

# **Estimate of the muon $g-2$ : Update of the HLS Model Prediction**

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**LPNHE Paris 6/7**

**Mainz WG  $g-2$  Theory Initiative (June 2018)**

# OUTLINE

- Why an Update ?
- $e^+e^- \rightarrow K^+K^-$ , within HLS : **CMD2/BaBar/CMD3**
- **Analysis of  $e^+e^- \rightarrow K^+K^-$ : CMD2**
- **Analysis of  $e^+e^- \rightarrow K^+K^-$ : BaBar**
- **Analysis of  $e^+e^- \rightarrow K^+K^-$ : CMD3**
- Modified  $(g-2)_\mu$  estimates
- Conclusions

# Update of the HLS Results

- **Main Focus : the  $K^+K^-$  Sector**

- **Use of Exponentiated Coulomb Term &  $FSR_K$**

A.Hoefer *et al.* EPJC 24 (2002) 51

- **Reexamining the  $K^+K^-$  Data within HLS :**

K+K- : CMD2 Phys. Lett. B669 (2018) 217

K+K- : BaBar Phys. Rev. D88 (2013) 032013

K+K- : CMD3 Phys. Lett. B779 (2018) 64

$K_L K_S$ : CMD3 Phys. Lett. B760 (2016) 314

# Data Handled via **Global Fits**

- Almost all ( $\approx 50 - 60$ ) existing data samples used
- **$e^+e^-$  annihil.** &  **$\tau$  spectra** &  **$\pi$  & K spacelike FF's**
- Devoted treatment of **normalization errors**  
(**unbiasing method**)

M. Benayoun *et al.* EPJ C75 (2015) 613

- **Update : Exponentiated Coulomb Term & FSR<sub>K</sub>**
- **Main focus of this update : the  $K^+ K^-$  sector**

$\pi^+ \pi^-$  : KLOE 2 JHEP 1803 (2018) 17

# Revisiting $e^+e^- \rightarrow K^+ K^-$ within BHLS2

- Method : **global fit** of the HLS channels/Data

$e^+e^- \rightarrow \pi\pi / K_L K_S / \pi\gamma / \eta\gamma / \pi\pi\pi$  &  $\tau \rightarrow \pi\pi \nu_\tau$   
&  $PV\gamma, P\gamma\gamma$  decays &  $\pi^+ & K^+$  FF's at  $s < 0$

Together with CMD2/CMD3/BaBar  $e^+e^- \rightarrow K^+ K^-$   
data samples **each in turn**

Presently : discard SND  $K^+ K^-$  spectra ( $\approx 7\%$  syst)

$\rho$  &  $\omega$  background beneath the  $\phi$  : 100% controlled  
by the  $e^+e^- \rightarrow \pi^+ \pi^- (K_L K_S)$  data/channel

# Systematics in $e^+e^- \rightarrow K^+ K^-$ Data

## Systematics (1.00 to 1.03 GeV Region)

- CMD2 :  $\sigma = 2.2 \%$
- BaBar :  $\sigma = (E < 1.01 \text{ GeV}) 0.82 \%$   $(E > 1.01 \text{ GeV}) 0.72 \%$
- CMD3 :  $\sigma = 2.0 \%$

**1/ Systematics : point to point correlated**  
(normalization uncertainties)

**2/ Normalization correction :**

sampled on a gaussian pdf  $D_{\text{norm}}(m=0, \sigma)$

**3/ Each Experiment : one sampling  $\lambda$  of  $D_{\text{norm}}$ .**

# Scale Uncertainty & Correction

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- **Normalization correction :**

$$\lambda = \left\{ M^T V^{-1} [ m - M ] \right\} / \left\{ M^T V^{-1} M + \frac{1}{\sigma^2} \right\}$$

Unbiasing data **m** vs Model /Ref. **M** :  $m \rightarrow \boxed{m - \lambda M}$

**V** = covariance of **uncorrelated** errors

- **$(m \text{ vs } M)$  Consistency Condition** :  $\lambda / \sigma < 3$

# CMD2 $e^+e^- \rightarrow K^+ K^-$ & global FIT

CMD2  $K^+ K^-$  : (incl. 1995)

Rescalings : 1.005/1.013

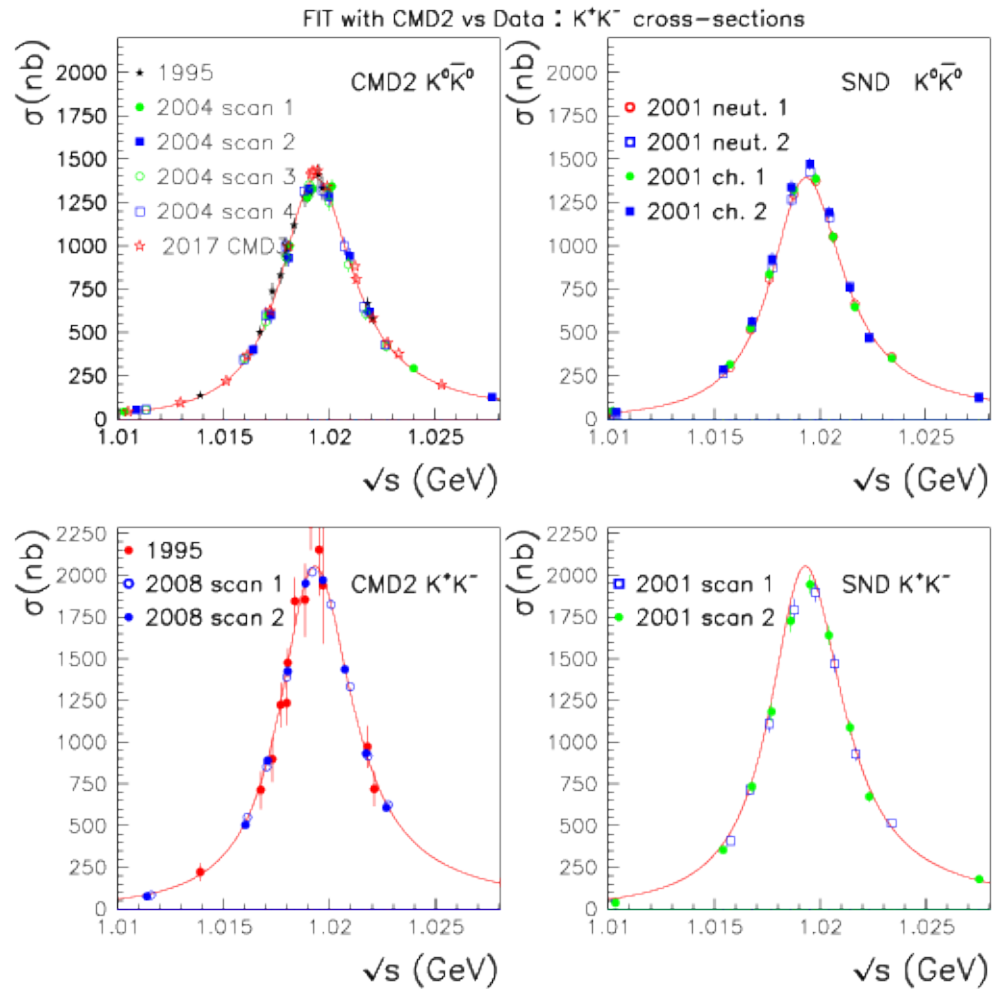
•  $\chi^2/N_p = 31.28/36$  (95.4%)

photon VP from DN

FSI (FSR+Coulomb): FJ

$E < 1.05$  GeV :

$a_\mu(K^+ K^-) = 17.28 \pm 0.29$





# $e^+e^- \rightarrow K^+ K^-$ : CMD3 & BaBar vs CMD2

Bulk of data samples for several channels  
in the  $\phi$  region ( $K_L K_S / \pi\gamma / \eta\gamma / \pi\pi\pi$ ) almost solely  
collected by CMD2 & SND :

Natural choice for Reference Energy : CMD2 & SND

Assumption on Energy calibrations around the  $\phi$  :

$$\triangleright E_{\text{CMD2}} = E_{\text{CMD3}} + \delta E_{\text{CMD3}}$$

$$\triangleright E_{\text{CMD2}} = E_{\text{BaBar}} + \delta E_{\text{BaBar}}$$

Energy shifts  $\delta E_{\text{CMD3}}$  &  $\delta E_{\text{BaBar}}$  submitted to fit

# BaBar $e^+e^- \rightarrow K^+ K^-$ & global fit

BaBar Data :

Fit up to 1.025 GeV

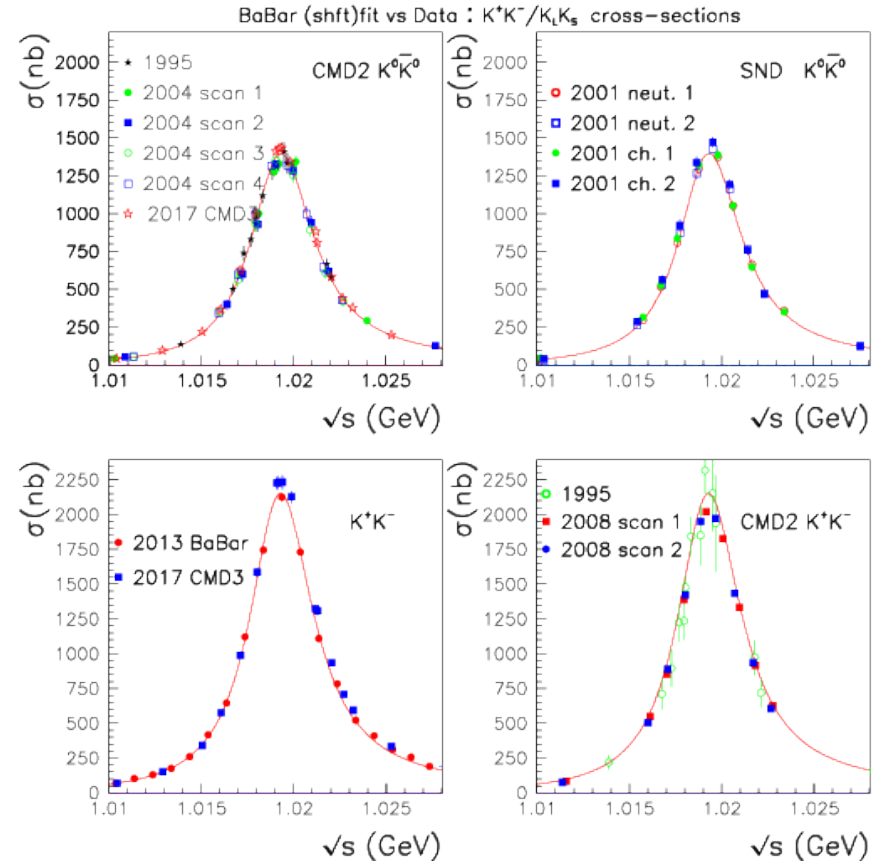
DN's VP:

- $\chi^2/N_p = 47.5/27$  (89.2%)

$E < 1.05$  GeV :

- $a_\mu(K^+ K^-) = 18.12 \pm 0.21$

$$a_\mu(\text{BaBar } < 1.06) = 18.64 \pm 0.16_{\text{stat}} \pm 0.13_{\text{syst}}$$



# BaBar $e^+e^- \rightarrow K^+ K^-$ & global fit II

(Meas-Fit)/Fit

**BaBar Data :**

**Fit up to 1.025 GeV**

**Unshifted Data :**

$\chi^2/N_p = 83.0/27$  (43.2%)

