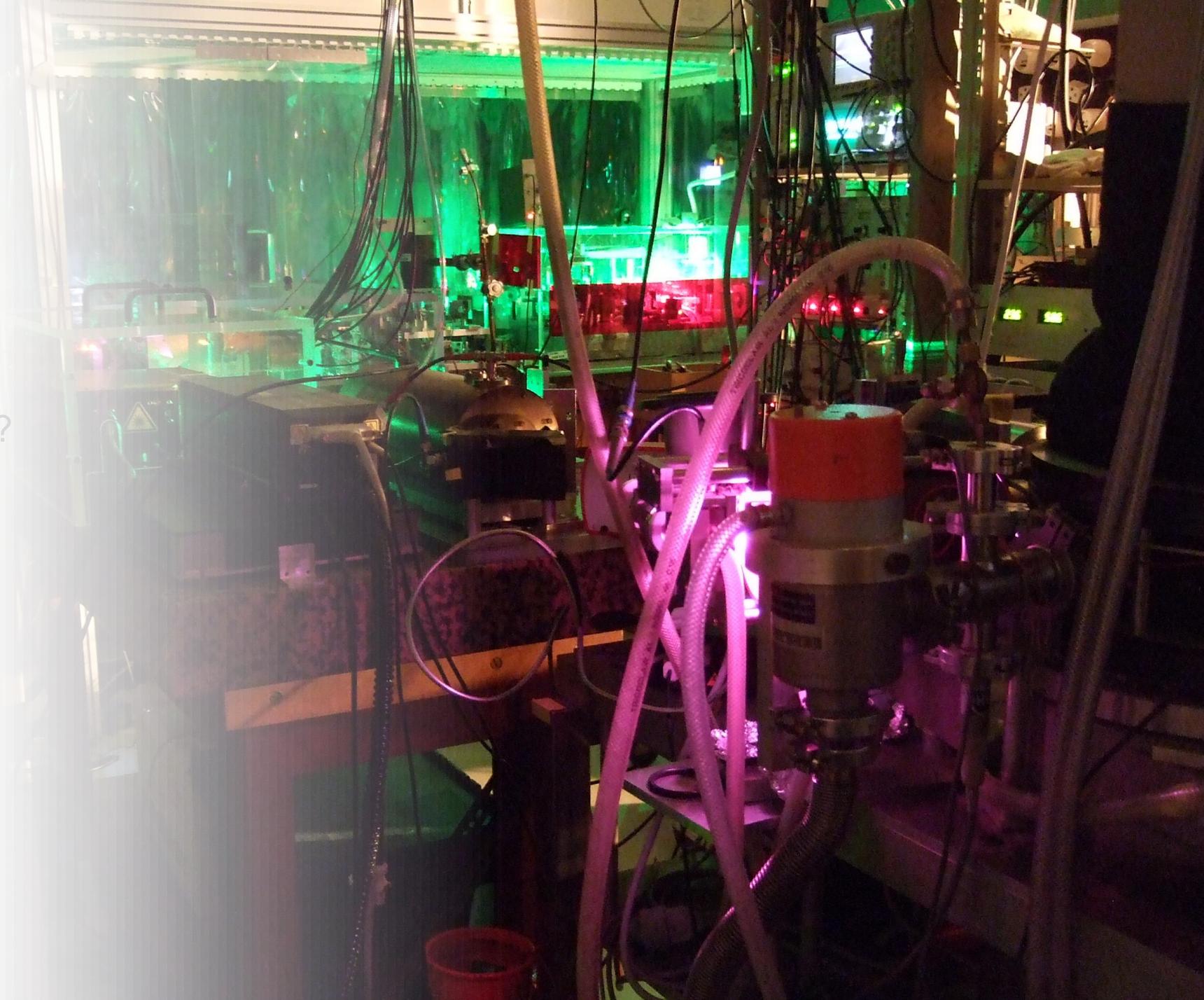




Thanks for your attention

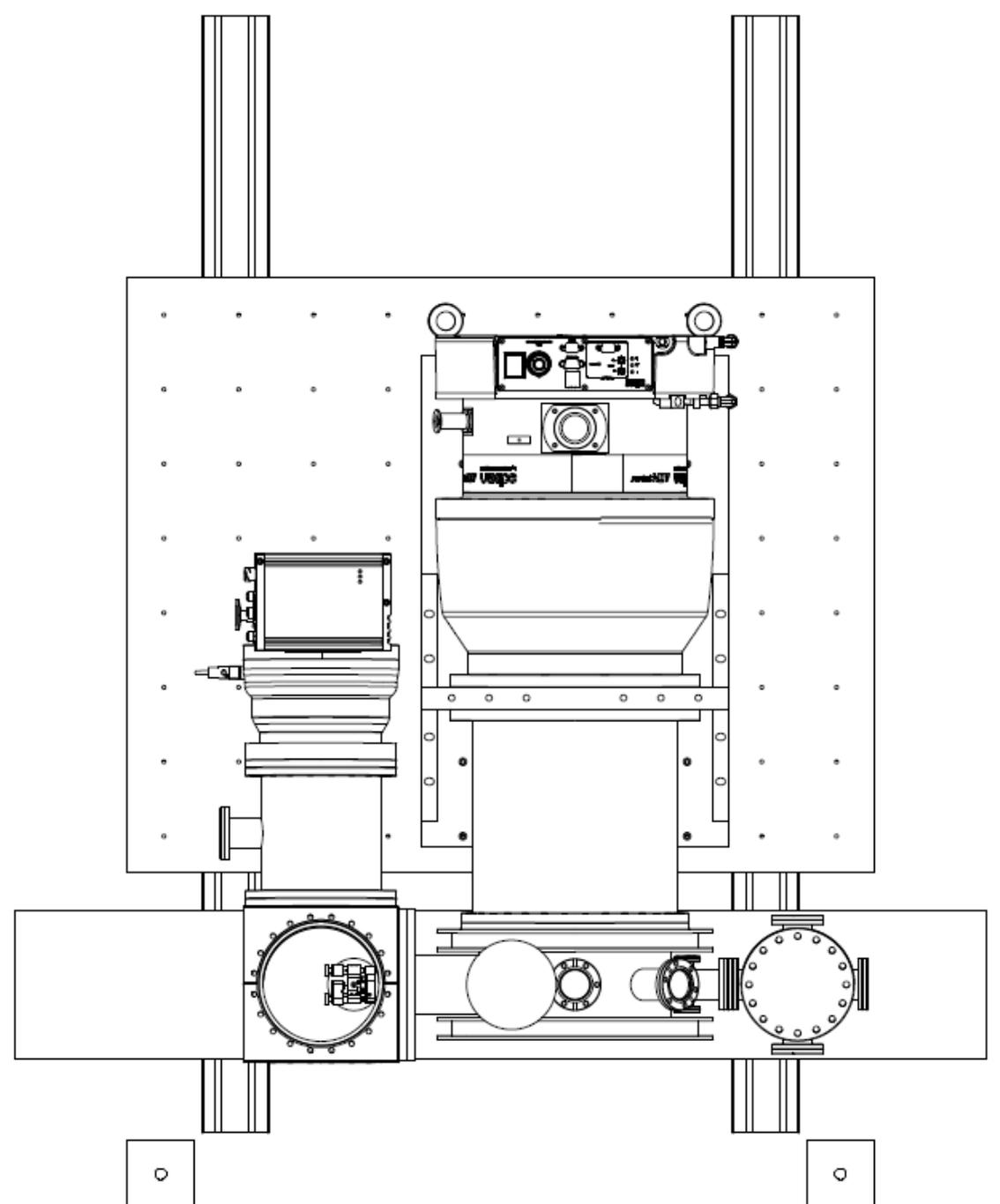
# Overview

- The experiment
- Dealing with systematics
- A new systematics effect ?
- Preliminary result ?
- What next ?



