

# Towards High-Precision Spectroscopy of the 1S–2S Transition in He<sup>+</sup>

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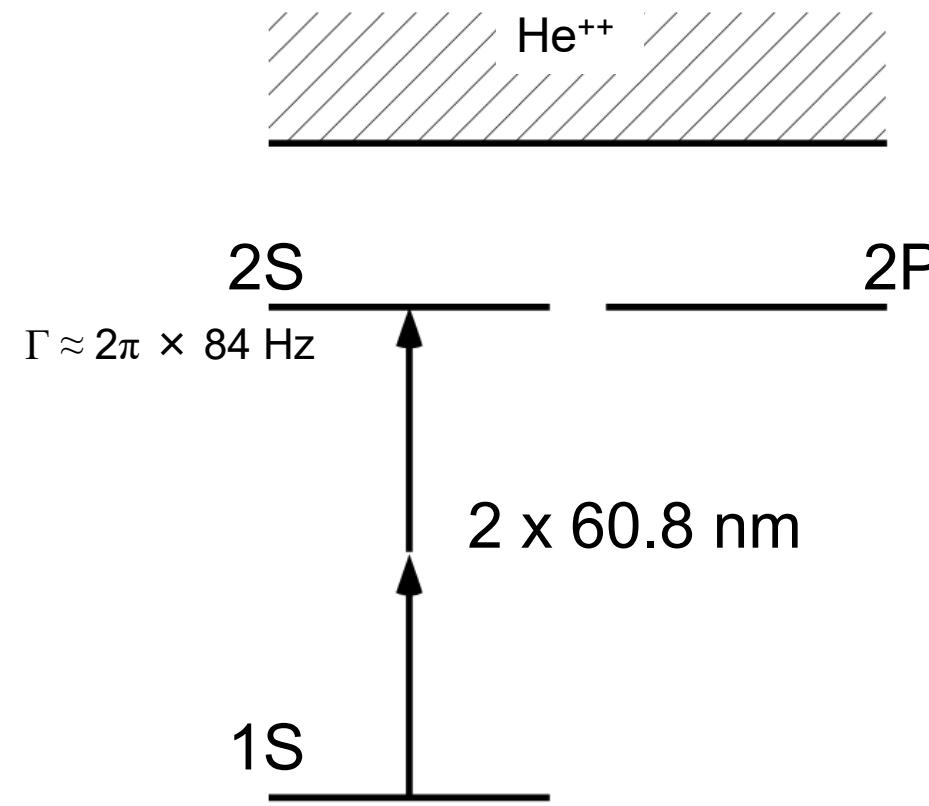
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MAX PLANCK INSTITUTE  
OF QUANTUM OPTICS

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# Spectroscopy of the 1S–2S Transition in He<sup>+</sup>



Difference from Hydrogen spectroscopy:

- Trapped and cooled target → smaller systematics
- QED test for Z=2 system
- Frequency comb spectroscopy in XUV (extreme ultraviolet) wavelength

# QED for $\text{He}^+$

$$E_{nlj} = R_\infty f_{nlj} \left( \alpha, \frac{m_N}{m_e}, \frac{r}{\lambda_C} \right)$$

Fine structure constant

Rydberg constant

State energy in SI units ( $\text{m}^{-1}$ )

Nucleus/electron mass ratio

Normalized nuclear charge radius

Parameters	Contribution to Uncertainty ( $\text{He}^+ 1S2S$ )
Rydberg constant $R_\infty$	$2 \times 10^{-12}$ (19 kHz)
Nuclear charge radius $\frac{r}{\lambda_C}$	$6 \times 10^{-12}$ (61 kHz)
Nucleus/electron mass $\frac{m_N}{m_e}$	$5 \times 10^{-15}$ (45 Hz)
Fine structure constant $\alpha$	$2 \times 10^{-14}$ (166 Hz)

Hydrogen spectroscopy  
 $\mu\text{-H} + \text{H}$  (1s-2s), CODATA2018

$\mu\text{-He}$  spectroscopy  
J. Krauth et al. Nature (2021)

Mass ratio measurements  
in Penning trap

Atomic recoil (atom interferometer)  
g-2 measurement

45% larger uncertainty  
In Pachucki et al.  
arXiv (2023)

# Uncertainty due to QED theory for He<sup>+</sup> 1S-2S transition

[1] Yerokhin. A and Shabaev V. M , Phys. Rev. A 93, 062514 (2016).

[2] Karshenboim, S. G et al. , Physics Letters B 795, 432–437 (2019)

[3] Karshenboim, S. G et al. , Phys. Rev. A 100, 032515 (2019).

[4] Karshenboim, S. G et al. , Phys. Rev. A 98, 022522 (2018)

[5] V. Yerokhin et al. , Annalen der Physik 531, 1800324 (2019).

[6] A.Czarnecki et al. , Phys. Rev. A 94, 060501 (2016).

[7] S. Laporta, Physics Letters B 800, 135137 (2020).

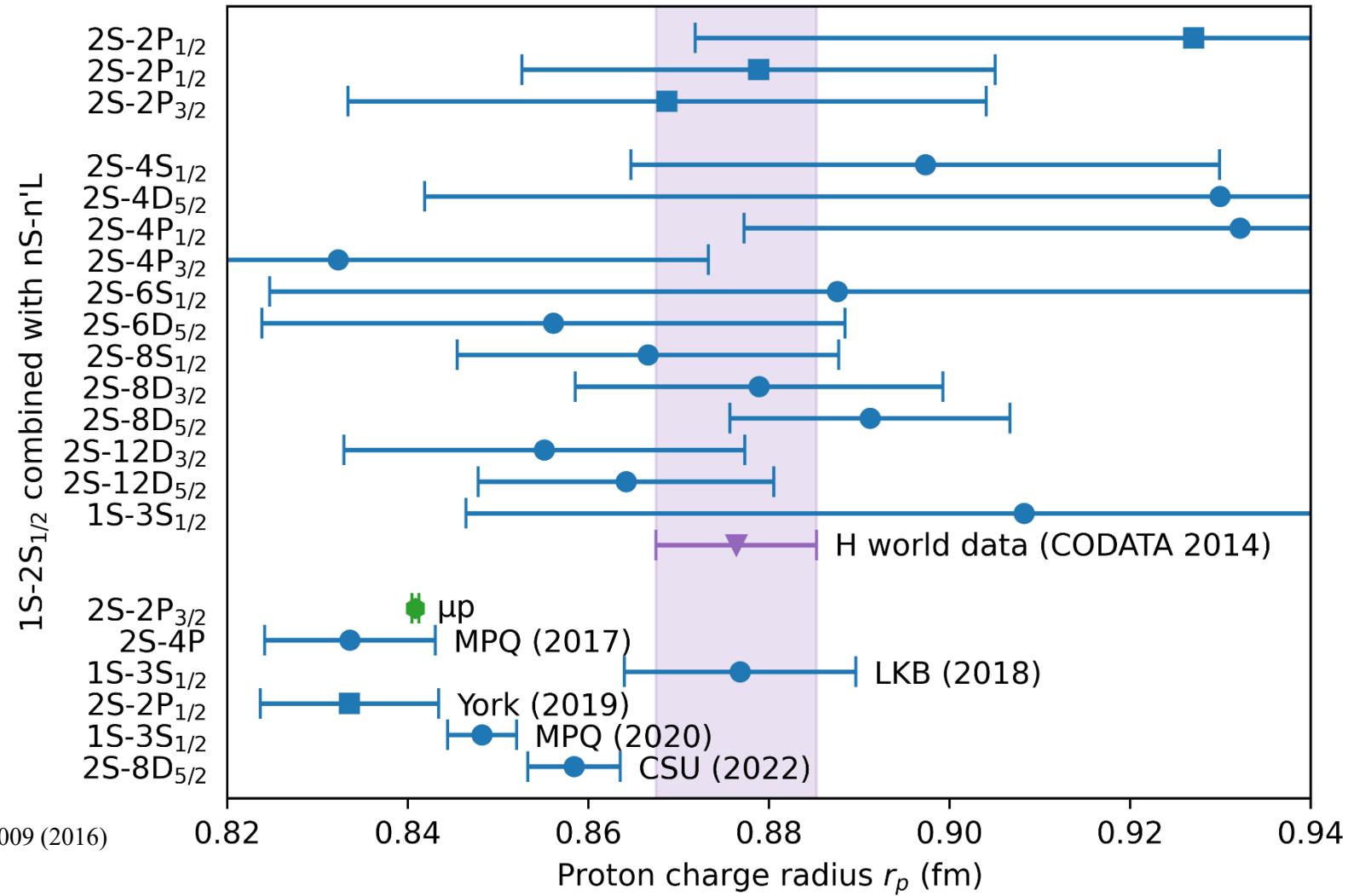
[8] Karshenboim, S. G and Shelyuto V. A. , Phys. Rev. A 100, 032513 (2019)

Contribution	Uncertainty (in kHz)	
Relativistic Recoil	0	[1]
Nuclear Polarizability	3	
Self Energy	0	
Vacuum Polarization	1	
Two Photon Corrections	32	[2-6]
Three Photon Corrections	9.5	[2,3]
Four Photon Corrections	0	[7]
Nuclear Size Correction to Self Energy	0	
Nuclear Size Correction to Vacuum Polarization	0	
Radiative Recoil	12	[8]
Nuclear Self Energy	0.6	
<b>Total</b>	<b>~36 kHz</b>	

Total QED uncertainty : 36 kHz (4x10<sup>-12</sup>) (preliminary)

