

# The muon $g - 2$ : a new data-based analysis

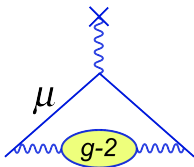
Alex Keshavarzi

with Daisuke Nomura, Thomas Teubner (KNT18)

[arXiv:1802.02995, **accepted for publication in Phys. Rev. D (in press)**]

Muon  $g - 2$  Theory Initiative Workshop, JGU Mainz

20th June 2018

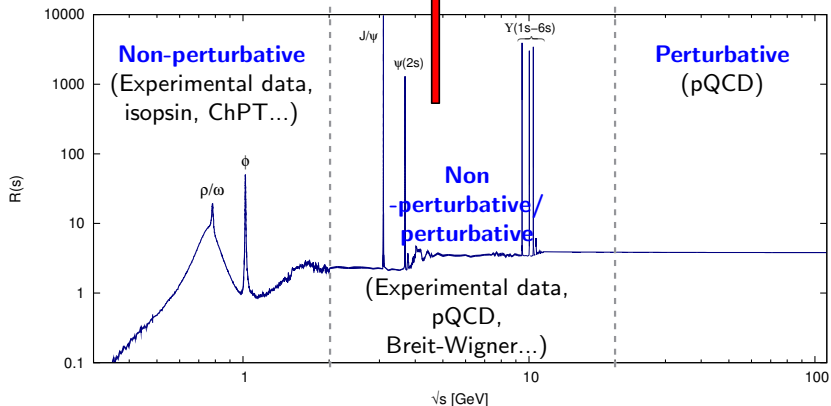


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LIVERPOOL



## Hadronic cross section input

$$a_{\mu}^{\text{had, LO VP}} = \frac{\alpha^2}{3\pi^2} \int_{s_{th}}^{\infty} \frac{ds}{s} R(s) K(s), \text{ where } R(s) = \frac{\sigma_{\text{had}, \gamma}^0(s)}{4\pi\alpha^2/3s}$$



**Must build full hadronic cross section/ $R$ -ratio...**

# Building the hadronic $R$ -ratio

$$\underline{m_\pi \leq \sqrt{s} \leq 2 \text{ GeV}}$$

- Input **experimental hadronic cross section data\***
- **Combine all available data** in exclusive hadronic final states ( $\pi^+\pi^-$ ,  $K^+K^-$ , ...)
- **Sum  $\sim 35$  exclusive channels**
- **Detailed data analysis**
- Robust **treatment of experimental errors**
- Estimate missing data input (**isospin relations, ChPT...**)

$$\underline{2 \leq \sqrt{s} \leq 11.2 \text{ GeV}}$$

- Can **use experimental inclusive  $R$  data\*** or pQCD
- Must use **data at quark flavour thresholds**
- **Combine all available  $R$  data**
- Robust treatment of **experimental errors**
- **Include narrow resonances**

$$\underline{11.2 \leq \sqrt{s} < \infty \text{ GeV}}$$

- **Calculate  $R$  using pQCD (rhad)**

\* $\sigma_{\text{had}}$  experiments

- KLOE
- BaBar
- SND
- CMD-(2/3)
- KEDR
- BESIII

Question: for reliable precision, how are data correlated and how should those correlations be implemented?

# Data combination: setup

⇒ Re-bin data into *clusters*

→ Scan cluster sizes for preferred solution (error,  $\chi^2$ , check by sight...)

⇒ Correlated data beginning to dominate full data compilation...

→ Non-trivial, energy dependent influence on both mean value and error estimate

## KNT18 prescription

- Construct full covariance matrices for each channel & entire compilation  
⇒ Framework available for inclusion of any and all inter-experimental correlations
- If experiment does not provide matrices...
  - Statistics occupy diagonal elements only
  - Systematics are 100% correlated
- If experiment does provide matrices...
  - Use all information provided
- Use correlations to full capacity

# Data combination consideration

## Question:

What are the **main points of concern** when combining experimental data to evaluate  $a_{\mu}^{\text{had, VP}}$ ?

⇒ When **combining data**...

- ...how to best **combine large amounts of data** from different experiments
- ...the **correct implementation of correlated uncertainties** (statistical and systematic)
- ...finding a **solution that is free from bias**

d'Agostini bias [Nucl.Instrum.Meth. A346 (1994) 306-311]

$$\begin{aligned} x_1 &= 0.9 \pm \delta x_1 \\ x_2 &= 1.1 \pm \delta x_2 \end{aligned} \quad C_{\text{sys}} = \begin{pmatrix} p^2 x_1^2 & p^2 x_1 x_2 \\ p^2 x_2 x_1 & p^2 x_2^2 \end{pmatrix} \quad \Rightarrow \bar{x} \simeq 0.98 \text{ (systematic bias)}$$

(Normalisation uncertainties defined by data)

Effect worsened with full,  
iterative data combination

# Data combination consideration

## Question:

What are the **main points of concern** when combining experimental data to evaluate  $a_{\mu}^{\text{had, VP}}$ ?

⇒ When **combining data**...

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Fixed matrix method [R. D. Ball et al. [NNPDF Collaboration], JHEP 1005 (2010) 075.]

$$\begin{aligned} x_1 &= 0.9 \pm \delta x_1 \\ x_2 &= 1.1 \pm \delta x_2 \end{aligned} \quad C_{\text{sys}} = \begin{pmatrix} p^2 \bar{x}^2 & p^2 \bar{x}^2 \\ p^2 \bar{x}^2 & p^2 \bar{x}^2 \end{pmatrix} \quad \Rightarrow \bar{x} = 1.00 \text{ (systematic bias)}$$

(Normalisation uncertainties defined by estimator)

⇒ **Redefinition repeated at each stage of iterative data combination**

# Linear $\chi^2$ minimisation [KNT18: arXiv:1802.02995, PRD (in press)]

⇒ Clusters are defined to have **linear cross section**

→ **Fix covariance matrix with linear interpolants** at each iteration  
(extrapolate at boundary)

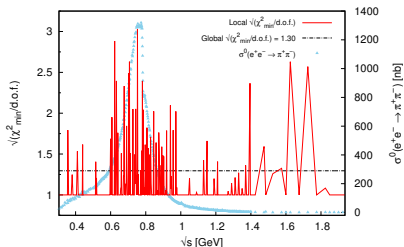
$$\chi^2 = \sum_{i=1}^{N_{\text{tot}}} \sum_{j=1}^{N_{\text{tot}}} (R_i^{(m)} - \mathcal{R}_m^i) \mathbf{C}^{-1}(i^{(m)}, j^{(n)}) (R_j^{(n)} - \mathcal{R}_n^j)$$

⇒ **Through correlations and linearisation**, result is the minimised solution of all available uncertainty information

→ ... through a method that has been **shown to avoid d'Agostini bias**

⇒ The **flexibly of the fit** to vary due to the energy dependent, correlated uncertainties benefits the combination

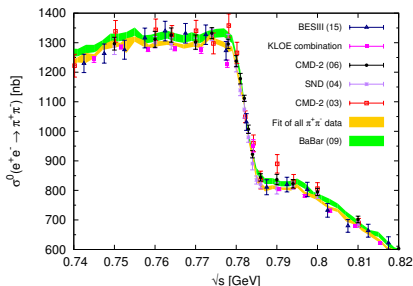
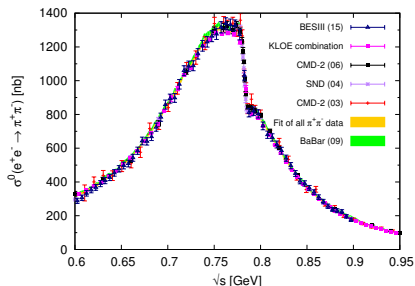
→ ... and any data tensions are reflected in a **local and global  $\chi_{\text{min}}^2/\text{d.o.f.}$  error inflation**



