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# **Low-energy collisions between two indistinguishable tritium-bearing hydrogen molecules: HT+HT and DT+DT**

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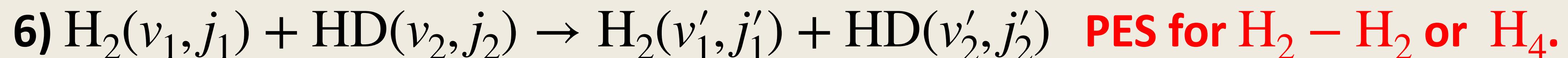
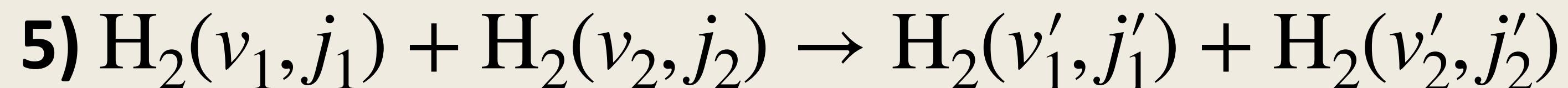
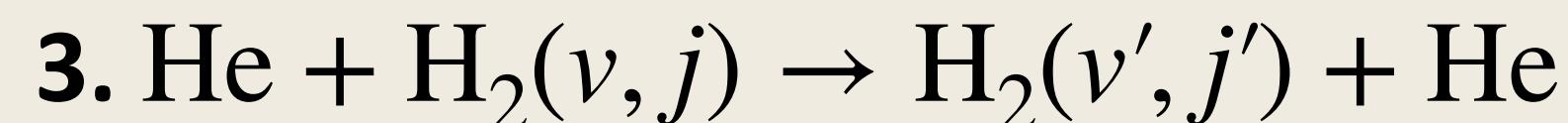
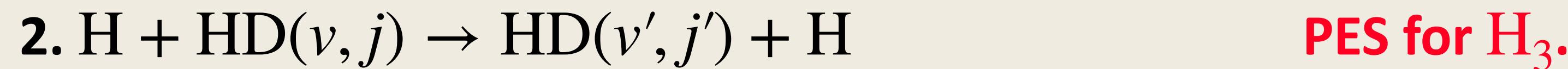
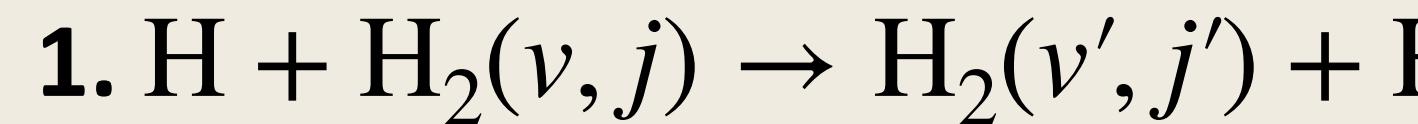
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1. In the interstellar medium (**ISM**), in cool stellar and in various planetary atmospheres, astronomers extensively observe and identify different atoms and molecules.
  2. These molecules are detected through emission and adsorption of electromagnetic radiation.
  3. The radiation and its interpretation provide astronomers with valuable information about astronomical objects.

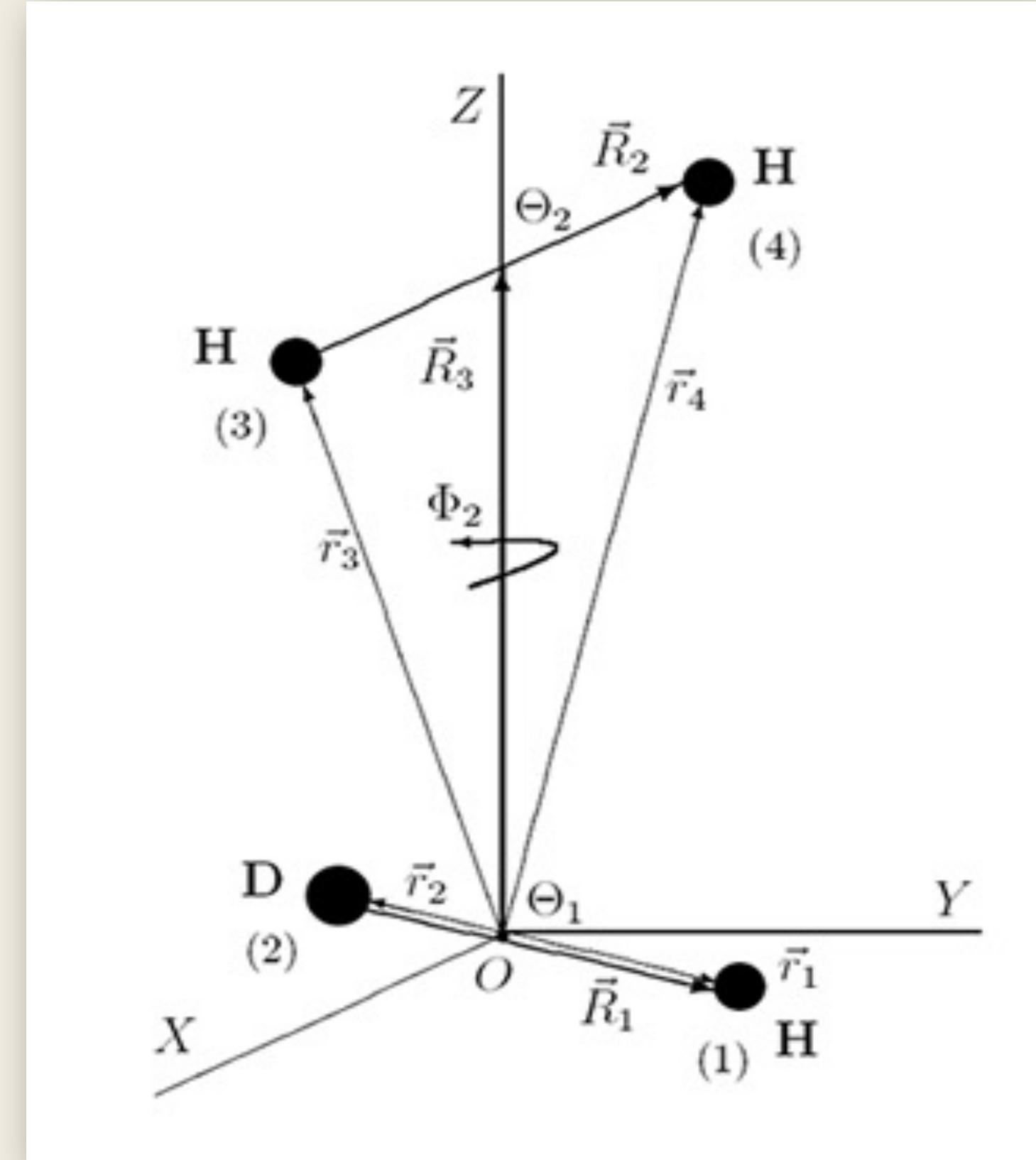
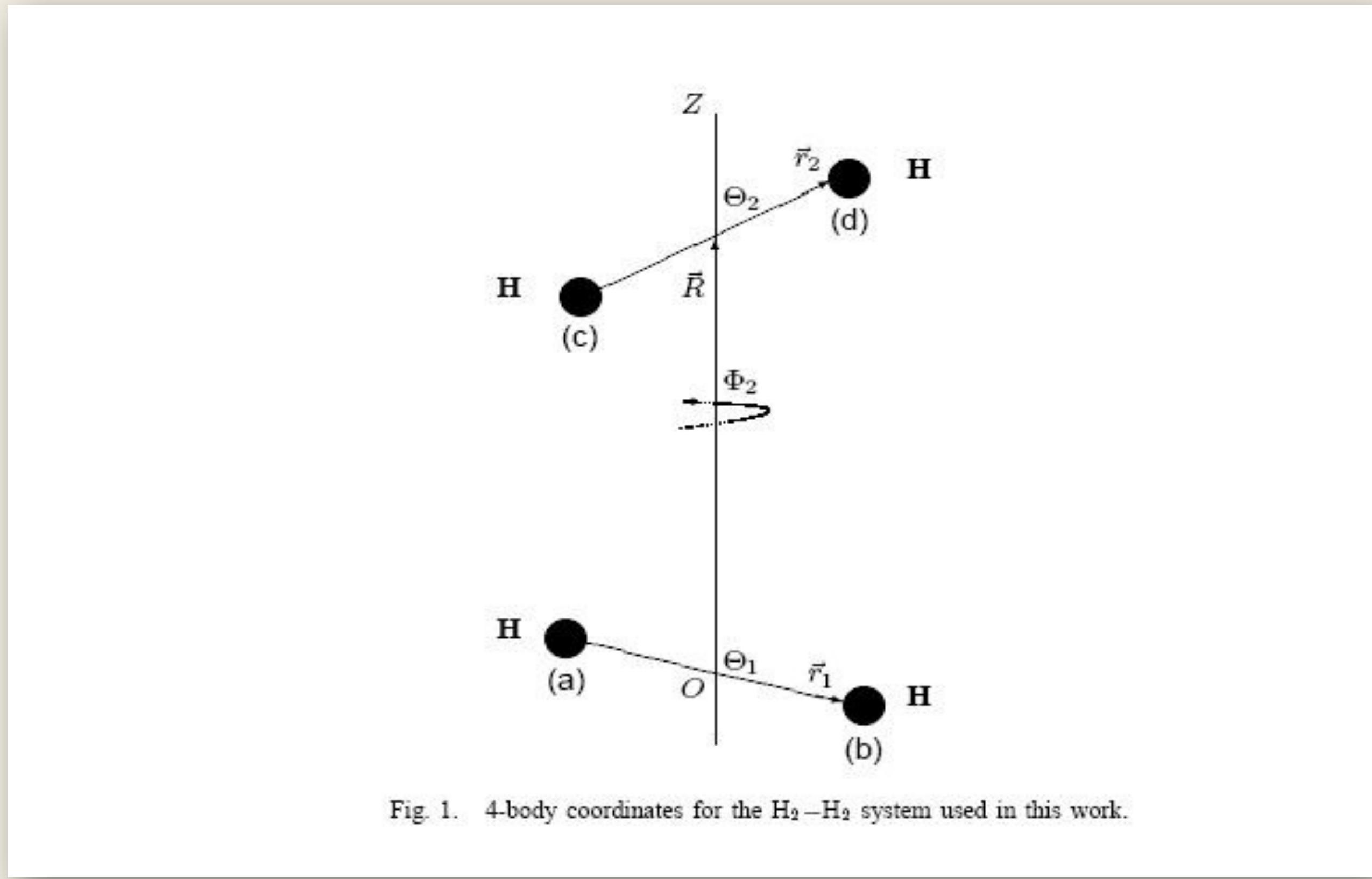
1. In order to obtain correct interpretation of different physical and chemical processes in the areas of interest in ISM it is important to know: **cross sections and thermal rate coefficients** of all microscopical processes in the population of emitting and adsorbing atoms and molecules.