# The new H beam experiment

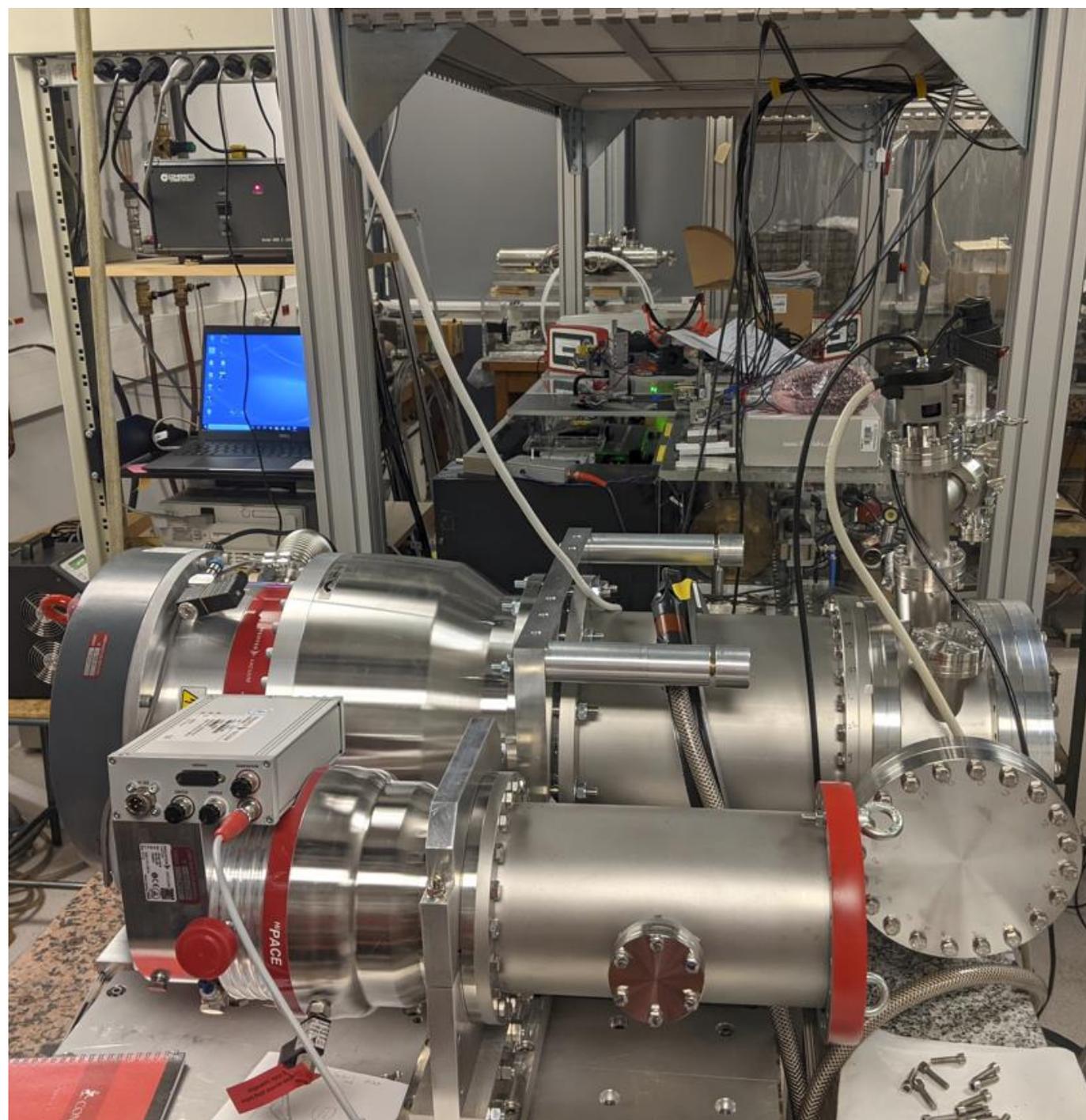
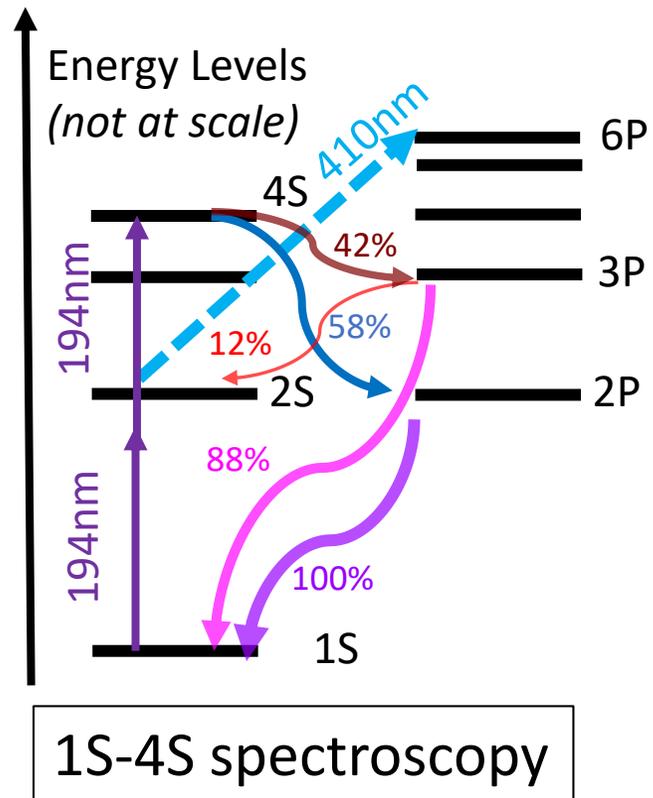
- Based on dry pumps ! ✓ .
- New design to decoupled the mirrors' cavity of the 205 nm from the vibrations of the turbo pumps ✓ .
- Better power supplies for the coils and compensating Earth-B field coils ✓ .
- Wider laser frequencies scans to investigate better the pedestal signal ✓ .
- Better laser stability ✓ .
- New design of the