# $\pi^+\pi^-$ channel [KNT18: arXiv:1802.02995, PRD (in press)]

⇒  $\pi^+\pi^-$  accounts for over 70% of  $a_\mu^{\text{had, LO VP}}$

→ Combines 30 measurements totalling 999 data points



⇒ Correlated & experimentally corrected  $\sigma_{\pi\pi(\gamma)}^0$  data now entirely dominant

$$a_\mu^{\pi^+\pi^-} [0.305 \leq \sqrt{s} \leq 1.937 \text{ GeV}] = 502.97 \pm 1.14_{\text{stat}} \pm 1.59_{\text{sys}} \pm 0.06_{\text{VP}} \pm 0.14_{\text{fsr}}$$

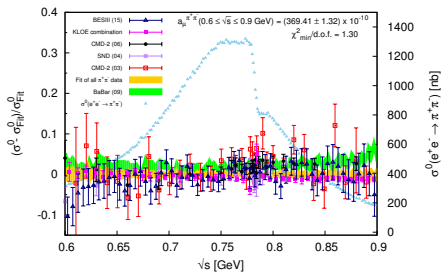
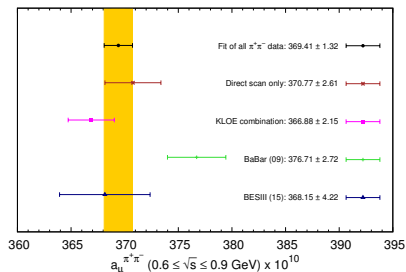
$$= 502.97 \pm 1.97_{\text{tot}} \quad \text{HLMNT11: } 505.77 \pm 3.09$$

⇒ 15% local  $\chi_{\text{min}}^2/\text{d.o.f.}$  error inflation due to tensions in clustered data



# $\pi^+\pi^-$ channel [KNT18: arXiv:1802.02995, PRD (in press)]

- ⇒ Tension exists between **BaBar data** and all other data in the dominant  $\rho$  region.
- Agreement between other radiative return measurements and direct scan data largely **compensates this**.

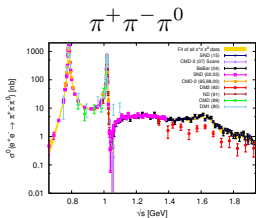


Compared to  $a_\mu^{\pi^+\pi^-} = 502.97 \pm 1.97$ :  $\Rightarrow a_\mu^{\pi^+\pi^-}$  (**BaBar data only**) =  $513.2 \pm 3.8$ .

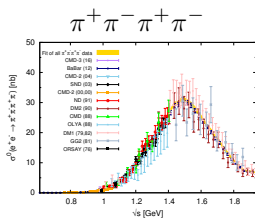
Simple **weighted average of all data**  $\Rightarrow a_\mu^{\pi^+\pi^-}$  (Weighted average) =  $509.1 \pm 2.9$ .  
(i.e. - no correlations in determination of mean value)

**BaBar data dominate** when no correlations are taken into account for the mean value  
Highlights **importance of fully incorporating all available correlated uncertainties**

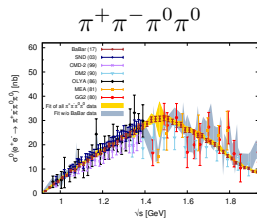
# Other notable exclusive channels [KNT18: arXiv:1802.02995, PRD (in press)]



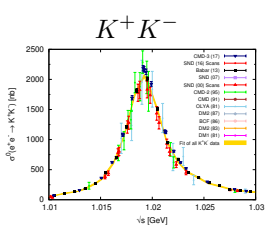
HLMNT11:  $47.51 \pm 0.99$   
 KNT18:  $47.92 \pm 0.89$



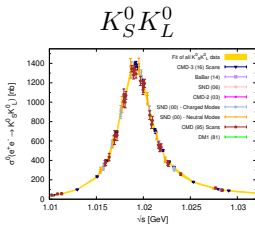
HLMNT11:  $14.65 \pm 0.47$   
 KNT18:  $14.87 \pm 0.20$



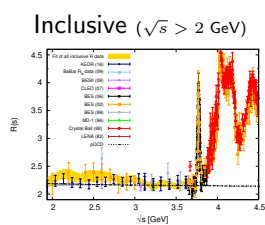
HLMNT11:  $20.37 \pm 1.26$   
 KNT18:  $19.39 \pm 0.78$



HLMNT11:  $22.15 \pm 0.46$   
 KNT18:  $23.03 \pm 0.22$



HLMNT11:  $13.33 \pm 0.16$   
 KNT18:  $13.04 \pm 0.19$



HLMNT11:  $41.40 \pm 0.87$   
 KNT18:  $41.27 \pm 0.62$

# KNT18 $a_\mu^{\text{had, VP}}$ update [KNT18: arXiv:1802.02995, PRD (in press)]

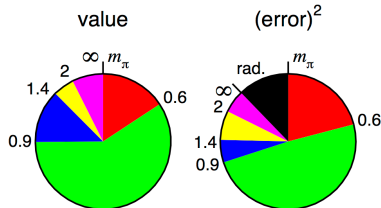
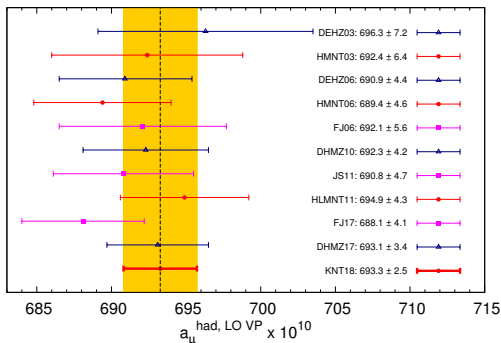
$$\text{HLMNT(11): } 694.91 \pm 4.27$$



$$\begin{aligned} \text{This work: } a_\mu^{\text{had, LO VP}} &= 693.27 \pm 1.19_{\text{stat}} \pm 2.01_{\text{sys}} \pm 0.22_{\text{VP}} \pm 0.71_{\text{fsr}} \\ &= 693.27 \pm 2.34_{\text{exp}} \pm 0.74_{\text{rad}} \\ &= 693.27 \pm 2.46_{\text{tot}} \end{aligned}$$

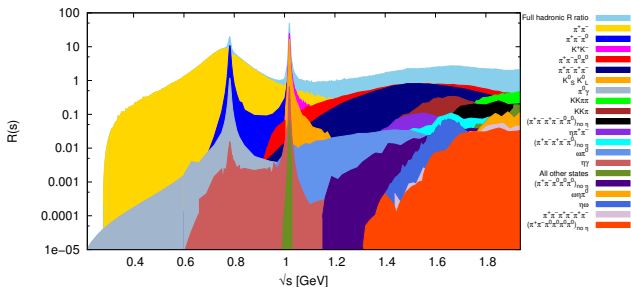
$$a_\mu^{\text{had, NLO VP}} = -9.82 \pm 0.04_{\text{tot}}$$

⇒ Accuracy better than 0.4%  
(uncertainties include all available correlations and local  $\chi^2$  inflation)