$$\Lambda_{\text{HD}}(T) = n_{j_1}(\text{H}_2)n_{j_2}(\text{HD}) \times k_{j_1 j_2 \rightarrow j'_1 j'_2}(T) \times h\nu_{j_2 j'_2}$$

$$\Lambda'_{\text{HD}} = \Lambda_{\text{HD}}(T) - \Lambda_{\text{HD}}(T_{CMB}), \quad T < 100 \text{ K}$$

# Few - Body Atomic Collisions





$$V(\vec{r}_1, \vec{r}_2, \vec{R}) = V(r_1, \theta_1, r_2, \theta_2, R, \phi_2) = V(\theta_1, \theta_2, R, \phi_2)$$

# A Pure Quantum-Mechanical Approach: *rigid rotor model*

$$(H_0 + V(\vec{r}_1, \vec{r}_2, \vec{R}) - E)\Psi(\vec{r}_1, \vec{r}_2, \vec{R}) = 0$$

$$\left( \frac{\hat{P}_{\vec{R}}^2}{2M_{12}} + \frac{\hat{L}_{\hat{r}_1}^2}{2\mu_1 r_1^2} + \frac{\hat{L}_{\hat{r}_2}^2}{2\mu_2 r_2^2} + V(\vec{r}_1, \vec{r}_2, \vec{R}) - E \right) \Psi(\hat{r}_1, \hat{r}_2, \vec{R}) = 0$$

$$M_{12} = 2m_h(m_h + m_d)/(3m_h + m_d)$$

$$\mu_1 = m_h/2\mu_2 = m_h m_d/(m_h + m_d)$$

$$\Psi(\hat{r}_1, \hat{r}_2, \vec{R}) = \sum_{JMj_1j_2j_{12}L} \frac{U_{j_1j_2j_{12}L}^{JM}(R)}{R} \phi_{j_1j_2j_{12}L}^{JM}(\hat{r}_1, \hat{r}_2, \vec{R})$$

$$\phi_{j_1j_2j_{12}L}^{JM}(\hat{r}_1, \hat{r}_2, \vec{R}) = \sum_{m_1m_2m_{12}m} C_{j_1m_1j_2m_2}^{j_{12}m_{12}} C_{j_{12}m_{12}lm}^{JM} Y_{j_1m_1}(\hat{r}_1) \\ Y_{j_2m_2}(\hat{r}_2) Y_{Lm}(\hat{R})$$

$$\left( \frac{d^2}{dR^2} - \frac{L(L+1)}{R^2} + k_\alpha^2 \right) U_\alpha^{JM}(R) = 2M_{12} \sum_{\alpha'} \int < \phi_\alpha^{JM}(\hat{r}_1, \hat{r}_2, \vec{R}) | V(\vec{r}_1, \vec{r}_2, \vec{R}) | \phi_{\alpha'}^{JM}(\hat{r}_1, \hat{r}_2, \vec{R}) > U_{\alpha'}^{JM}(R) d\hat{r}_1 d\hat{r}_2 d\hat{R}$$

$$\alpha \equiv (j_1 j_2 j_{12} L)$$

$$U_{\alpha}^J \underset{R \rightarrow +\infty}{\sim} \delta_{\alpha\alpha'} e^{-i(k_{\alpha\alpha} R - (l\pi/2))} - \sqrt{\left(\frac{k_{\alpha\alpha}}{k_{\alpha\alpha'}}\right)} S_{\alpha\alpha'}^J e^{-i(k_{\alpha\alpha'} R - (l'\pi/2))}$$

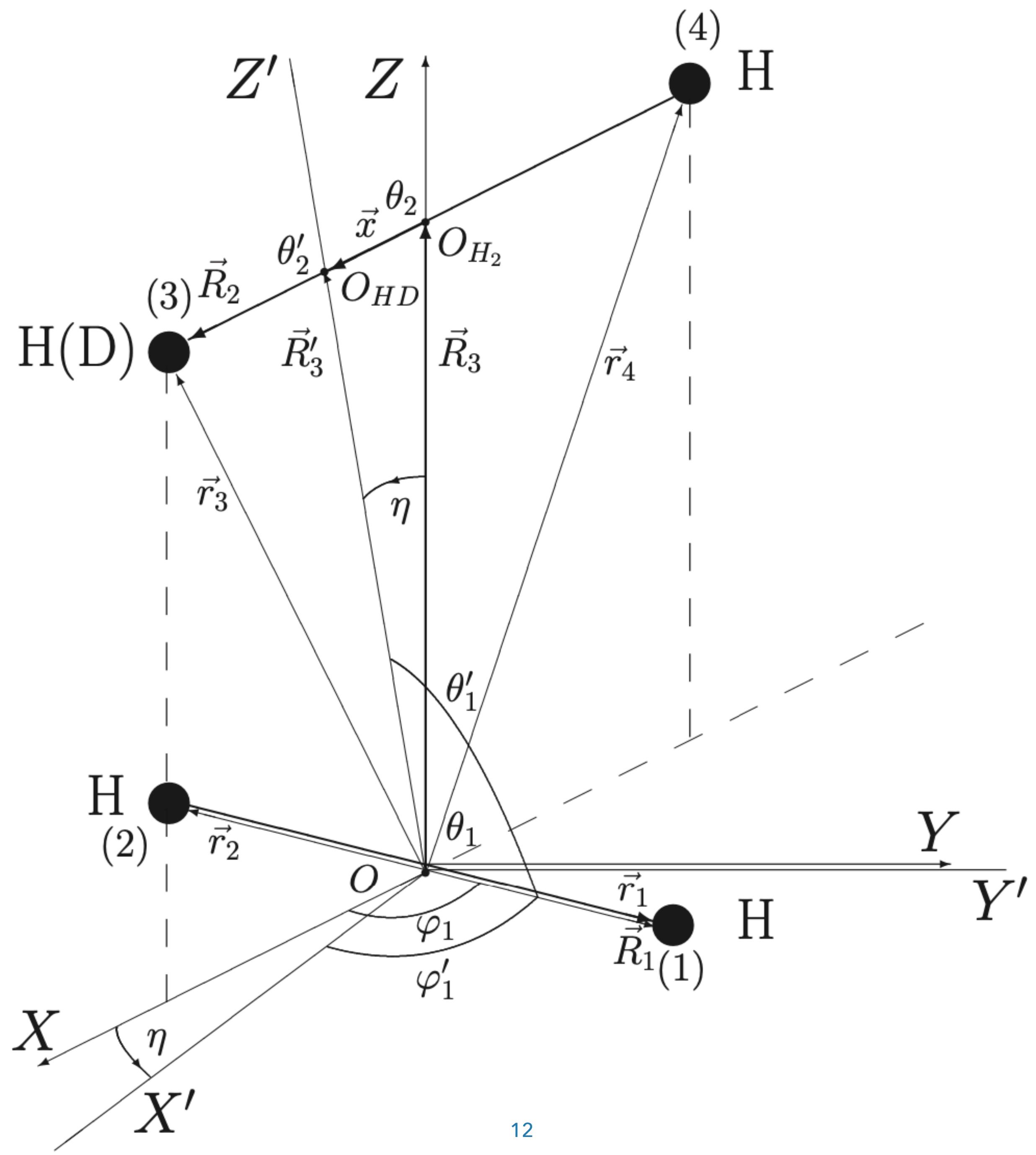
$$k_{\alpha\alpha'} = \sqrt{2M_{12}(E + E_{\alpha} - E_{\alpha'})}$$

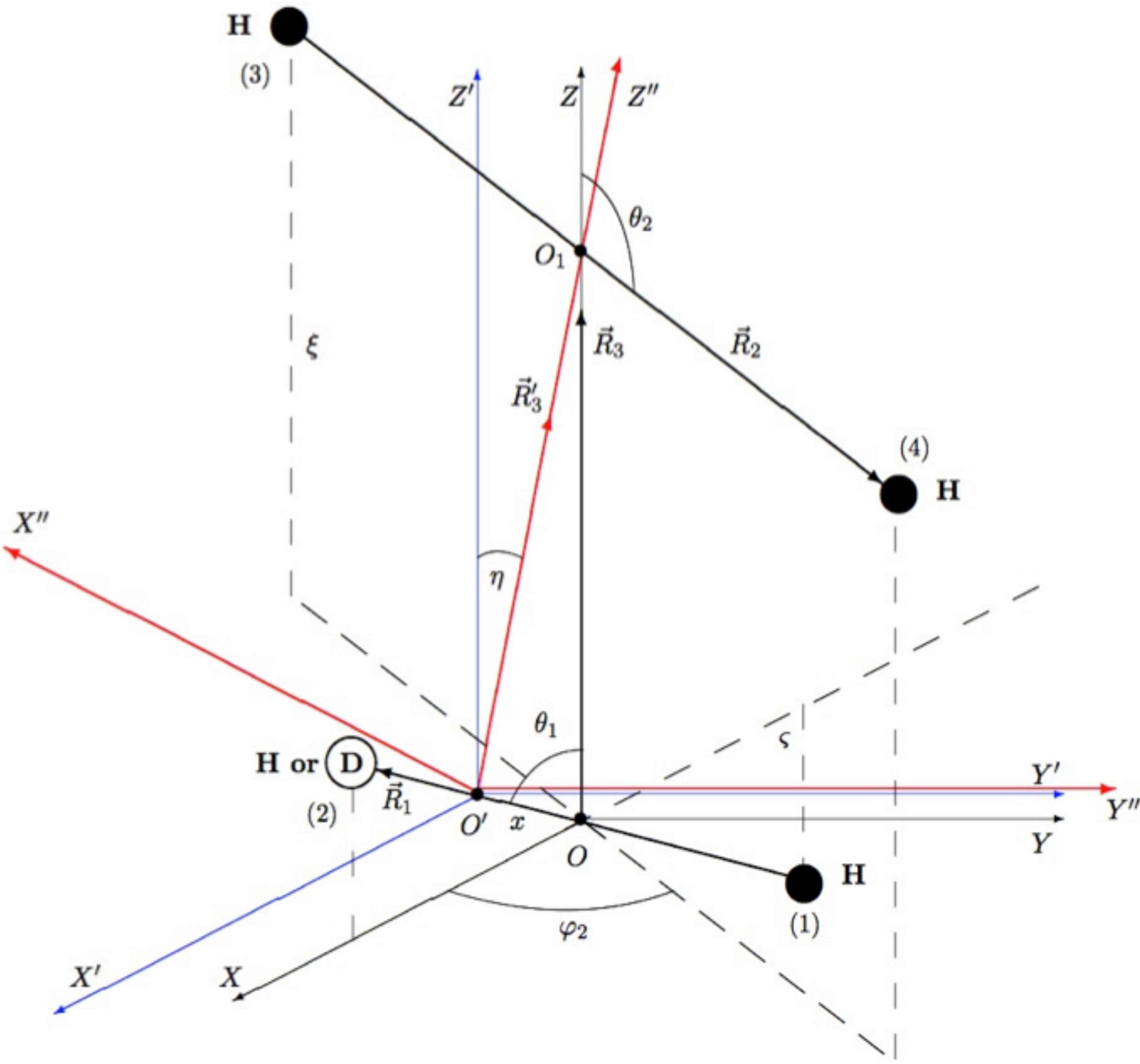
$$\sigma(j'_1, j'_2; j_1 j_2, \epsilon) = \frac{\pi}{(2j_1 + 1)(2j_2 + 1)k_{\alpha\alpha'}} \sum_{J j_{12} j'_{12} L L'} (2J + 1) |\delta_{\alpha\alpha'} -$$

