



Future Facilities and Directions

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# Map of Mainz

Gut/nb N-UIC/Caf

# DARK MESA

Kurfürstliches Schloss

the niu Mood 😜

**B&B HOTEL Mainz-Hbf** Naturhistorisches Landesmuseum Mainz Museum Mainz Am Fort Consenheim Taiyo Finest Sushi & Asian Fusion Rheingoldhalle Mainzer Taubertsbergbad Stadion am Bruchweg 8 🕀 🕥 台 SCHMID Mainzer Dom HARTENBERG-MÜNCHFELD Praxis am Holztur me & all hotel Mainz la kus Schugt, MS **Erbacher Hof** Am besten bewertet Saarstraße Eskute E-Bike Shop straße St. Stephan Mainz Hege Johannes Saarstra Gutenberg-Universität CineStar Mainz Bastion Zitadelle auf dem Jakobsberg Mainz reriusweg Botaris (1)St (1)

OBERSTADT Maps from https://google.de/maps/

Drususwall

# Institute of Nuclear Physics

nonz-institut viainz

Mainzer Mikrotron

M. Biroth for the MAGIX Collaboration, Mainz – mbiroth@uni-mainz.de

DARK MESA

ser-Hüsch-Weg

# Electron Accelerator MAMI



- DARK
- Former A4 experiment measured parity violating asymmetries



Former A4 experiment measured parity violating asymmetries •



High-power beam dump

Thickness: 2 m  $\sim$  26  $X_0$ 

...and after clearing the hall...



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# Electron Accelerator MESA



25 MeV superconducting cavities

Mainz Energy-recovering Superconducting Accelerator

Energy Recovery Linac (ERL) mode:

- Energy:  $E_e \leq 105$  MeV in 2 turns
- Recovery:  $\delta E_e \leq 100 \text{ MeV}$
- Current:  $I_e > 1 \text{ mA}$

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P2 exp. t beam dump

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P2 exp. + Vecu Rich portfolio of dark sector experiments

# WIMPs and Light Dark Matter

Weakly Interacting Massive Particles



- Matching relic abundance for the electroweak mass scale
- WIMPs require only SM interaction
- No positive evidence after LHC and galactic DM searches

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Light Dark Matter



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- Thermal relic targets exist for the MeV-GeV mass scale
- LDM requires a beyond SM force
- Rich phenomenology of portals: vector, higgs, neutrino, axion

- Dark photon  $\gamma'$  is the vector mediator
- Coupling to ordinary photons by kinetic mixing  $\varepsilon = g_{\rm D}/g_{\rm SM} = \sqrt{\alpha_{\rm D}/\alpha}$
- Decay modes to both sectors



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- MESA in the ERL mode
- Supersonic gas jet target
- Magnetic spectrometers



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Invisible decay: 
$$m_{\gamma \prime} > 2m_{\chi}$$



• Scintillator telescope layer for PID











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P2 experiment (runtime 20,000 h)

- MESA in the EB mode
- Liquid hydrogen target  $60 \text{ cm} \rightarrow dE/dx \sim 17 \text{ MeV}$
- Dumped electrons:  $6.74 \times 10^{22}$

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**Direct LDM measurement!** 





# Simulation: Dark Photons

**GEANT4** simulation:

- 155 MeV electron beam interacts with P2 target
- Residual beam is stopped within one radiation length X<sub>0</sub>
- On average 1 hard photon and 3 charged particles per electron are produced:
  - Dark Bremsstrahlung
  - Dark annihilation

 $e^+ + e^- \rightarrow \gamma' + \gamma$ 





#### M. Biroth for the MAGIX Collaboration, Mainz – mbiroth@uni-mainz.de

# Simulation: Dark Matter Pairs

MadGraph simulation:

- Calculation of physical dark photon decays into DM pairs  $\chi \overline{\chi}$
- Assumptions:  $m_{\gamma \prime} := 3 \ m_{\chi}$  ,  $\alpha_{\rm D} := 0.5$

Back in GEANT4:

- Dark matter particles knock out electrons with energies as low as  $E_{\rm th} \sim 10 {\rm ~MeV}$
- A Calorimeter is perfectly suited to detect low-energetic electrons
- Different Cherenkov glasses were tested for their sensitivity







# Detector Sensitivity Study

PbF<sub>2</sub> and the Pb-glass SF5 offer proper electron sensitivity and neutron insensitivity



#### Electron sensitivity study

Electron beam test, read: Mirco Christmann, et al.