# Proton radius puzzle

Known to  $4 \times 10^{-15}$



P. Mohr *et al.*, Rev. Mod. Phys. **88**, 035009 (2016)

R. Pohl *et al.*, Nature **466**, 213 (2010)

A. Beyer *et al.*, Science **358**, 79 (2017)

H. Fleurbaey *et al.*, Phys. Rev. Lett. **120**, 183001 (2018)

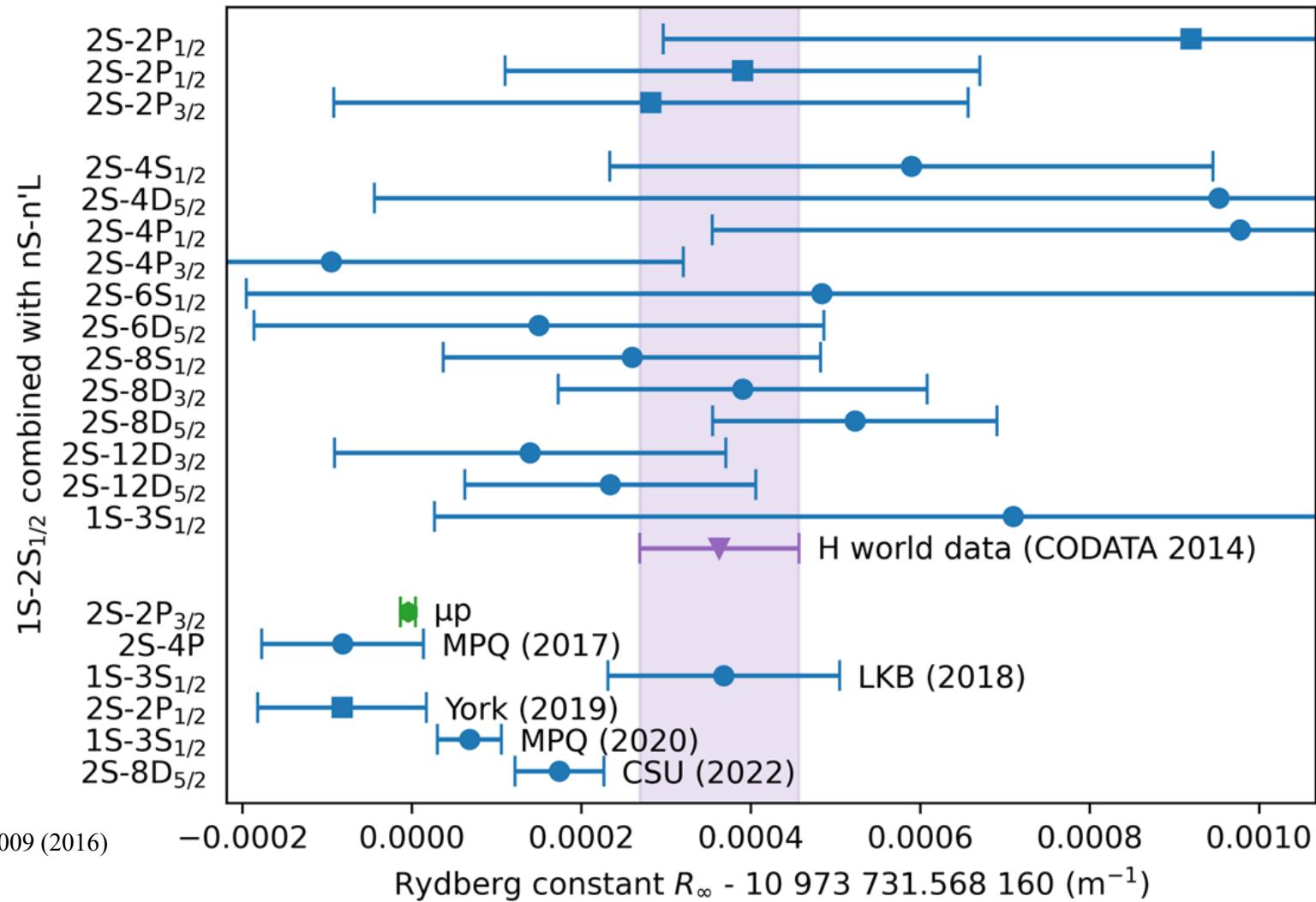
N. Bezuginov *et al.*, Science **365**, 1007 (2019)

A. Grinin *et al.*, Science **370**, 1061 (2020)

A. D. Brandt *et al.*, Phys. Rev. Lett. **128**, 023001 (2022)

# Rydberg constant puzzle

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# $\text{He}^+$ spectroscopy for testing QED

- Measuring the 1S-2S transition

better than 400 kHz ( $4 \times 10^{-11}$ )

————→ Determine  $R_\infty$  and contribute to the proton charge radius puzzle

better than 60 kHz ( $6 \times 10^{-12}$ ) [90 kHz from Pachucki et al. arXiv (2023)]

————→ Determine charge radius, which can be compared with  $\mu\text{-He}$  measurement

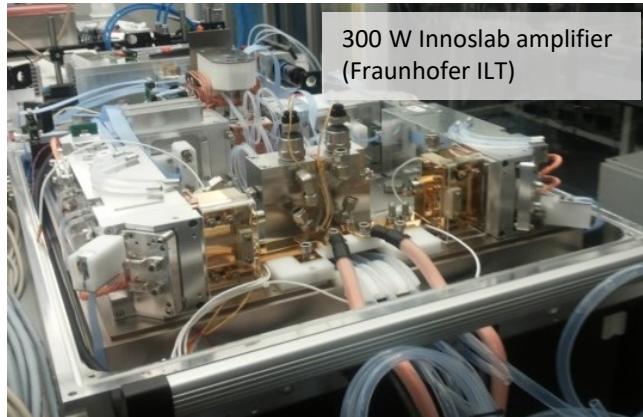
better than 36 kHz ( $4 \times 10^{-12}$ )

————→ QED theory uncertainty will be a limitation

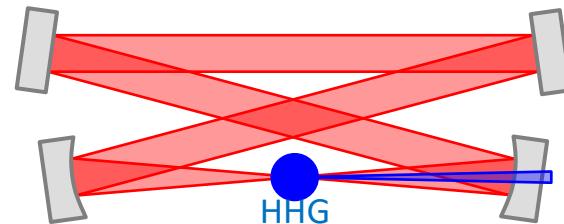
- Measuring other transitions in  ${}^4\text{He}^+$  —————→ QED test more independent from  $\mu\text{-He}$  and Hydrogen
- Measuring  ${}^3\text{He}^+$  —————→ Charge radii difference, which can be compared with neutral He spectroscopy

# $\text{He}^+$ spectroscopy setup: overview

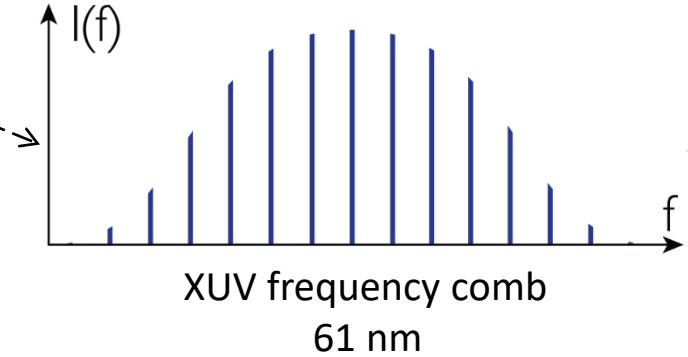
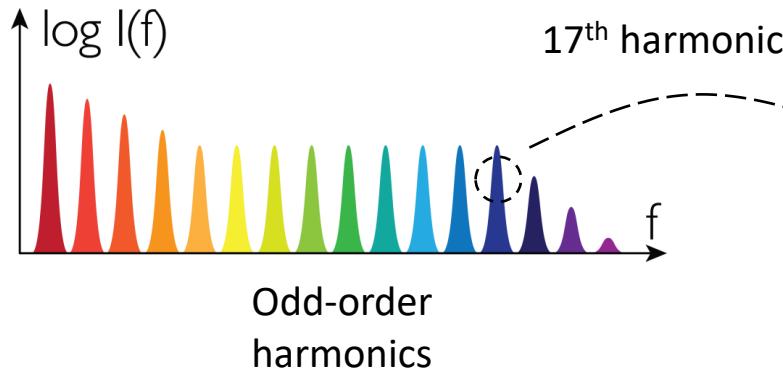
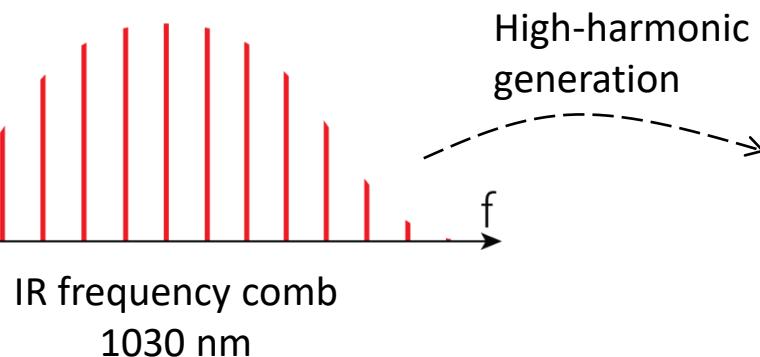
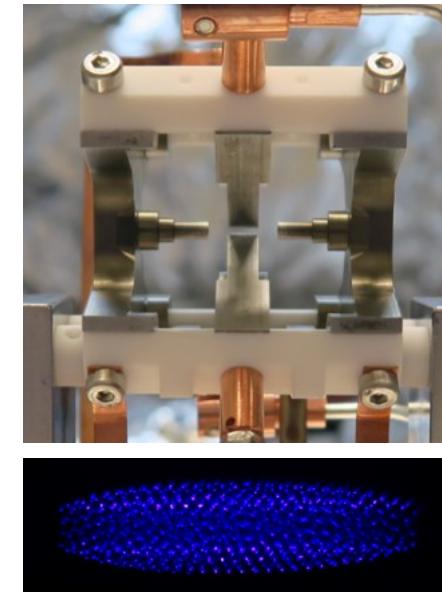
High-power IR frequency comb  
Yb:KYW laser + amplifiers @1030 nm