Vue isométrique  
Echelle : 1:3

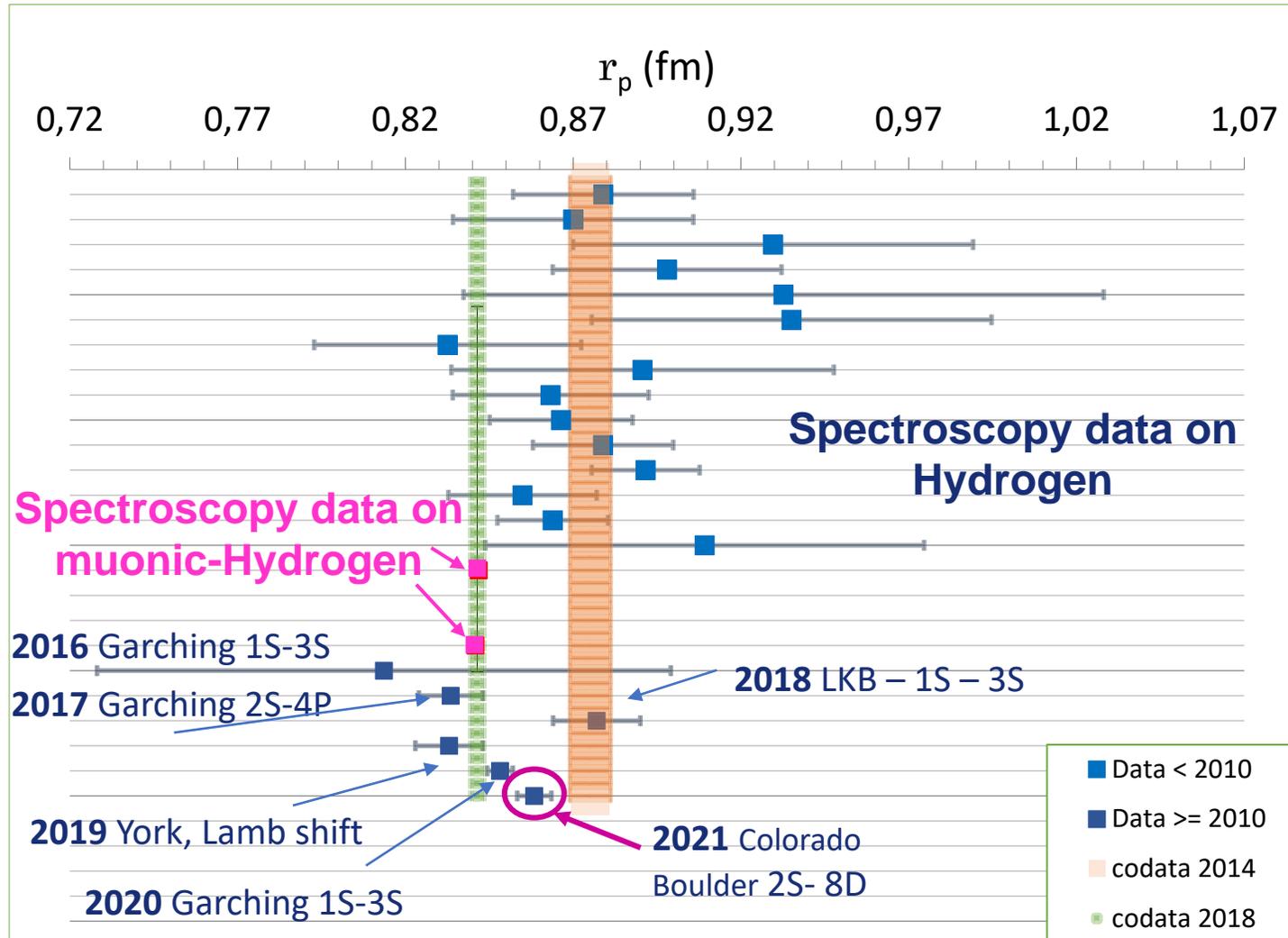
# The new H beam experiment

- For 1S-3S then 1S-4S

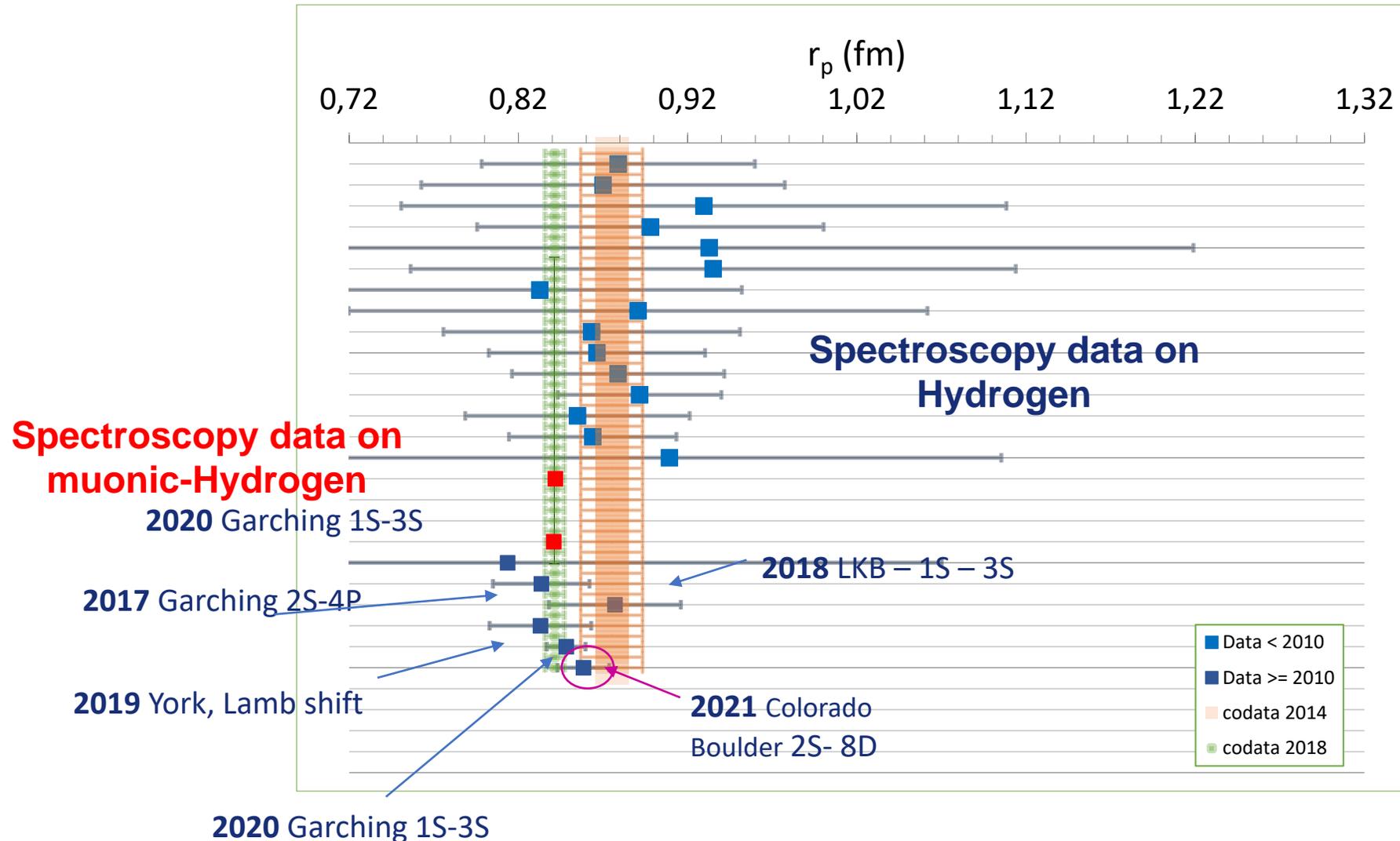