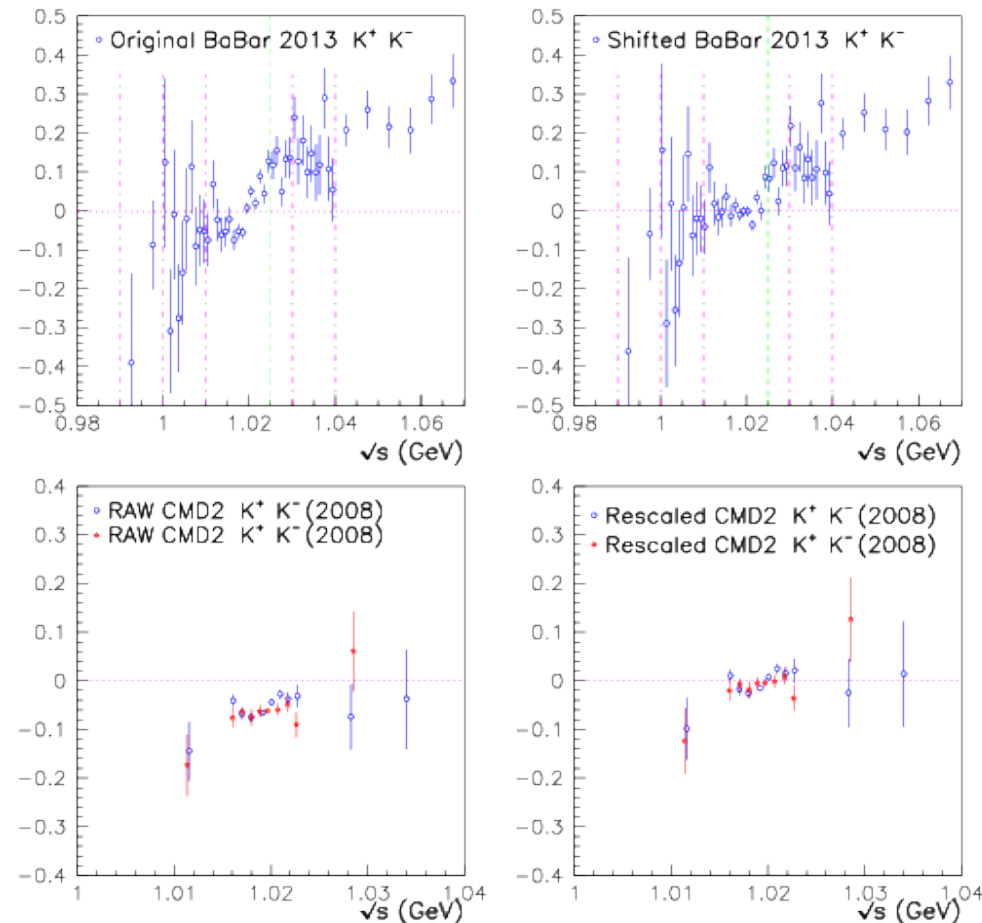
**Shifted Data :**

$\chi^2/N_p = 47.5/27$  (89.2%)

**CMD2 Rescaling :  $\approx 1.05$**

**only a  $\approx 2.2 \sigma$  effect**

BaBar Fit vs Data : Uncorrected Normalized Residuals



# BaBar $e^+e^- \rightarrow K^+ K^-$ & global fit III

BaBar Data :

Fit up to 1.025 GeV

DN's VP:

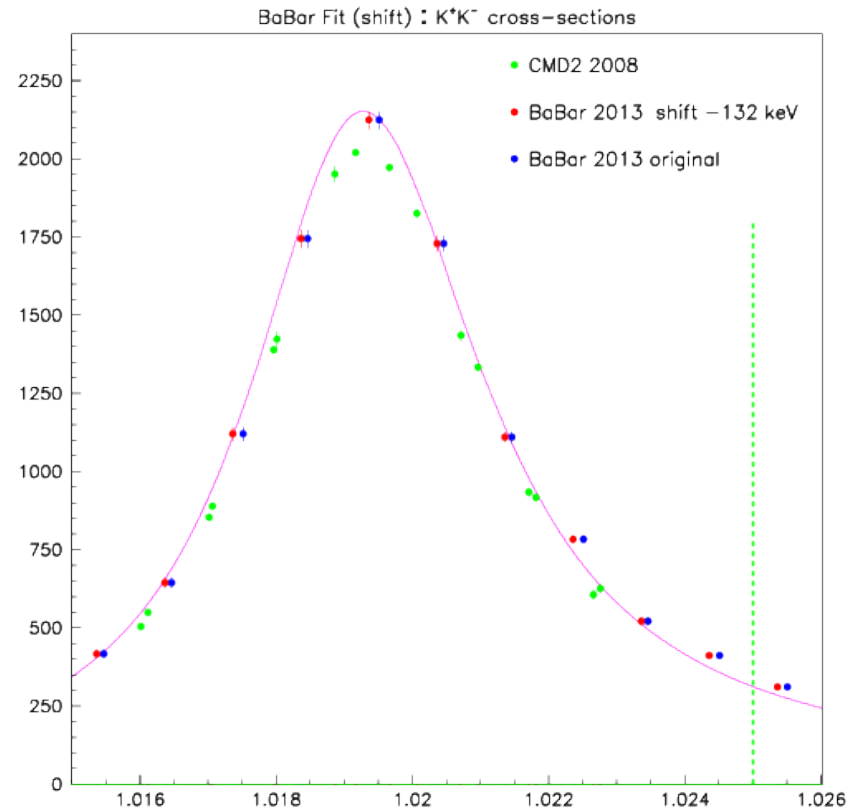
- $\delta E_{\text{BaBar}} = -132.0 \pm 19.1 \text{ keV}$

FJ's VP:

- $\delta E_{\text{BaBar}} = -121.0 \pm 16.2 \text{ keV}$

FI's VP:

- $\delta E_{\text{BaBar}} = -135.1 \pm 21.2 \text{ keV}$



# BaBar $e^+e^- \rightarrow K^+ K^-$ & global fit III

BaBar Data :

Fit up to 1.025 GeV

DN's VP:

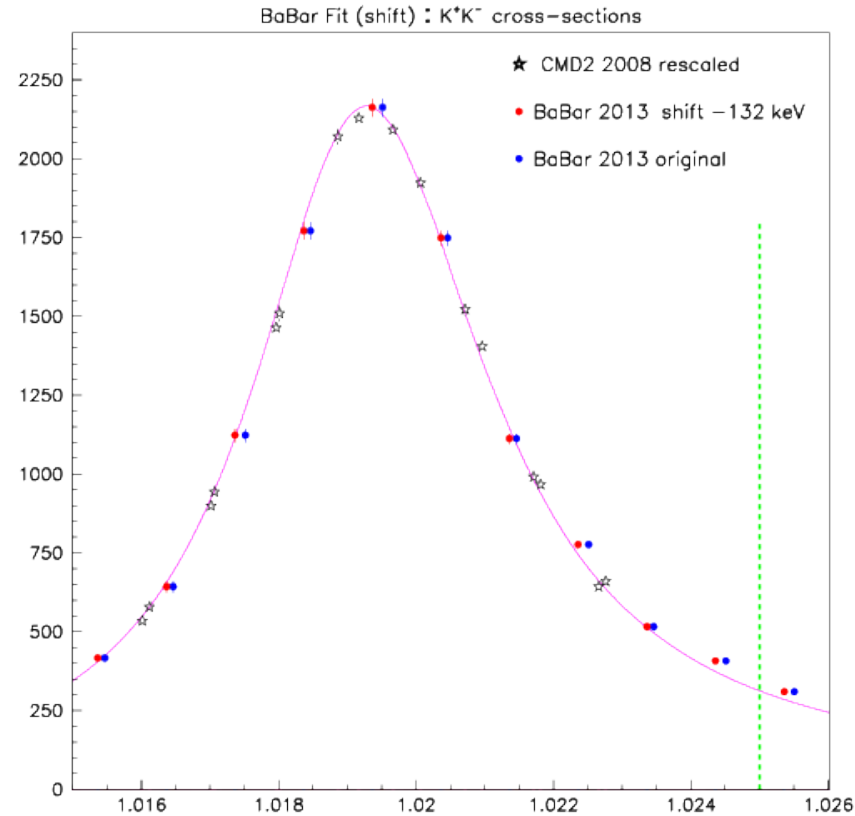
- $\delta E_{\text{BaBar}} = -132.0 \pm 19.1 \text{ keV}$

FJ's VP:

- $\delta E_{\text{BaBar}} = -121.0 \pm 16.2 \text{ keV}$

FI's VP:

- $\delta E_{\text{BaBar}} = -135.1 \pm 21.2 \text{ keV}$



# $a_\mu$ from BHLS2 Fits (update I)

HLS channels :

FSI updated

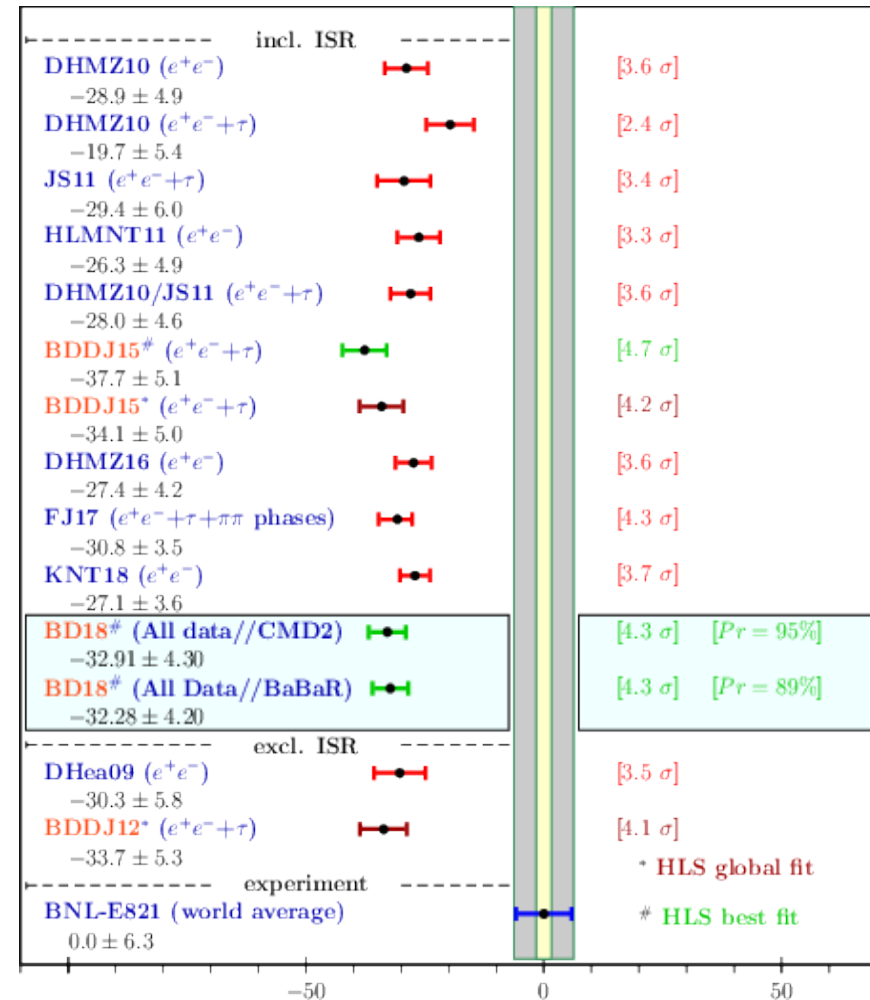
Non HLS channels :

HVP update from FJ

+ ADD. SYST.