Frequency comb at 60.8 nm  
Intracavity high-harmonic generation  
17<sup>th</sup> harmonic

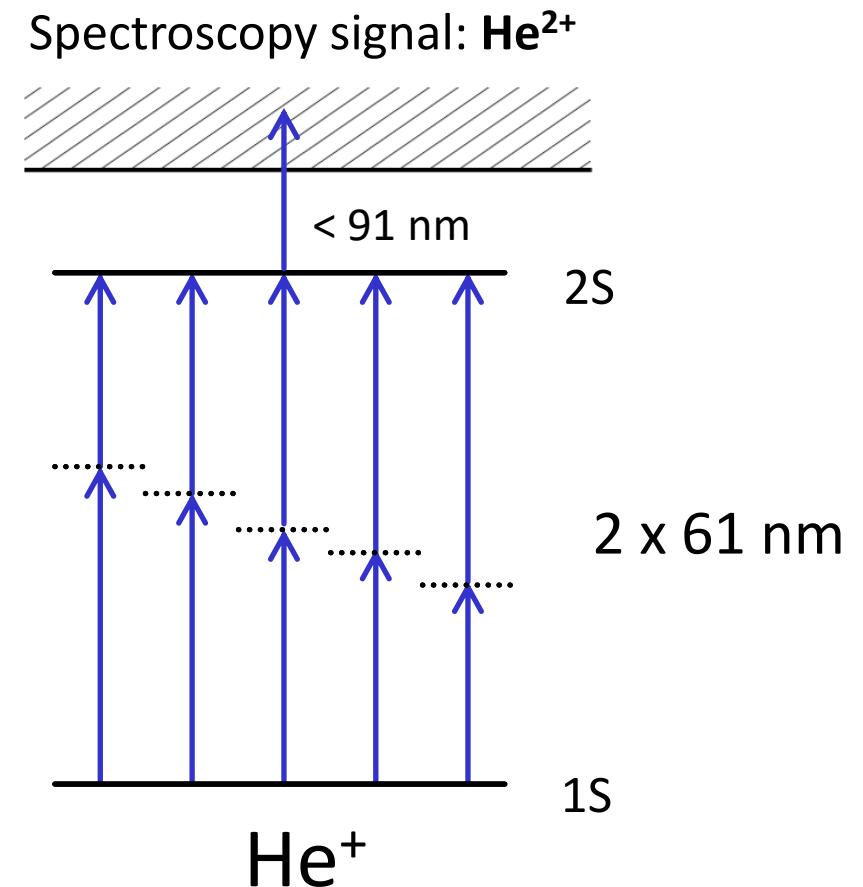
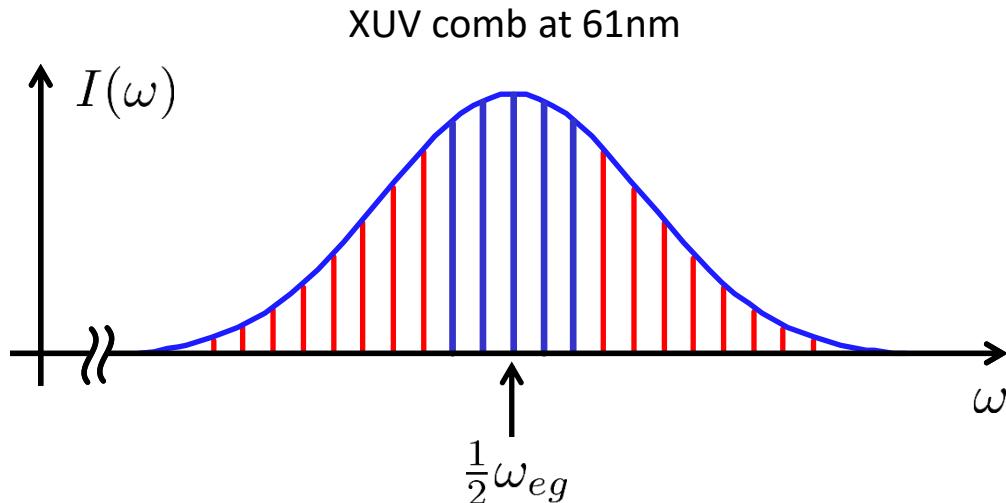


Trapped  $\text{He}^+$  ions



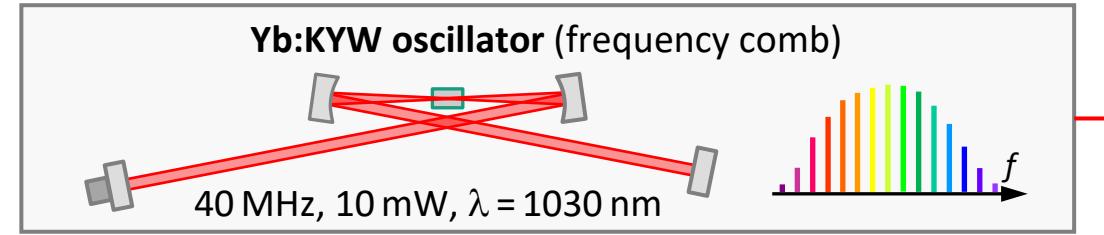
# Direct frequency comb spectroscopy

- Line width of a single laser mode
- Efficient nonlinear frequency conversion due to high peak power of pulsed laser
- As efficient as CW laser of same average power