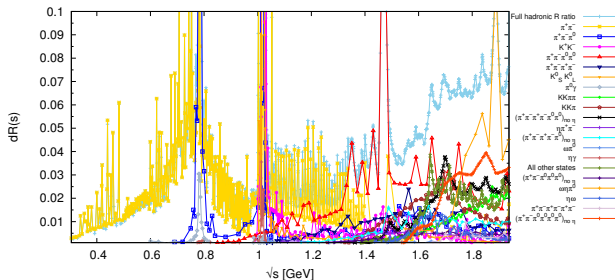
⇒  $2\pi$  dominance

# Contributions below 2GeV [KNT18: arXiv:1802.02995, PRD (in press)]

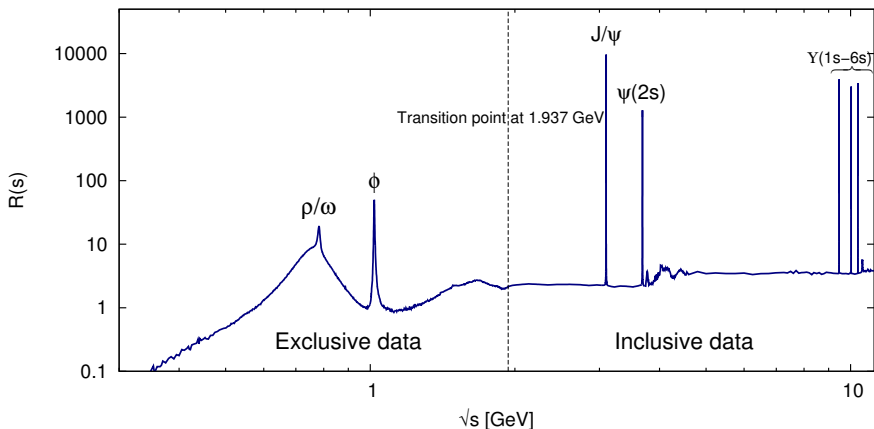


→ Dominance of  $2\pi$  below 0.9 GeV evident for both cross section and uncertainty

→ Large improvement to cross section and uncertainty from new  $4\pi$  data



# $R(s)$ for $m_\pi \leq \sqrt{s} \leq 11.2$ GeV [KNT18: arXiv:1802.02995, PRD (in press)]



⇒ Full KNT18 compilation data set for hadronic  $R$ -ratio now available...

⇒ **...complete with full covariance matrix**

# KNT18 $a_\mu^{\text{SM}}$ update [KNT18: arXiv:1802.02995, PRD (in press)]

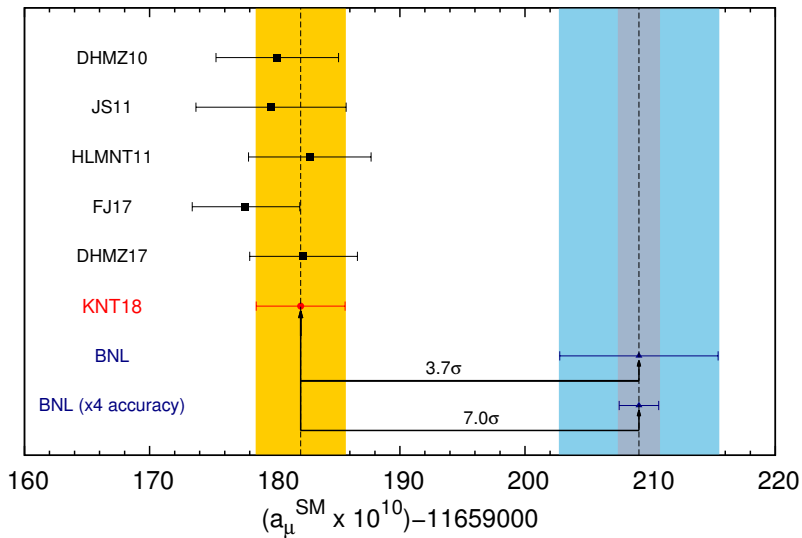
	<u>2011</u>	→	<u>2017</u>	
QED	11658471.81 (0.02)	→	11658471.90 (0.01)	[arXiv:1712.06060]
EW	15.40 (0.20)	→	15.36 (0.10)	[Phys. Rev. D 88 (2013) 053005]
LO HLbL	10.50 (2.60)	→	9.80 (2.60)	[EPJ Web Conf. 118 (2016) 01016]
NLO HLbL			0.30 (0.20)	[Phys. Lett. B 735 (2014) 90]

	<u>HLMNT11</u>	→	<u>KNT18</u>	
LO HVP	694.91 (4.27)	→	693.27 (2.46)	this work
NLO HVP	-9.84 (0.07)	→	-9.82 (0.04)	this work
NNLO HVP			1.24 (0.01)	[Phys. Lett. B 734 (2014) 144]

Theory total	11659182.80 (4.94)	→	11659182.05 (3.56)	this work
Experiment			11659209.10 (6.33)	world avg
Exp - Theory	26.1 (8.0)	→	27.1 (7.3)	this work

$\Delta a_\mu$	3.3 $\sigma$	→	3.7 $\sigma$	this work
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# KNT18 $a_\mu^{\text{SM}}$ update [KNT18: arXiv:1802.02995, PRD (in press)]



# Conclusions

- ⇒ Accuracy of  $a_\mu^{\text{SM}}$  limited by hadronic contributions
- ⇒ Hadronic VP contributions can be determined from dispersion relations and hadronic cross section
- ⇒ Must build hadronic  $R$ -ratio from experimental data
- ⇒ New data combination method + new data yields improvements in all channels due to increased fit flexibility
- ⇒ Correlations have large effect on mean value and uncertainty and all available information should be correctly incorporated
- ⇒  $a_\mu^{\text{had,LOVP}}$  accuracy better than 0.4%
- ⇒ Improvement in HVP yields  $g - 2$  discrepancy of  $3.7\sigma$
- ⇒ Overall HVP uncertainty now better than HLbL uncertainty

**Thank you**



# Extra Slides

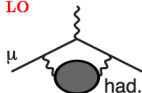
# Hadronic contributions

⇒ Uncertainty on  $a_\mu^{\text{SM}}$  dominated by hadronic contributions

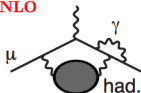
→ **Non-perturbative, low energy region** of hadronic resonances

$$a_\mu^{\text{had}} = a_\mu^{\text{had,VP LO}} + a_\mu^{\text{had,VP NLO}} + a_\mu^{\text{had,Light-by-Light}}$$

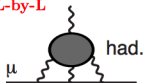
LO



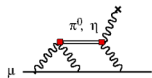
NLO



L-by-L



e.g.



⇒ LbL contributions ( $\mathcal{O}(\alpha^3)$ ), so far **only fully determined using model calculations**

→ **Difficult to quantify/control uncertainties** from models

→ **Huge progress from lattice and dispersive** approaches

→ So far, **no indication of unpleasant surprises**

→ But, **big improvements expected** in near future Phys. Rev. D 94 (2016) 053006.

