

PRISMA+

MPA

MAINZ PHYSICS
ACADEMY

XENON

First results from XENONnT

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Patras Workshop 2022 JGU-Mainz



Bundesministerium
für Bildung
und Forschung



Studienstiftung
des deutschen Volkes



XENON Collaboration:



- 27 institutes
- 167 members

XENONnT experiment:

XENON1T  XENONnT upgrades:

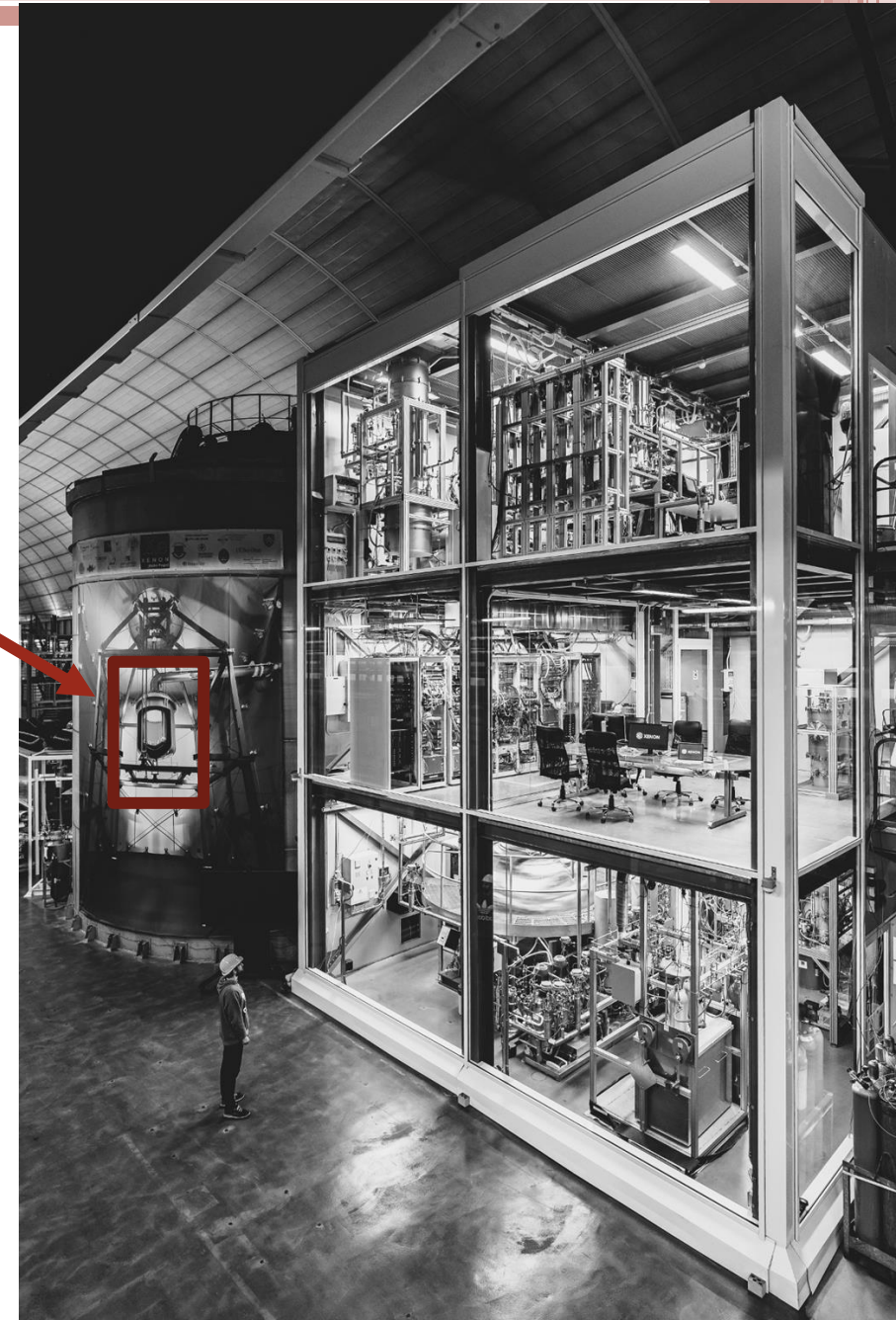
- Larger TPC and inner cryostat
- New purification and distillation system
- Additional water Cherenkov neutron-veto
- New calibration systems and techniques
- New analysis software package STRAXEN and triggerless data acquisition
- Improved cleanliness and radiopurity



XENONnT experiment:

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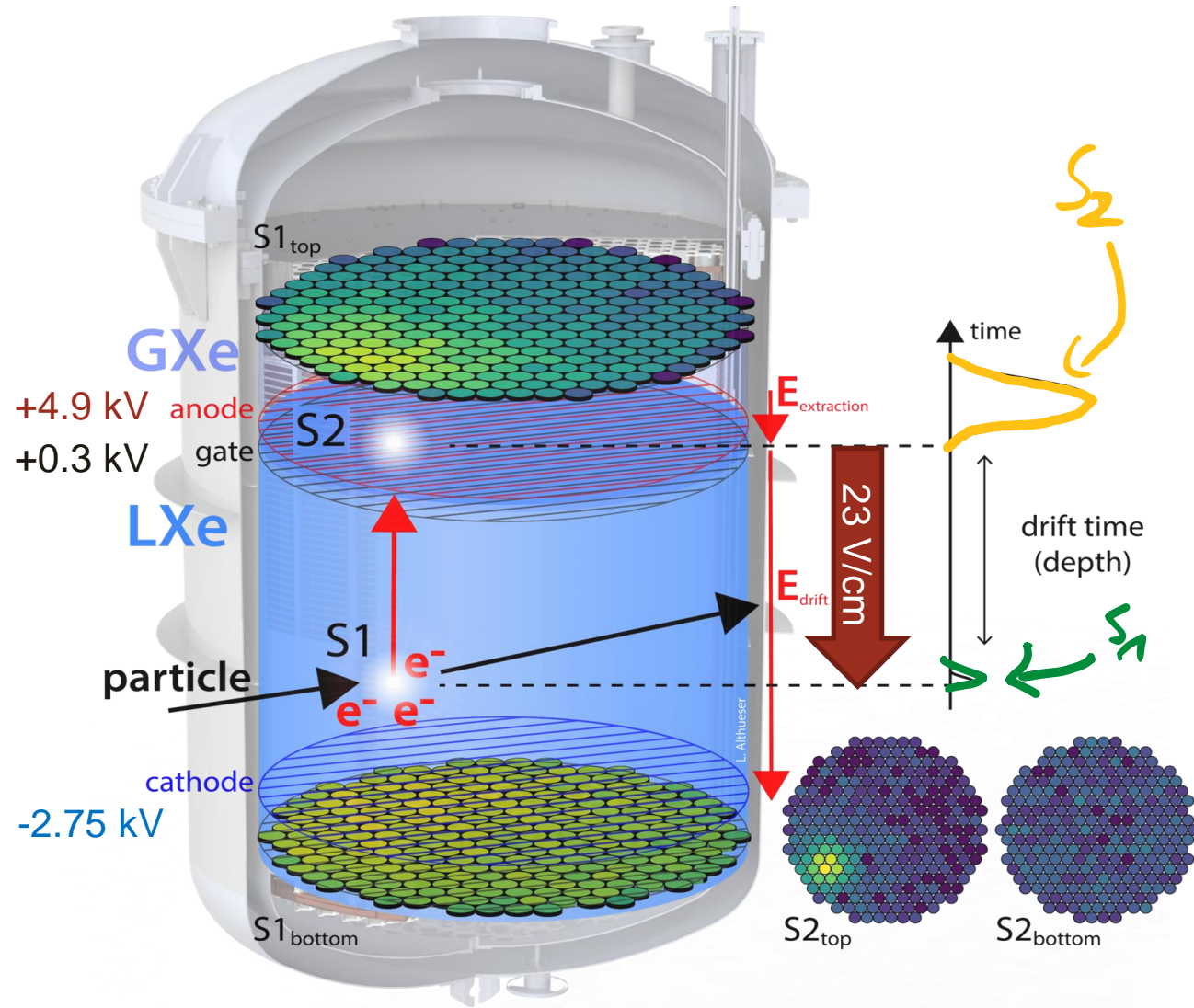