BHLS

BHLS2



$$[a_\mu - a_\mu(Exp)] \times 10^{10}$$

$$[a_\mu(Exp) \times 10^{10}] = 11659209.1$$

# Global fit : CMD2 & BaBar $e^+e^- \rightarrow K^+K^-$

Data :

BaBar  $E < 1.025$  GeV

CMD2 (2 scans)

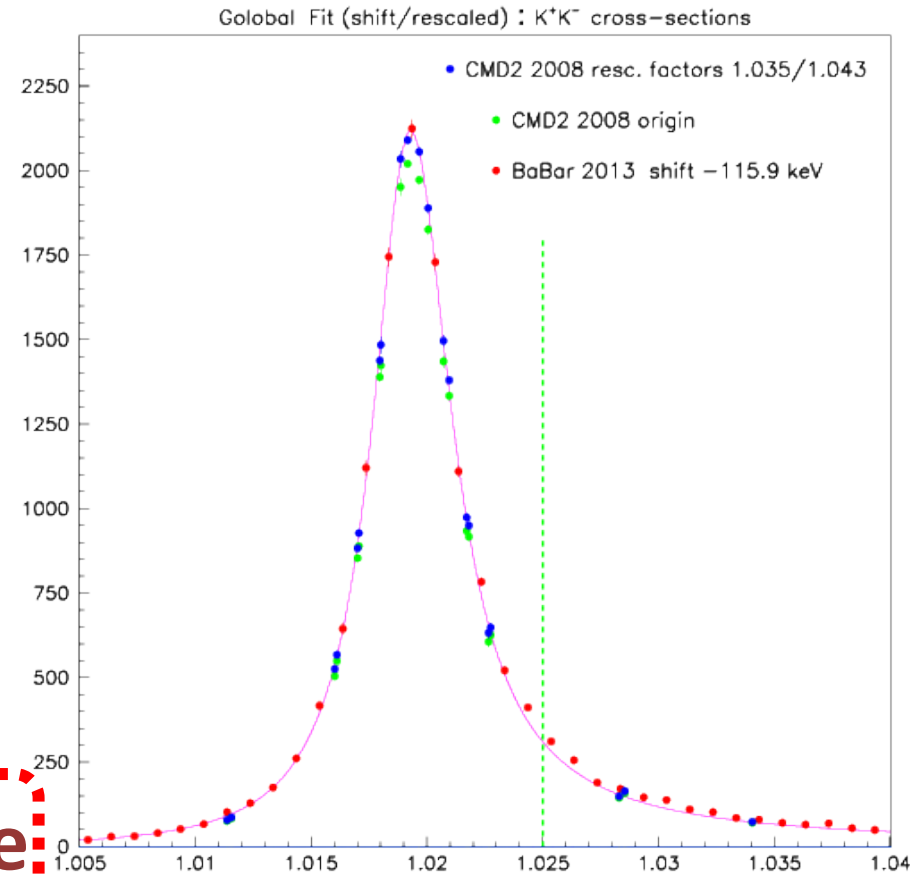
Fit with DN's VP: (85.8%)

•  $\chi^2/N_p(\text{CMD2})=36.4/36$

•  $\chi^2/N_p(\text{BaBar})=49.7/27$

Reminder:

$\rho$ - $\omega$  bkg (&  $\phi$  pole) not free



# Global fit : CMD2 & BaBar $e^+e^- \rightarrow K^+K^-$

BaBar & CMD2 Data :

$$\delta E_{\text{BaBar}} = -115.9 \pm 17.4 \text{ keV}$$

(BaBar Fit  $\phi$  mass vs CMD2:

$$\Delta m = (-)99 \pm 8_{\text{CMD2}} \pm 13_{\text{BaBar}} \text{ keV}$$

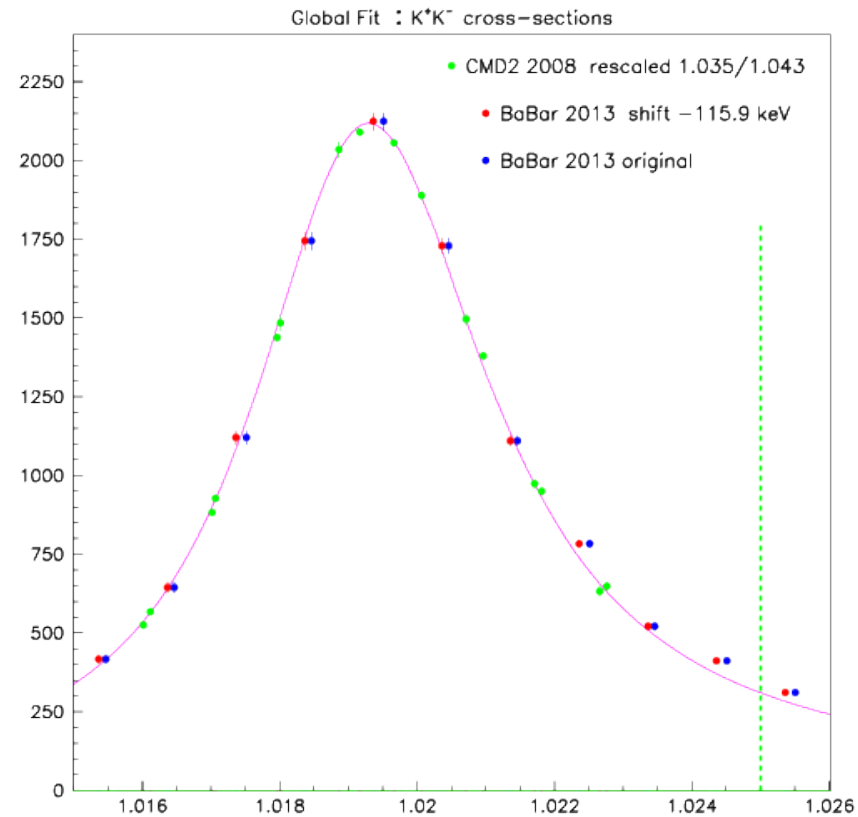
CMD2 Rescaling factor :

$$1.041 \pm 0.005$$

Still a  $\approx 2 \sigma$  effect !

$E < 1.05 \text{ GeV}$  :

- $a_\mu(K^+ K^-) = 17.83 \pm 0.20$





# Global fit : CMD2 & BaBar $e^+e^- \rightarrow K^+K^-$

BaBar & CMD2 Data :

$$\delta E_{\text{BaBar}} = -115.9 \pm 17.4 \text{ keV}$$

CMD2 Rescaling factor :

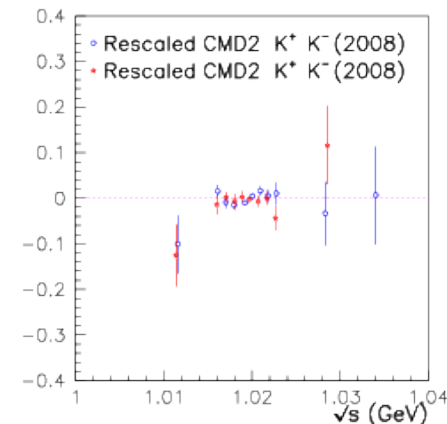
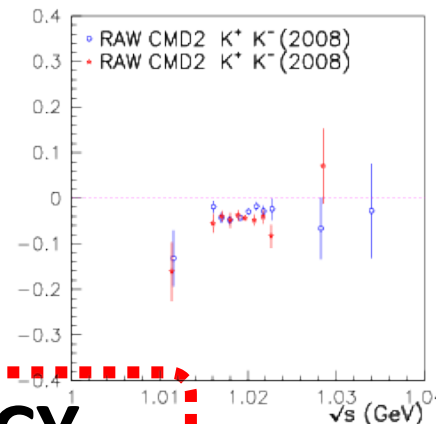
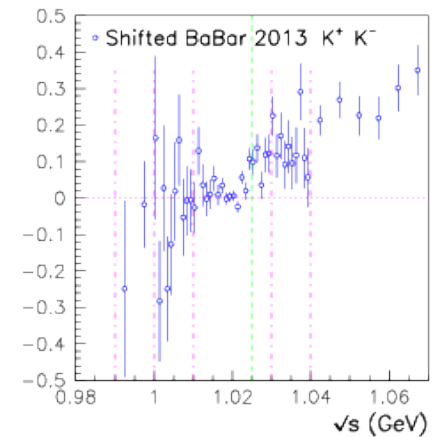
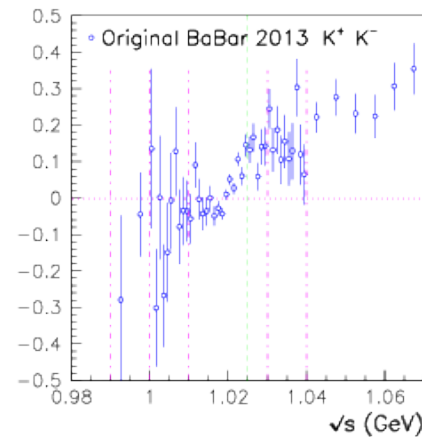
$$1.041 \pm 0.005$$

Still a  $\approx 2. \sigma$  effect !

Fit Prob. 85.8%

**NO OBVIOUS INCONSISTENCY**

Global Fit vs Data : Normalized Residuals



# $a_\mu$ from BHLS2 Fits (update II)

HLS channels :

FSI updated

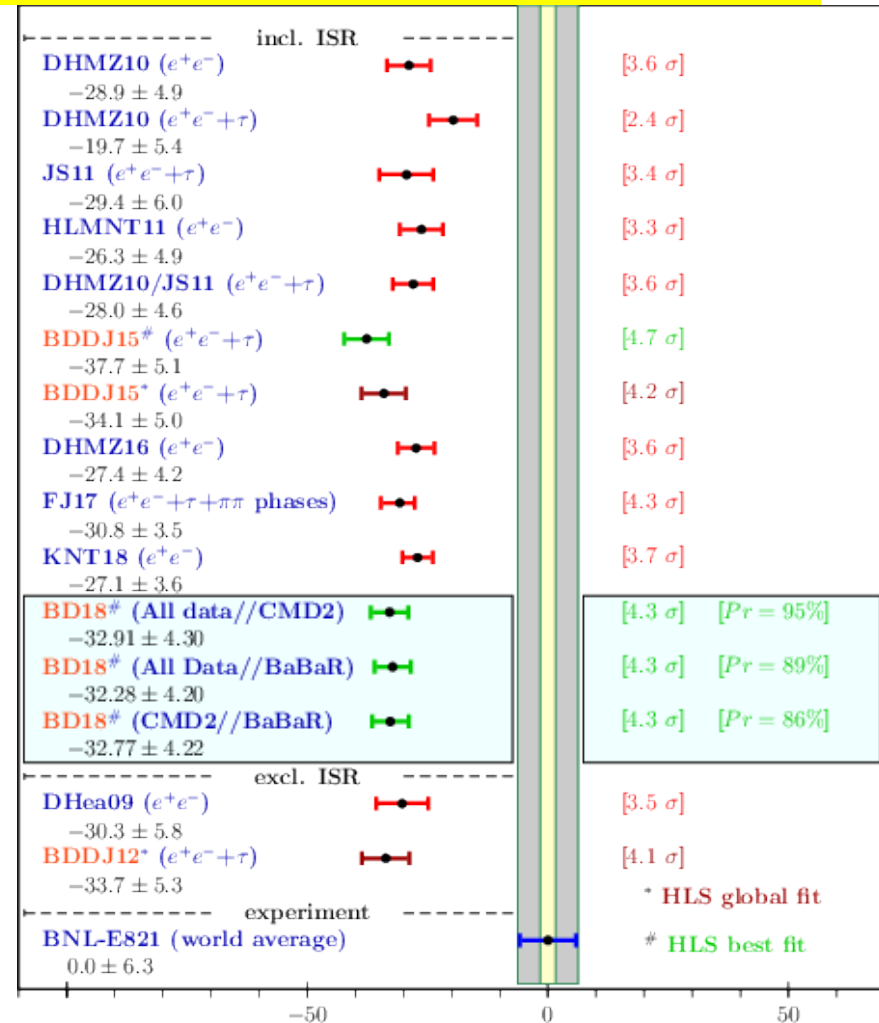
Non HLS channels :

HVP update from FJ

+ ADD. SYST.