Back up

# The proton radius puzzle ? 1 sigma

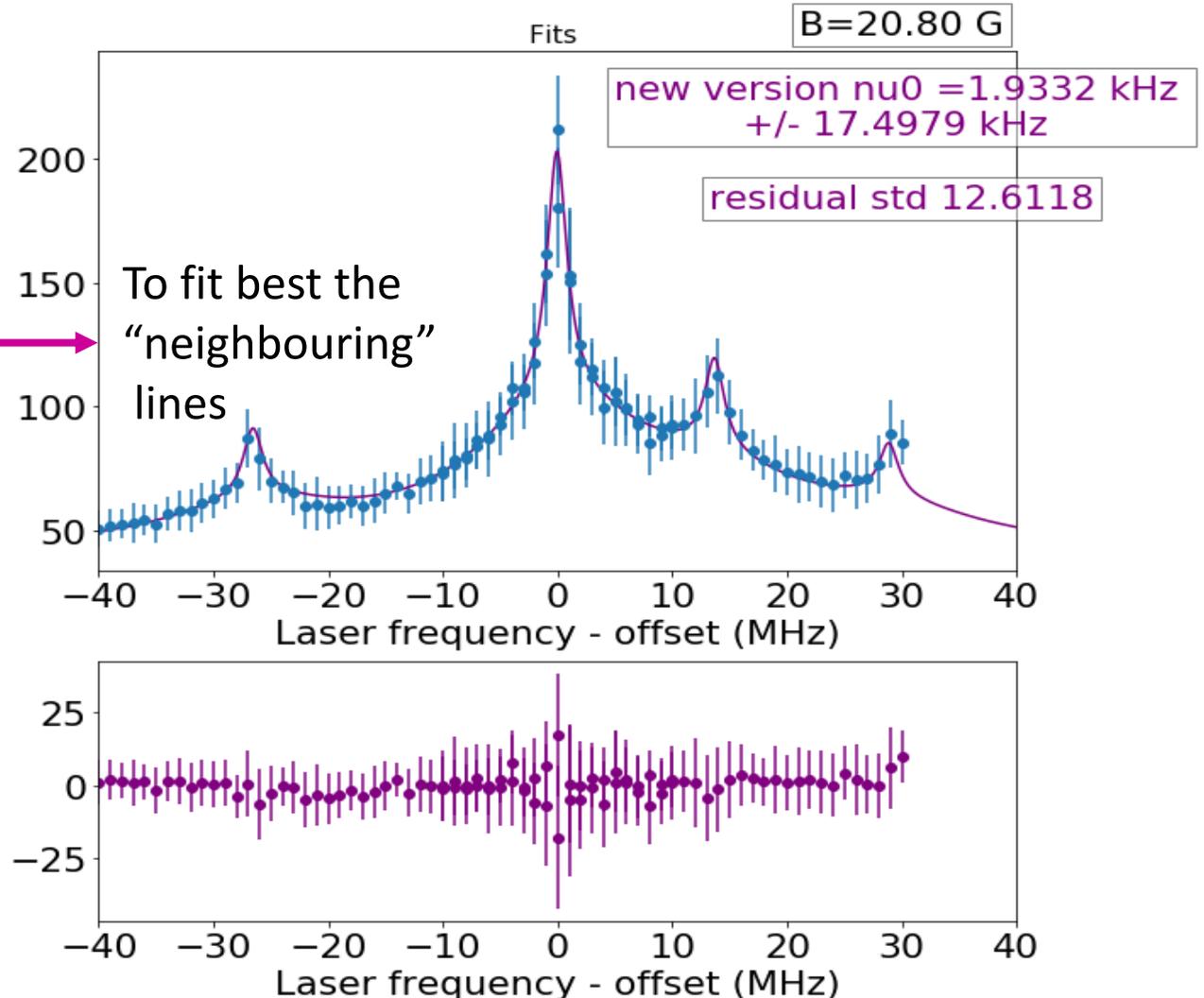
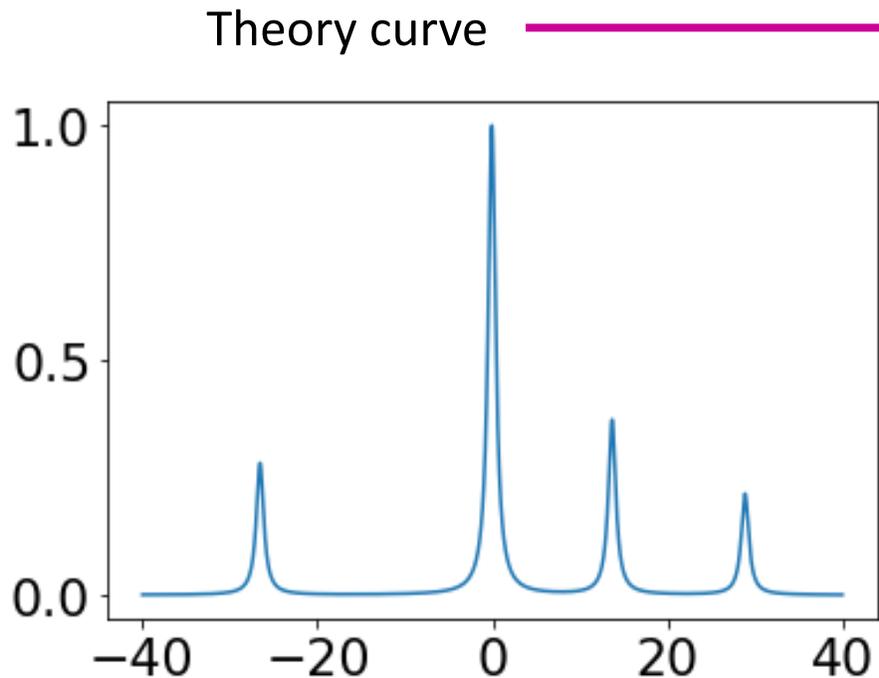