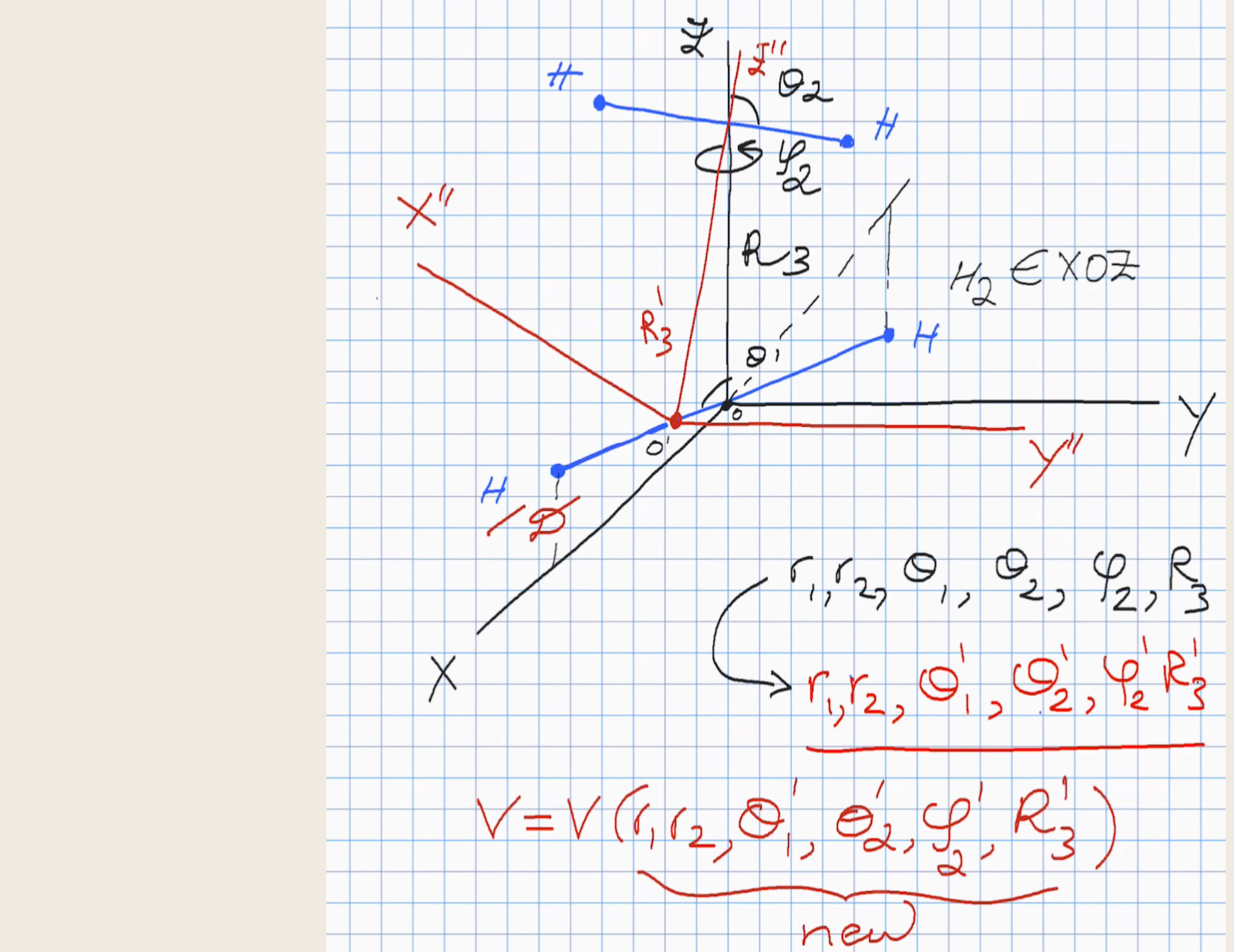
$$S^J(j'_1, j'_2, j'_{12} L'; j_1, j_2, j_{12}, L; E)|^2$$

$$\epsilon = E - B_1 j_1 (j_1 + 1) - B_2 j_2 (j_2 + 1)$$

$$k_{j_1 j_2 \rightarrow j'_1 j'_2}(T) = \sqrt{\frac{8k_B T}{\pi M_{12}}} \frac{1}{(k_B T)^2} \int_{\epsilon_s}^{\infty} \sigma_{j_1 j_2 \rightarrow j'_1 j'_2}(\epsilon) e^{-\epsilon/k_B T} \epsilon d\epsilon$$







$$\cos \eta = \frac{R_3'^2 + R_3^2 - x^2}{2R_3'R_3}$$

$$\sin \eta = \frac{x}{R_3' \sin \theta_1}$$

$$\begin{aligned}\theta'_1 &= \arccos(\cos \theta_1 \cos \eta + \sin \theta_1 \sin \eta), \\ \theta'_2 &= \arccos(\cos \theta_2 \cos \eta + \sin \theta_2 \sin \eta \cos \Phi_2), \\ \varphi'_2 &= \arccos \left( \cot \phi_2 \cos \eta + \frac{\cot \theta_1 \sin \eta}{\sin \phi_2} \right).\end{aligned}$$

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# RESULTS

$\text{HD} + \text{ortho-H}_2$  ,  $\text{HD} + \text{para-H}_2$  ,  $\text{H}_2 + \text{H}_2$

Original  $\text{H}_2\text{-H}_2$  PESs: DJ (2000), **BMKP (2002)**, R. Hinde (2008).

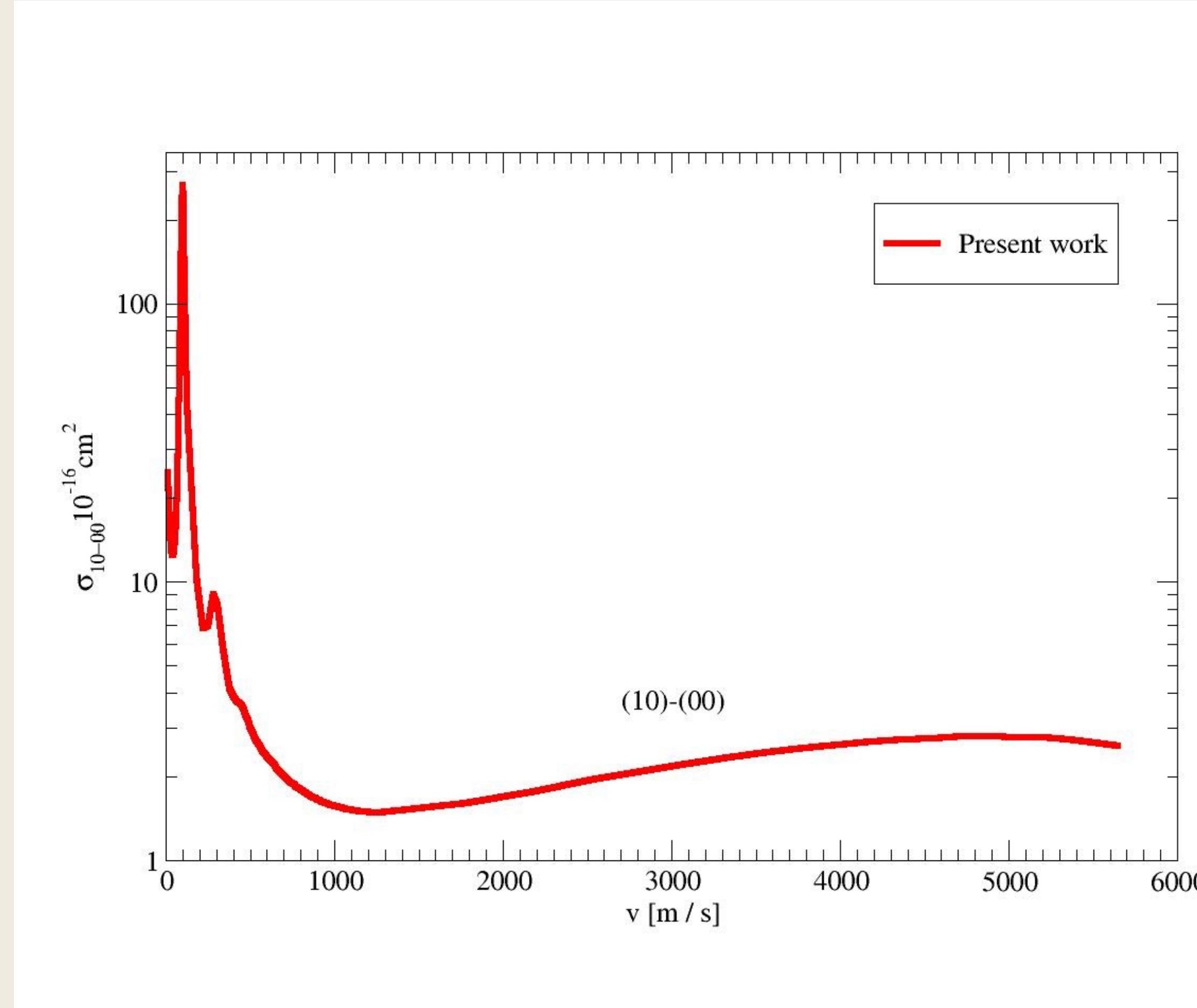
- 1) **A. I. Boothroyd, P. G. Martin, W. J. Keogh, and M. J. Peterson, J. Chem. Phys. 116, 666 (2002).**
- 2) **J. Schaefer, Astron. Astrophys. Suppl. Ser. 85 (1990) 1101.**
- 3) D. R. Flower, E. Roueff, Mon. Not. R. Astron. Soc. 309 (1999) 833.
- 4) D. R. Flower, J. Phys. B: At. Mol. Opt. Phys. 32 (1999) 1755.
- 5) D. R. Flower, Mon. Not. R. Astron. Soc. 318 (2000) 875.