BHLS

BHLS2



$[a_\mu - a_\mu(Exp)] \times 10^{10}$   
 $[a_\mu(Exp) \times 10^{10} = 11659209.1]$

# CMD3 $e^+e^- \rightarrow K^+ K^-$ & global fit I

**CMD3 Data :**

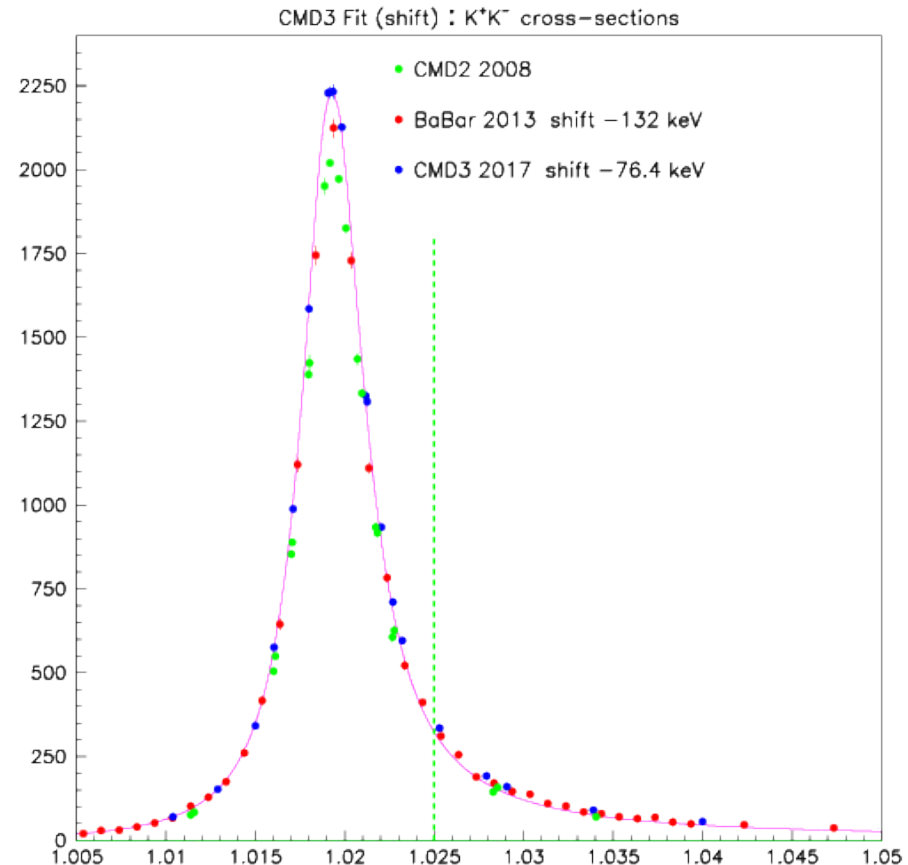
**Fit up to 1.025 GeV**

**Most likely fit : FJ's VP**

- $\chi^2/N_p = 29.7/16$  (87.7%)
- $\delta E_{\text{CMD3}} = -76.4 \pm 9.9$  keV

**$E < 1.05$  GeV :**

- $a_\mu(K^+ K^-) = 18.75 \pm 0.16$
- $a_\mu(\text{CMD3} < 1.06) = 19.33 \pm 0.40$



# CMD3 $e^+e^- \rightarrow K^+ K^-$ & global fit II

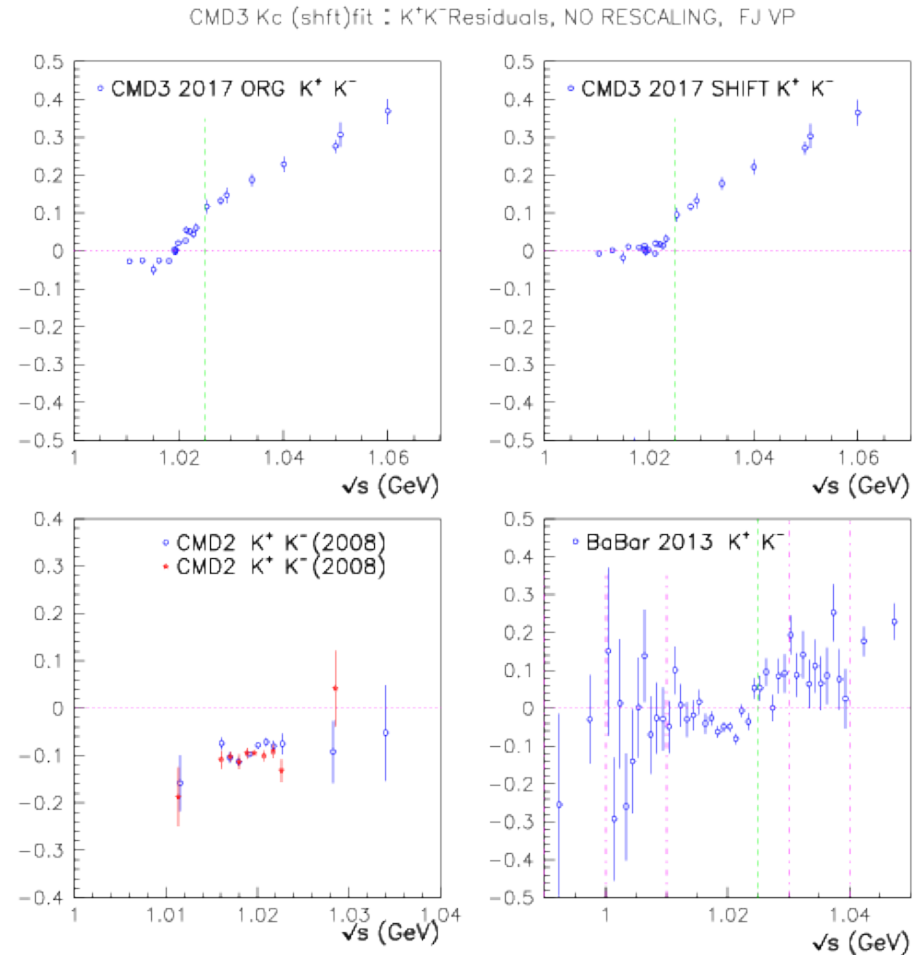
CMD3 Data :

Fit up to 1.025 GeV

FJ's VP:

- CMD2 scale  $\approx 1.10$   
( $\approx 4.5 \sigma_{\text{CMD2}}$ )
- BaBar scale  $\approx 1.034$   
( $\approx 4. \sigma_{\text{BaBar}}$ )

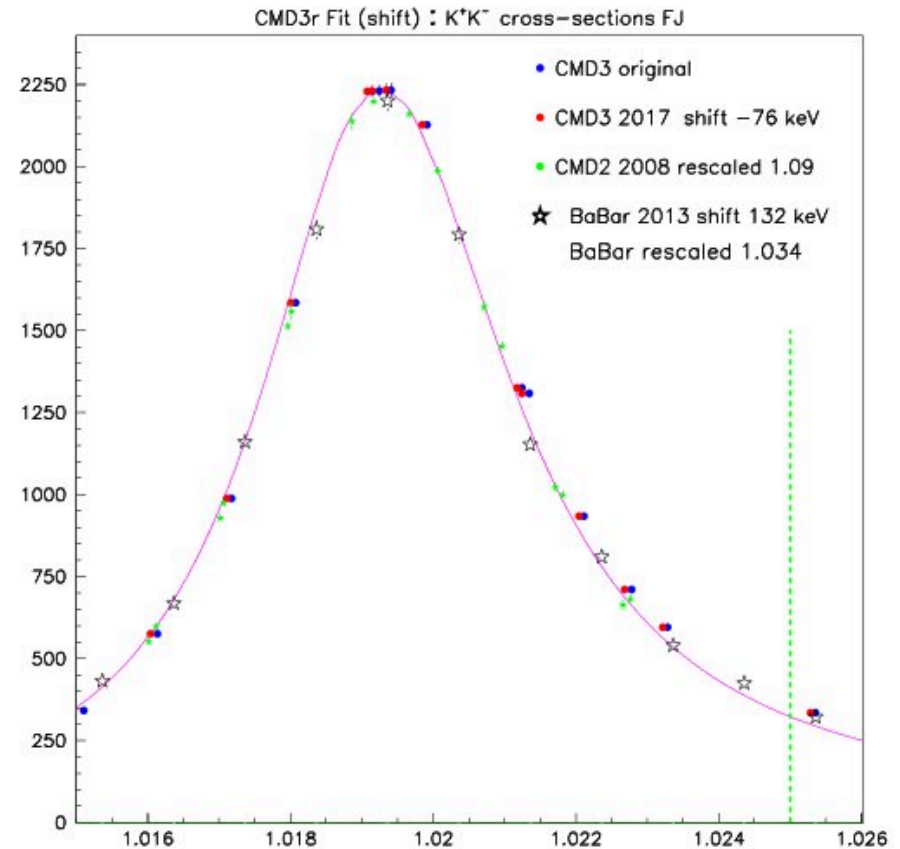
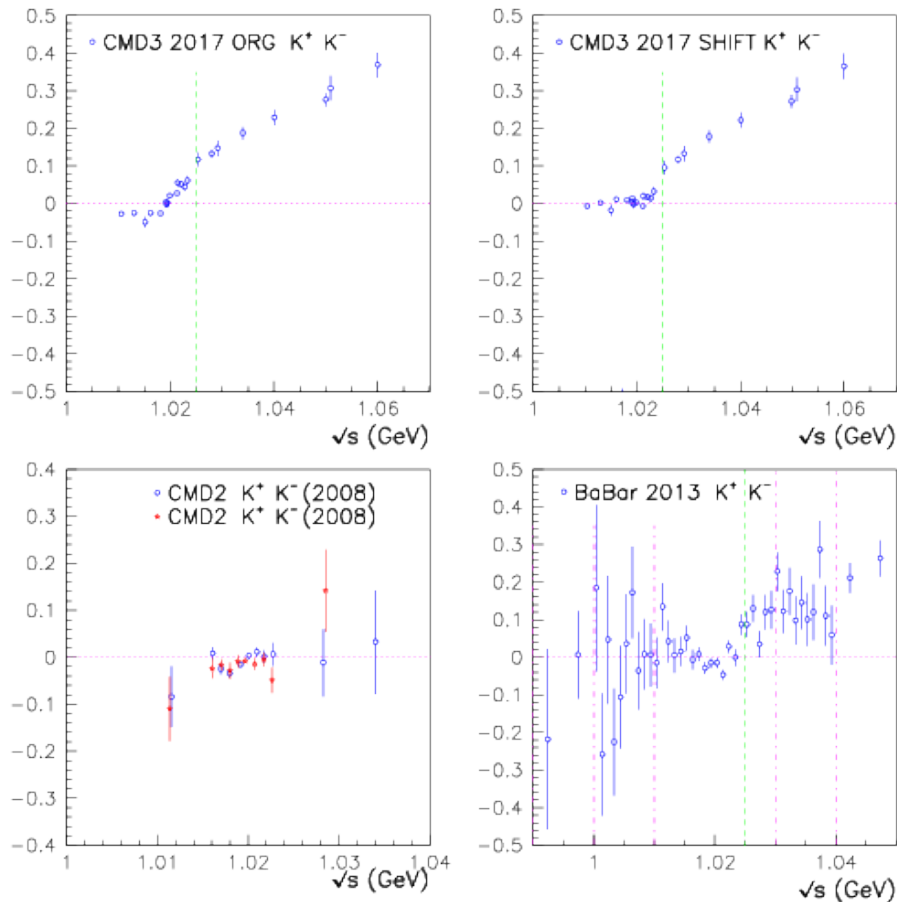
Confirm Inconsistency