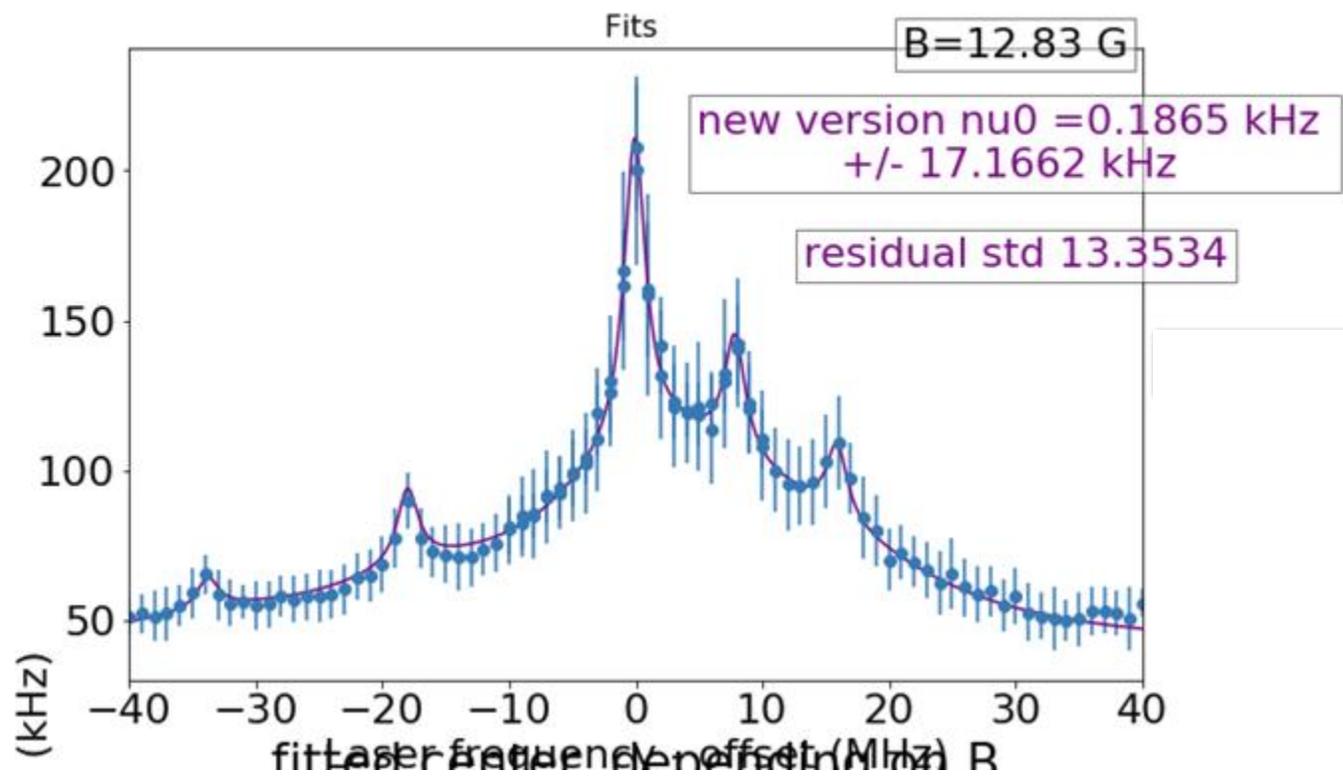
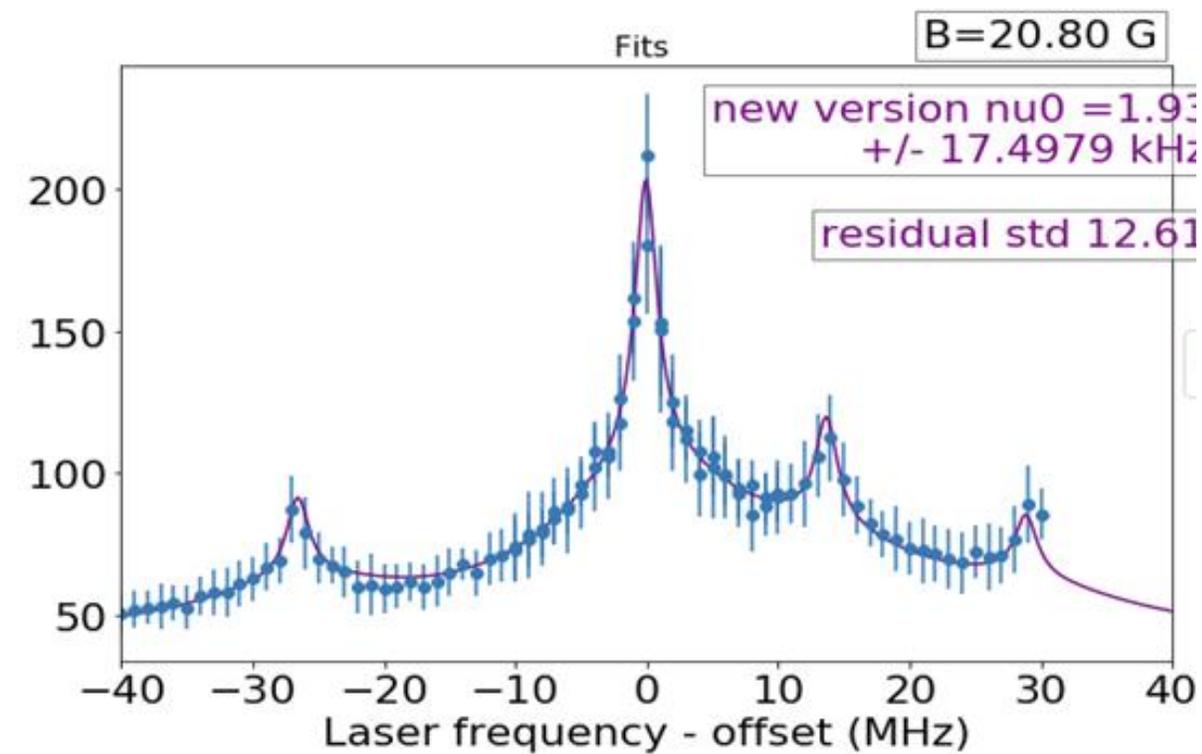
# The proton radius puzzle ? 3 sigmas