⇒ **LO LbL**, updated 'Glasgow consensus' estimate:  $a_\mu^{\text{had,LO LbL}} = (9.8 \pm 2.6) \times 10^{-10}$

→ **NLO LbL** estimated to be  $a_\mu^{\text{had,NLO LbL}} = (0.3 \pm 0.2) \times 10^{-10}$  Phys. Lett. B 735 (2014) 90.

$$a_\mu^{\text{had,LbL}} = (10.1 \pm 2.6) \times 10^{-10}$$

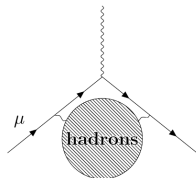
# $a_\mu^{\text{had,VP}}$ : theoretical setup

⇒ We want to calculate the **leading order hadronic vacuum polarisation (HVP) contribution**

1) Feynman rules for **HVP insertion to photon propagator**:

$$\mu \text{ wavy } q \text{ --- } \text{hadrons} \text{ --- } q \text{ wavy } \nu = \frac{-ig^{\mu\alpha}}{(q^2 - i\varepsilon)} (-ie)i\Pi_{\alpha\beta}(q^2)(-ie) \frac{-ig^{\beta\nu}}{(q^2 - i\varepsilon)}$$

$\Pi_{\alpha\beta}(q^2)$



2) Employ **analyticity**:

$$\mu \text{ wavy } q \text{ --- } \text{hadrons} \text{ --- } q \text{ wavy } \nu = \frac{ie^2 g_{\mu\nu}}{(q^2 - i\varepsilon)^2} \frac{q^4}{\pi} \int_{s_{th}}^{\infty} ds \frac{\text{Im } \Pi(s)}{s(s - q^2 - i\varepsilon)}$$

$\Pi_{\alpha\beta}(q^2)$

3) **Insert to vertex correction**, solve for  $a_\mu$ :  $a_\mu^{\text{had,LOVP}} = \frac{\alpha}{\pi^2} \int_{s_{th}}^{\infty} \frac{ds}{s} \text{Im } \Pi_{\text{had}}(s) K(s)$

4) Utilise **optical theorem**:

$$\text{Im} \left| \text{wavy } \gamma \text{ --- } \text{had} \text{ --- } \gamma \text{ wavy} \right| \Leftrightarrow \left| \text{wavy } \gamma \text{ --- } \text{had} \text{ --- } \text{hadrons} \right|^2$$

$\text{Im } \Pi_{\text{had}}(q^2)$        $\sim \sigma_{\text{had}}(q^2)$

5) Arrive at **equation for  $a_\mu^{\text{had,LOVP}}$** :

$$a_\mu^{\text{had,LOVP}} = \frac{1}{4\pi^3} \int_{s_{th}}^{\infty} ds \sigma_{\text{had},\gamma}^0(s) K(s)$$

$\sigma_{\text{had},\gamma}^0 =$  **bare cross section, FSR included**

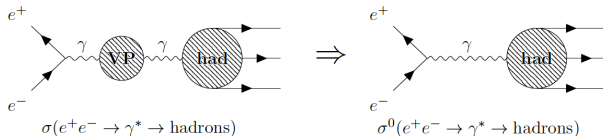
⇒ **Similar dispersion integrals for NLO and NNLO HVP**

# $\sigma_{\text{had},\gamma}^0$ : vacuum polarisation corrections

⇒ Reconsider the **optical theorem**:  $\text{Im} \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right| \Leftrightarrow \left| \begin{array}{c} \gamma \\ \text{had} \\ \text{had} \\ \text{had} \\ \text{had} \end{array} \right|^2$   
 $\text{Im} \Pi_{\text{had}}(q^2)$   $\sim \sigma_{\text{had}}(q^2)$

⇒ Photon VP corresponds to higher order contributions to  $a_{\mu}^{\text{had, VP}}$

→ **Must subtract VP:**



⇒ Fully updated, self-consistent VP routine: [vp\_knt\_v3\_0], available for distribution

→ Cross sections undressed with **full photon propagator** (must include imaginary part),  $\sigma_{\text{had}}^0(s) = \sigma_{\text{had}}(s) |1 - \Pi(s)|^2$

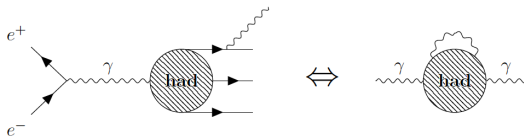
⇒ If correcting data, **apply corresponding radiative correction uncertainty**

→ Take  $\frac{1}{3}$  of total correction per channel as conservative extra uncertainty

# $\sigma_{\text{had},\gamma}^0$ : final state radiation corrections

⇒ Reconsider the **optical theorem**:  $\text{Im} \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right| \Leftrightarrow \left| \begin{array}{c} \gamma \\ \text{had} \\ \text{had} \\ \text{had} \\ \text{had} \end{array} \right|^2$   
 $\text{Im} \Pi_{\text{had}}(q^2)$   $\sim \sigma_{\text{had}}(q^2)$

⇒ Photon FSR formally higher order corrections to  $a_\mu^{\text{had, VP}}$



⇒ **Cannot be unambiguously separated, not accounted for in HO contributions**

→ Must be **included as part of 1PI hadronic blobs**

⇒ Experiment may cut/miss photon FSR → **Must be added back**

⇒ For  $\pi^+\pi^-$ , **sQED approximation** [Eur. Phys. J. C 24 (2002) 51, Eur. Phys. J. C 28 (2003) 261]

⇒ For **higher multiplicity states**,  
difficult to estimate correction

**Need new, more developed tools to increase precision here**

∴ **Apply conservative uncertainty** (e.g. - CARLOMAT 3.1 [Eur.Phys.J. C77 (2017) no.4, 254 ]?)

# Systematic bias and use of the data/covariance matrix

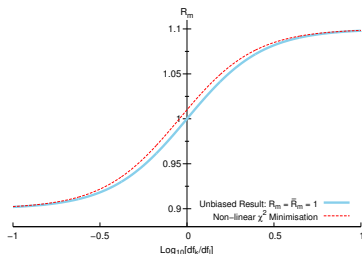
⇒ Data is re-binned using an **adaptive clustering algorithm**

⇒ Iterative fit of covariance matrix as defined by data → **D'Agostini bias**

[Nucl.Instrum.Meth. A346 (1994) 306-311]

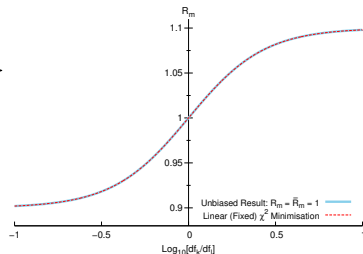
## HLMNT11

⇒ Non-linear  $\chi^2$  minimisation fitting  
nuisance parameters  
→ **Penalty trick bias**



## KNT18

⇒ **Fix the covariance matrix** in an  
iterative  $\chi^2$  minimisation  
→ **Free from bias**



**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**

# Fixing the covariance matrix [JHEP 1005 (2010) 075, Eur.Phys.J. C75 (2015), 613]

⇒ Apply a procedure to **fix the covariance matrix**

$$C_I(i^{(m)}, j^{(n)}) = C^{\text{stat}}(i^{(m)}, j^{(n)}) + \frac{C^{\text{sys}}(i^{(m)}, j^{(n)})}{R_i^{(m)} R_j^{(n)}} R_m R_n ,$$

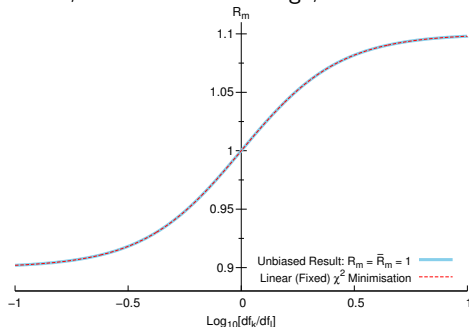
in an **iterative  $\chi^2$  minimisation** method that, to our best knowledge, is **free from bias**

⇒ Fixing with theory value **regulates influence**

⇒ Can be shown from toy models to be **free from bias**

⇒ **Swift convergence**

⇒ Comparison with past results shows **HLMNT11 estimates are largely unaffected**