# CMD3 $e^+e^- \rightarrow K^+ K^-$ & global fit II

## Applying Rescalings :

CMD3  $K^+ K^-$  (shift) fit :  $K^+ K^-$  Residuals - RESCALING - FJ VP



# $a_\mu$ from BHLS2 Fits (update III)

HLS channels :

FSI updated

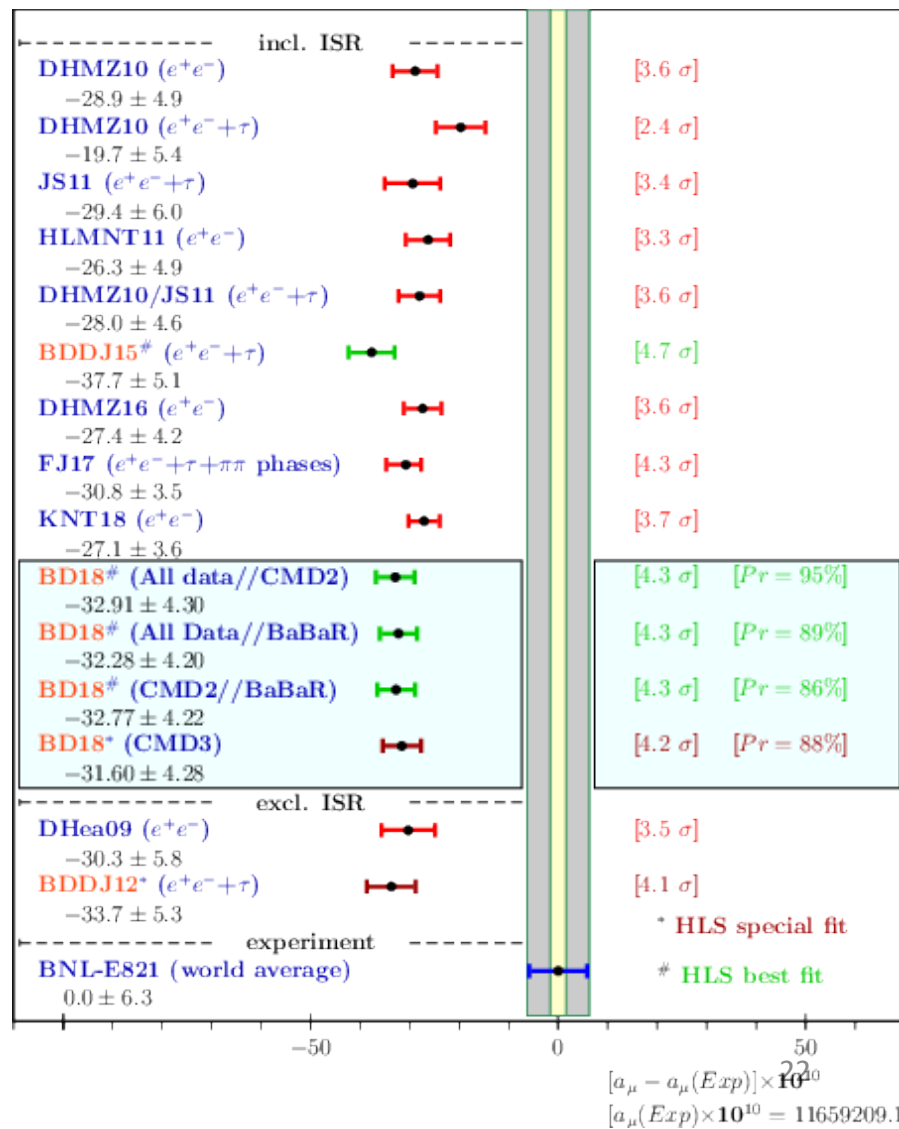
Non HLS channels :

HVP update from FJ

• + ADD. SYST.

BHLS

BHLS2



# Additional syst. to $a_\mu^{\text{HVP}}$

- Updated value for  $10^{10} a_\mu^{\text{HVP}}$ :

$$a_\mu = 11659176.8 \pm 4.20$$

with additional systematics

$$\left[ a_\mu^{\text{exp}} - a_\mu^{\text{th}} \right] = 32.28 \pm 0.5_{\phi\pi\pi} \pm 0.4_{\alpha_{em}} \pm 0.7_{\text{CMDx}} \pm 4.20_{\text{th}} \pm 6.30_{\text{exp}}$$

- $0.5_{\phi\pi\pi} + 0.4_{\alpha_{em}} + 0.7_{\text{CMDx}}$  may contribute as bias
- $\Delta a_\mu^{\text{HVP}} \geq 30.68 \pm 4.20 \pm 6.30$  (4.1  $\sigma$ )
- CMD2 update ?

# Conclusions

- CMD2 & BaBaR distant each from CMD3 by  $>4 \sigma$
- CMD2 distant from BaBaR only by  $\approx 2.2 \sigma_{\text{CMD2}}$
- CMD2 & BaBaR fit well within the HLS framework
- New Reference central value (using BaBaR) :

$$\left[ a_{\mu}^{\text{exp}} - a_{\mu}^{\text{th}} \right] = 32.28 \pm 0.5_{\phi\pi\pi} \pm 0.4_{\alpha_{em}} \pm 0.7_{\text{CMD}x} \pm 4.20_{\text{th}} \pm 6.30_{\text{exp}}$$

- Additional syst. account for the CMD2/3 fits
- CMD2 update : welcome!



# BACKUP

# The HLS Phenomenology

- Once broken, the **H**idden **L**ocal **S**ymmetry model:  
a unified **VMD** framework for

$e^+e^- \rightarrow \pi^+ \pi^- / K^+ K^- / K_L K_S / \pi^0 \gamma / \eta \gamma / \pi^+ \pi^- \pi^0$  &  $\tau \rightarrow \pi \pi \nu_\tau$   
&  $PV\gamma, P\gamma\gamma$  decays &  $(\eta / \eta' \rightarrow \gamma \pi \pi)$  &  $F_\pi / F_K$  at  $s < 0$

(HLS Correlated via their underlying physics)

[83% of the muon HVP]

- **Limit : One vector nonet (the best constrained)**
- ✓ **Up to the  $\approx \phi$  mass region ( $\approx 1.05$  GeV)**

# Scale Uncertainty

- Minimize :

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$$\chi^2 = [m - M - \lambda A]^T V^{-1} [m - M - \lambda A] + \lambda^2 / \sigma^2$$

- Solve for  $\lambda$  ( $A = m$  or  $M$ ) gives:

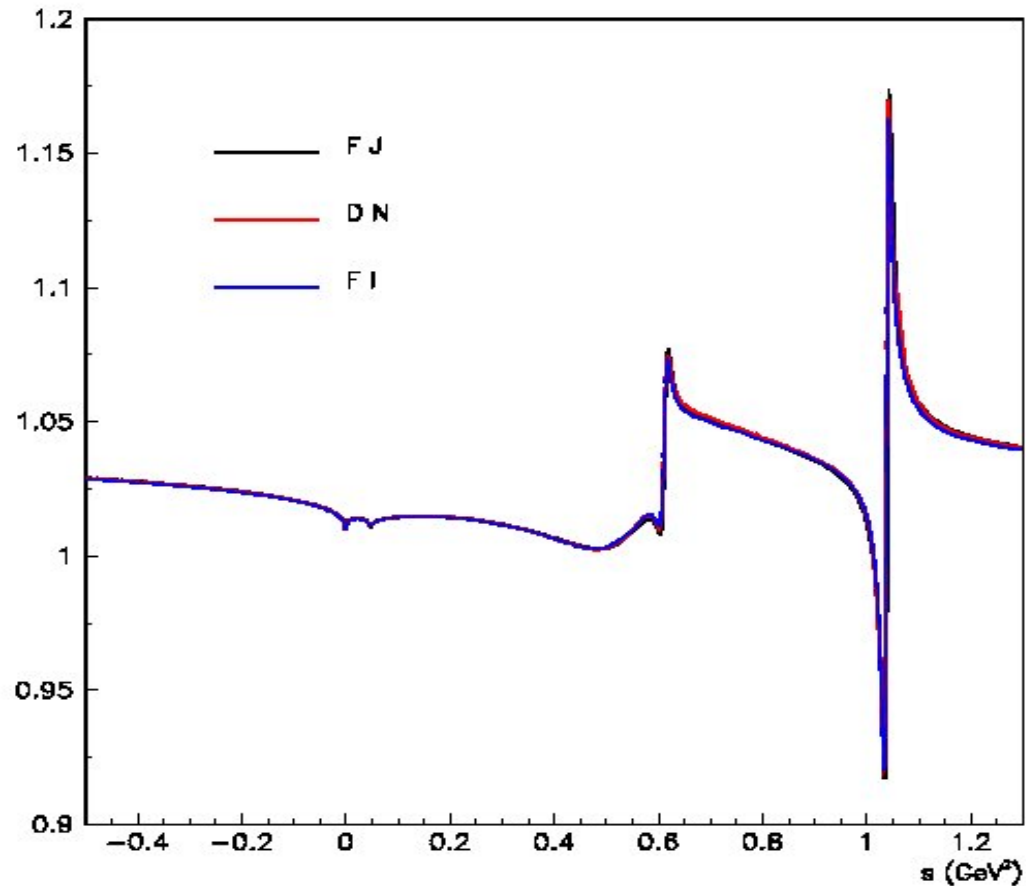
$$\chi^2 \equiv [m - M]^T [V + \sigma^2 A A^T]^{-1} [m - M]$$

- Unbiasing data vs Model  $m \rightarrow m - \lambda M$  :

$$\lambda = \left\{ A^T V^{-1} [m - M] \right\} / \left\{ A^T V^{-1} A + \frac{1}{\sigma^2} \right\}$$

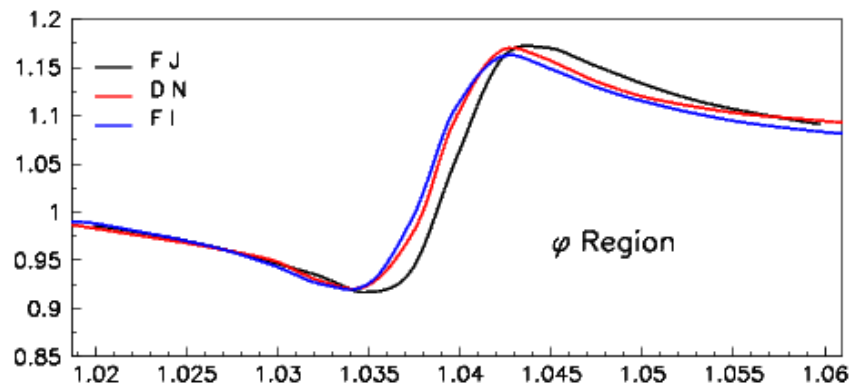
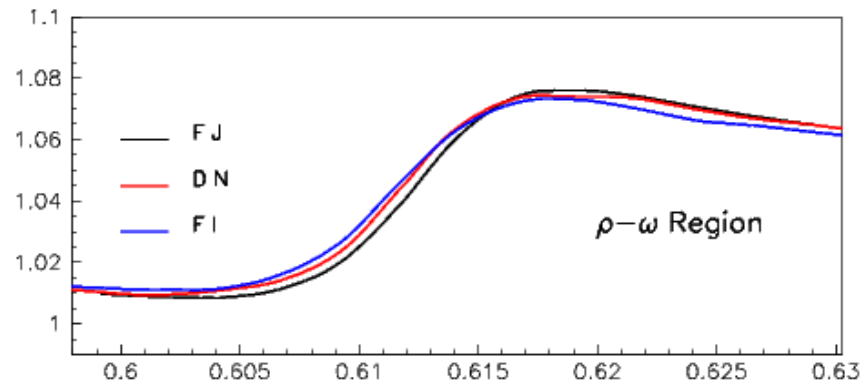
- **(m vs M) Consistency :  $\lambda / \sigma < 3$**