Liquid noble gas time projection chamber:

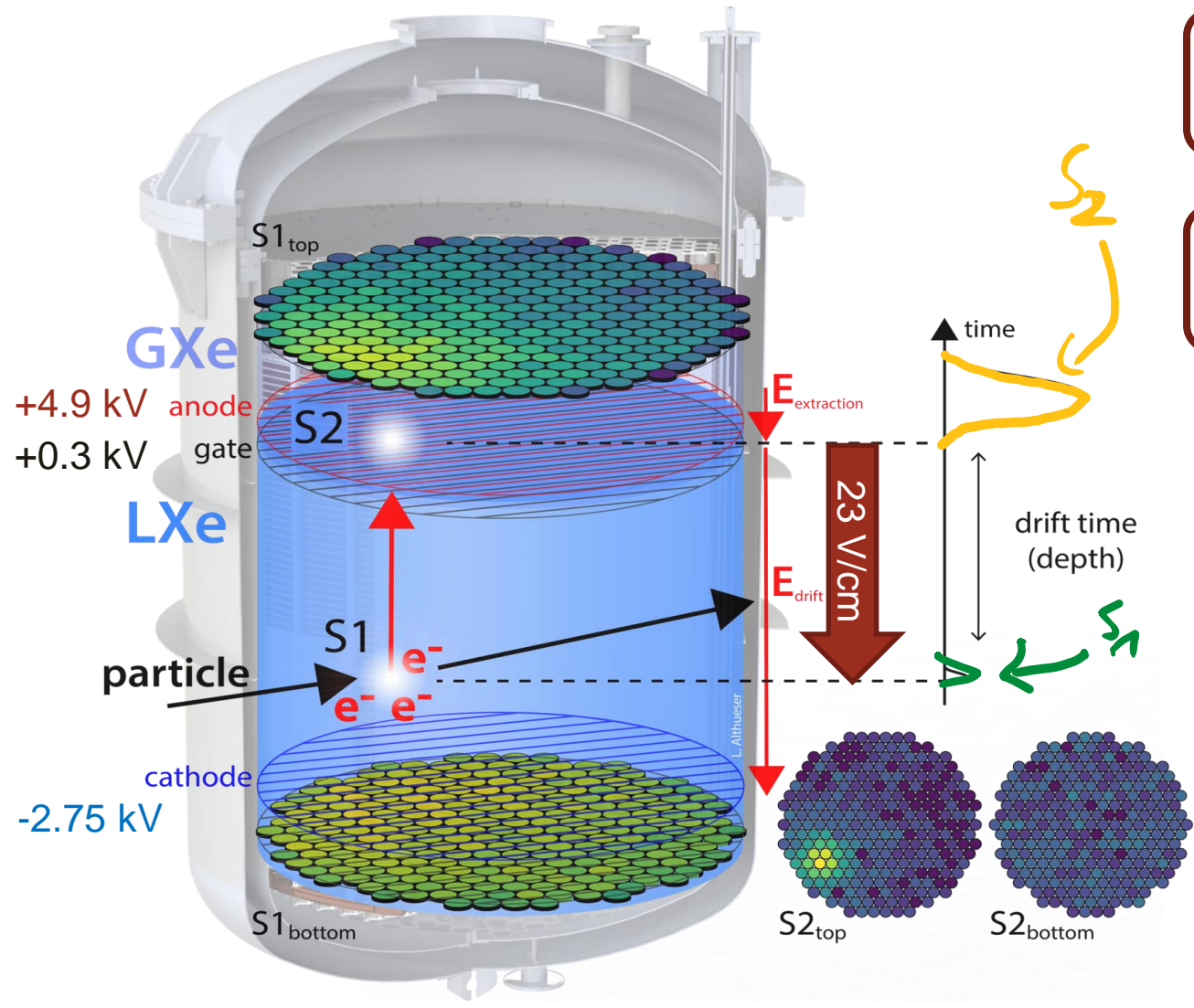


- 1.3 m x 1.5 m in height and diameter
- 494 3" PMTs (R11410-21) densely packed in a top and bottom array
- Cylinder walls made out of PTFE for high reflectivity
- 5 meshes to put electric fields and protect PMTs
- All materials carefully screened and selected

Liquid noble gas time projection chamber:



Liquid noble gas time projection chamber:



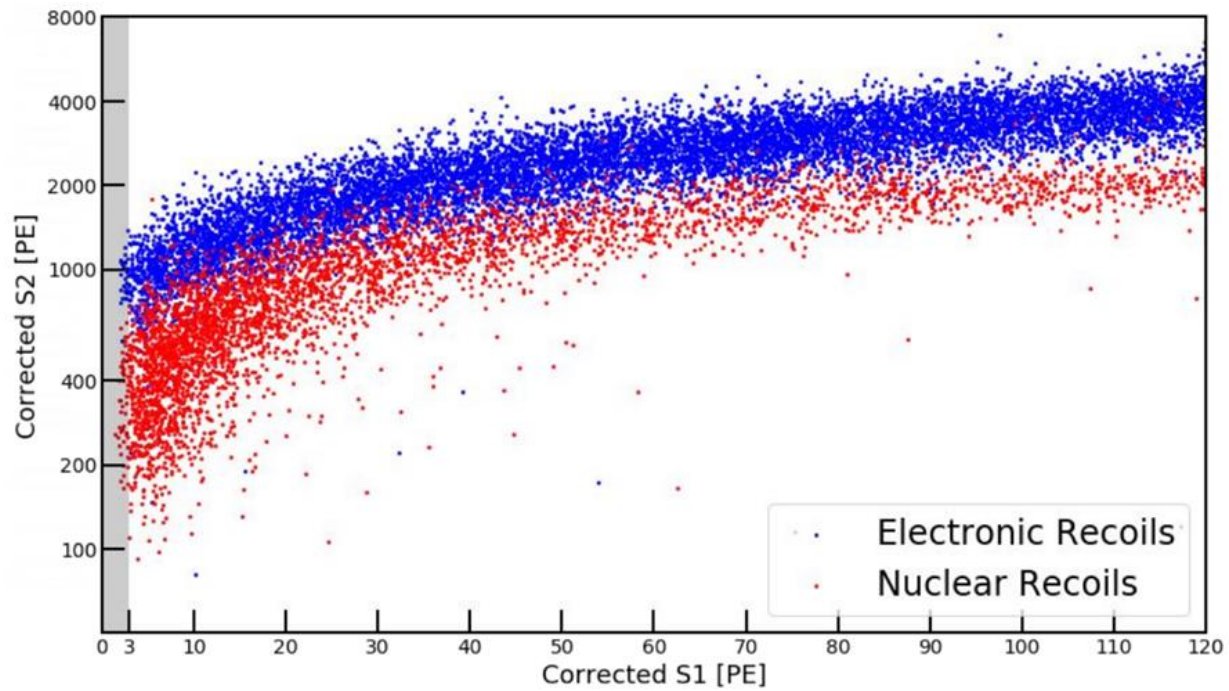
Electronic Recoils:



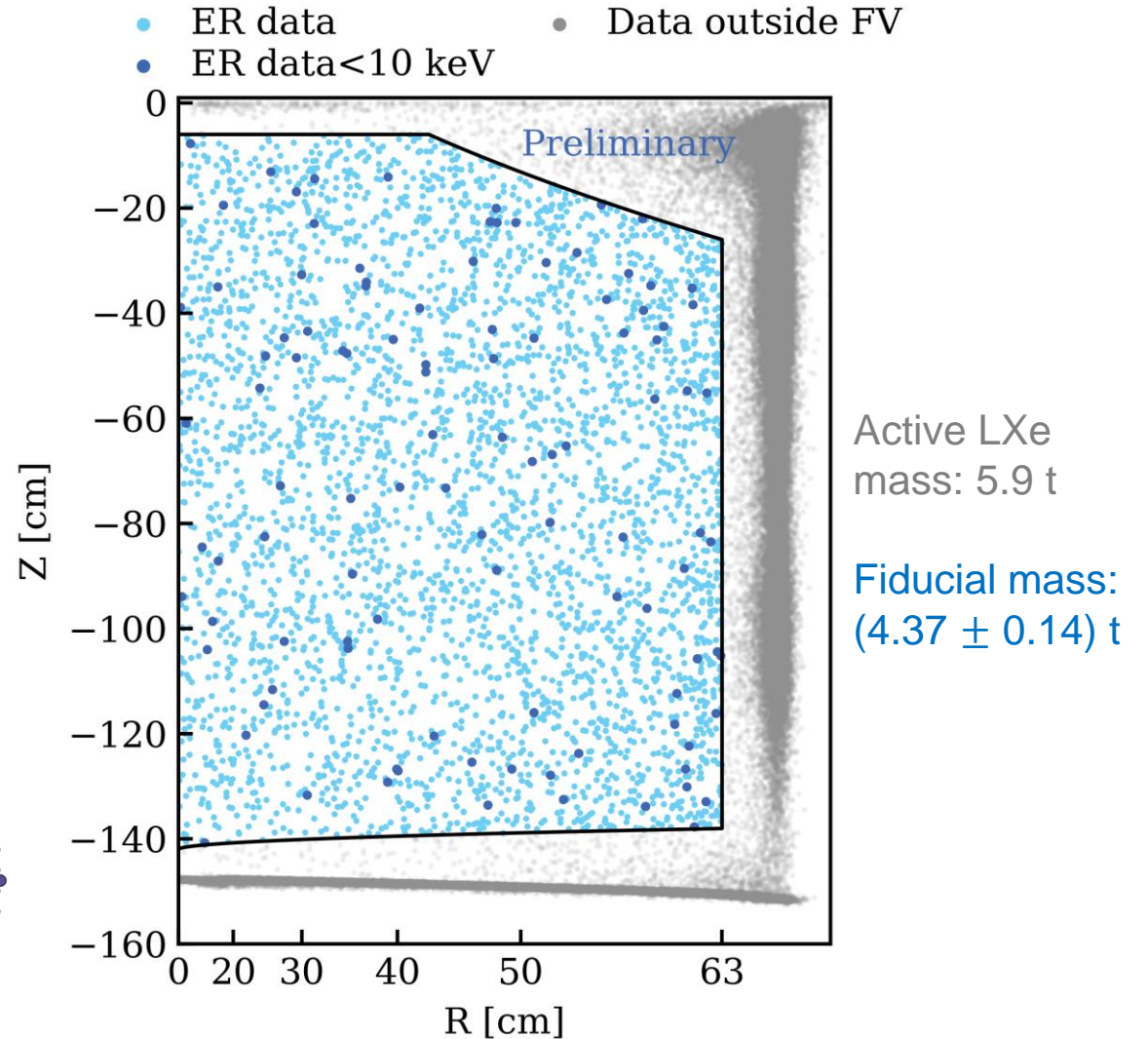
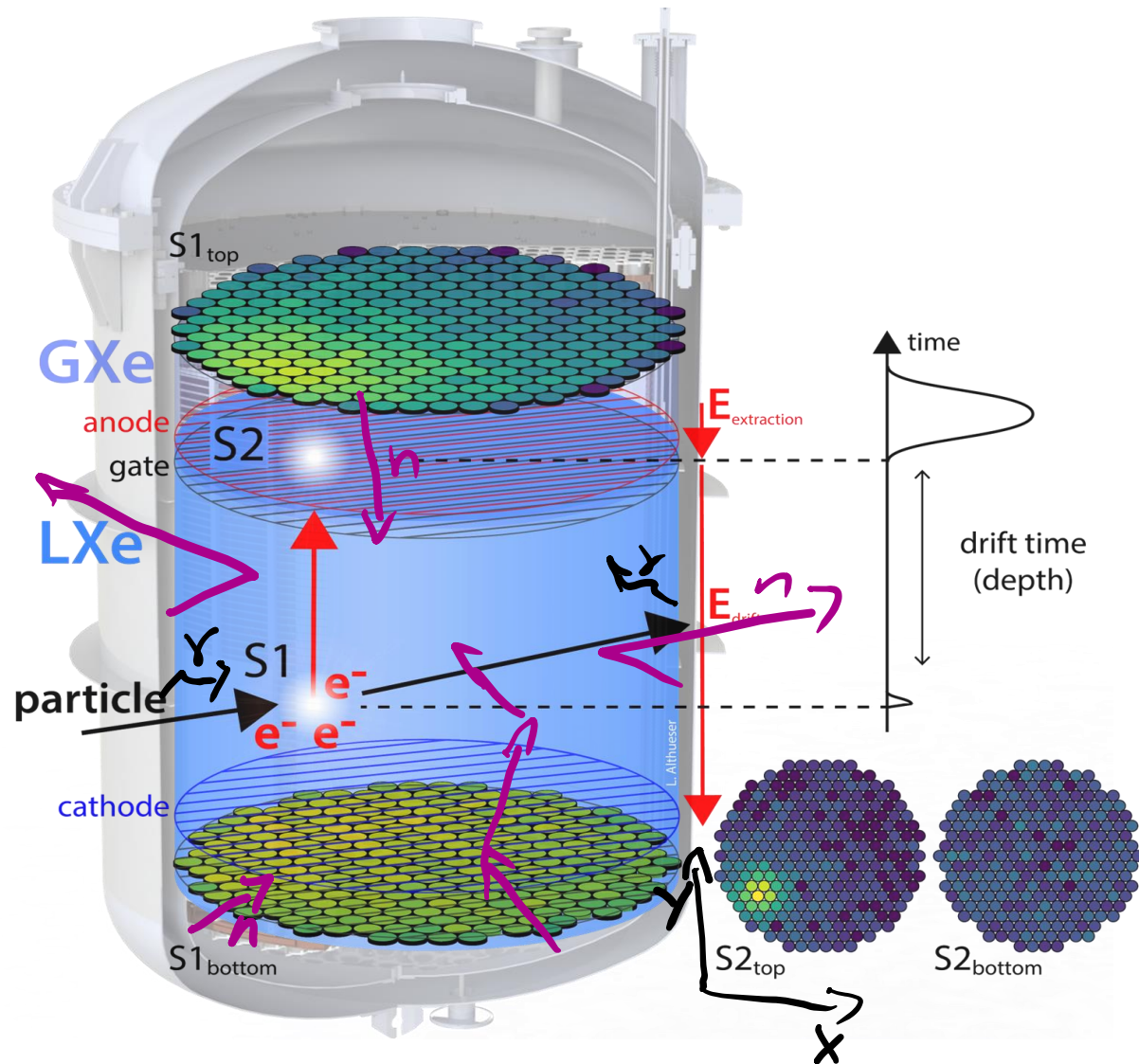
Nuclear Recoils:



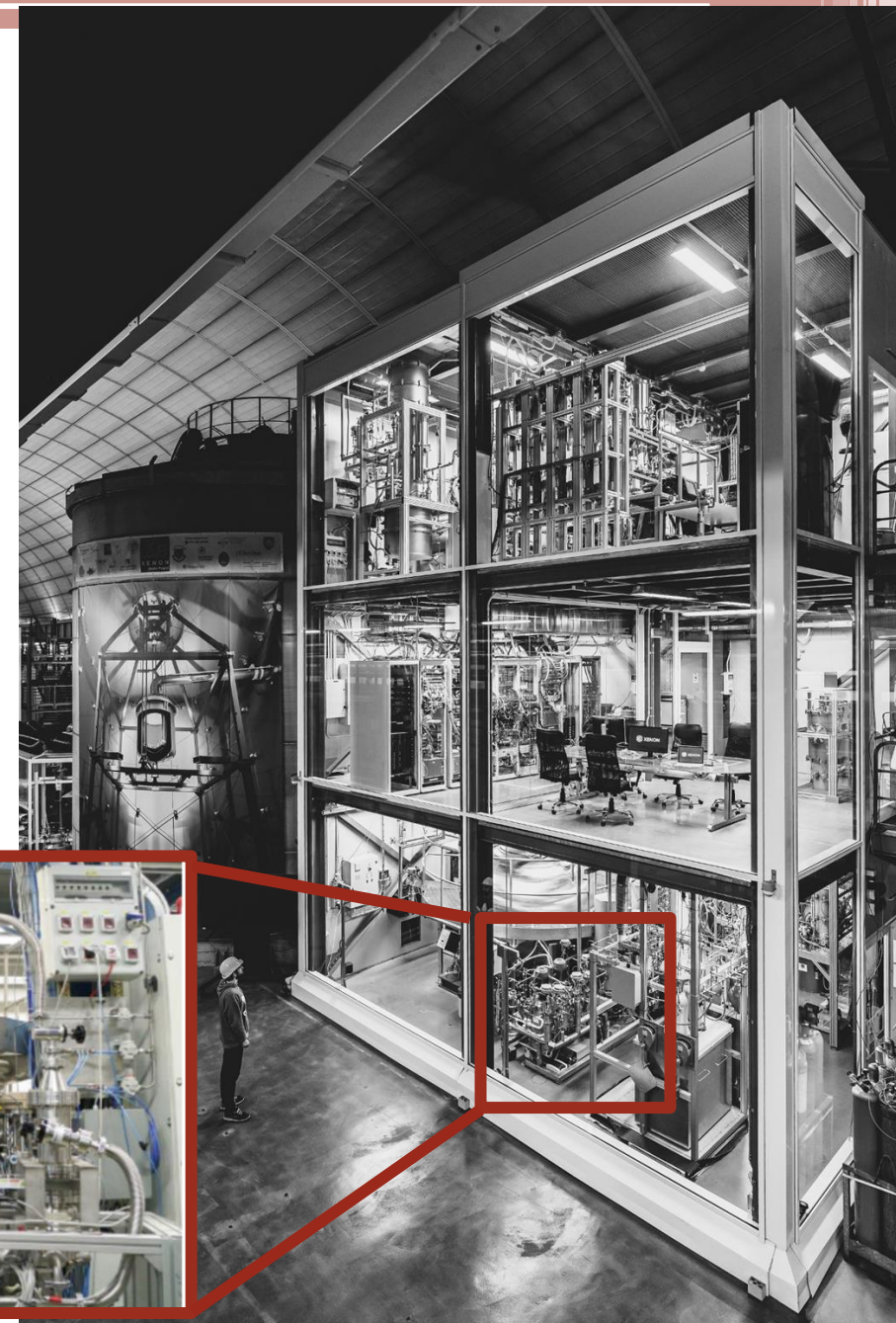
Ratio: $S2/S1$ (ER) > $S2/S1$ (NR)



Liquid noble gas time projection chamber:



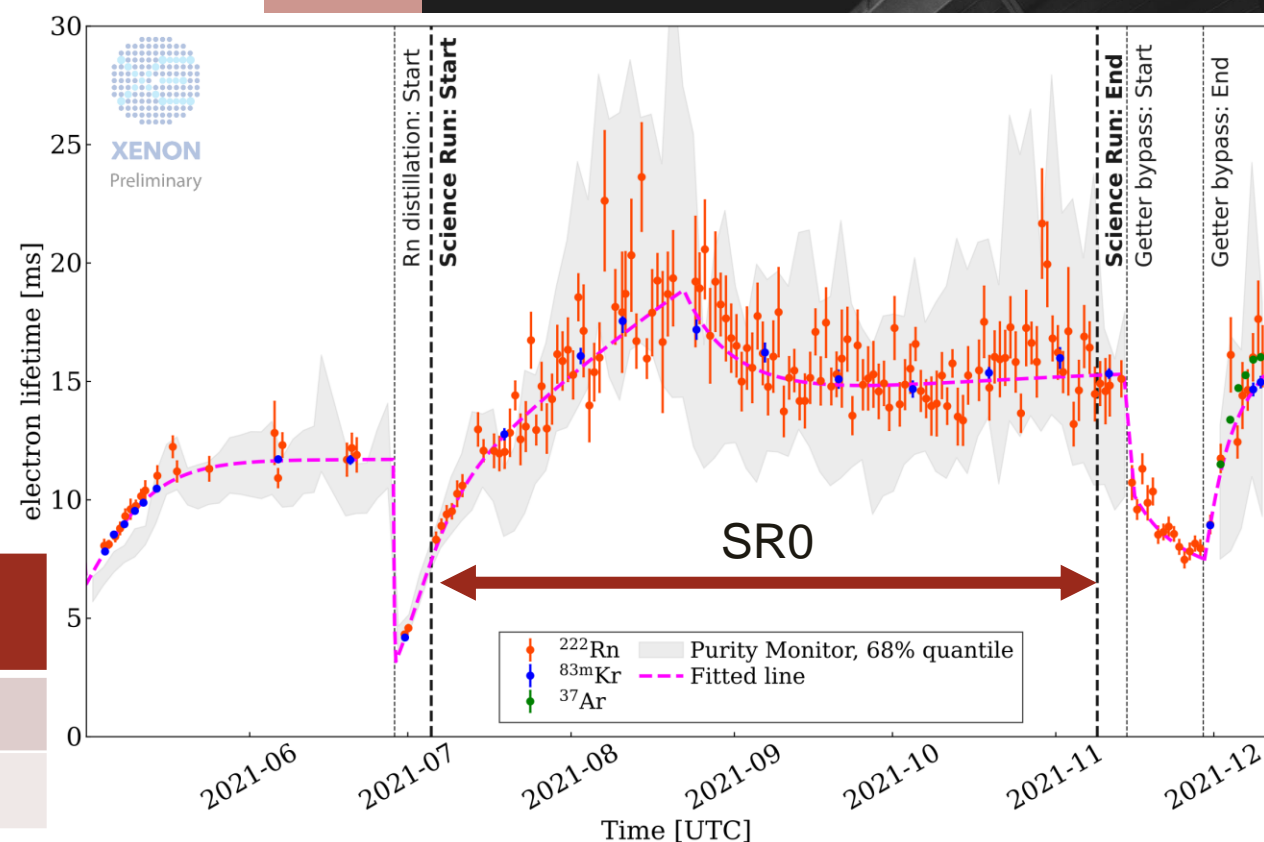
Liquid xenon purification



Liquid xenon purification

- Xenon purity important to drift electrons
- New liquid purification technique:
 - Replaceable filter units
 - Low radon emanation
 - 2 liters **liquid** / minute → 18 h for entire volume