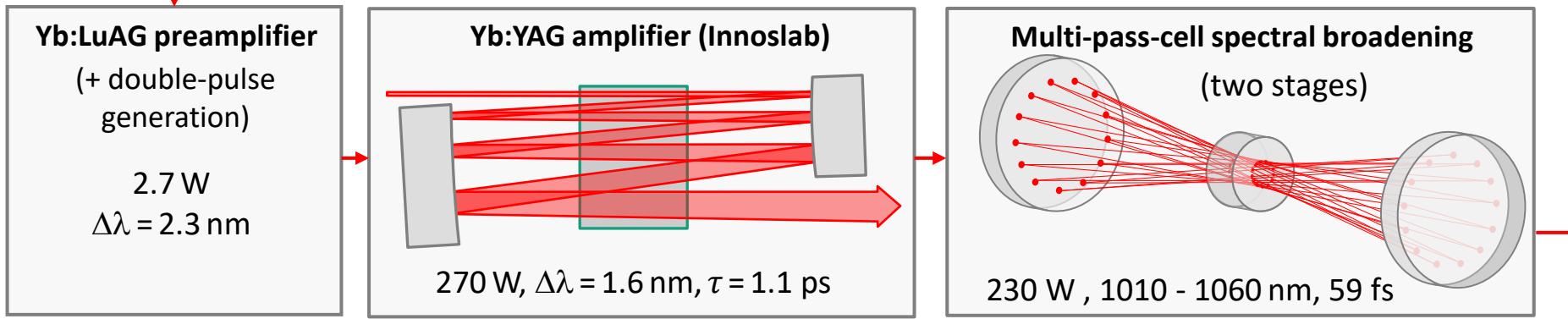


# 61 nm laser system

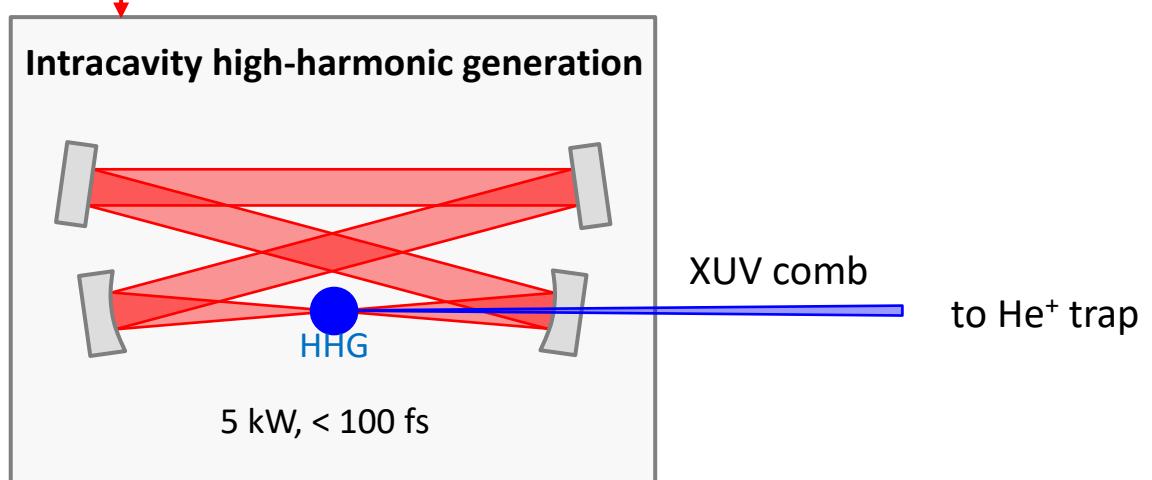
Seed laser



Amplification  
&  
compression



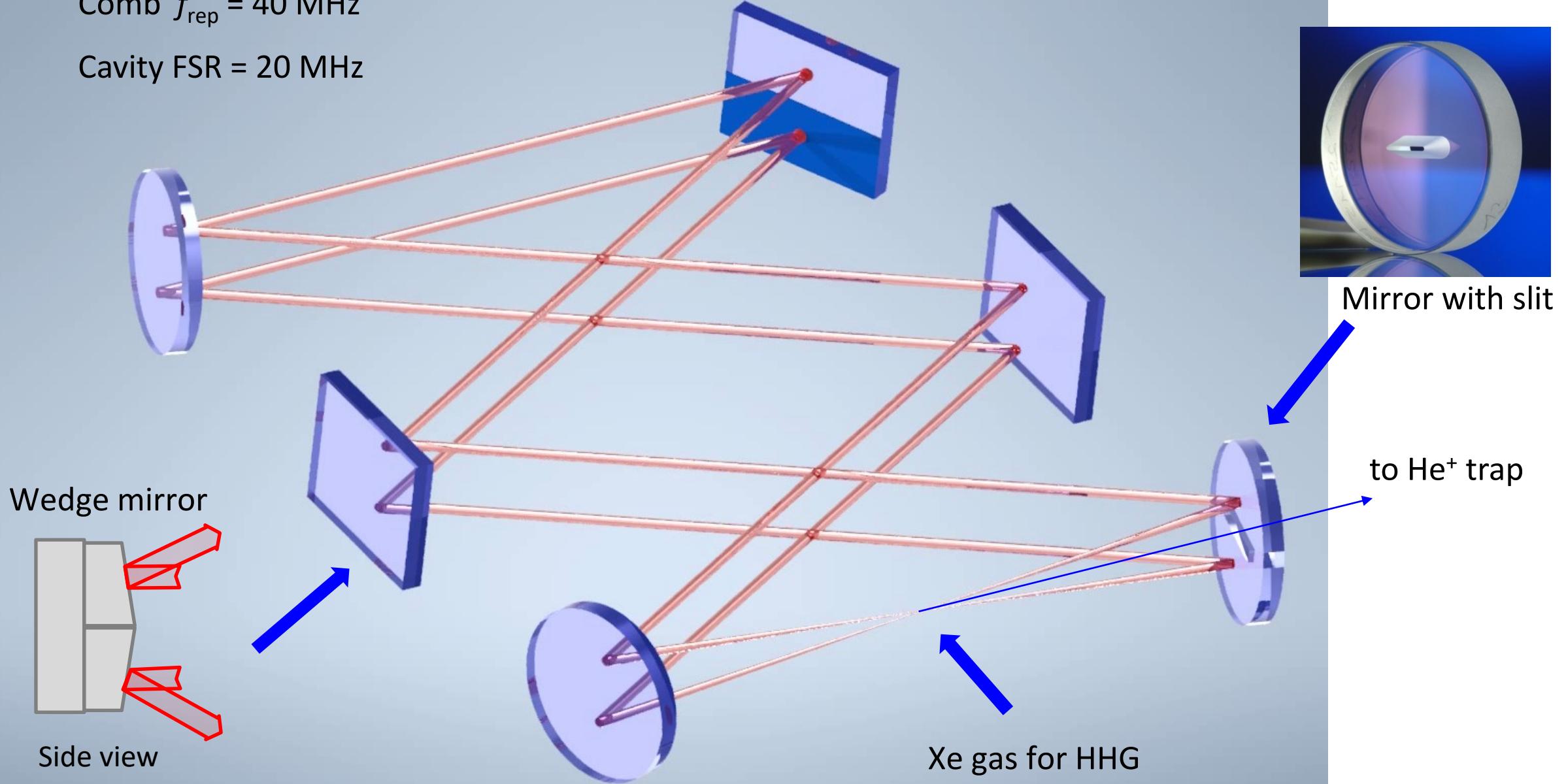
XUV comb  
generation



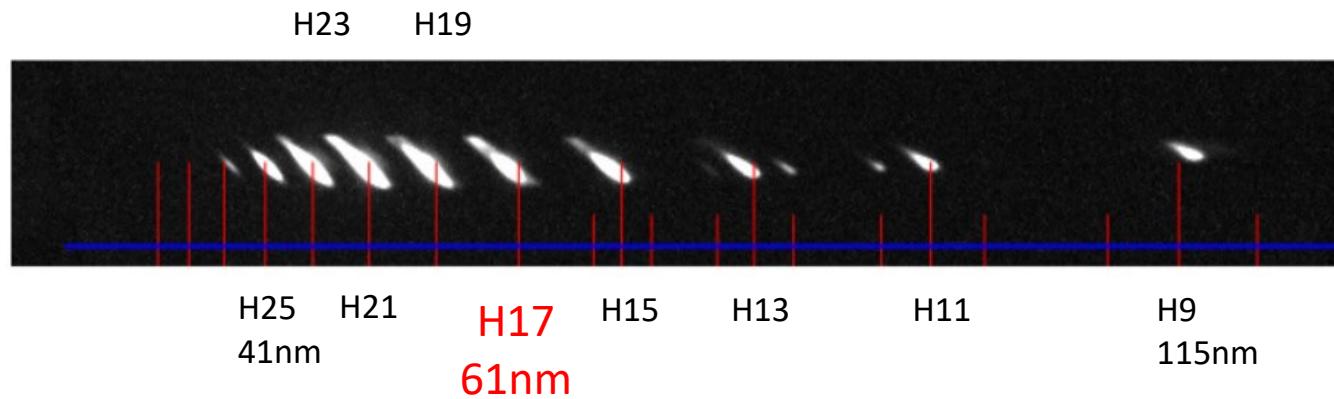
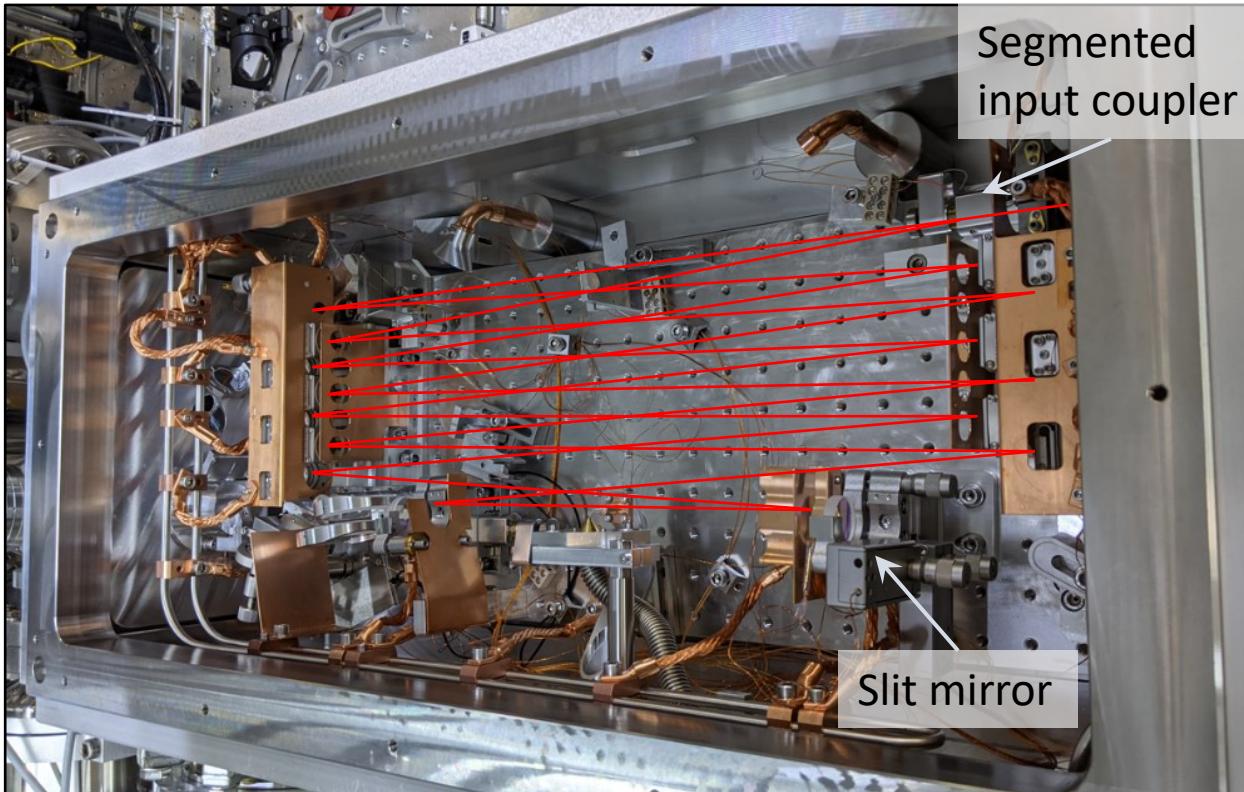
# Enhancement Resonator for HHG