- NIM A 958 (2020) 162398
- NIM A 960 (2020) 163665

Bachelor thesis of Paul Burger: Neutron sensitivity study with an AmBe source

#### Neutron sensitivity study





# Modular Calorimeter Concept





• Modular calorimeter concept

 $\frac{PbF_{2} \text{ module}}{5 \times 5 \text{ matrix of } PbF_{2}}$ Volume: 4 l

1000 crystals: A4@Mainz each  $(18 \times 3 \times 3)$  cm<sup>3</sup>



SF5 module 4×4 matrix of SF5 Pb-glasses Volume: 91

2000 glasses: WA98@CERN each  $(46 \times 3.5 \times 3.5)$  cm<sup>3</sup>

# Active Volume Expansion

 Phase A: 1 PbF<sub>2</sub> module, active volume of 0.004 m<sup>3</sup>



 Phase B: 30 PbF<sub>2</sub> and 64 SF5 modules, active volume 0.7 m<sup>3</sup>



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- Phase C: Long term plans
  - DarkMESA DRIFT: 1 m<sup>3</sup> negative ion TPC as complimentary detection method
  - Calorimeter upgrade: Volume expansion up to 9 m<sup>3</sup>

# DarkMESA Expected Reach



Phase	Detector	Period	Time	EOT
А	Prototype	14. year	2,200 h	$7.42 \cdot 10^{21}$
В	PbF <sub>2</sub> , SF5	46. year	6,600 h	$2.22 \cdot 10^{22}$
С	+TPC	712. year	13,200 h	$4.45 \cdot 10^{22}$

Simulation by M. Christmann, Saskia Plura

Limits given as  $y \propto m_{\chi}^2 \langle \sigma_A v \rangle$  $\sigma_A \sim \text{cross section of } \chi \overline{\chi} \to SM$ 



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Main Competitors:

- LDMX
- Future NA64
- Belle II at high masses



#### DARK Phase-A Calorimeter Response

#### Phase A calorimeter (5x5 PbF<sub>2</sub>):



Response to 4 different cosmic tracks:









#### Bachelor thesis of Jonas Pätschke

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-0.004

# Calorimeter Crystal Calibration

Gain calibration of the PMTs

• Curve-fitting of charge spectra under short laser pulses

Light output calibration of the crystals:

• Sensitivity to particle tracks with respect to the Cherenkov angle:

 $\cos\theta = 1/n\beta$ 

• Beam test data is under analysis

PhD thesis of Matteo Lauß





Bachelor thesis of Jonas Pätschke

# Cosmic Particles Veto Concept DARK

- Veto layers from double plastic scintillator-absorber sandwiches
- Heavy absorber shields electrons and converts photons, neutrons
- Muons produce a signal in opposite veto layers
- DM is assumed to pass the veto detector



Beam test data under analysis



PhD thesis of Matteo Lauß

# Phase-A Veto Readout





Further reading: M. Lauß, et al., NIM A 1012 (2021) 165617

# Detector Signal Processing



#### Narrow calorimeter PMT signal:

work

25 50

-0.02

Amplitude [V] -0.04 -0.06

-0.08

-0.10

-0.12

175

0.00

-0.05

-0.10

-0.15 -upplitude -0.20 -

-0.25

-0.30

-0.35

50 100 150

200 250

Time [ns]

300 350

ò

Broad Veto SiPM signal:

- Signals are sampled for • documentation purposes
- Ongoing study of pulse ٠ shaping to match ADC characteristics
- PANDA SADC
  - Channel count: 64
  - Voltage resolution: 14 bit
  - Sampling frequency: 80 MHz = 1/12.5 ns

See: PhD thesis of E.J.O. Noll (2020) Digital Signal Processing for the Measurement of Particle Properties with the PANDA Electromagnetic Calorimeter







Additional information:

- S. Plura (today 16:51) Search for Light Dark Matter with the DarkMESA Experiment
- S. Schlimme (Thursday 11:00) The MESA physics programme



## Thermal Dark Matter

- Early universe  $T > m_{\chi}$ : Dark Matter (DM) is hold in thermal equilibrium by interaction with Standard Model (SM) matter
- During cooling down  $T < m_{\chi}$ : Number density dropped  $n \propto e^{-m_{\chi}/T}$
- Finally  $T \ll m_{\chi}$ : Number density froze out at  $n_{\rm FO} \approx H/\langle \sigma_{\rm A} v \rangle$
- Increased annihilation cross-section  $\sigma_{\rm A}$  would lead to decreased relic abundance and later freeze-out



Hubble constant *H*, thermally averaged annihilation cross-section  $\langle \sigma_A v \rangle$ 

### WIMPs Expected Reach



#### chamber filled with Carbon Disulfide at 0.05 bar, see: D.P.

Snowden-Ifft, et al., arXiv:1809:06809

- Lowers the nuclear recoil detection threshold:  $E_{\rm th} \sim 35 \; {\rm keV}$
- Requires Gadolinium-doped scintillators as cosmic Neutron veto detectors
  - Started cooperation with Philip Cole, Lamar University, Texas, US

# DarkMESA DRIFT

Negative ion time projection





# Cosmic Particles Background

Simulation of cosmogenic background based on the Cosmic-Ray shower librarY •

102 10 Cosmogenic **mu plus** flux 10 MeV Cosmogenic neutron flux 10 10 Cosmogenic flux (s<sup>-1</sup> m<sup>-2</sup>) Cosmogenic flux (s<sup>-1</sup> m<sup>-2</sup>) in hall  $\sim 7 s^{-1} m^{-2}$ 10 10 10-5  $10^{-1}$ 10-10  $10^{-2}$  $10^{-10}$  $10^{-2}$ 10-4 10<sup>4</sup> 10<sup>6</sup> 10<sup>8</sup>  $10^{2}$  $10^{8}$ Energy (MeV) Energy (MeV)