# The photon HVP functions



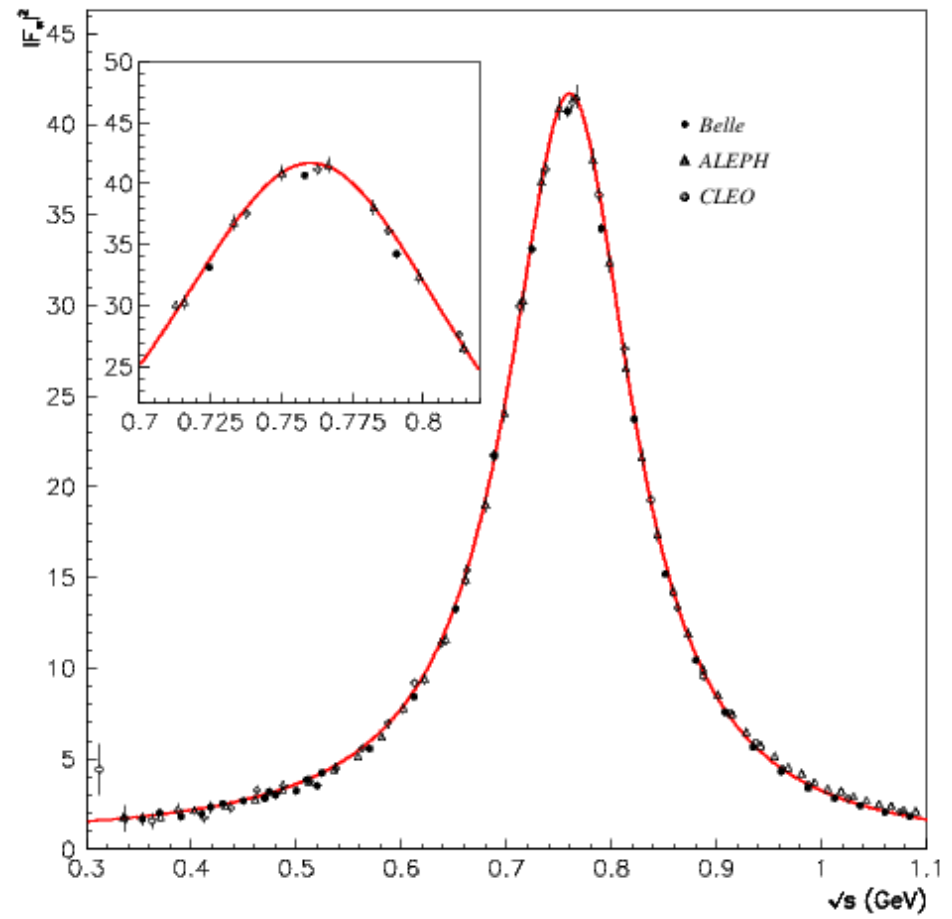
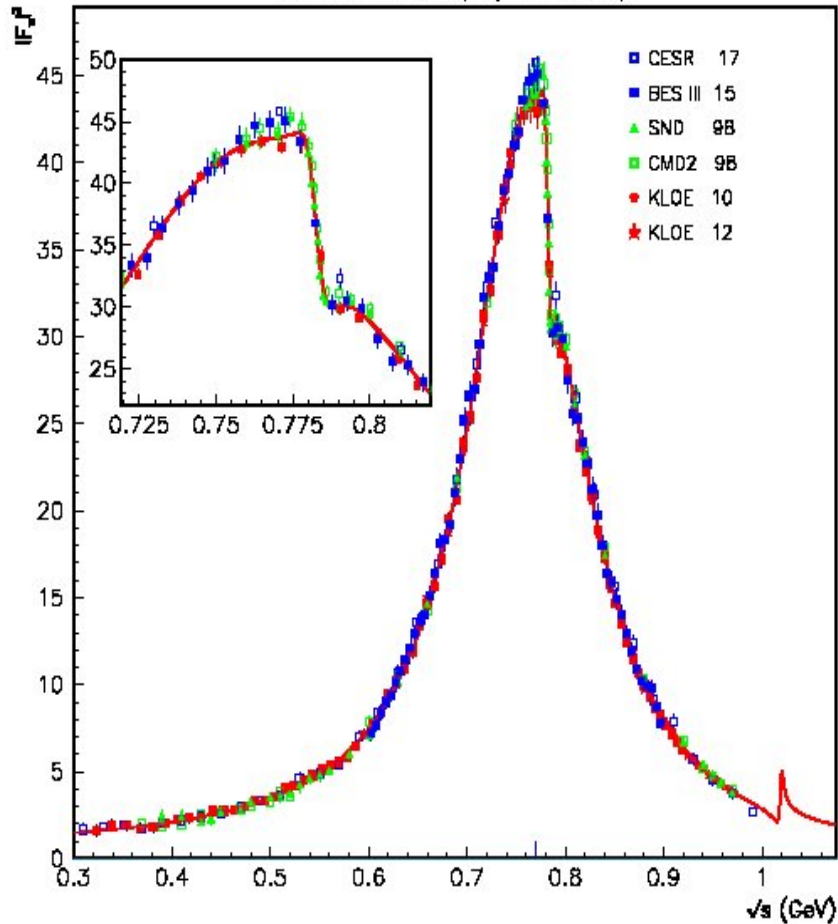
# Uncertainty on photon HVP

- No **unique** way to treat narrow resonances

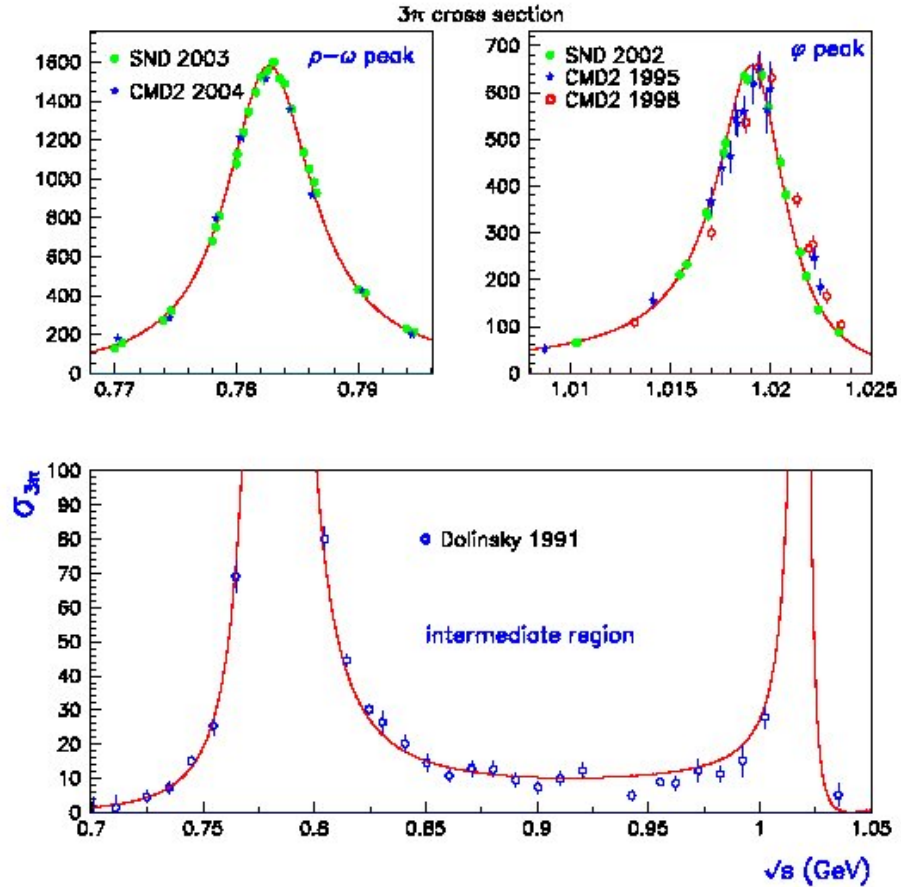


# $\pi\pi$ Data

Pion Form Factor (original spectra)



# 3 pion Data

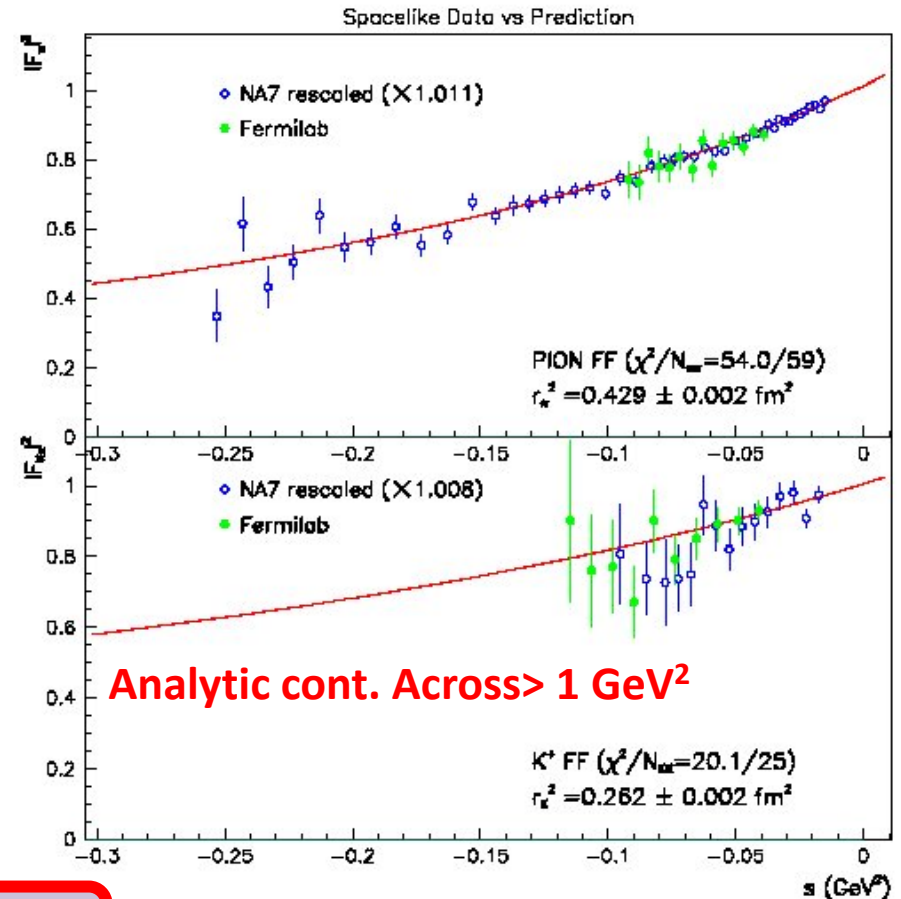


# Spacelike Regions : BHLS2 Predictions

- NA7 scale err :  $\sigma=0.9\%$

$R_\pi^2$ (fm <sup>2</sup> )	
NA7	$0.440 \pm 0.007$ (scale)
Fermilab	$0.440 \pm 0.031$
SELEX	$0.42 \pm [0.06]_{\text{stat}} \pm [0.06]_{\text{syst}}$
1706.04020	$0.432 \pm 0.004$
BHLS 2018	$0.430 \pm 0.001$

$R_k^2$ (fm <sup>2</sup> )	
NA7	$0.336 \pm 0.046$ (scale)
Fermilab	$0.281 \pm 0.053$
BHLS 2018	$0.261 \pm 0.001$