Comb  $f_{\text{rep}} = 40 \text{ MHz}$

Cavity FSR = 20 MHz



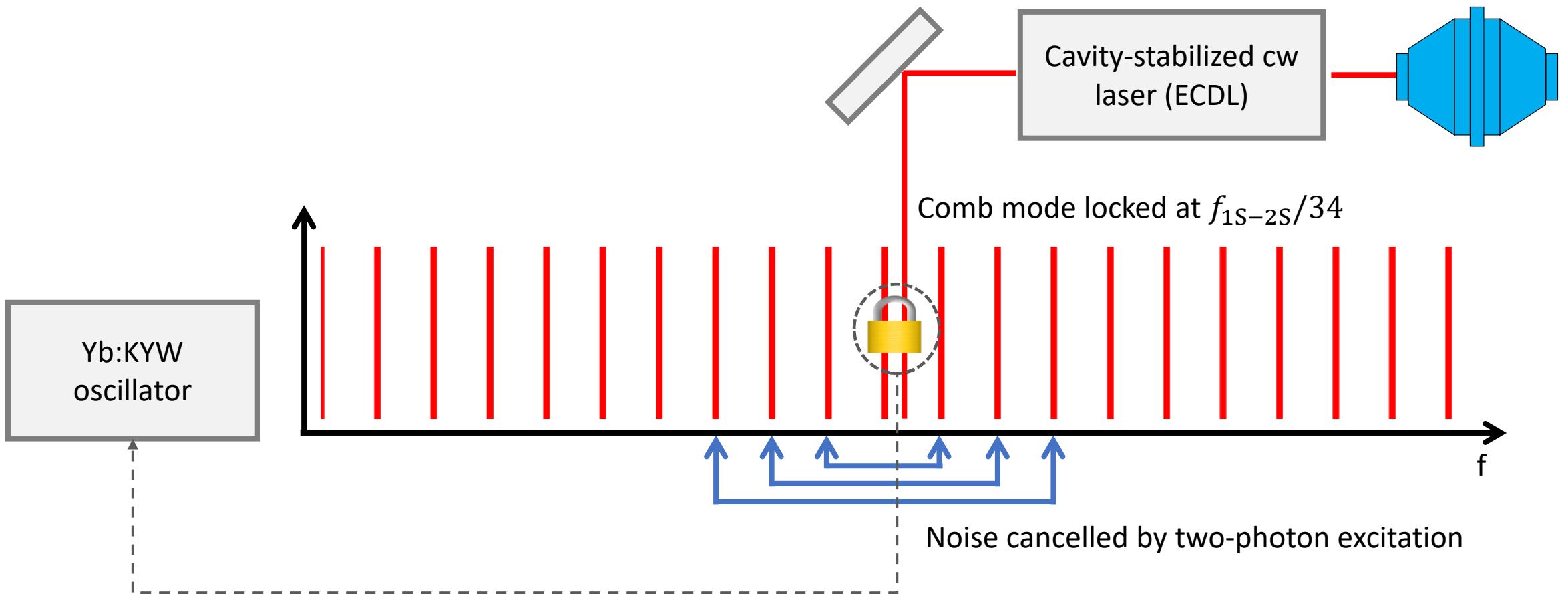
# Enhancement Resonator for HHG



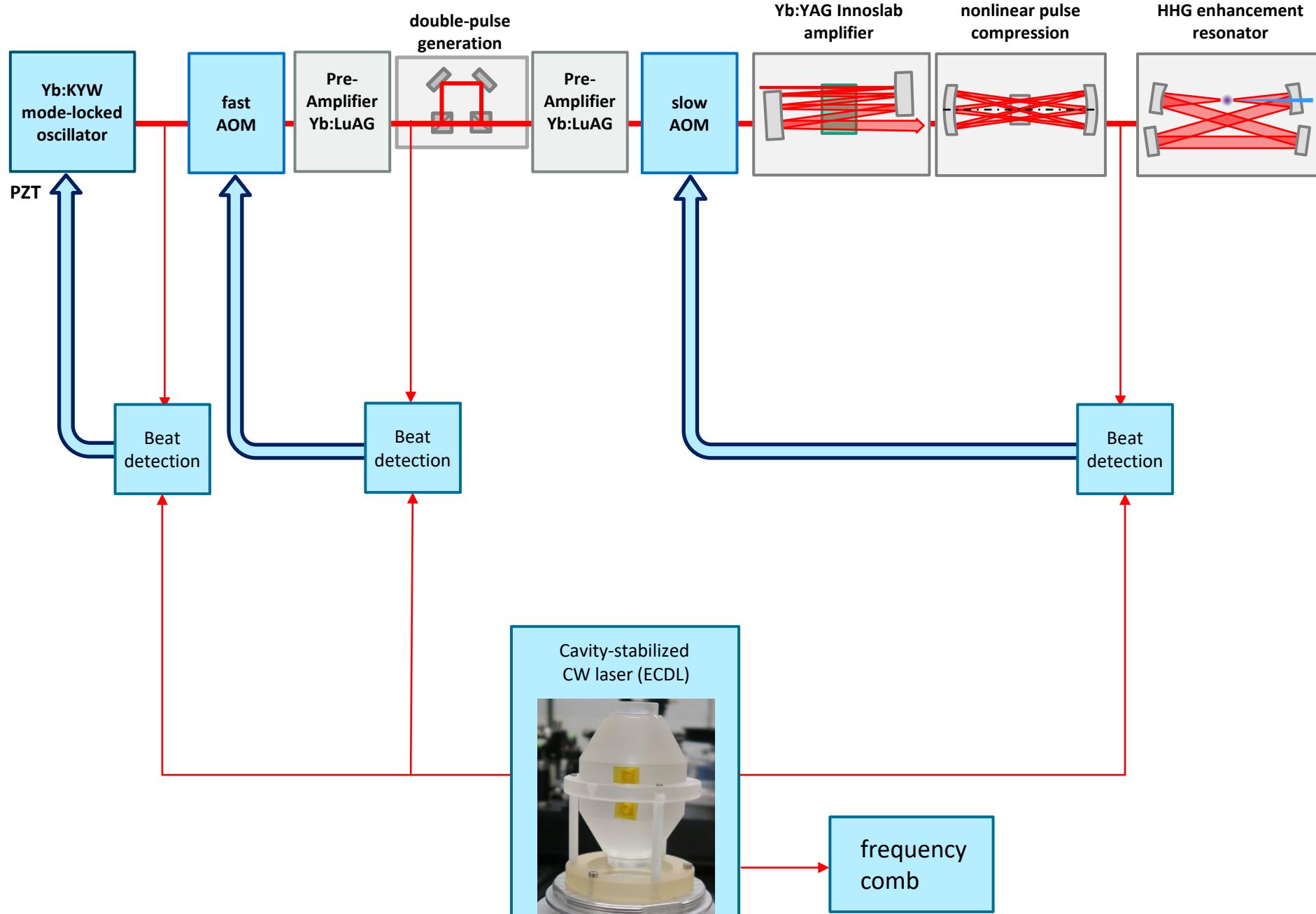
~30  $\mu\text{W}$  at 60.8 nm (preliminary)

→ ~ 0.1 Hz signal rate (ionization)