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## **RECENT RESULTS: HD + H<sub>2</sub>, H<sub>2</sub> + H<sub>2</sub>**

- 1) Balakrishnan N, Croft J F E, Yang B H, Forrey R C, and Stancil P C, **ApJ** 2018 866 : 95.
- 2) Wan Y, Balakrishnan N, Yang B H, Forrey R C, and Stancil P C, **MNRAS** 2019 488, 381.
- 3) Wan Yier, **PhD Dissertation 2019**, The University of Georgia, Athens, Georgia:  
[http://getd.libs.uga.edu/pdfs/wan\\_yier\\_201908\\_phd.pdf](http://getd.libs.uga.edu/pdfs/wan_yier_201908_phd.pdf).

**HD ( $j_1=1$ ) + *para* - H<sub>2</sub> ( $j_2=0$ ): **(10) - (00)****  
**with BMKP PES (CPL 436 (2007) 19-24).**



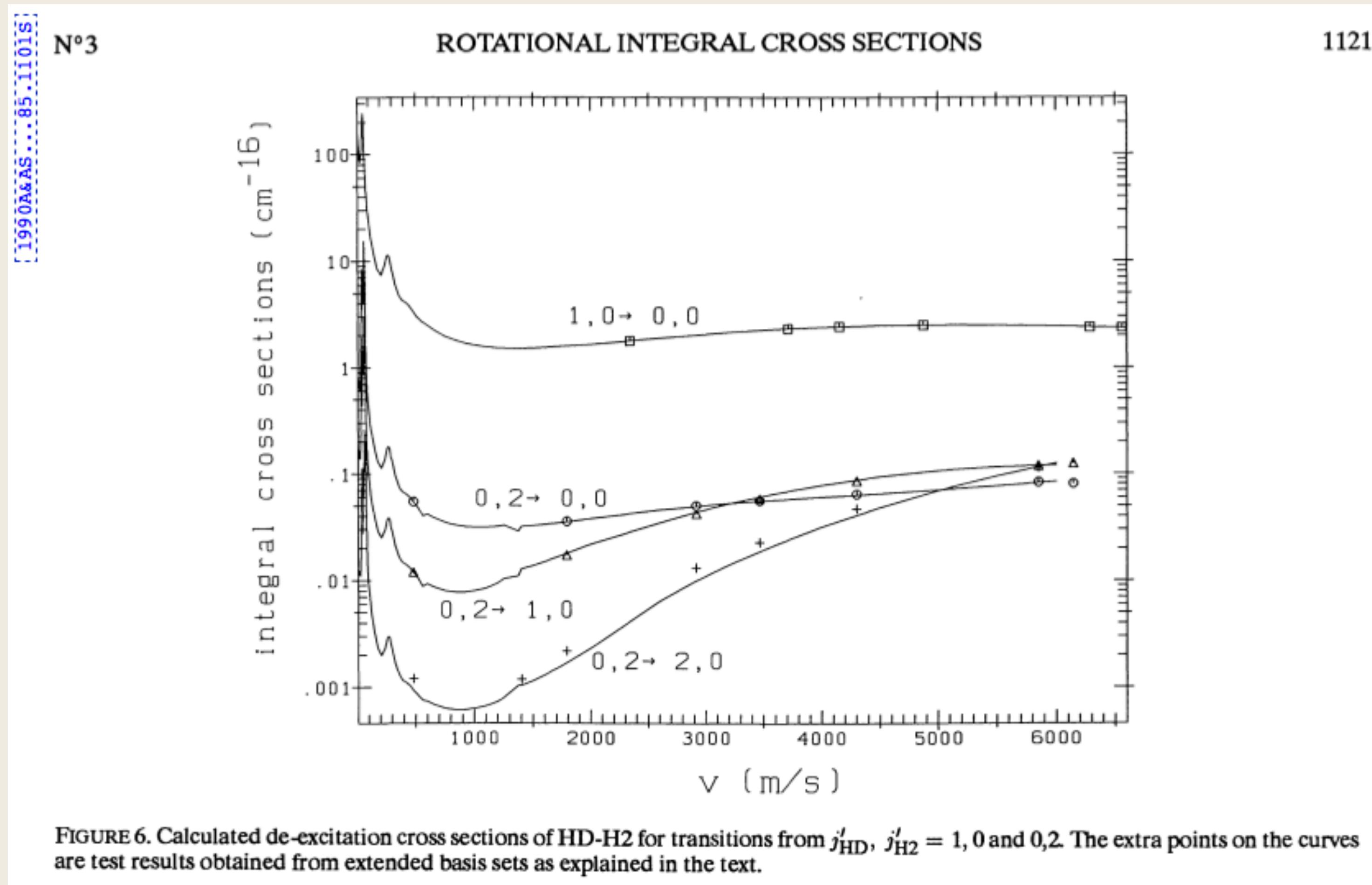
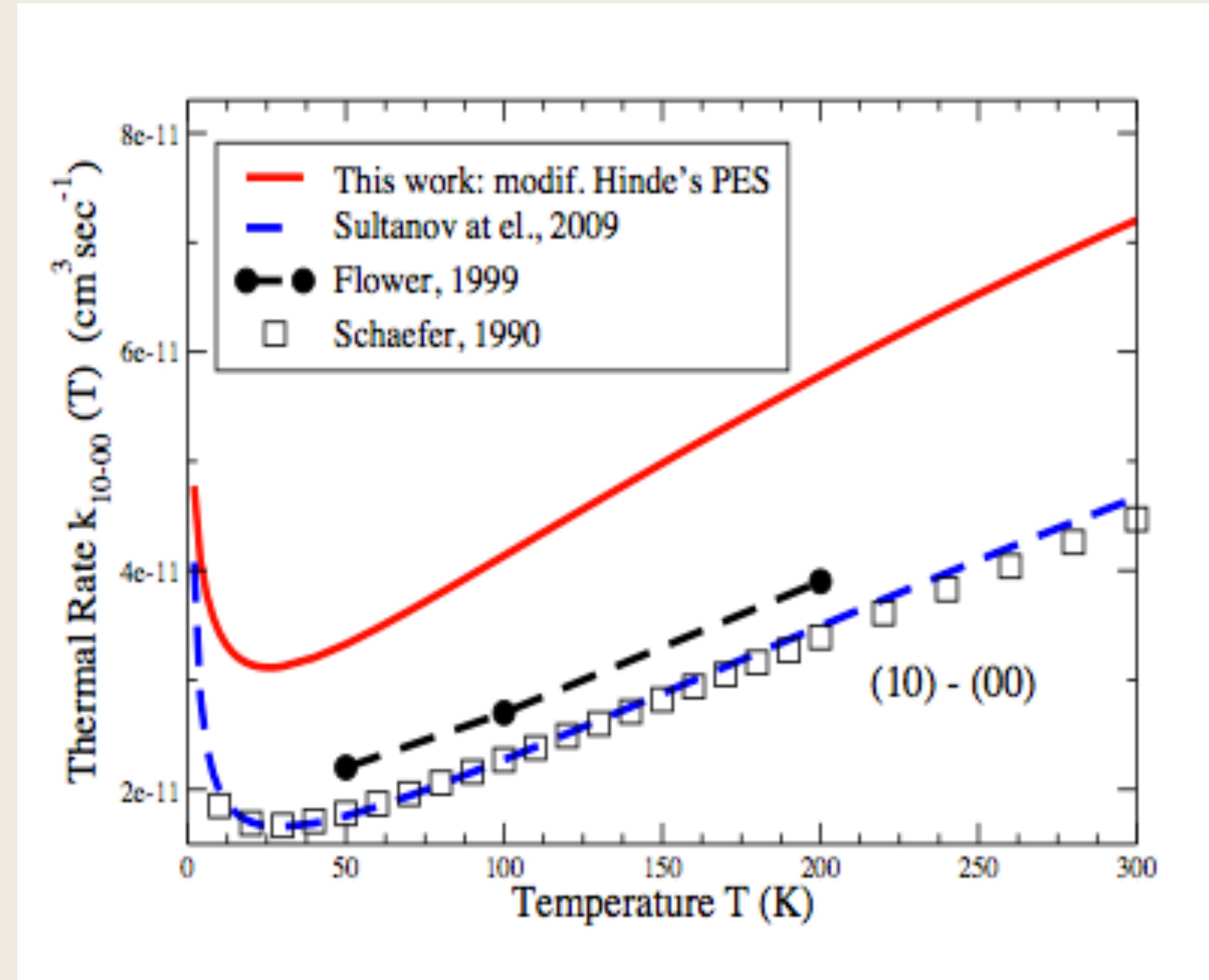
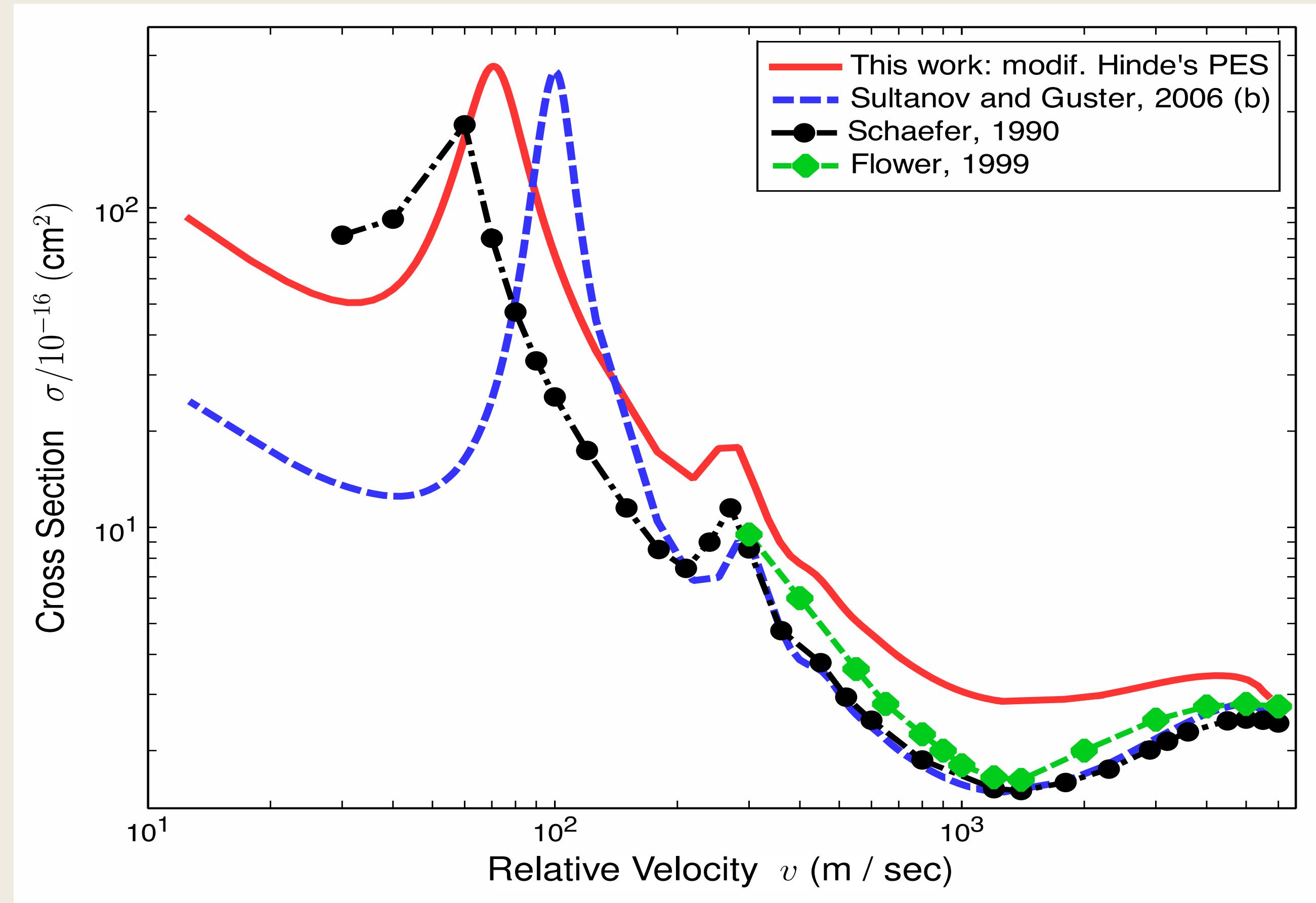