**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**

## Properties of a covariance matrix

Any covariance matrix,  $\mathcal{C}_{ij}$ , of dimension  $n \times n$  must satisfy the following requirements:

- As the diagonal elements of any covariance matrix are populated by the corresponding variances, all the diagonal elements of the matrix are positive. Therefore, the trace of the covariance matrix must also be positive

$$\text{Trace}(\mathcal{C}_{ij}) = \sum_{i=1}^n \sigma_{ii} = \sum_{i=1}^n \text{Var}_i > 0$$

- It is a symmetric matrix,  $\mathcal{C}_{ij} = \mathcal{C}_{ji}$ , and is, therefore, equal to its transpose,  $\mathcal{C}_{ij} = \mathcal{C}_{ij}^T$
- The covariance matrix is a positive, semi-definite matrix,

$$\mathbf{a}^T \mathcal{C} \mathbf{a} \geq 0 ; \mathbf{a} \in \mathbf{R}^n,$$

where  $\mathbf{a}$  is an eigenvector of the covariance matrix  $\mathcal{C}$

- Therefore, the corresponding eigenvalues  $\lambda_{\mathbf{a}}$  of the covariance matrix must be real and positive and the distinct eigenvectors are orthogonal

$$\mathbf{b}^T \mathcal{C} \mathbf{a} = \lambda_{\mathbf{a}}(\mathbf{b} \cdot \mathbf{a}) = \mathbf{a}^T \mathcal{C} \mathbf{b} = \lambda_{\mathbf{b}}(\mathbf{a} \cdot \mathbf{b})$$

$$\therefore \text{if } \lambda_{\mathbf{a}} \neq \lambda_{\mathbf{b}} \Rightarrow (\mathbf{a} \cdot \mathbf{b}) = 0$$

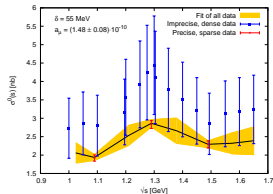
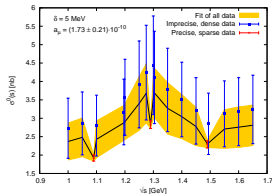
- The determinant of the covariance matrix is positive:  $\text{Det}(\mathcal{C}_{ij}) \geq 0$



# Clustering data

⇒ Re-bin data into *clusters*

Better representation of data combination through adaptive clustering algorithm



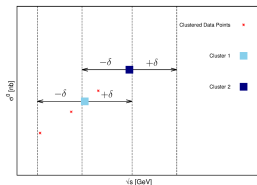
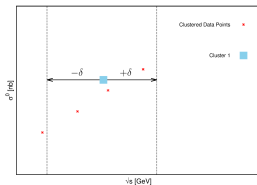
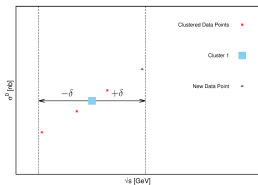
→ More and more data ⇒ risk of **over clustering**

⇒ loss of information on resonance

→ Scan cluster sizes for **optimum solution** (error,  $\chi^2$ , check by sight...)

⇒ Scanning/**sampling by varying bin widths**

→ Clustering algorithm now **adaptive to points at cluster boundaries**



# Correlation and covariance matrices

⇒ **Correlated data** beginning to **dominate** full data compilation...

→ Non-trivial, **energy dependent influence** on both **mean value and error estimate**

## KNT18 prescription

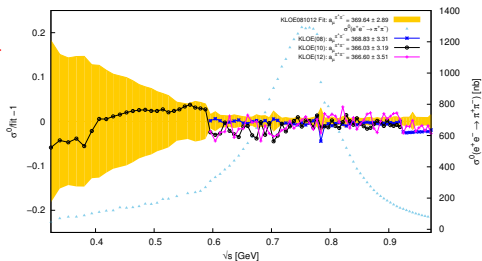
- Construct full covariance matrices for each channel & entire compilation  
⇒ **Framework available for inclusion of any and all inter-experimental correlations**

- If experiment does not provide matrices...  
→ Statistics occupy diagonal elements only  
→ Systematics are 100% correlated

- If experiment does provide matrices...  
→ **Matrices must satisfy properties of a covariance matrix**

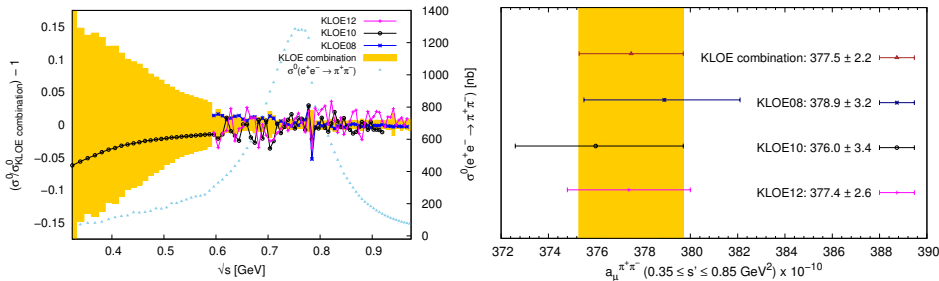
e.g. - KLOE  $\pi^+\pi^-\gamma(\gamma)$  combination covariance matrices update

⇒ **Originally, NOT a positive semi-definite matrix**



# The resulting KLOE $\pi^+\pi^-\gamma(\gamma)$ combination [JHEP 1803 (2018) 173.]

⇒ Combination of KLOE08, KLOE10 and KLOE12 gives 85 distinct bins between  $0.1 \leq s \leq 0.95 \text{ GeV}^2$



→ Covariance matrix now correctly constructed

⇒ a **positive semi-definite matrix**

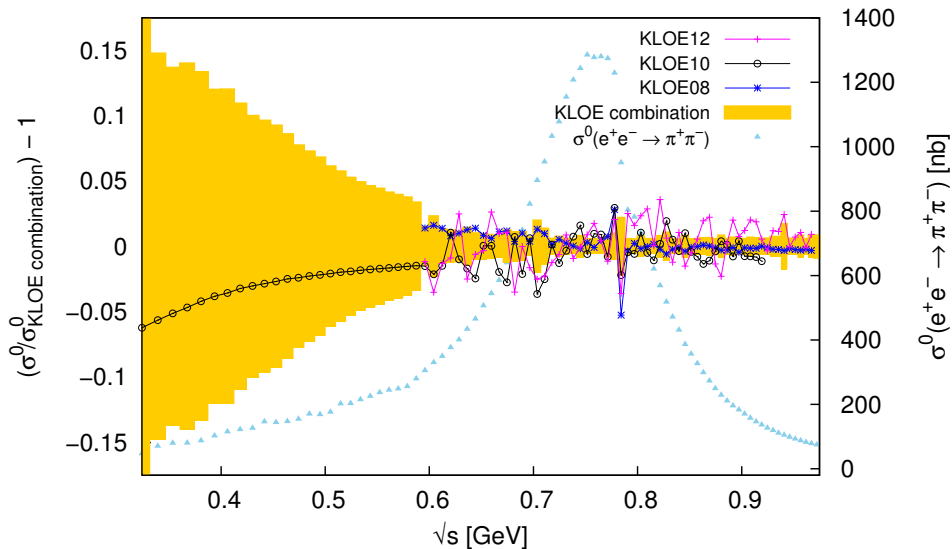
→ **Non-trivial influence of correlated uncertainties** on resulting mean value

$$a_\mu^{\pi^+\pi^-} (0.1 \leq s' \leq 0.95 \text{ GeV}^2) = (489.9 \pm 2.0_{\text{stat}} \pm 4.3_{\text{sys}}) \times 10^{-10}$$

→ All previous **combinations issues** now eliminated...