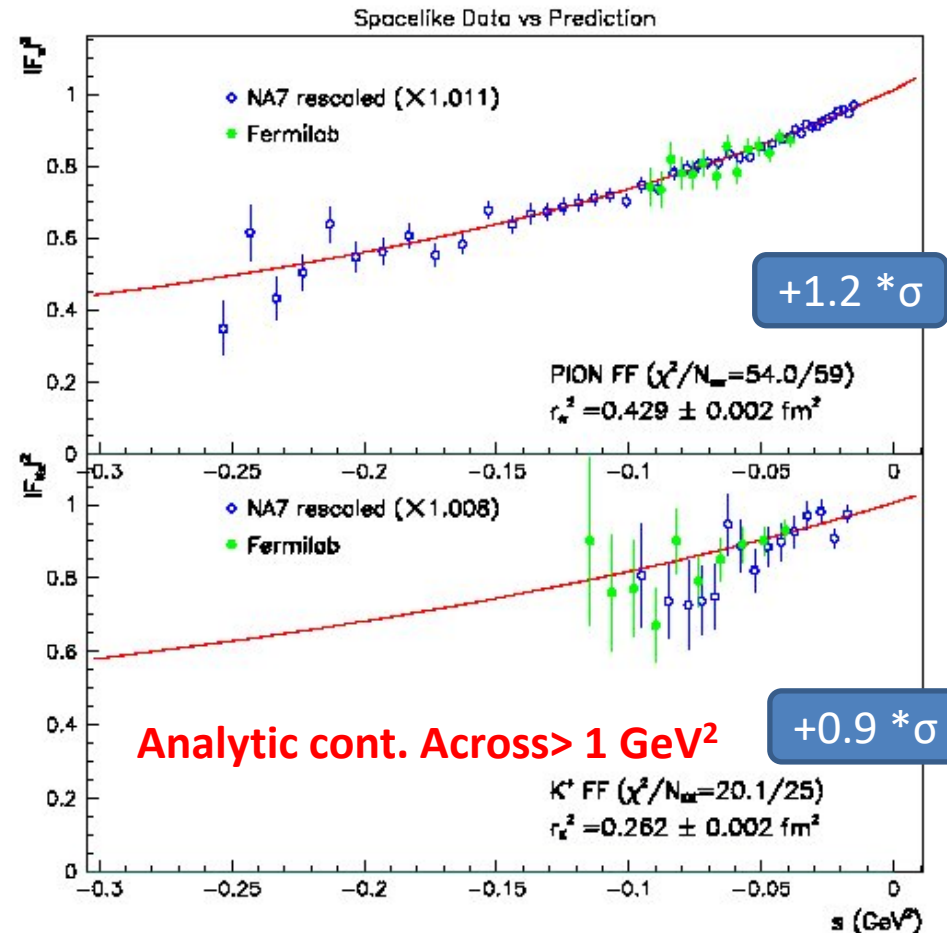
# Spacelike Regions : BHLS2 Predictions

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$R_K^2$ (fm <sup>2</sup> )	
NA7	0.336 ± 0.046 (scale)
Fermilab	0.281 ± 0.053
BHLS 2018	0.261 ± 0.001

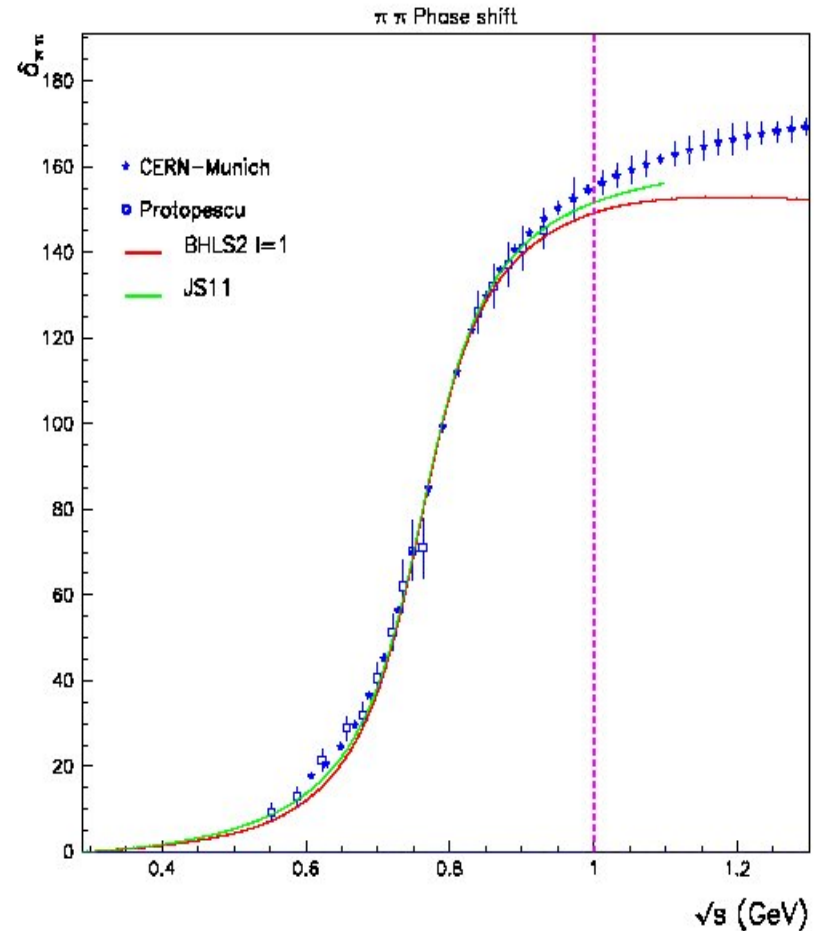
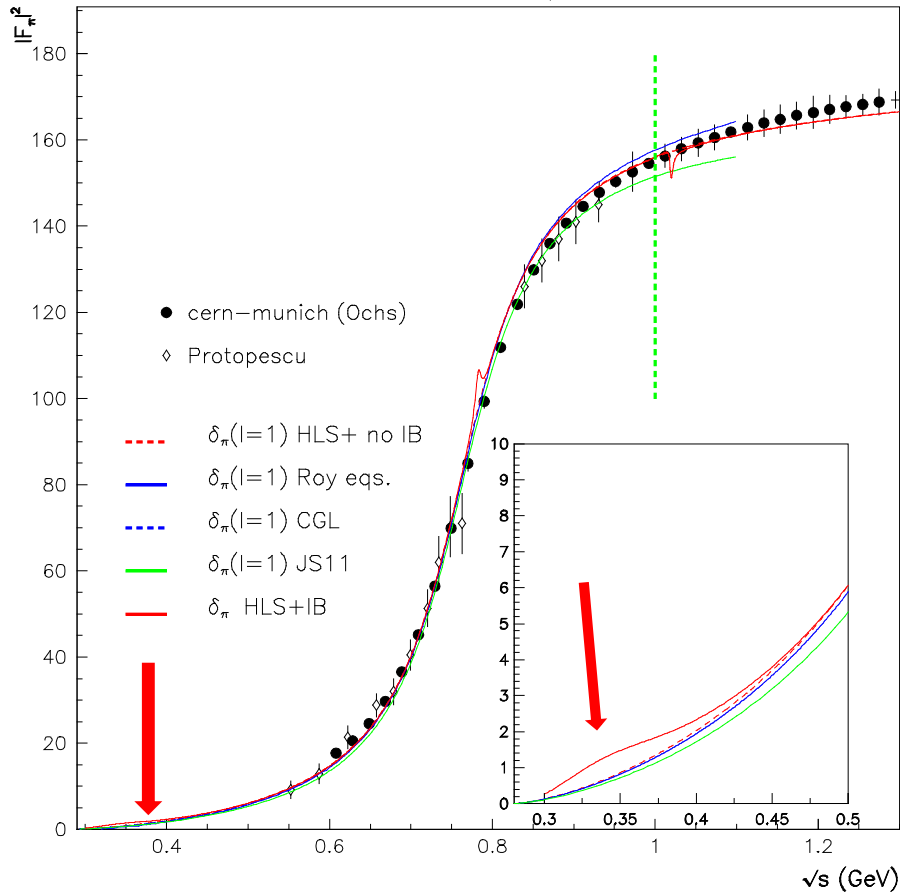
≈ Unbiased FF's



# $\pi\pi$ Phase shift BHLS vs BHLS2

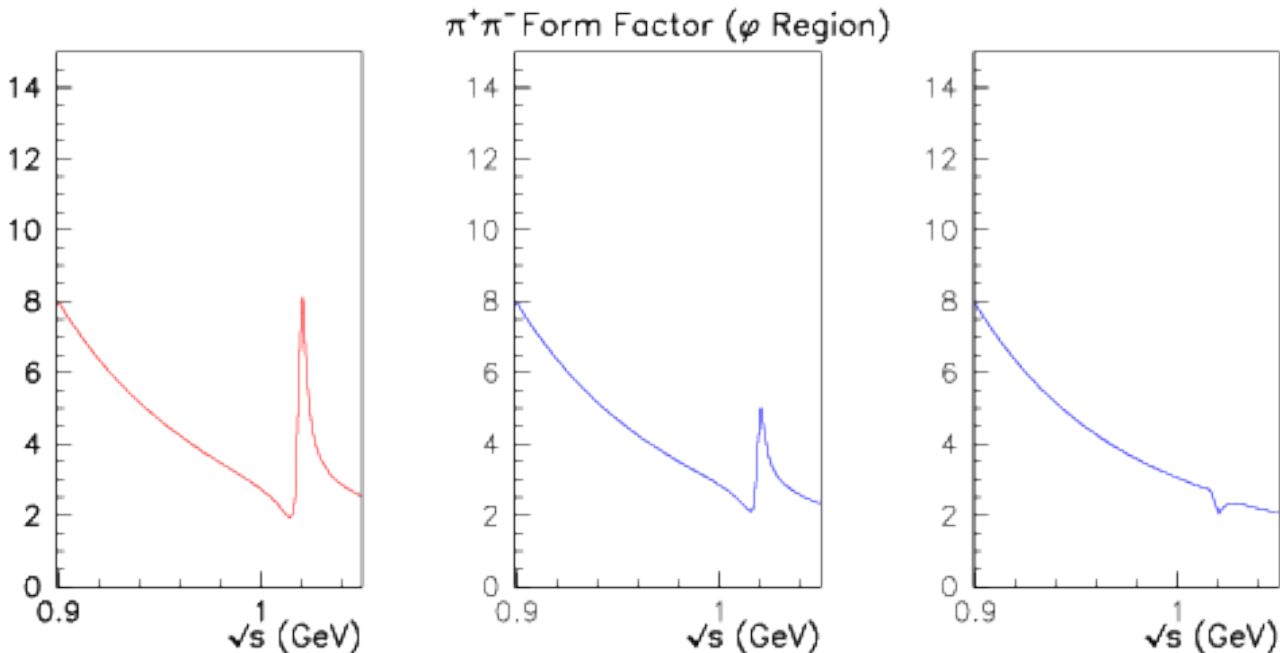
$$\left[ \frac{m_{\omega}^{HK}}{m_{\rho}^{HK}} \right]^2 = 1.00$$