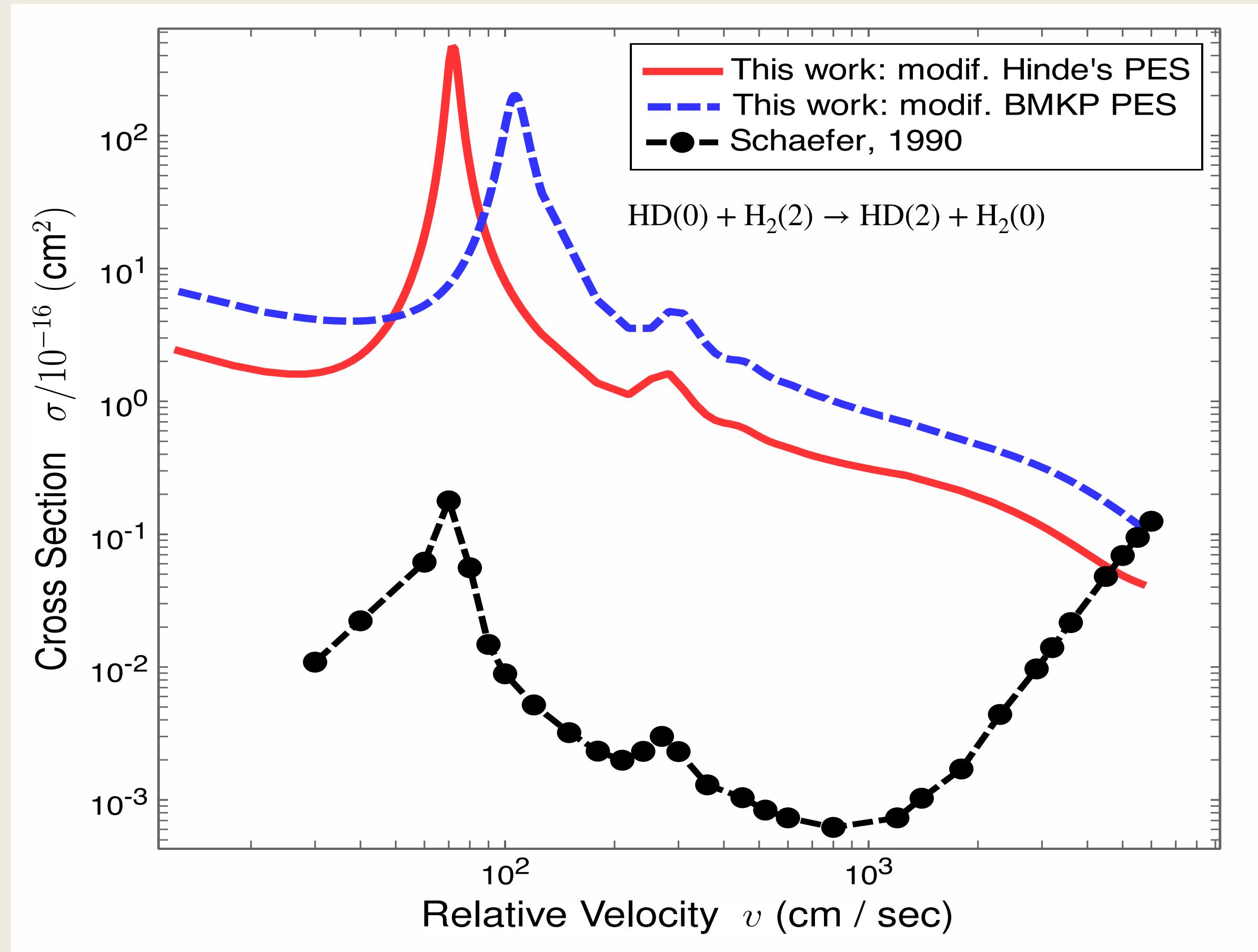
FIGURE 6. Calculated de-excitation cross sections of HD-H<sub>2</sub> for transitions from  $j'_{\text{HD}}$ ,  $j'_{\text{H}_2} = 1, 0$  and 0,2. The extra points on the curves are test results obtained from extended basis sets as explained in the text.