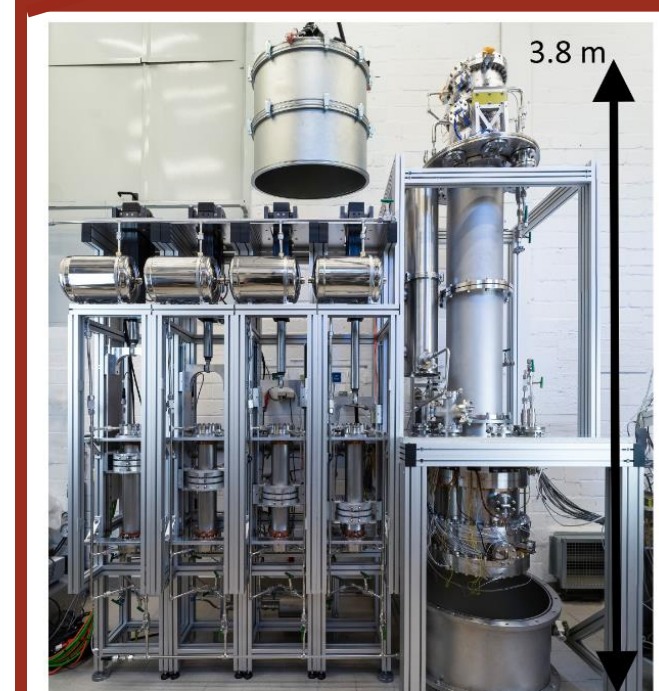
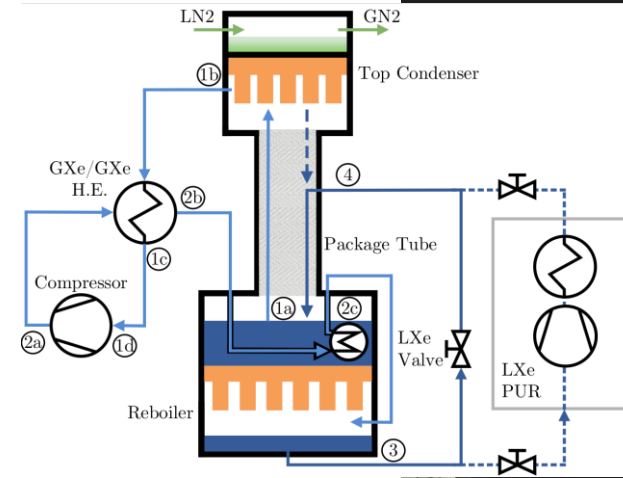
	Full drift time:	Electron lifetime:	Electron survival (@full drift length):
1T	0.67 ms	0.65 ms	30 %
nT	2.2 ms	>10 ms	86 % @ 15 ms



G. Plante, E. Aprile, J. Howlett, Y. Zhang
arXiv:2205.07336 [physics.ins-det]



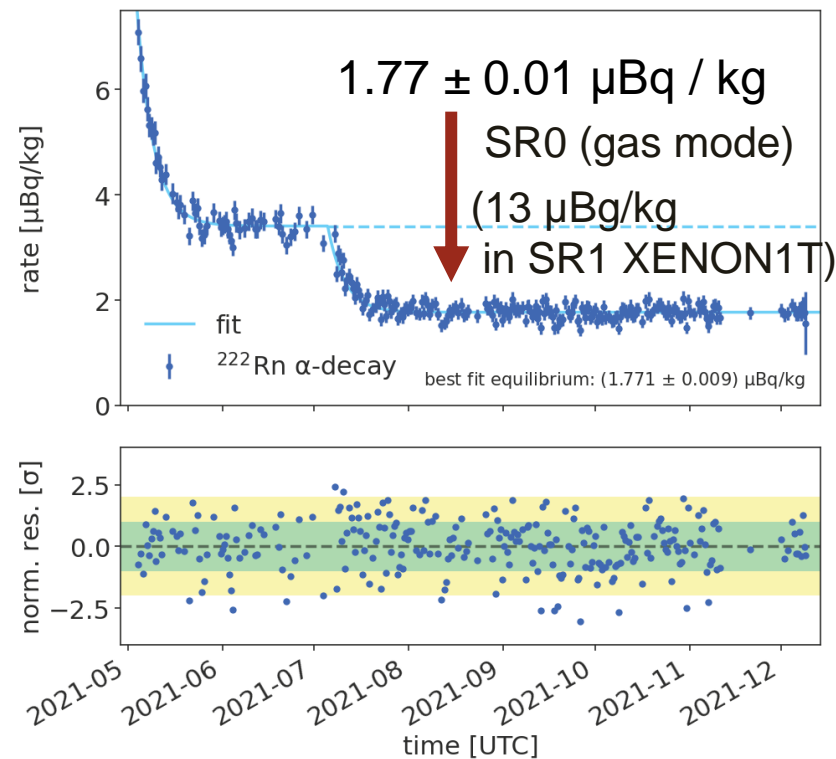
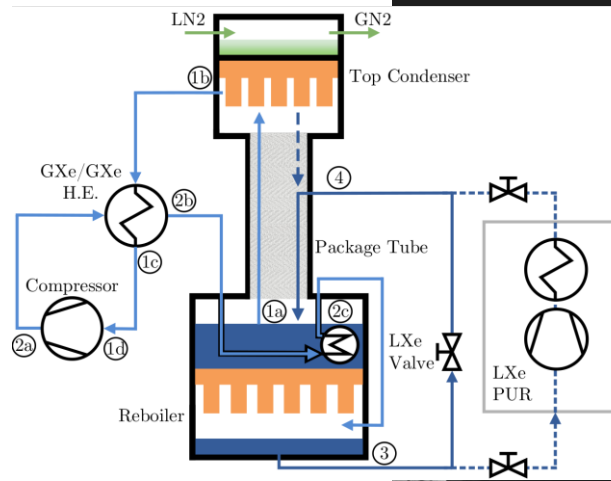
Radon distillation:



Radon distillation:

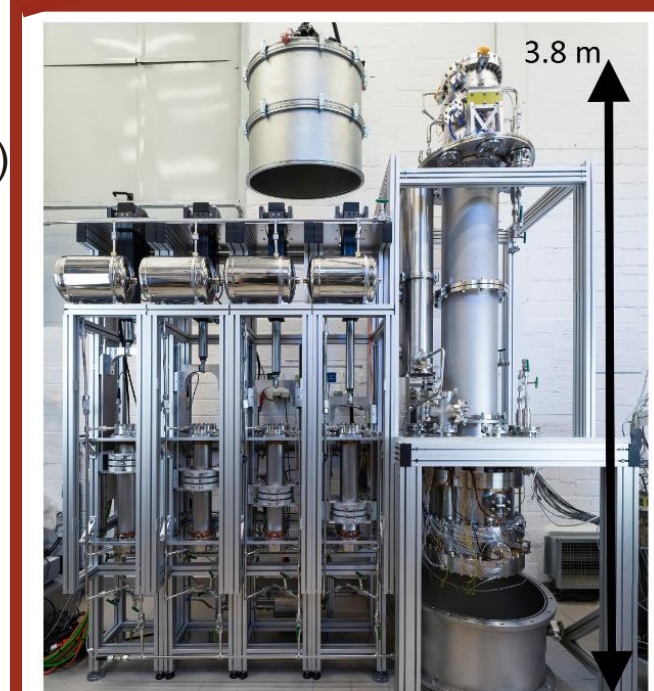
- Rn222 dominant background in 1T/nT
- Added radon distillation together with radon free magnetic-piston pumps
- In SR0 operated in gas-mode
- Preliminary:

We are able to reach values $< 1 \mu\text{Bq/kg}$ for SR1 in liquid + gas mode

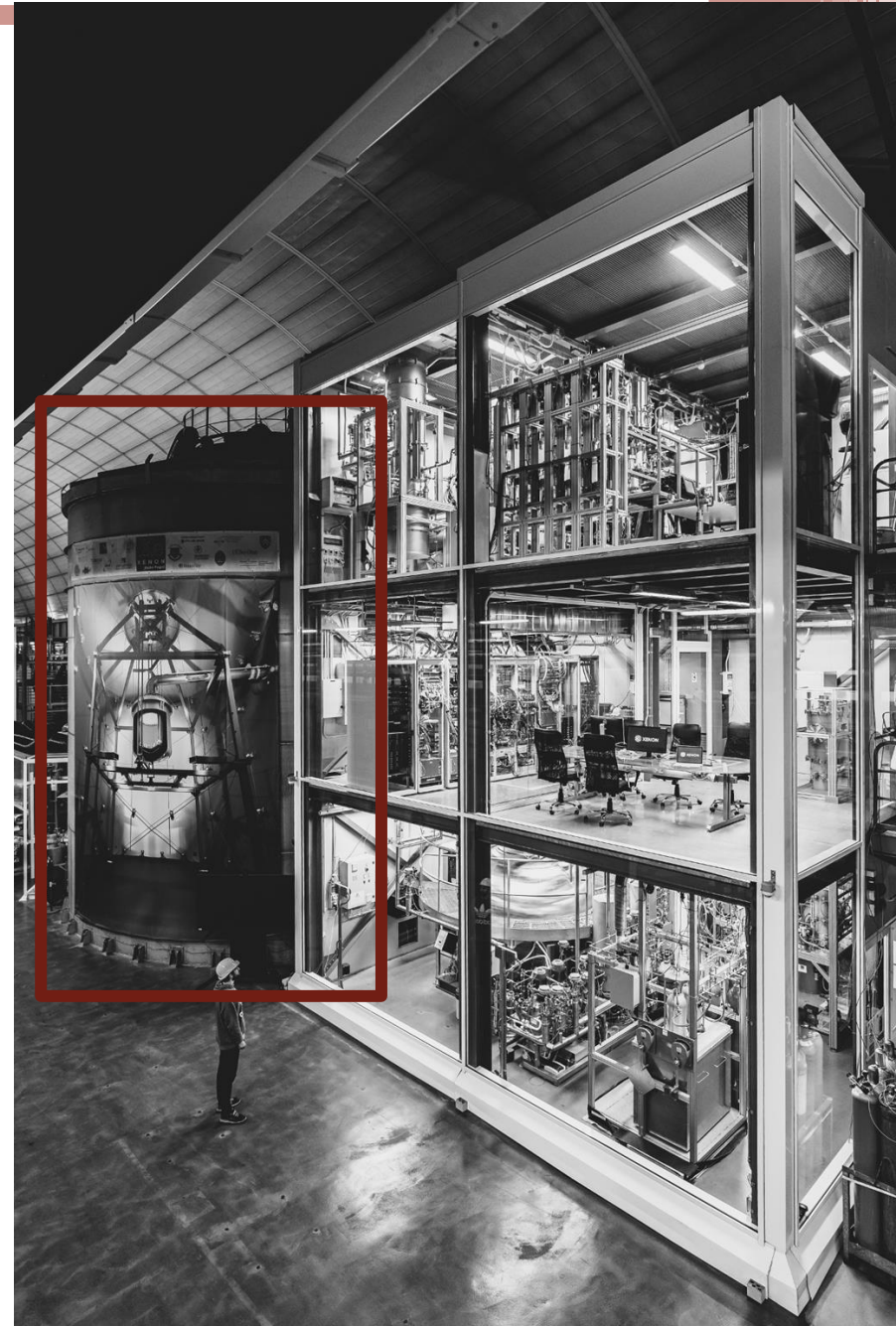


M. Murra, D. Schulte, C. Huhmann,
C. Weinheimer

[arXiv:2205.11492](https://arxiv.org/abs/2205.11492) [physics.ins-det]



XENONnT neutron-veto:



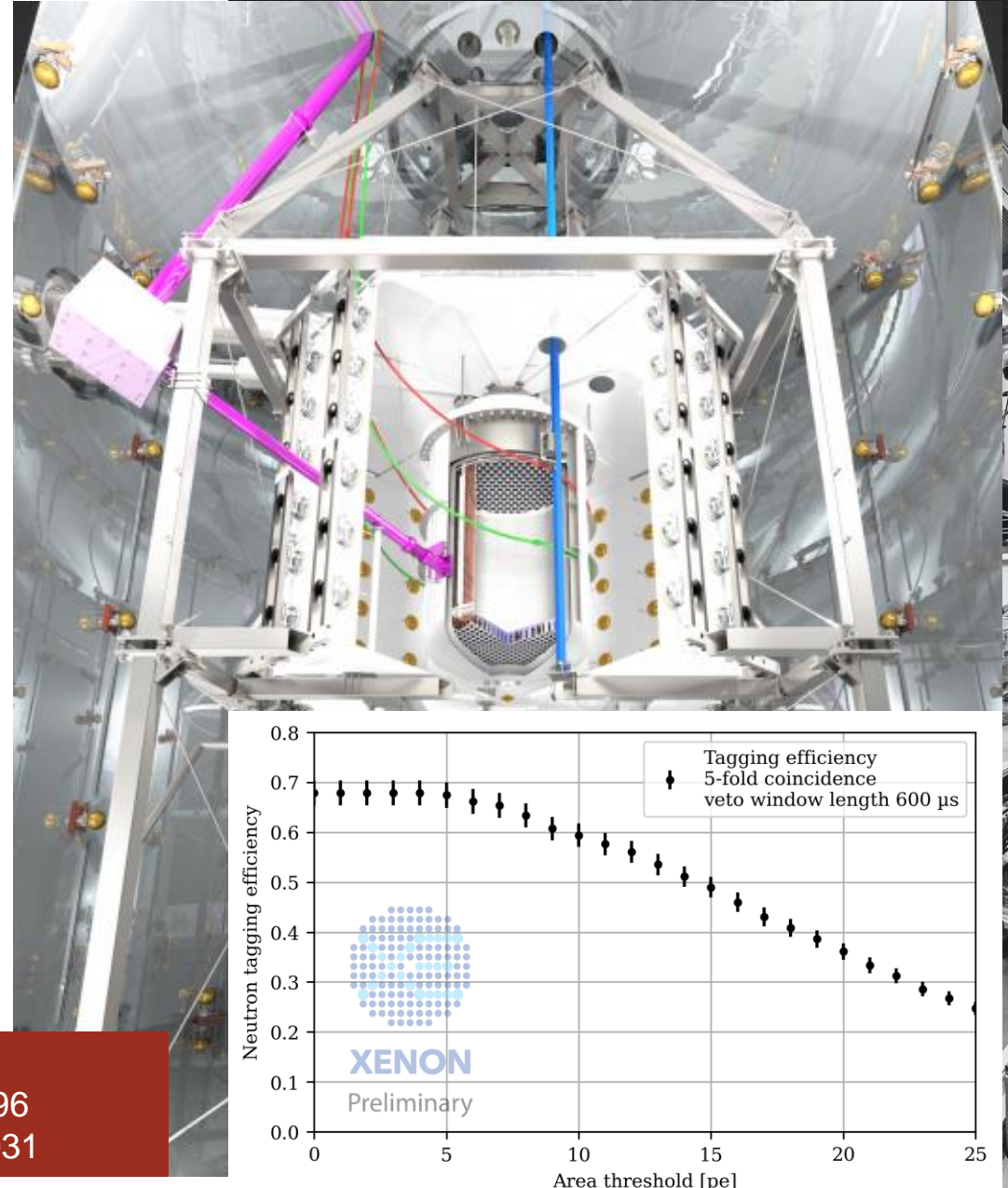
XENONnT neutron-veto:

- Added new water Cherenkov neutron-veto into muon-veto water tank
- 120 8" PMTs are watching the TPC cryostat
- Highly reflective ePTFE and ultra-pure water to maximize light collection efficiency
- Using tagged neutrons to calibrate the NR response as well as the neutron-veto tagging efficiency.