...and **consistency between measurements and combination**

# The KLOE $\pi^+\pi^-\gamma(\gamma)$ combination [JHEP 1803 (2018) 173.]

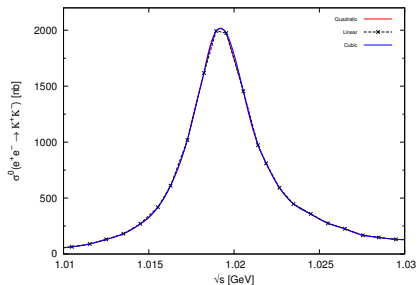
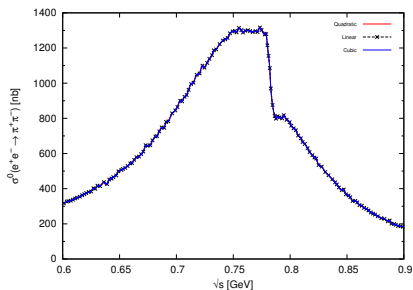


# Integration

⇒ Trapezoidal rule integral

→ Consistency with linear cluster definition

→ High data population ∴ **Accurate estimate from linear integral**



→ Higher order polynomial integrals give **(at maximum)** differences of  $\sim 10\%$  of error

⇒ Estimates of error non-trivial at **integral borders**

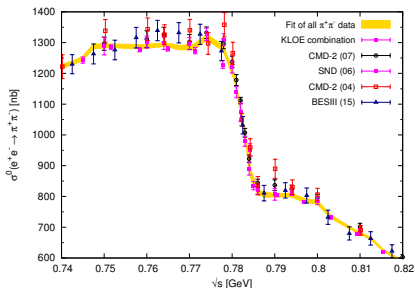
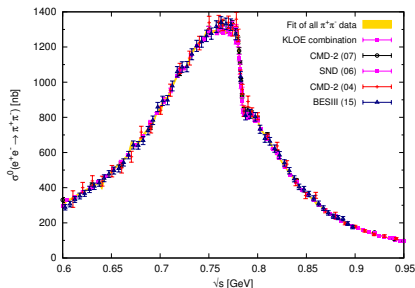
→ **Extrapolate/interpolate covariance matrices**

## $2\pi$ without BaBar

⇒  $2\pi$  data combination **stable without BaBar data**

→ **Other data saturate the effect of BaBar**

→ **Differences with and without BaBar are now fairly small**



In range  $0.32 \leq \sqrt{s} \leq 2$  GeV :

$$\Rightarrow \text{All data: } a_{\mu}^{\pi^+\pi^-} = 501.45 \pm 1.95 ; \sqrt{\chi^2_{\min}/\text{d.o.f.}} = 1.29$$

$$\Rightarrow \text{No BaBar: } a_{\mu}^{\pi^+\pi^-} = 500.28 \pm 2.67 ; \sqrt{\chi^2_{\min}/\text{d.o.f.}} = 1.35$$

## $2\pi$ CLEO-c data [Phys.Rev. D97 (2018) 032012]

⇒ New  $2\pi$  data from CLEO-c should be **used with caution**

- Two measurements taken at different COM energies ( $\psi(3770)/\psi(4170)$ ) have **very different cross sections**
- **Large statistical and systematic errors** compared to other radiative return sets
- VP correction has been applied with **FJ03VP (needs updated version)** and **only subtracts real part**
- The values for  $a_{\mu}^{\pi^+\pi^-}$  **given in the paper only calculated using weighted average**
  - **Systematics will be highly correlated and should be incorporated**
- The authors have fitted the data to **Gounaris-Sakurai parametrisation**
  - **Unreliable representation of cross section at high  $s$**
- The authors find (with FJ03VP):

$$a_{\mu}^{\pi^+\pi^-}(\psi(3770)) = 489.6 \pm 4.5_{\text{stat}}, \quad a_{\mu}^{\pi^+\pi^-}(\psi(4170)) = 503.6 \pm 5.9_{\text{stat}}$$

$$a_{\mu}^{\pi^+\pi^-}(\text{Weighted average}) = 500.4 \pm 3.6_{\text{stat}} \pm 7.5_{\text{sys}}$$

→ I find (with KNT18VP):

$$a_{\mu}^{\pi^+\pi^-}(\psi(3770)) = 499.6 \pm 4.5_{\text{stat}} \pm 7.5_{\text{sys}}, \quad a_{\mu}^{\pi^+\pi^-}(\psi(4170)) = 504.3 \pm 5.9_{\text{stat}} \pm 7.6_{\text{sys}}$$

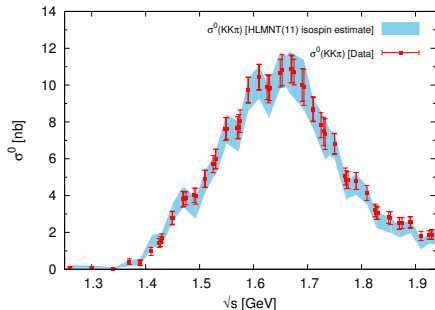
$$a_{\mu}^{\pi^+\pi^-}(\text{Fit - w/o correlated systematics}) = 500.9 \pm 4.0_{\text{stat}} \pm 5.9_{\text{sys}}$$

$$a_{\mu}^{\pi^+\pi^-}(\text{Fit - with correlated systematics}) = 500.7 \pm 4.0_{\text{stat}} \pm 8.3_{\text{sys}}$$

# $KK\pi$ , $KK\pi\pi$ and isospin

⇒ New data for  $KK\pi$  and  $KK\pi\pi$   
 removes reliance on isospin (only  $K_S^0 \cong K_L^0$ )

$KK\pi$   
 $K_S^0 K_L^0 \pi^0$  [Phys.Rev. D95 (2017), 052001, arXiv:1711.07143]

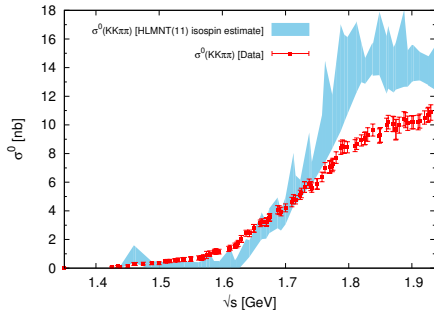


HLMNT11:  $2.65 \pm 0.14$

KNT18:  $2.71 \pm 0.12$

⇒ **But**, still reliant on isospin estimates for  $\pi^+\pi^-3\pi^0$ ,  $\pi^+\pi^-4\pi^0$ ,  $KK3\pi\dots$

$KK\pi\pi$   
 $K_S^0 K_L^0 \pi^+\pi^-$  [Phys.Rev. D80 (2014), 092002]  
 $K_S^0 K_S^0 \pi^+\pi^-$  [Phys.Rev. D80 (2014), 092002],  
 $K_S^0 K_L^0 \pi^0\pi^0$  [Phys.Rev. D95 (2017), 052001]  
 $K_S^0 K^\pm \pi^\mp \pi^0$  [Phys.Rev. D95 (2017), 092005]



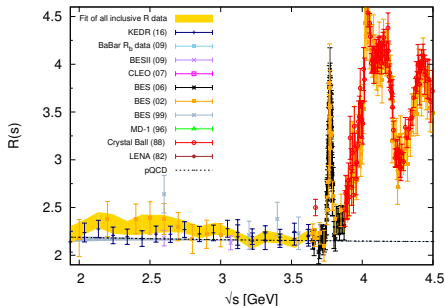
HLMNT11:  $2.51 \pm 0.35$

KNT18:  $1.93 \pm 0.08$

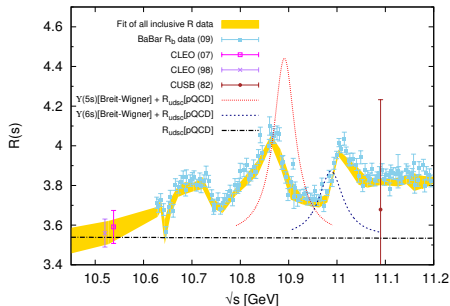


# Inclusive

⇒ **New KEDR inclusive  $R$  data** [Phys.Lett. B770 (2017) 174-181, Phys.Lett. B753 (2016) 533-541] and **BaBar  $R_b$  data** [Phys. Rev. Lett. 102 (2009) 012001].