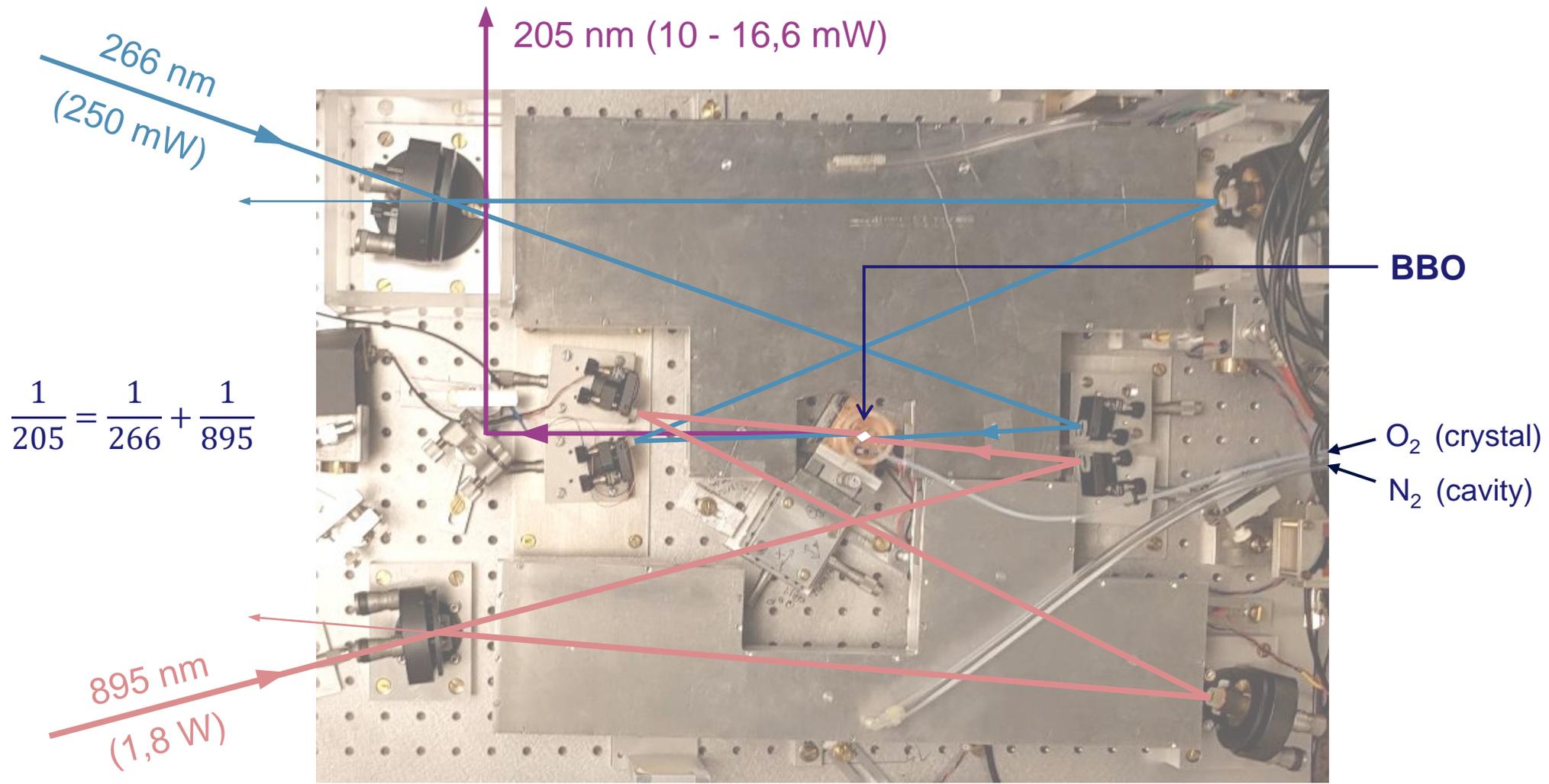
# Calibration of B fields – 2 methods

- 1) Over one Zeeman shifted line
- 2) With neighbouring lines at low B field

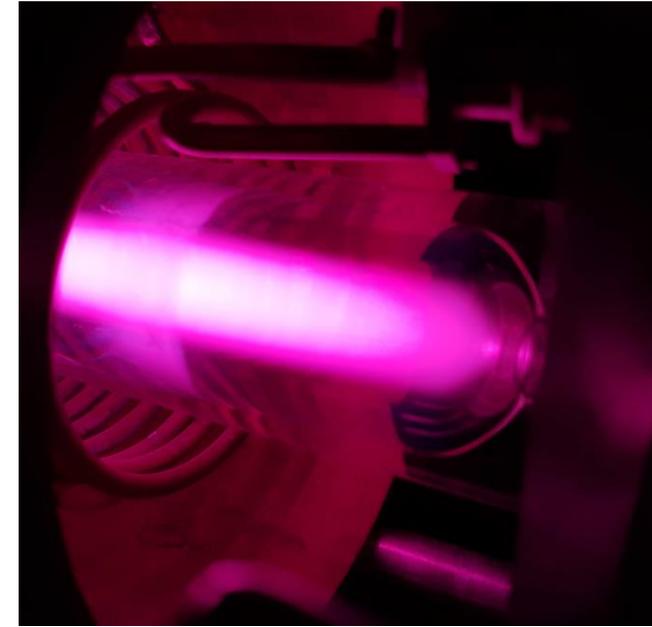
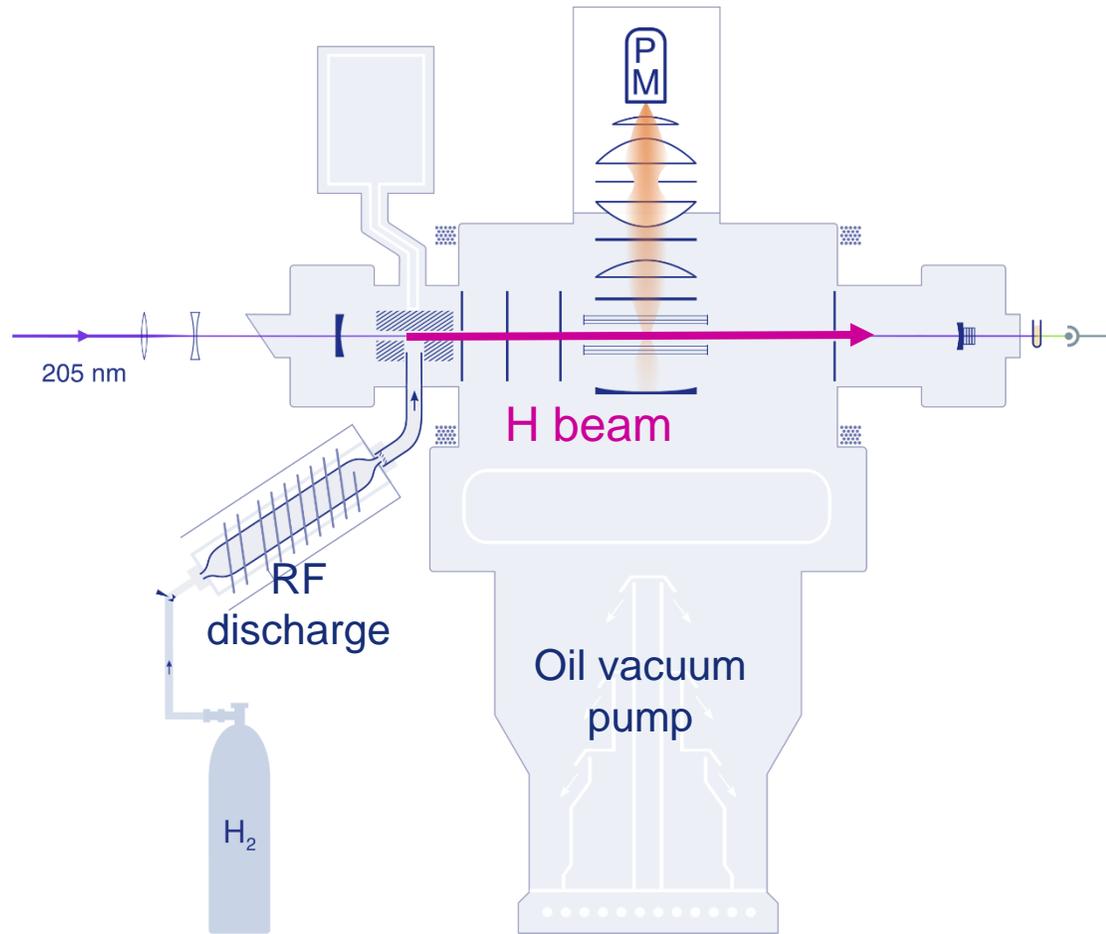