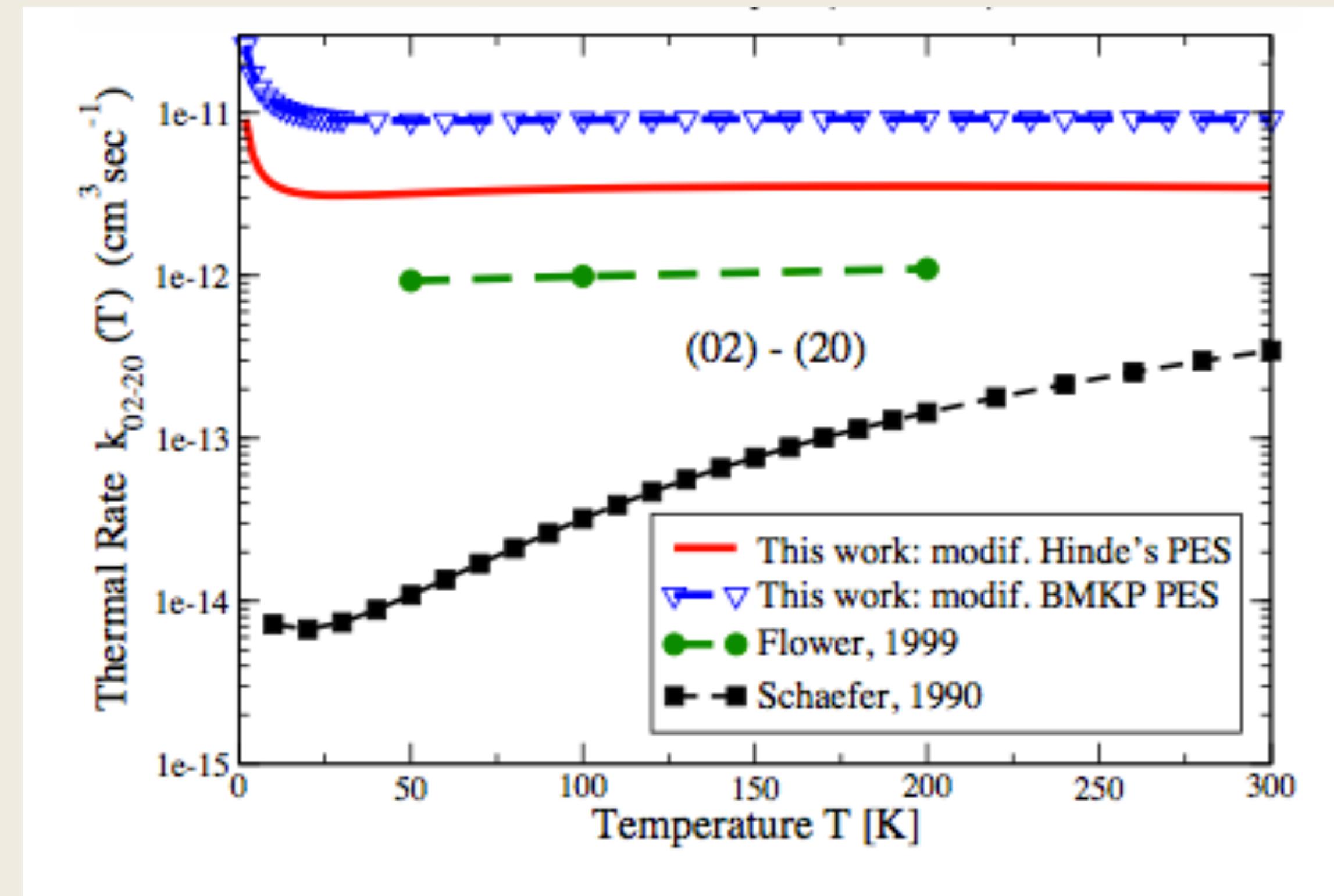


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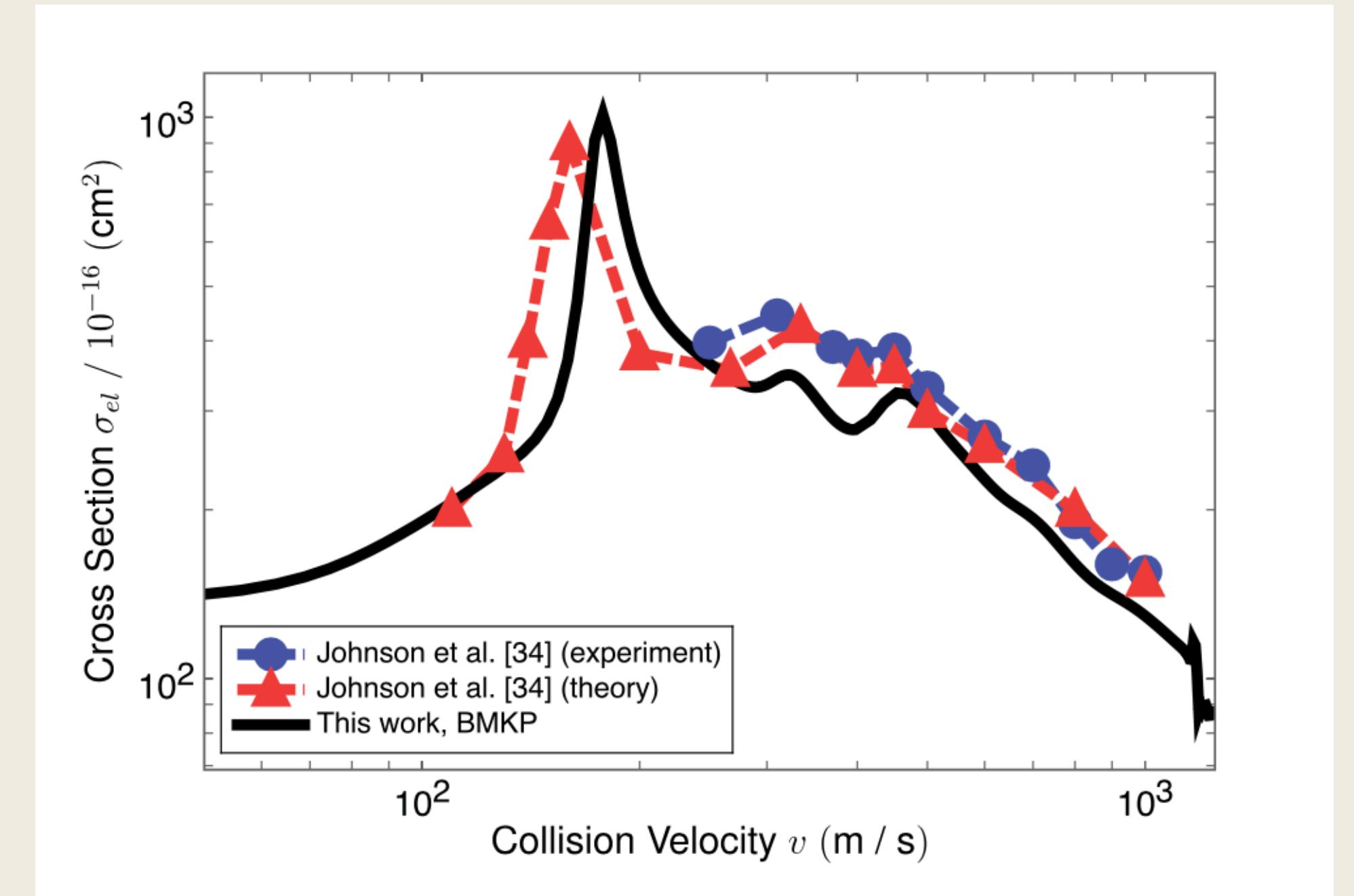
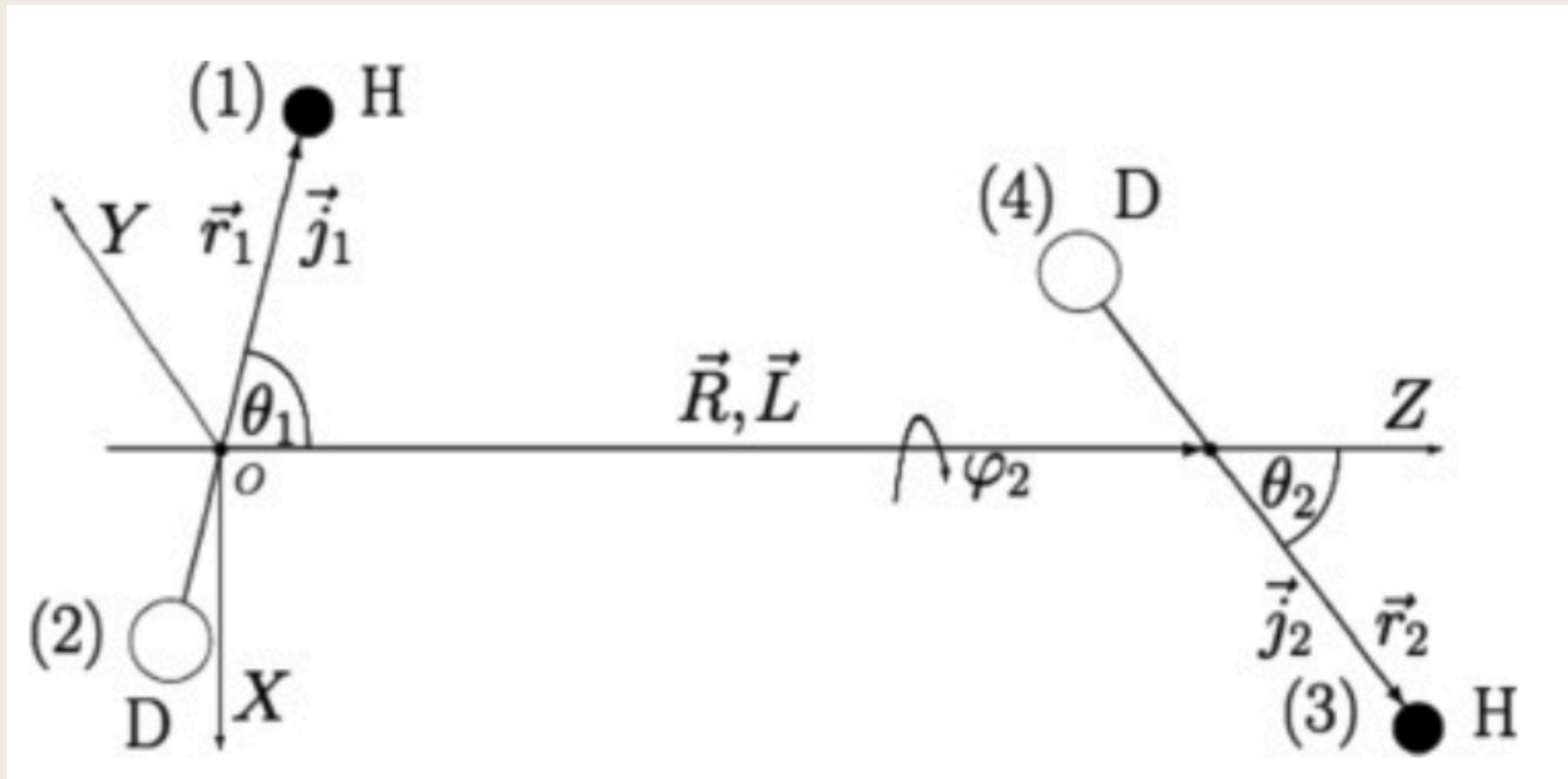
(J. Phys. B 49 (2016) 015203)