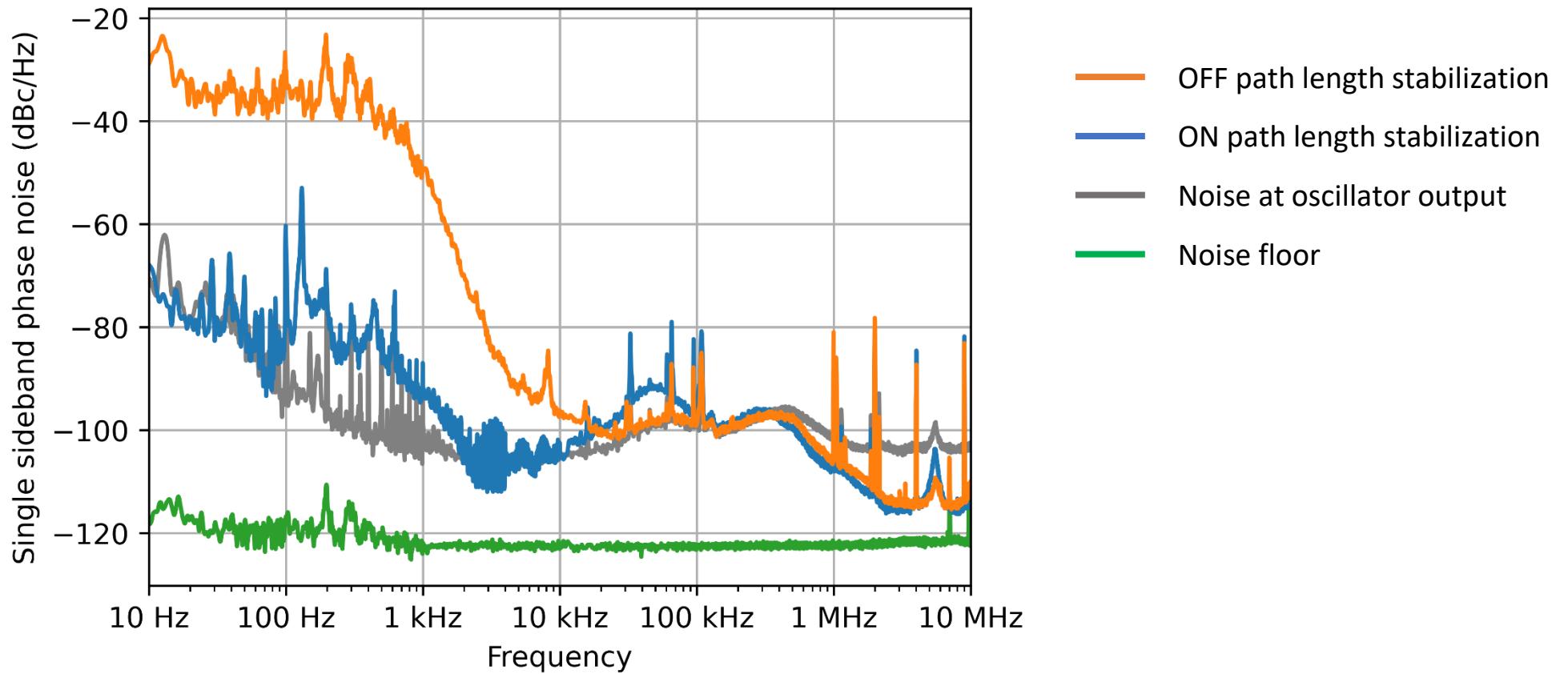
# Comb stabilization



# Stabilization of frequency comb

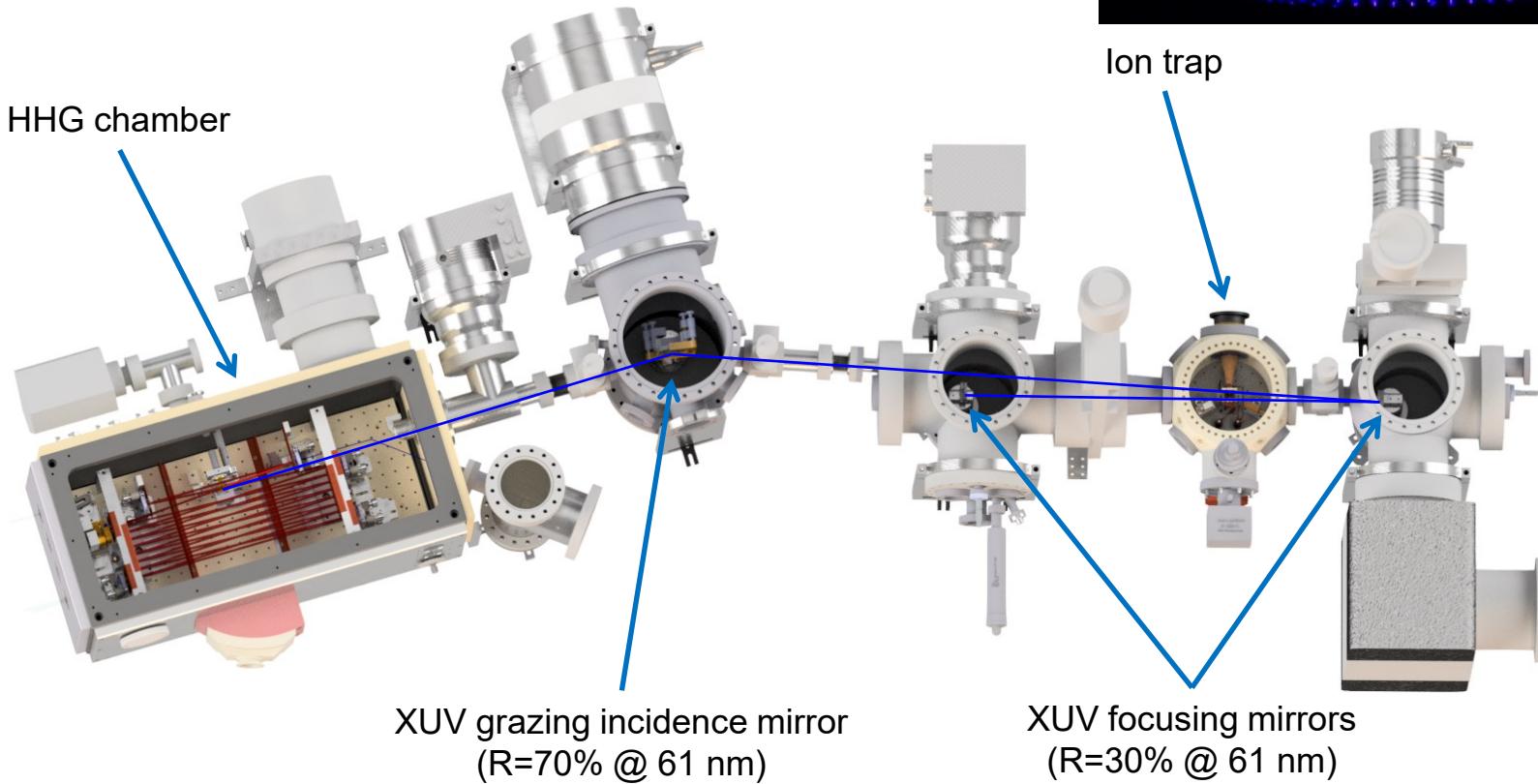


# Stabilization of frequency comb

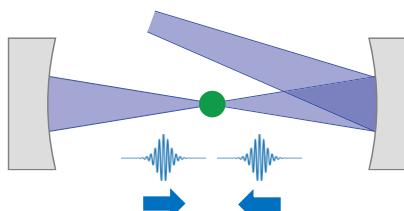


- 80% of power in carrier at XUV
- Less than 20% of signal loss due to phase noise

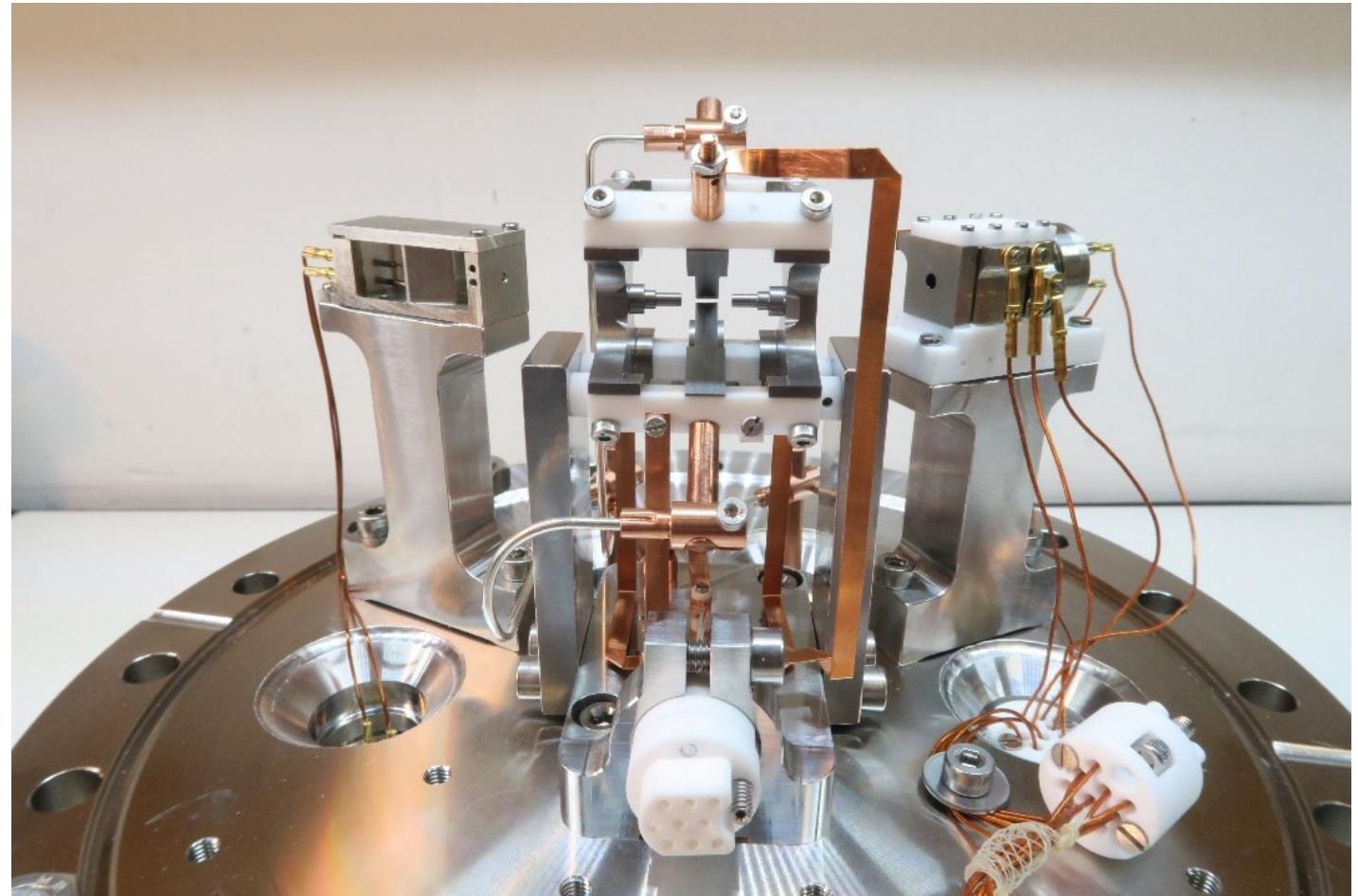
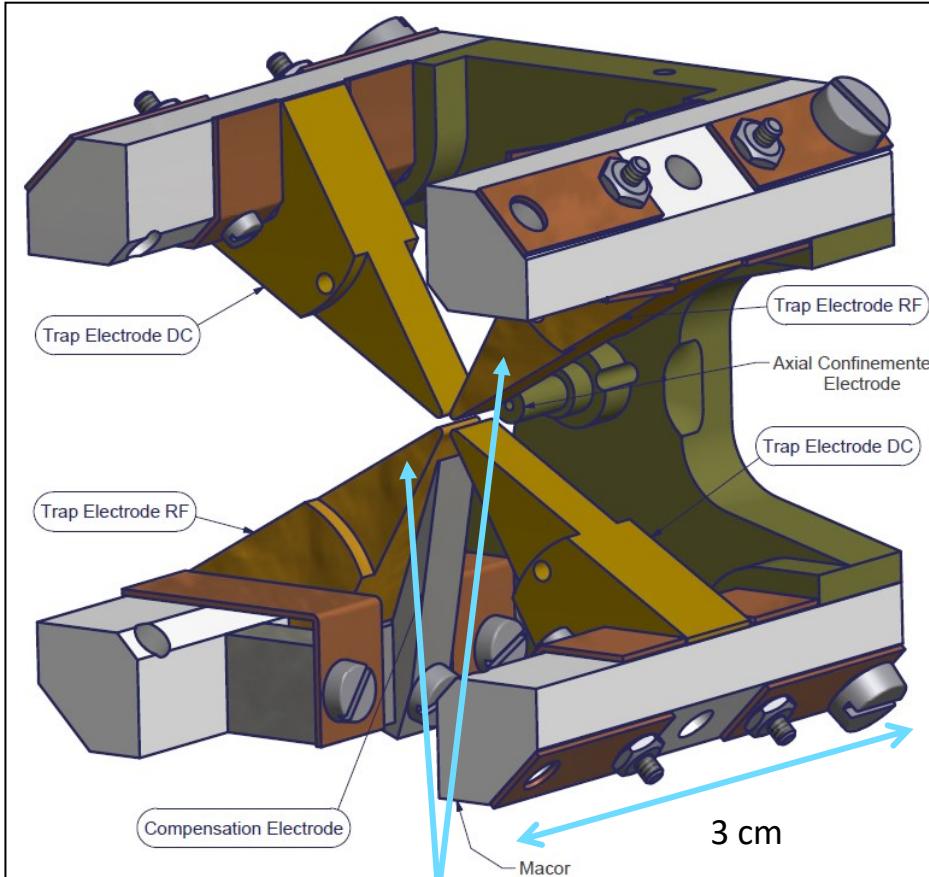
# XUV beam delivery to trapped $\text{He}^+$