KEDR data improves the inclusive data combination below  $c\bar{c}$  threshold



$R_b$  resolves the resonances of the  $\Upsilon(5S - 6S)$  states.

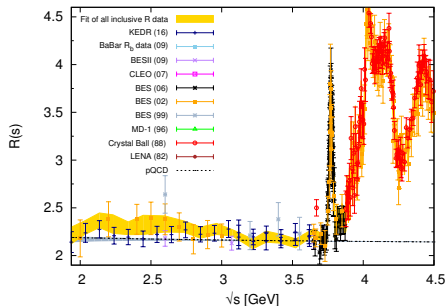
⇒ **Choose to adopt entirely data driven estimate from threshold to 11.2 GeV**

$$a_\mu^{\text{Inclusive}} = 43.67 \pm 0.17_{\text{stat}} \pm 0.48_{\text{sys}} \pm 0.01_{\text{vp}} \pm 0.44_{\text{fsr}} = 43.67 \pm 0.67_{\text{tot}}$$

# KEDR update of $R(s)$ with covariance matrix

⇒ **New precise KEDR update** [arXiv:1805.06235] with **systematic covariance matrix for all measurements provided by experiment**

### KNT18



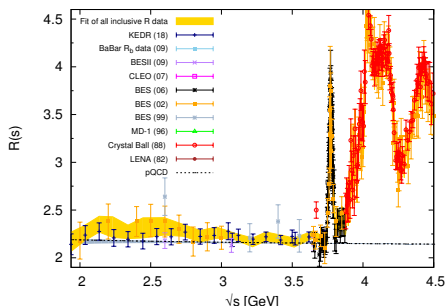
$$a_{\mu}^{\text{Inclusive}} = 43.67 \pm 0.51$$

$$\sqrt{\chi^2_{\text{min}}/\text{d.o.f.}} = 1.44$$

Note: Uncertainties quoted here do not include radiative correction uncertainties

⇒ **Observe very small changes** due to including correlations (slightly closer to pQCD)

### KNT18 + new KEDR data



$$a_{\mu}^{\text{Inclusive}} = 43.54 \pm 0.51$$

$$\sqrt{\chi^2_{\text{min}}/\text{d.o.f.}} = 1.47$$

# Exclusive/inclusive transition point

⇒ New KEDR data allow **reconsideration of exclusive/inclusive transition point**

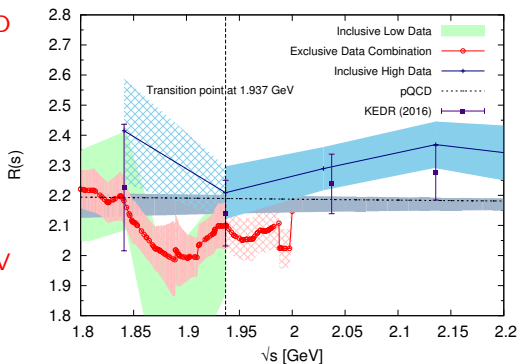
→ KNT18 aim to **avoid use of pQCD** and **keep a data-driven analysis**

→ **Disagreement** between sum of **exclusive states** and **inclusive data/pQCD**

→ New  $\pi^+\pi^-\pi^0\pi^0$  data result in **reduction of the cross section**

→ Previous transition point at **2 GeV** **no longer the preferred choice**

→ More natural choice for this transition point at **1.937 GeV**



Input	$a_\mu^{\text{had, LO VP}} [1.841 \leq \sqrt{s} \leq 2.00 \text{ GeV}] \times 10^{10}$
Exclusive sum	$6.06 \pm 0.17$
Inclusive data	$6.67 \pm 0.26$
pQCD	$6.38 \pm 0.11$
Exclusive ( $< 1.937 \text{ GeV}$ ) + inclusive ( $> 1.937 \text{ GeV}$ )	$6.23 \pm 0.13$

# $\chi_{\min}^2/\text{d.o.f}$ comparison with HLMNT11

Channel	This work (KNT18)	HLMNT11
$\pi^+\pi^-$	1.3	1.4
$\pi^+\pi^-\pi^0$	2.1	3.0
$\pi^+\pi^-\pi^+\pi^-$	1.8	1.7
$\pi^+\pi^-\pi^0\pi^0$	2.0	1.3
$(2\pi^+2\pi^-\pi^0)_{\text{no } \eta}$	1.0	1.2
$(2\pi^+2\pi^-\pi^0)_{\text{no } \eta\omega}$	3.5	4.0
$K^+K^-$	2.1	1.9
$K_S^0K_L^0$	0.8	0.8

**Table:** Comparison of the global  $\sqrt{\chi_{\min}^2/\text{d.o.f}}$  for the leading and major sub-leading channels determined in the HLMNT11 analysis and in this work (KNT18). The first column indicates the final state or individual contribution, the second column gives the KNT18 value, the third column states the HLMNT11 value and the last column gives the difference between the two numbers.

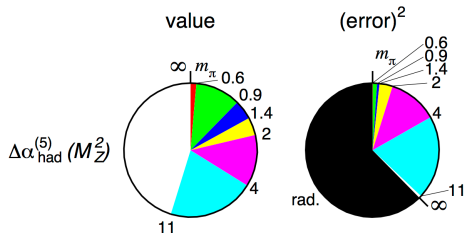
# KNT18 $\alpha(M_Z^2)$ update

$$\underline{\Delta\alpha^{(5)}(M_Z^2)}$$

$$\text{HLMNT11: } (276.26 \pm 1.38_{\text{tot}}) \times 10^{-4}$$

↓

$$\begin{aligned} \text{This work: } \Delta\alpha_{\text{had}}^{(5)}(M_Z^2) &= (276.11 \pm 0.26_{\text{stat}} \pm 0.68_{\text{sys}} \pm 0.14_{\text{vp}} \pm 0.82_{\text{fsr}}) \times 10^{-4} \\ &= (276.11 \pm 0.73_{\text{exp}} \pm 0.84_{\text{rad}}) \times 10^{-4} \\ &= (276.11 \pm 1.11_{\text{tot}}) \times 10^{-4} \end{aligned}$$