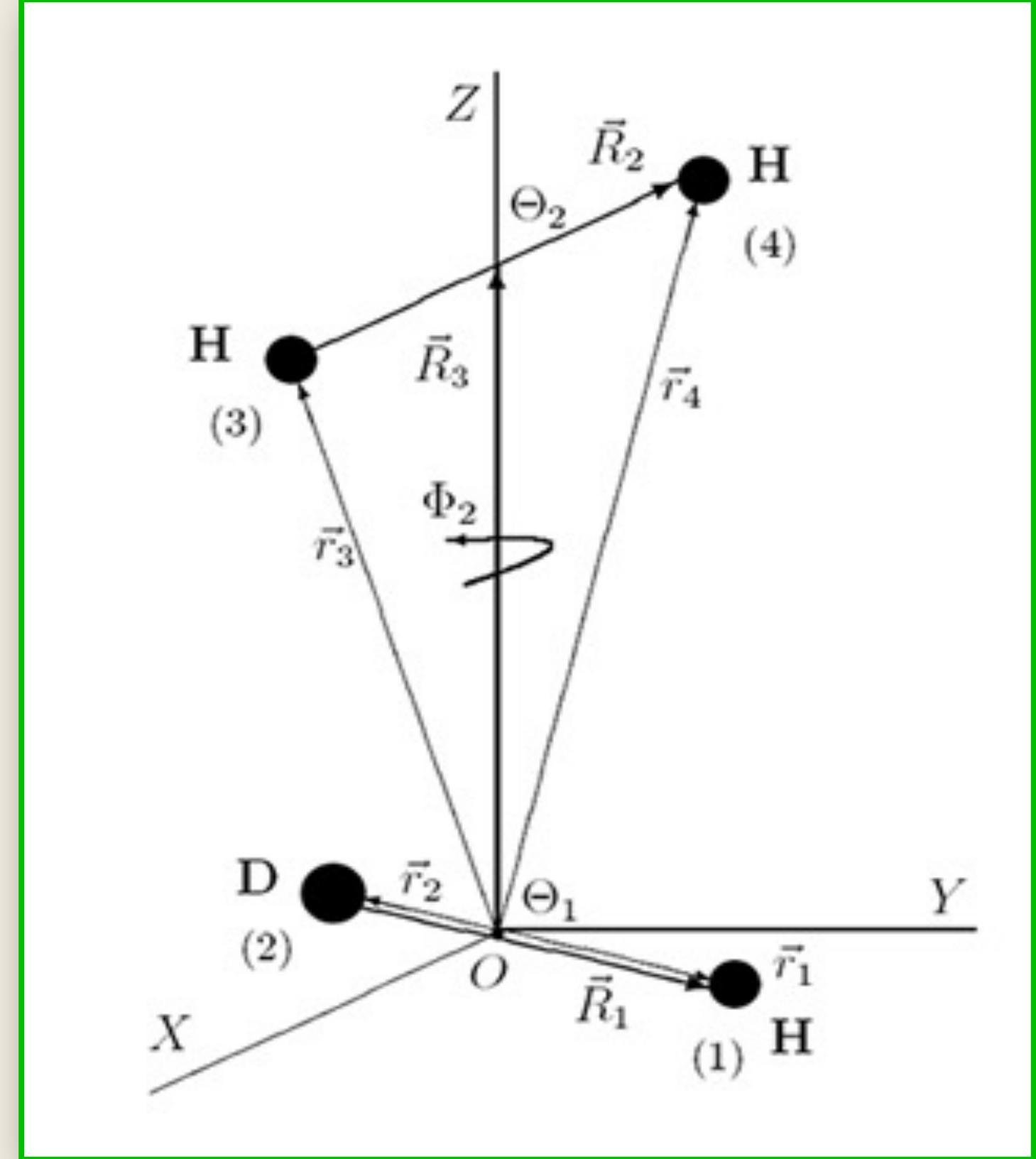
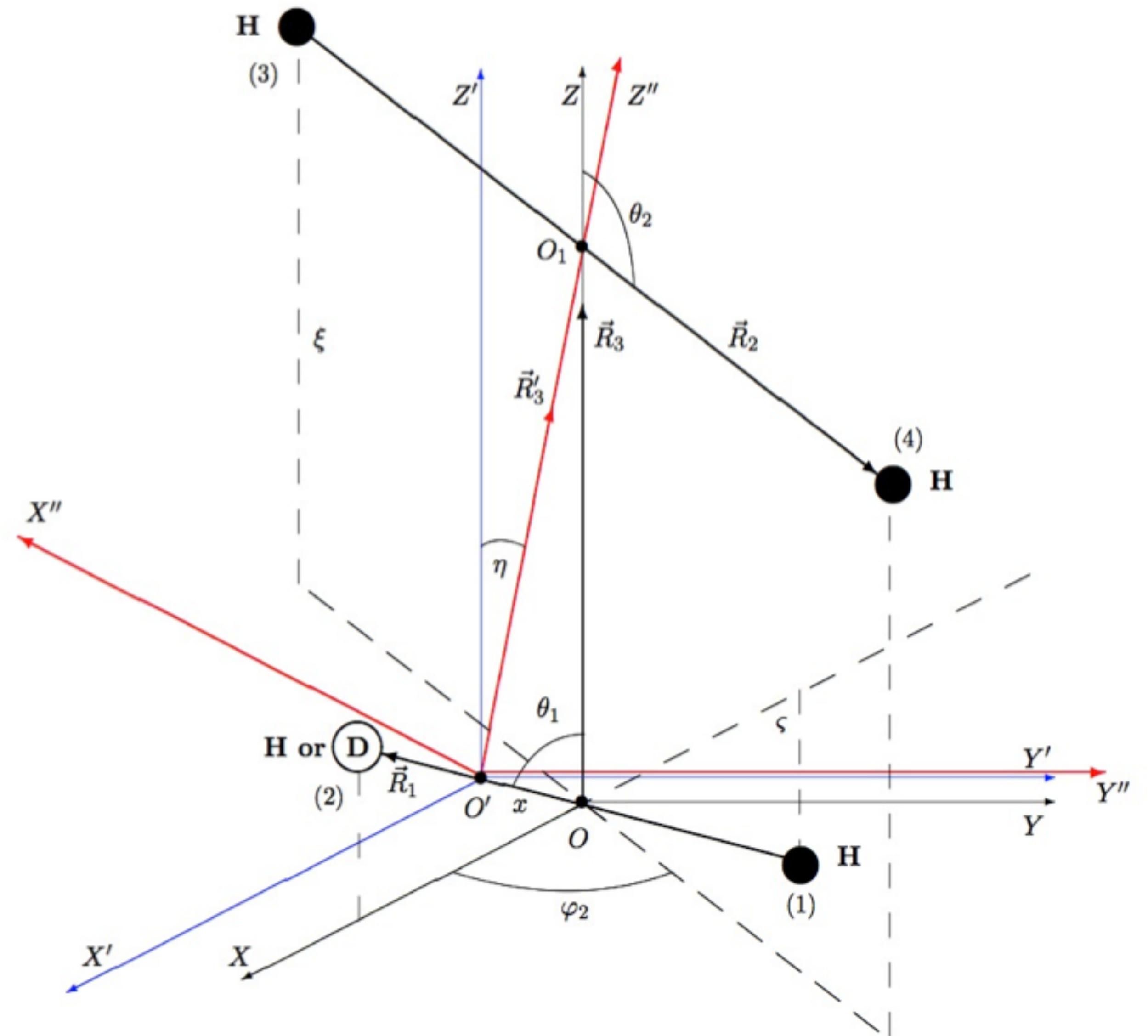