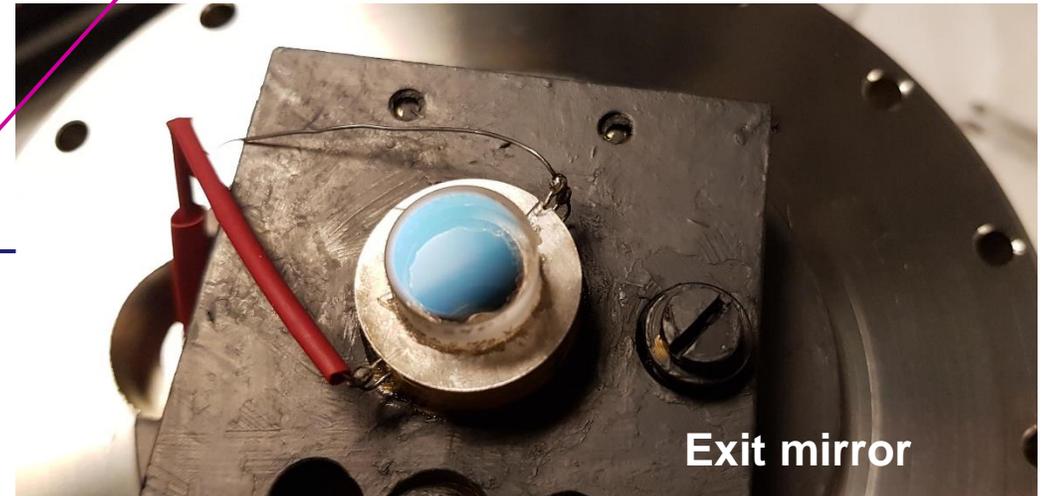
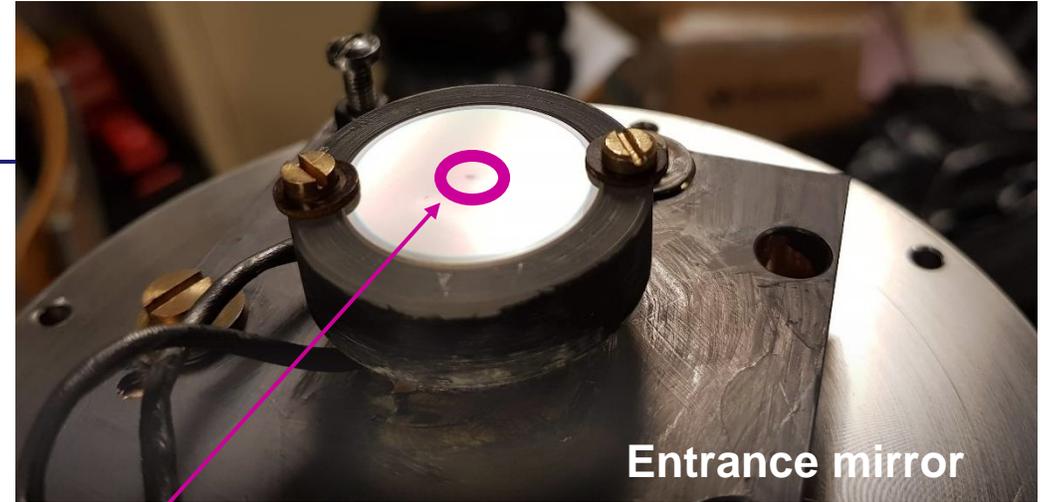
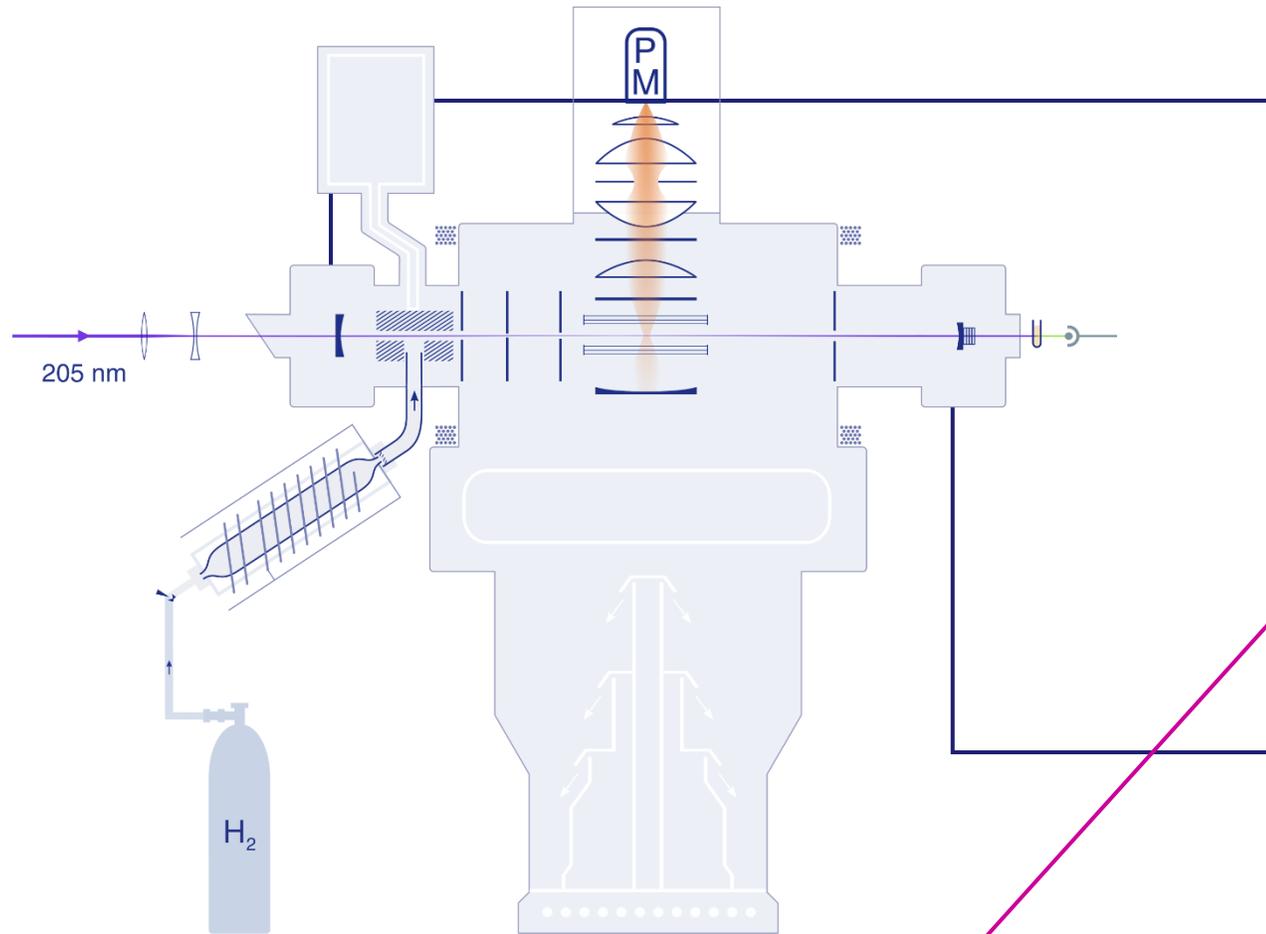
# CW 205nm laser generation