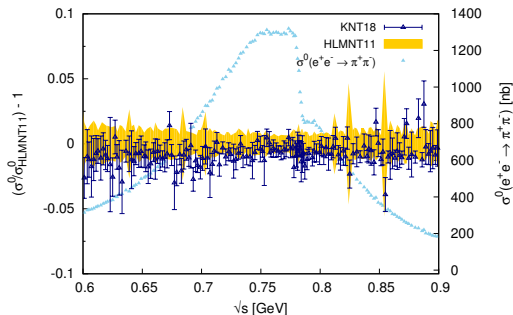


$$\Rightarrow \alpha^{-1}(M_Z^2) = 128.946 \pm 0.015$$

# Comparison with HLMNT11

Channel	This work (KNT18)	HLMNT11	Difference
$\pi^+\pi^-$	$502.99 \pm 1.97$	$505.77 \pm 3.09$	-2.78
$\pi^+\pi^-\pi^0$	$47.82 \pm 0.89$	$47.51 \pm 0.99$	0.31
$\pi^+\pi^-\pi^+\pi^-$	$15.17 \pm 0.21$	$14.65 \pm 0.47$	0.52
$\pi^+\pi^-\pi^0\pi^0$	$19.80 \pm 0.79$	$20.37 \pm 1.26$	-0.57
$K^+K^-$	$23.05 \pm 0.22$	$22.15 \pm 0.46$	0.90
$K_S^0K_L^0$	$13.05 \pm 0.19$	$13.33 \pm 0.16$	-0.28
Inclusive channel	$41.27 \pm 0.62$	$41.40 \pm 0.87$	-0.13
<b>Total</b>	<b><math>693.27 \pm 2.46</math></b>	<b><math>694.91 \pm 4.27</math></b>	<b>-1.64</b>

- ⇒ Biggest difference in  $2\pi$  channel
  - large reduction in mean and uncertainty
- ⇒ Tensions with HLMNT11 analysis for both two-kaon channels
- ⇒ Overall agreement with HLMNT11
- ⇒ Notable improvement of about one third in uncertainty



# Comparison with other similar works

Channel	This work (KNT18)	DHMZ17	Difference
$\pi^+\pi^-$	$503.74 \pm 1.96$	$507.14 \pm 2.58$	-3.40
$\pi^+\pi^-\pi^0$	$47.70 \pm 0.89$	$46.20 \pm 1.45$	1.50
$\pi^+\pi^-\pi^+\pi^-$	$13.99 \pm 0.19$	$13.68 \pm 0.31$	0.31
$\pi^+\pi^-\pi^0\pi^0$	$18.15 \pm 0.74$	$18.03 \pm 0.54$	0.12
$K^+K^-$	$23.00 \pm 0.22$	$22.81 \pm 0.41$	0.19
$K_S^0K_L^0$	$13.04 \pm 0.19$	$12.82 \pm 0.24$	0.22
$1.8 \leq \sqrt{s} \leq 3.7 \text{ GeV}$	$34.54 \pm 0.56 \text{ (data)}$	$33.45 \pm 0.65 \text{ (pQCD)}$	1.09
Total	$693.3 \pm 2.5$	$693.1 \pm 3.4$	0.2

- ⇒ Total estimates from two analyses in very good agreement
- ⇒ Masks much larger differences in the estimates from individual channels
- ⇒ Unexpected tension for  $2\pi$  considering the data input likely to be similar
  - Points to marked differences in way data are combined
  - From  $2\pi$  discussion:  $a_{\mu}^{\pi^+\pi^-}$  (Weighted average) =  $509.1 \pm 2.9$
- ⇒ Compensated by lower estimates in other channels
  - For example, the choice to use pQCD instead of data above 1.8 GeV
- ⇒ FJ17:  $a_{\mu, \text{FJ17}}^{\text{had, LO VP}} = 688.07 \pm 41.4$ 
  - Much lower mean value, but in agreement within errors

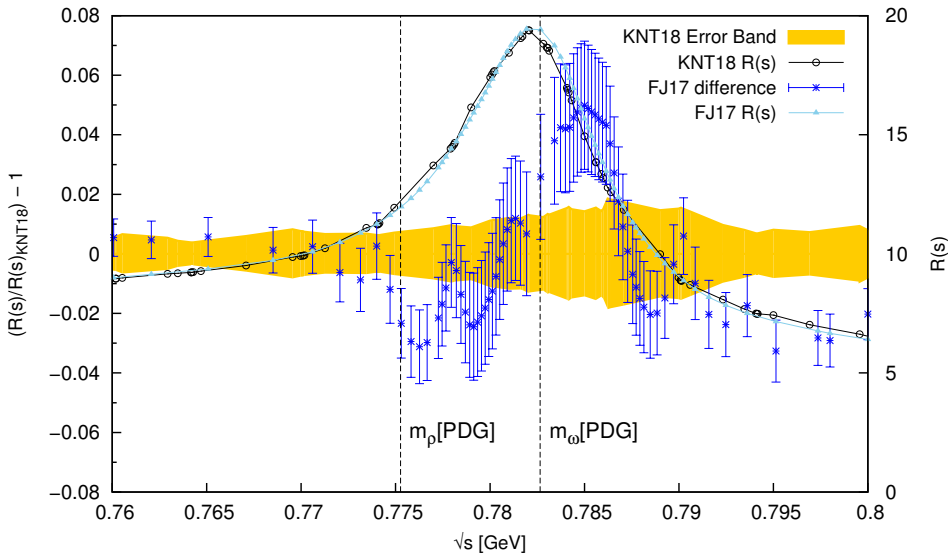
## Comparison tables

Channel	KNT18	DHMZ17	Difference
Data based channels ( $\sqrt{s} \leq 1.8$ GeV)			
$\pi^+\pi^-$	$503.74 \pm 1.96$	$506.70 \pm 2.58$	-2.96
$\pi^+\pi^-\pi^0$	$47.70 \pm 0.89$	$46.20 \pm 1.45$	1.50
$\pi^+\pi^-\pi^+\pi^-$	$13.99 \pm 0.19$	$13.68 \pm 0.31$	0.31
$\pi^+\pi^-\pi^0\pi^0$	$18.15 \pm 0.74$	$18.03 \pm 0.54$	0.12
$K^+K^-$	$23.00 \pm 0.22$	$23.06 \pm 0.41$	-0.06
$K_S^0K_L^0$	$13.04 \pm 0.19$	$12.82 \pm 0.24$	0.22
Total	$693.3 \pm 2.5$	$693.1 \pm 3.4$	0.2

Channel	KNT18	FJ17	Difference
Data based channels ( $0.318 \leq \sqrt{s} \leq 2$ GeV)			
$\pi^+\pi^-$	$501.68 \pm 1.71$	$502.16 \pm 2.44$	-0.48
$\pi^+\pi^-\pi^0$	$47.83 \pm 0.89$	$44.32 \pm 1.48$	3.51
$\pi^+\pi^-\pi^+\pi^-$	$15.17 \pm 0.21$	$14.80 \pm 0.36$	0.37
$\pi^+\pi^-\pi^0\pi^0$	$19.80 \pm 0.79$	$19.69 \pm 2.32$	0.11
$K^+K^-$	$23.05 \pm 0.22$	$21.99 \pm 0.61$	1.06
$K_S^0K_L^0$	$13.05 \pm 0.19$	$13.10 \pm 0.41$	-0.05
Total	$693.27 \pm 2.46$	$688.07 \pm 4.14$	5.20

Channel	KNT18	Benayoun et. al	Difference
Data based channels ( $\sqrt{s} \leq 1.05$ GeV)			
$\pi^+\pi^-$	$495.86 \pm 1.94$	$489.83 \pm 1.22$	6.03
$\pi^+\pi^-\pi^0$	$44.49 \pm 0.80$	$42.94 \pm 0.52$	1.55
$K^+K^-$	$18.12 \pm 0.18$	$17.18 \pm 0.25$	0.94
$K_S^0K_L^0$	$11.97 \pm 0.17$	$11.87 \pm 0.25$	0.10



Comparison of  $\rho - \omega$  with FJ17

Comparison of  $\phi$  with FJ17