Doppler-free  
Anti-collinear excitation

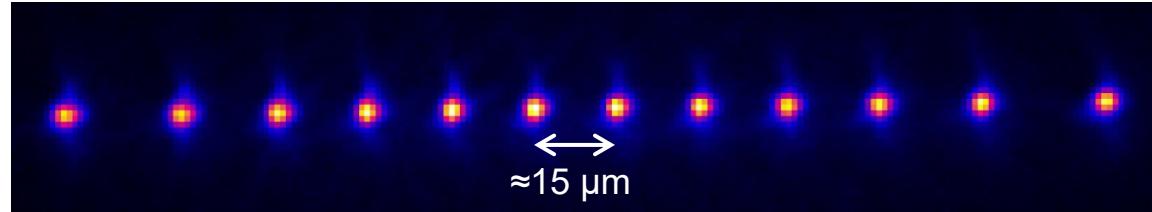


# $\text{Be}^+/\text{He}^+$ ion trap

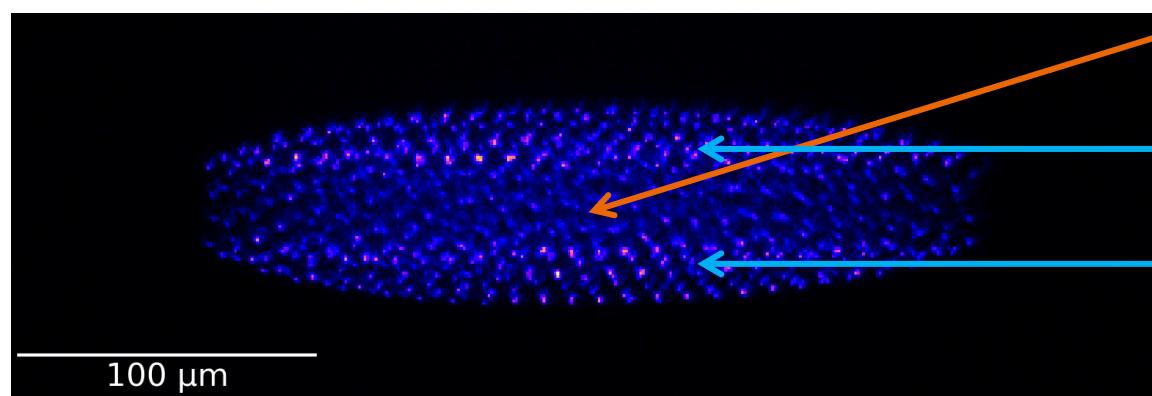
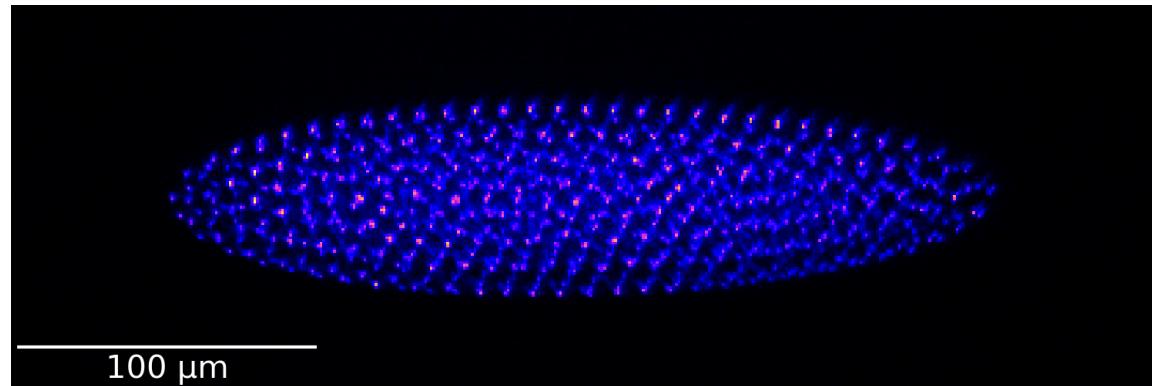


# Ion crystals and sympathetic cooling

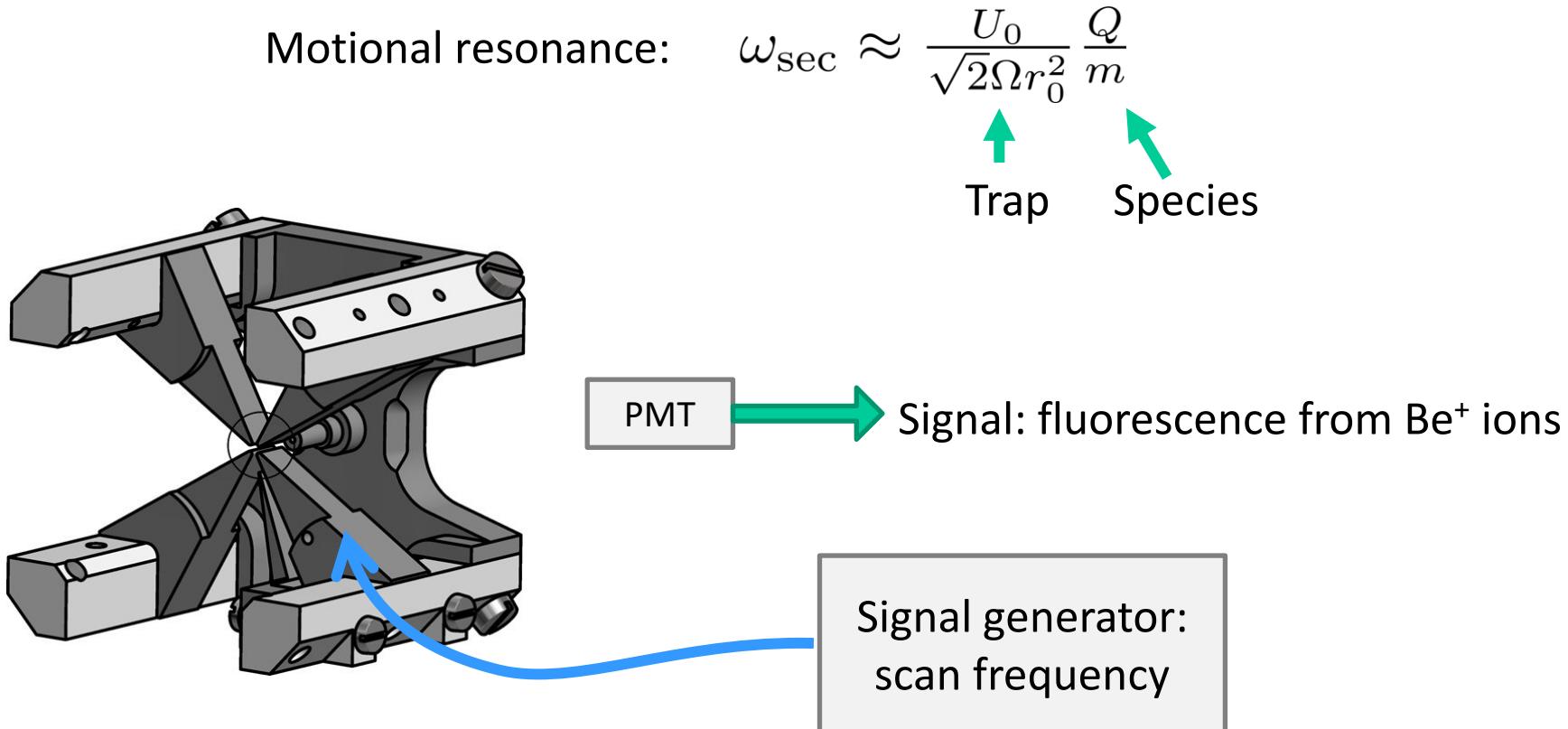
$\text{Be}^+$  laser cooled to a few mK



$\text{He}^+$  sympathetically cooled

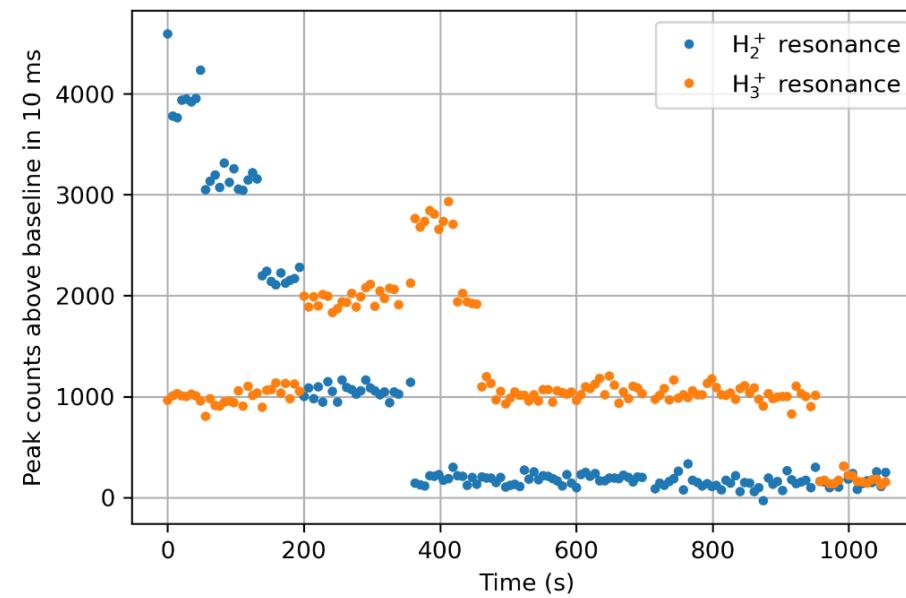
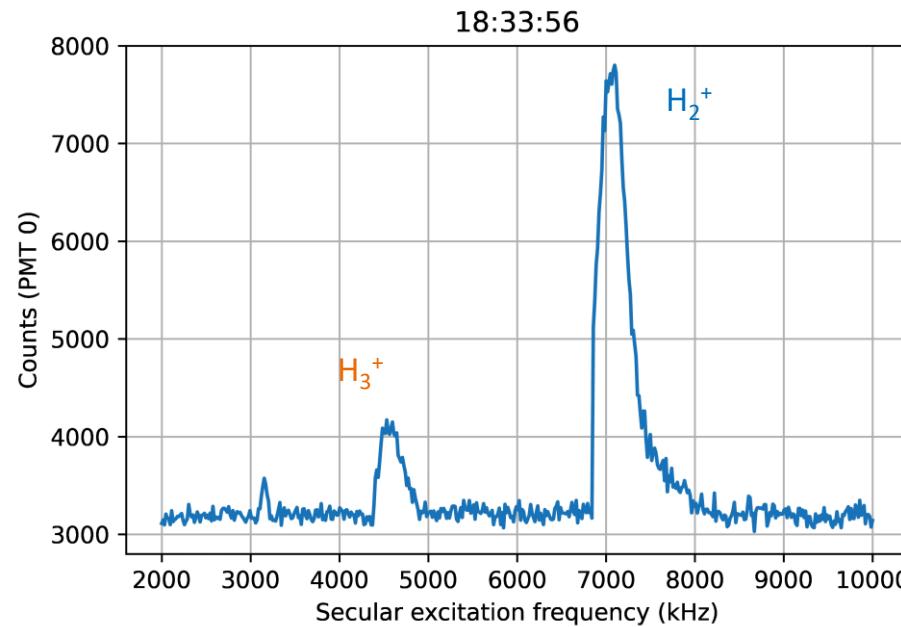
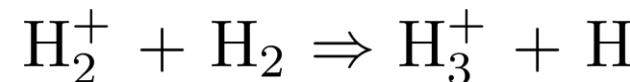