# Schematic of the «old » H beam experiment



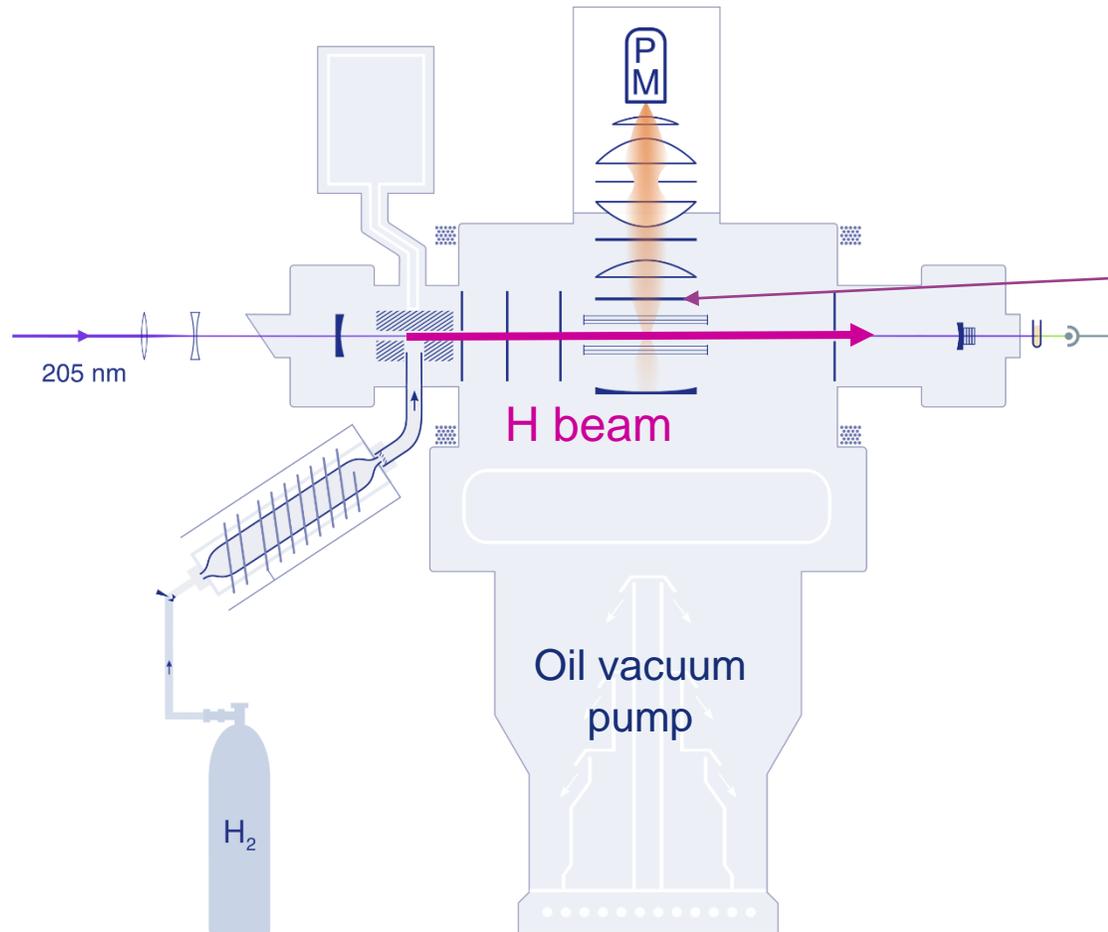
# Schematics of the «old » H beam experiment



Pollution of the mirrors  $\Rightarrow$  breaking vacuum every 2 days to clean

# Schematic of the «old » H beam experiment

B= 0.46G

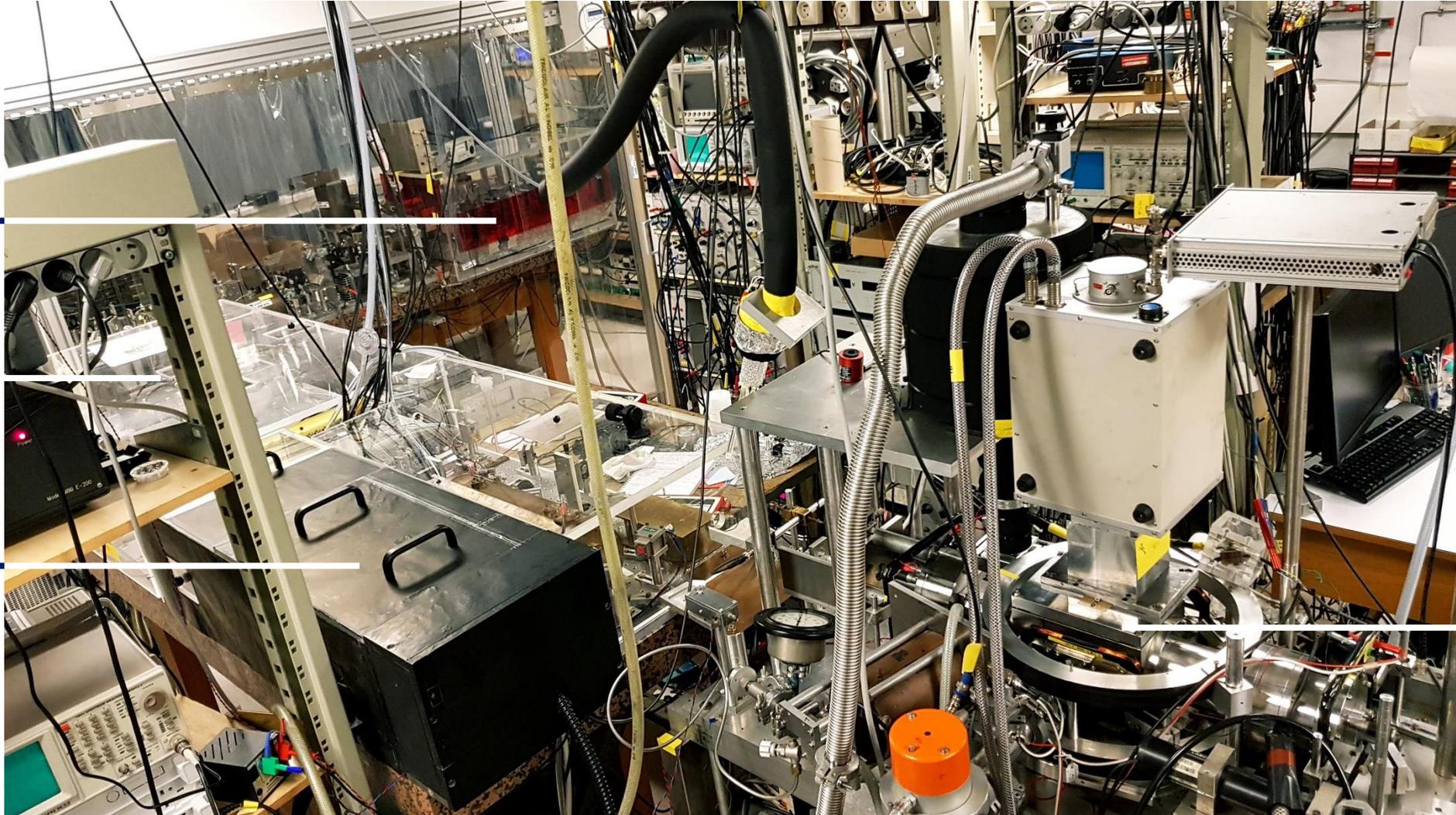


# With some light ON

TiSa  
895nm

205nm  
generation

Verdi  
266nm



(old) H  
beam

## Velocity distribution of the D beam