Delta m2=0 z Phase of the pion form factor



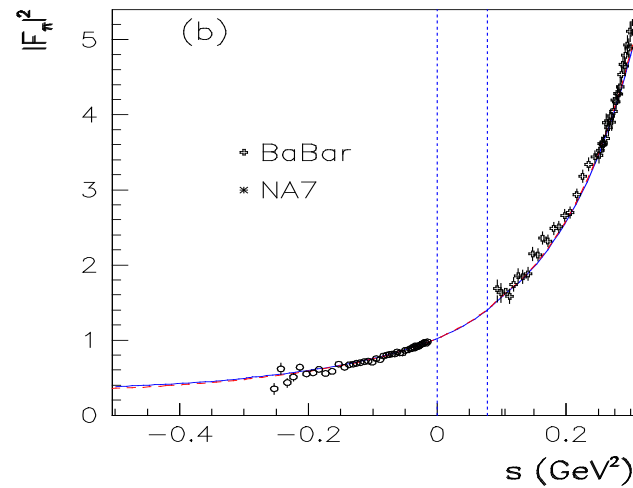
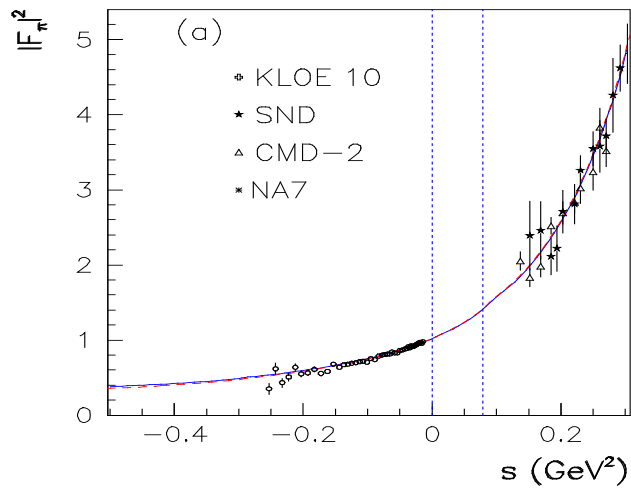
# Lack of constraints : $F_\pi$ at $\phi$

- **Missing** exp. spectrum for  $F_\pi(s)$  around  $\phi$
- Constraints :  $\text{Br}(\phi \rightarrow \pi^+ \pi^-)$  &  $\text{Br}(\phi \rightarrow e^+ e^-) * \text{Br}(\phi \rightarrow \pi^+ \pi^-)$

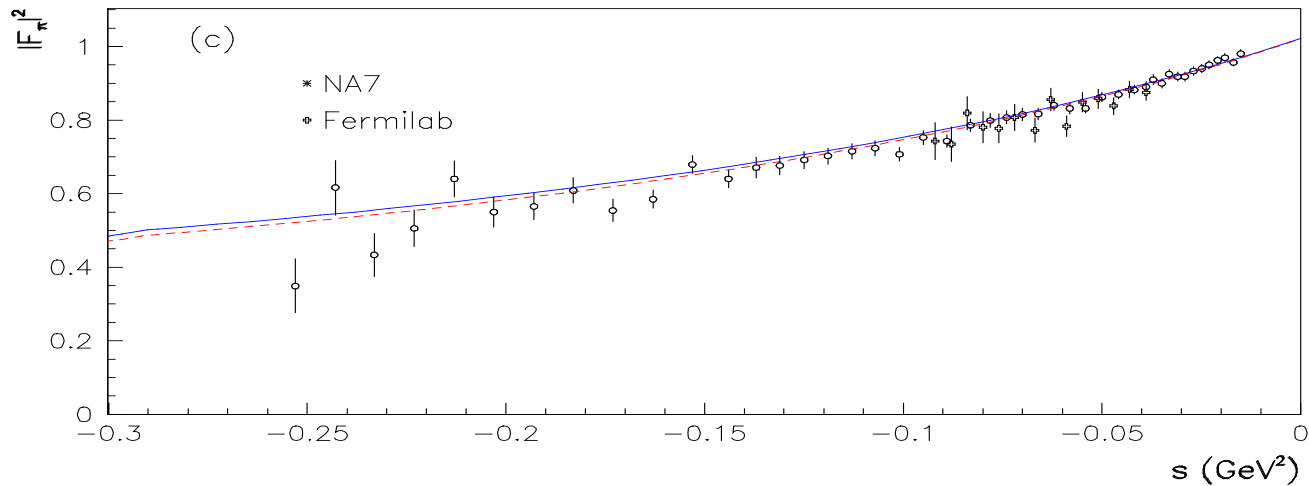


# BACKUP'S

# Spacelike & Threshold Regions

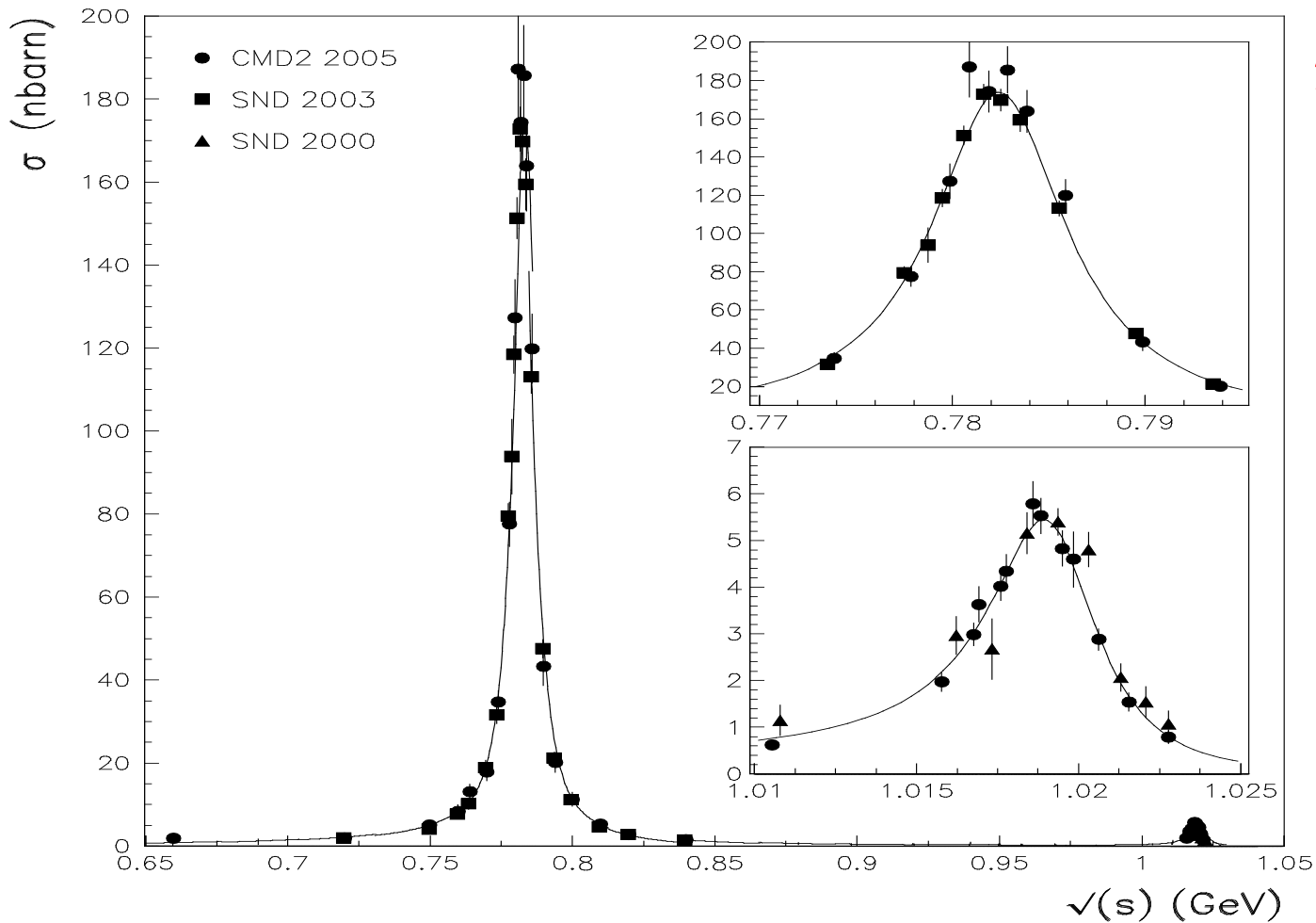


Connection at  
threshold



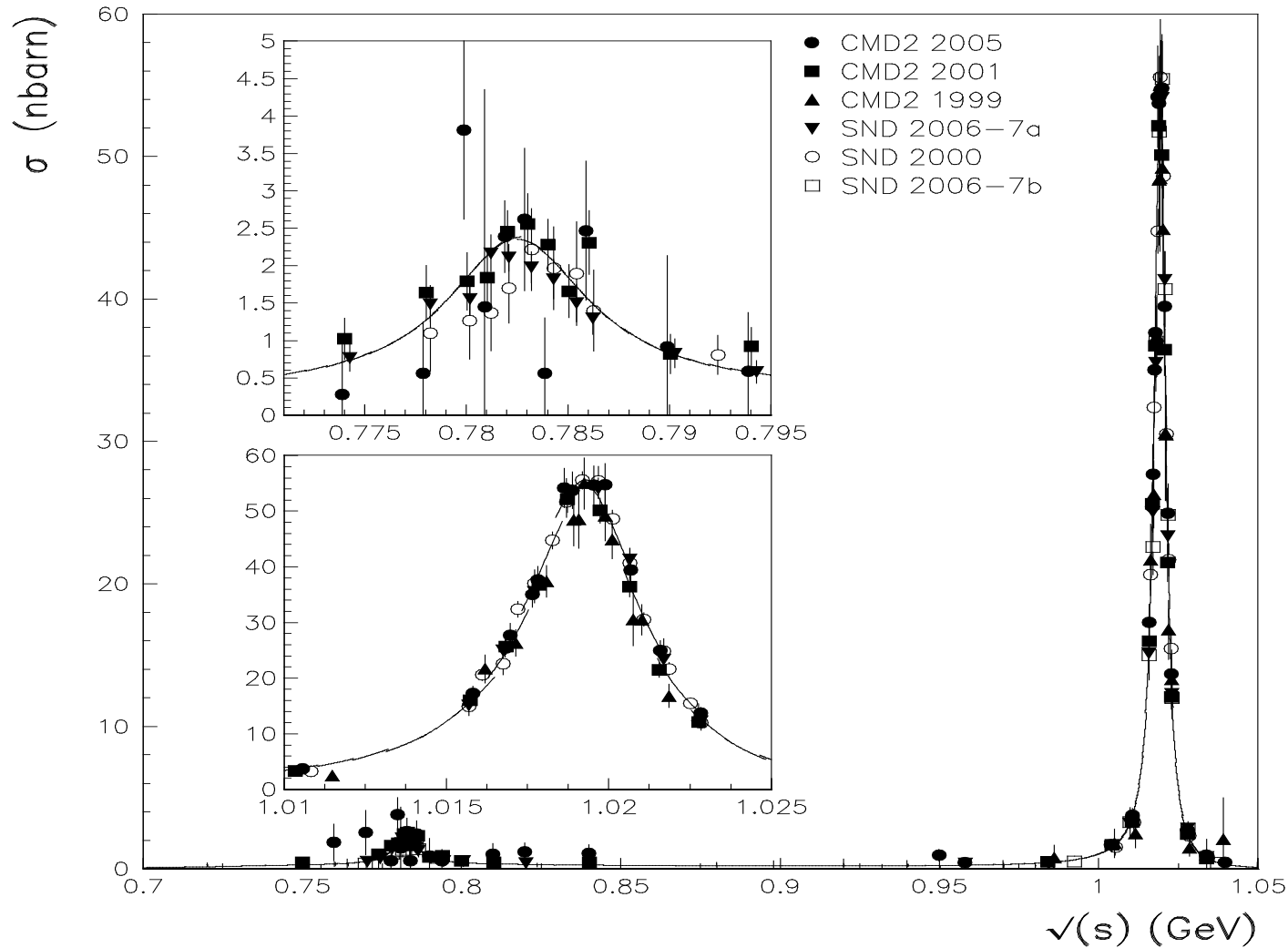
Extrapolation to  $s < 0$

# $e^+e^- \rightarrow \pi^0 \gamma$



$$\frac{\chi^2}{N_p} = \frac{68}{86}$$

# $e^+e^- \rightarrow \eta\gamma$



$$\frac{\chi^2}{N_p} = \frac{123}{182}$$