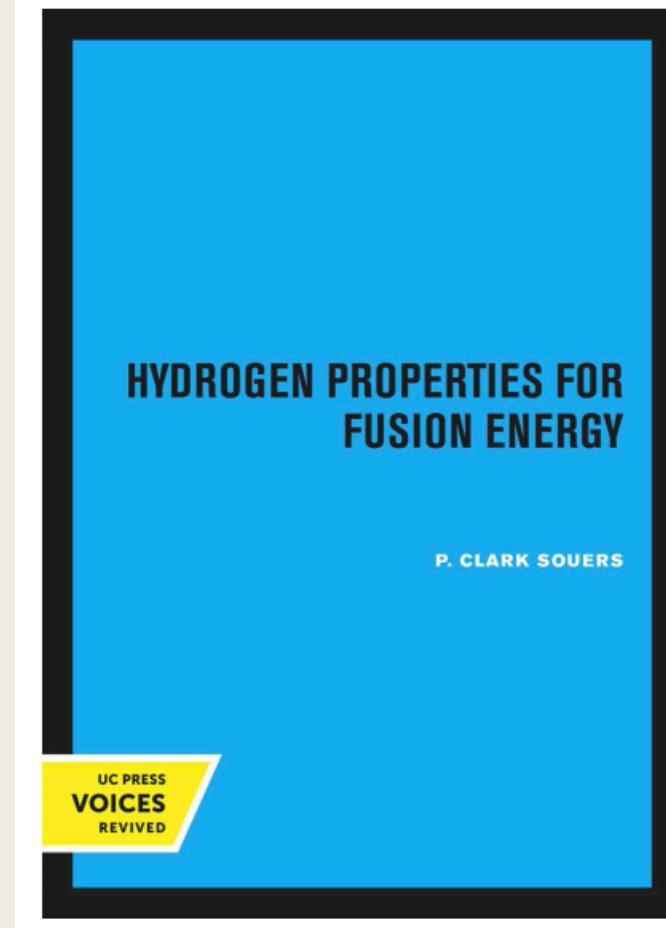


# HD + HD Collision





# HT + HT and DT + DT Collisions



🔓 Requires Authentication Published by University of California Press 1986

## Hydrogen Properties for Fusion Energy

P. Clark Souers

<https://doi.org/10.1525/9780520338401>



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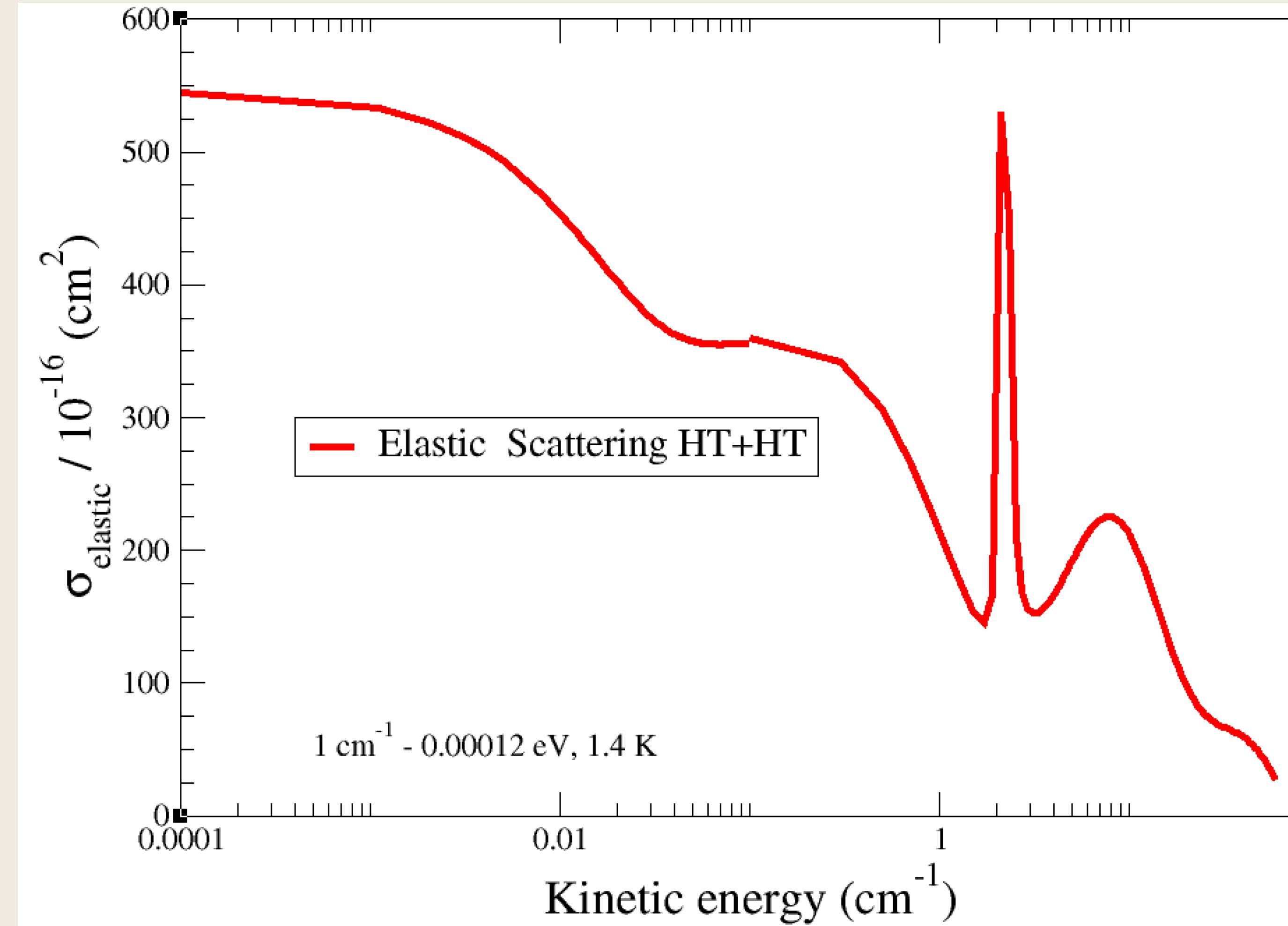
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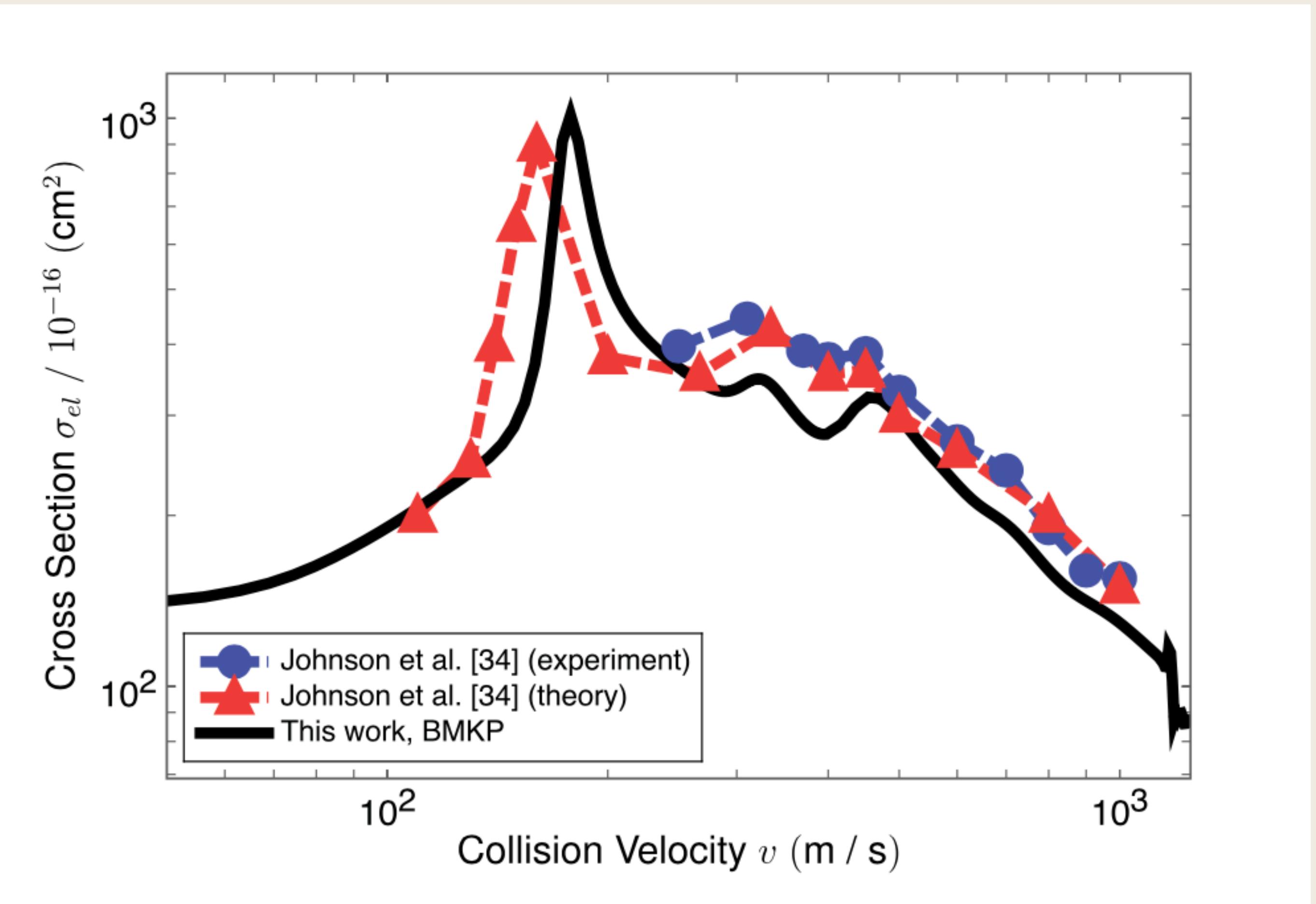
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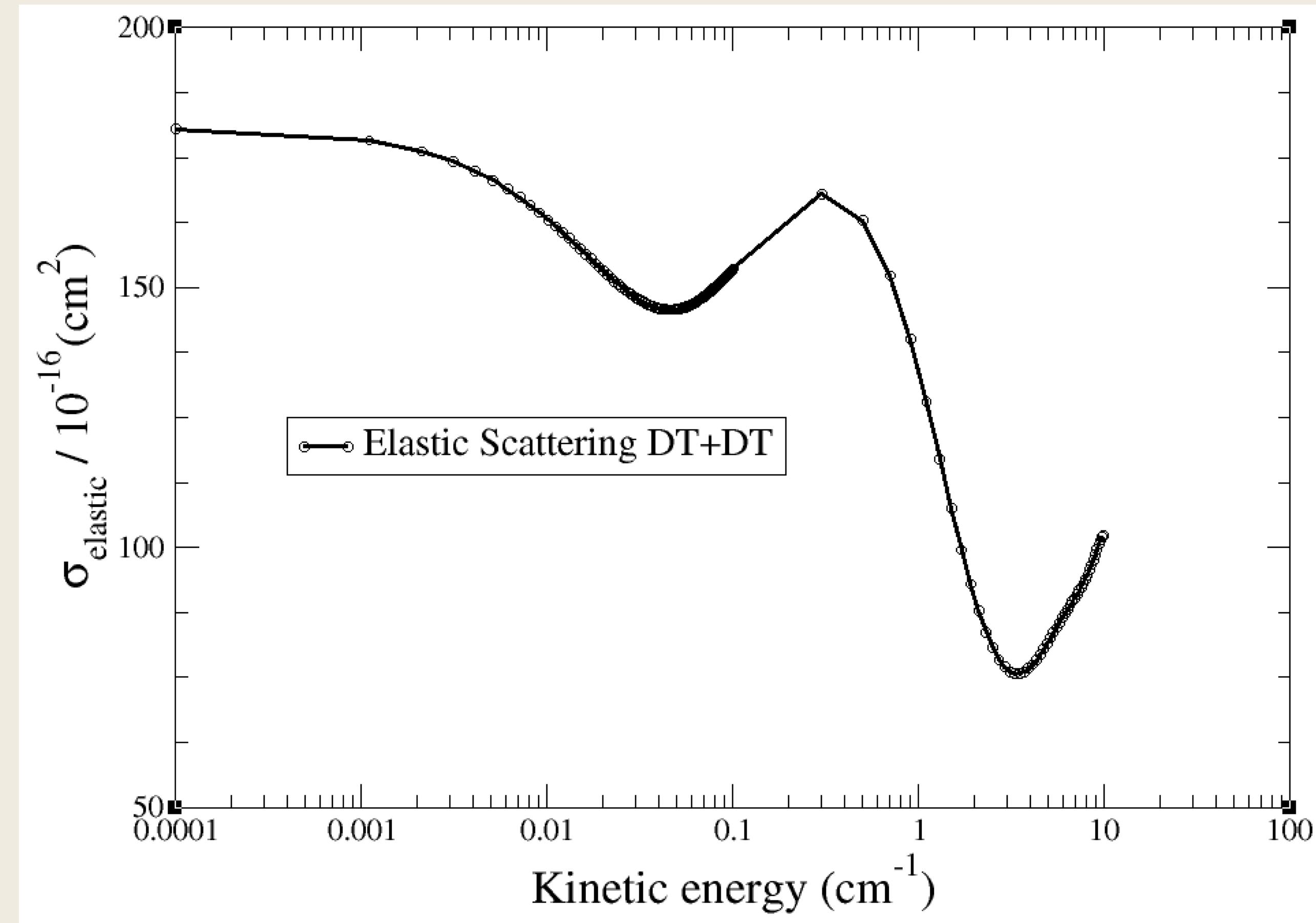
**Keywords:** physics; nuclear power; thermodynamics; energy technology; ceramic; low temperature physics; mechanical engineering; materials science

# HT + HT Elastic / Inelastic Scattering





# DT + DT    Elastic / Inelastic Scattering (calculations in progress...so far)



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# Thank You for Your Time & Attention!

- PLEASE LET ME KNOW IF YOU HAVE QUESTIONS!
- Thank you very much!
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