# Spectroscopy signal: Secular excitation detection



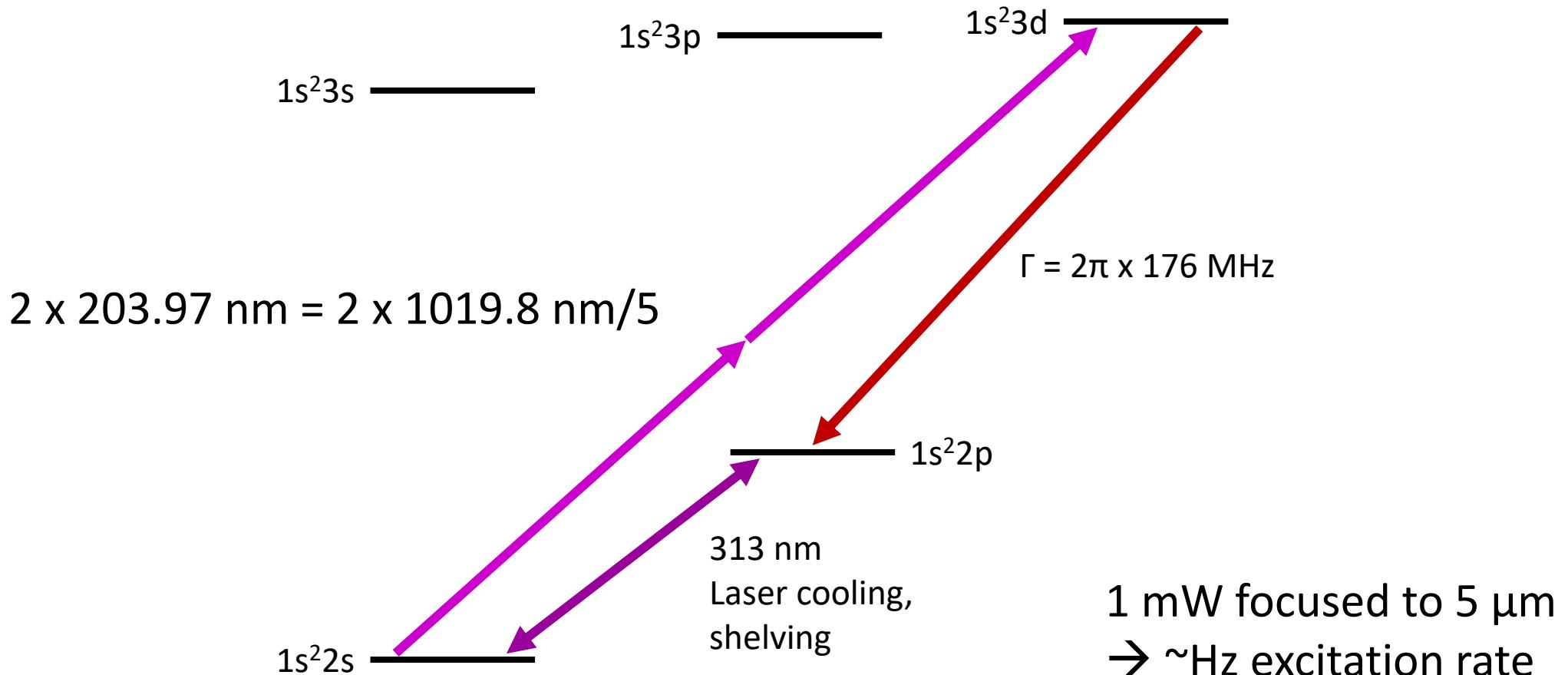
# Testing secular excitation detection

$\text{H}_2^+$ : Same charge-to-mass ratio as  $\text{He}^{2+}$

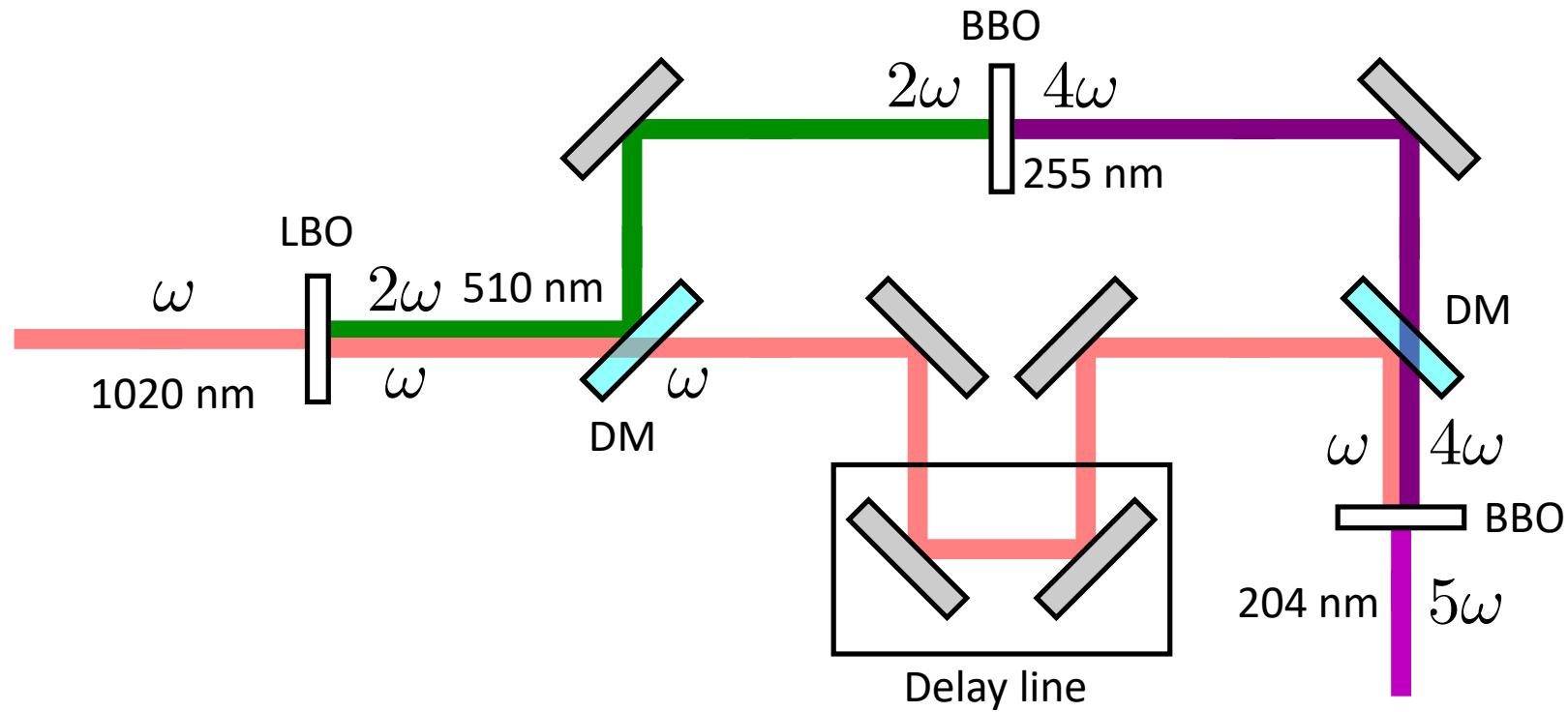


Single event sensitive detection

# Driving two-photon transition in Be<sup>+</sup>



# 5<sup>th</sup> harmonic generation in crystals



## Achieved:

- Generation of 61nm comb
- Trapped and cooled  $\text{He}^+$
- Demonstrated sensitive  $\text{He}^{2+}$  detection

## Plans :

- Optimization of 61nm comb  
(power, spectrum)
- Testing spectroscopy setup with easier transition  
(Two-photon transition at 204 nm with  $\text{Be}^+$ )
- Finding  $\text{He}^+$  1S-2S signal:  
QED prediction:  $\sim 70$  kHz  
Expected linewidth of signal:  $\sim 1$  kHz

# The He<sup>+</sup> team



Jorge  
Moreno



Fabian  
Schmid



Akira  
Ozawa



Johannes  
Weitenberg



Theodor  
W. Hänsch



Thomas  
Udem

Thank you for your attention!