Cosmogenic particles have to be vetoed with high efficiency

35

Neutron flux: Sea level (dashed), Hall (solid)



# Beam Neutrons Background

• FLUKA simulation of beam-induced neutrons



- No significant background
- Work in progress

# PbF<sub>2</sub> Module Assembly



#### 3D printed housing with locking points



#### First layer assembly



Bachelor thesis of Jonas Pätschke

#### Single-detector assembly



#### Complete prototype



# Veto Scintillator Properties



- Veto scintillators tested with cosmic particles and in the electron beam
- Wavelength shifting fibers increased light output but also the inhomogeneity
- Further reading: M. Lauß, et al., NIM A 1012 (2021) 165617
- Ongoing study of
  - Readout electronics
  - Attenuation length < 10 %/cm</li>
  - Homogeneity







PhD and master thesis of Matteo Lauß

# Detector Signal Shapes

• Waveforms of calorimeter and veto detectors are differently shaped



#### Narrow calorimeter PMT signal:

- For documentation purposes the waveforms of rare events should be stored
- Budged per channel is limited in case of a total count of 2000 in phase B



**Broad Veto SiPM signal:** 

# Detector Signal Processing

- PANDA SADC
  - Channel count: 64
  - Voltage resolution: 14 bit
  - Sampling frequency: 80 MHz = 1/12.5 ns
  - See: PhD thesis of E.J.O. Noll (2020) Digital Signal Processing for the Measurement of Particle Properties with the PANDA Electromagnetic Calorimeter
- Energy signals of the fast PMTs have to be bandwidth limited to match the slow sampling rate
- Investigation of the optimum time constant  $\tau$ :
  - Require sample in the rising edge (Phase error)

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Optimize the charge integral SNR







# P2 Experiment at MESA



The weak mixing angle: Measurement of the parity violating asymmetry in the elastic scattering of polarized electrons off unpolarized nuclei



# Competitors Overview

- Beam dump experiments similar to DarkMESA
  - BDX@JeffersonLAB, e<sup>-</sup>beam (future unclear)
  - SHiP@SPS/CERN, p beam (future unclear)
  - MiniBooNE@FermiLAB, ν beam (ongoing)
- Missing mass experiments similar to MAGIX
  - PADME@INFN,Frascati, e<sup>+</sup> beam on e<sup>-</sup> (ongoing)
  - DarkLight@ARIEL/TRIUMF, e<sup>-</sup> beam (30 MeV in 2024, 50 MeV in 2025)
- Missing momentum/energy experiments
  - LDMX@SLAC,Stanford, e<sup>-</sup> beam (4 GeV in 2025, 8 GeV in 2027)
  - NA64@SPS/CERN, e<sup>-</sup> beam (upgrade)
  - Lohengrin@ELSA,Bonn, e<sup>-</sup> beam (future)
- Collider experiments
  - FASER@LHC/CERN, pp collisions (ongoing)
  - BaBar, BES III, ISR in e<sup>+</sup>e<sup>-</sup> collisions (future: Belle II)

# Visible Decay Expected Reach

- Decay length  $l = \gamma c\tau \le 8 \text{ mm}$ requires vertex reconstruction
- Search for a sharp invariant mass peak:  $m_{\gamma'}^2 = (\mathbf{e^+} + \mathbf{e^-})^2$
- Simulation by H. Merkel
- Competitors:
  - DarkLight
  - FASER
  - NA64



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- Search for a sharp invariant mass peak:  $m_{\gamma'}^2 = (\mathbf{e^+} + \mathbf{e^-})^2$
- Simulation by H. Merkel
- Competitors:
  - DarkLight
  - FASER
  - NA64
- Sensitivity enables the search for the vector boson candidate X17 in excited states of He, Be and C





# Invisible Decay Expected Reach



- Missing mass:  $m_{\gamma\prime}^2 = (e^{o-} - e^- + p^o - p)^2$
- Simulation by P. Gülker
- Competitors:
  - LDMX@SLAC 8 GeV
  - Future NA64
  - Future Belle II at high masses



# Visible Decay at MAMI

Linear electron accelerator Mainz Mikrotron (MAMI):

- Energy:  $E_e \leq 1.6 \text{ GeV}$
- Current:  $I_e \leq 100 \ \mu A$



RTM2

- 3 magnetic spectrometers
- Mass resolution  $\delta m < 100 \, {\rm keV}/c^2$

10 m

#### Experiment:

- Beam:  $E_e \in [180, 855]$  MeV
- Target: tantalum foil

#### **Exclusion limits:**

- $m_{\gamma \prime} \in [40, 300] \,\mathrm{MeV}/c^2$
- $\varepsilon \ge 8 \cdot 10^{-7}$



Let. 112 (2014) 221802