➔ Preliminary tagging efficiency of $(67.5 \pm 2.9)\%$

- After SR0 loading with 0.2 % of $(\text{Gd}_2(\text{SO}_4)_3 \times 8(\text{H}_2\text{O}))$

➔ Expected tagging efficiency $\sim 87\%$

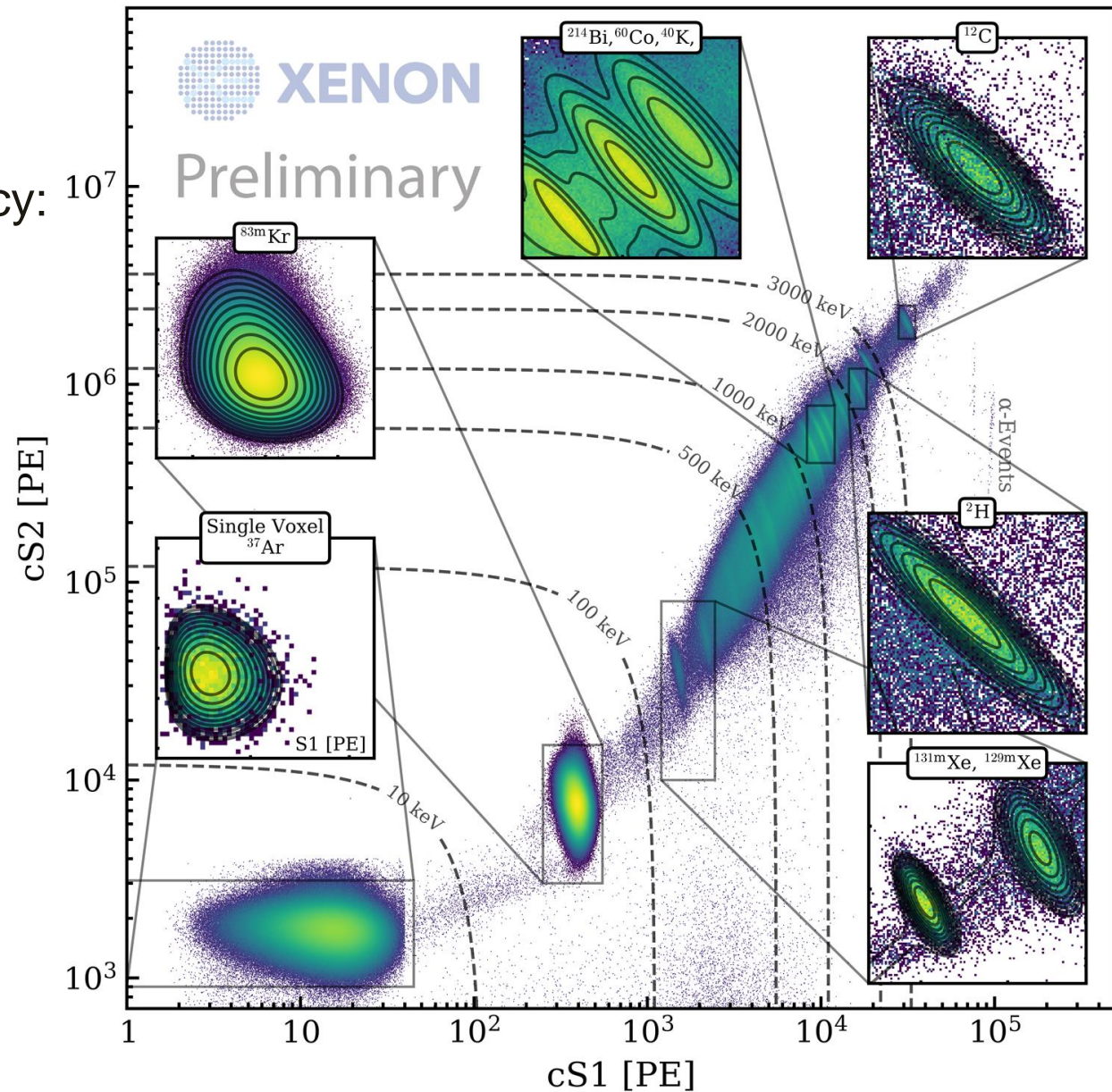
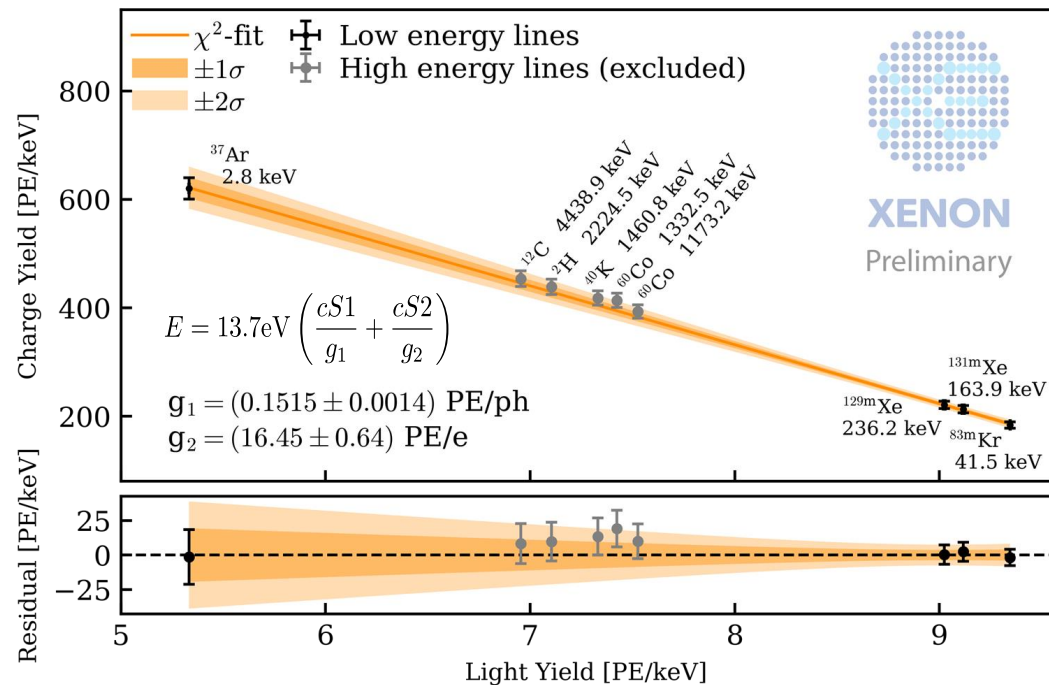


E. Aprile *et al*
arXiv:2007.08796
JCAP11(2020)031

Calibration of XENONnT:

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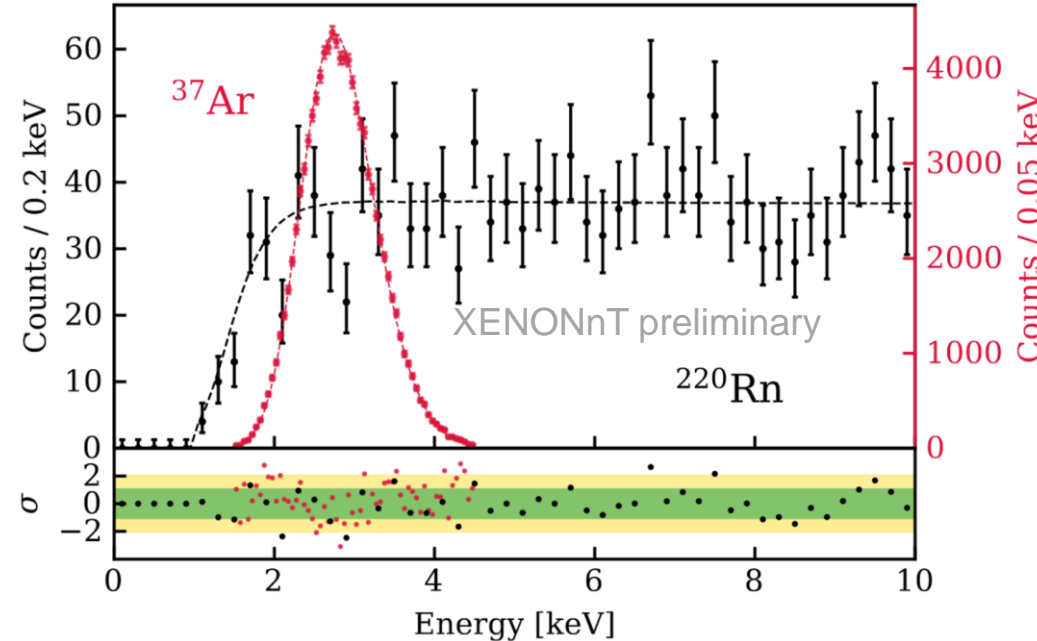
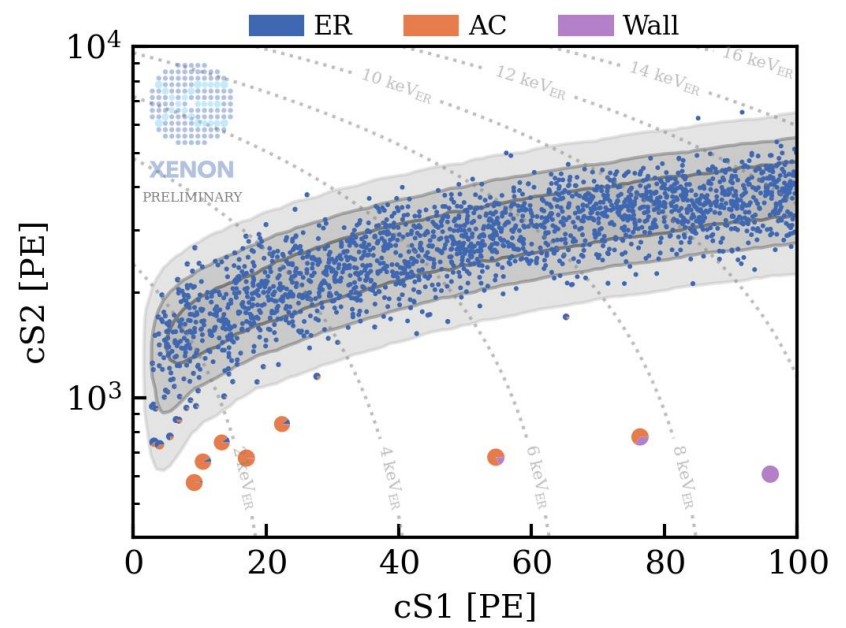
- Weekly PMT calibrations via LEDs
- Calibration of detector response and efficiency:
 - Use internal source: ^{37}Ar , $^{83\text{m}}\text{Kr}$, $^{129\text{m}}\text{Xe}$, $^{131\text{m}}\text{Xe}$
 - Bias is used as systematic uncertainties
- Bi-weekly $^{83\text{m}}\text{Kr}$ and materials background α and γ are used for stability monitoring
- $^{83\text{m}}\text{Kr}$ used to validate drift field



Calibration of XENONnT:

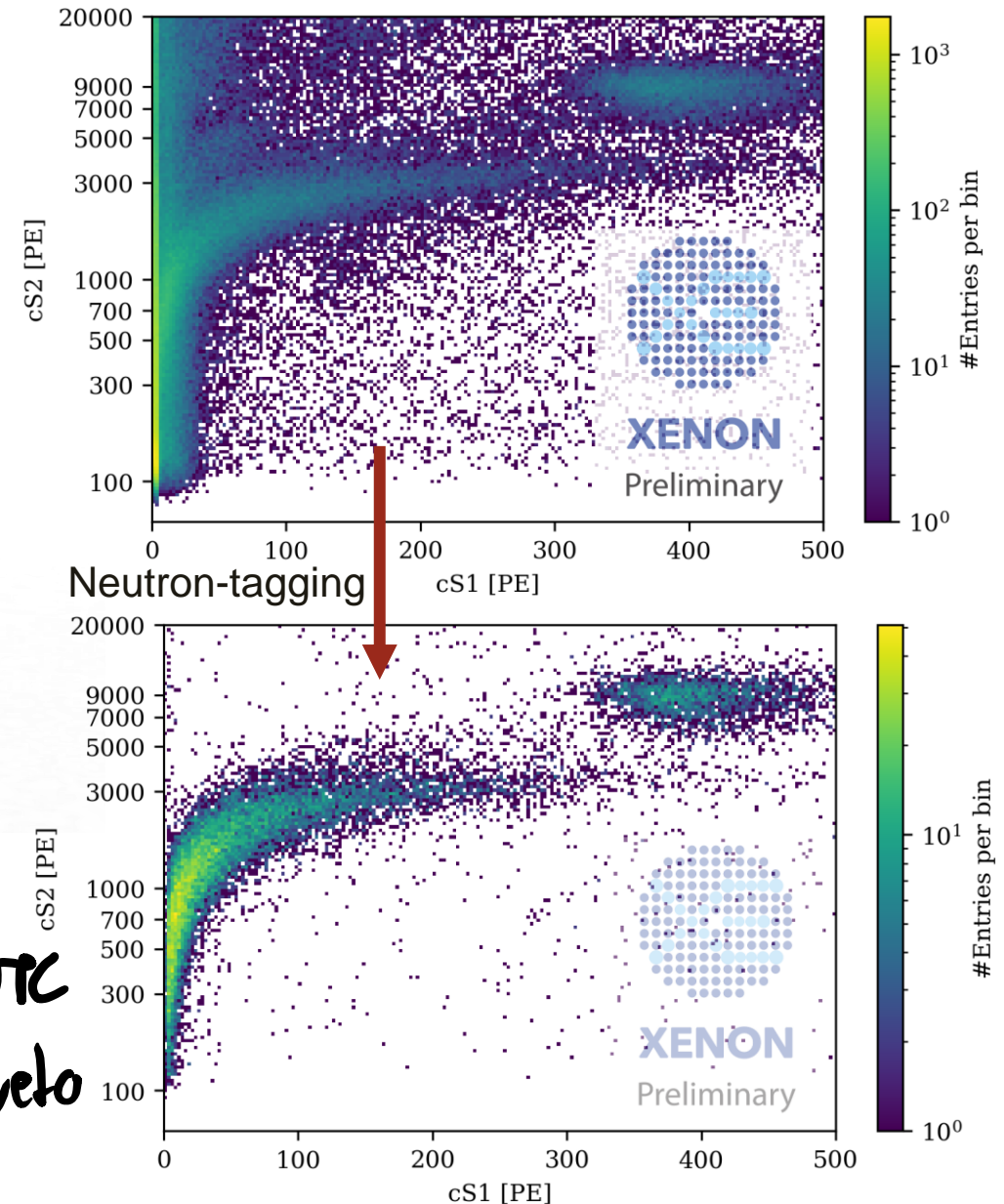
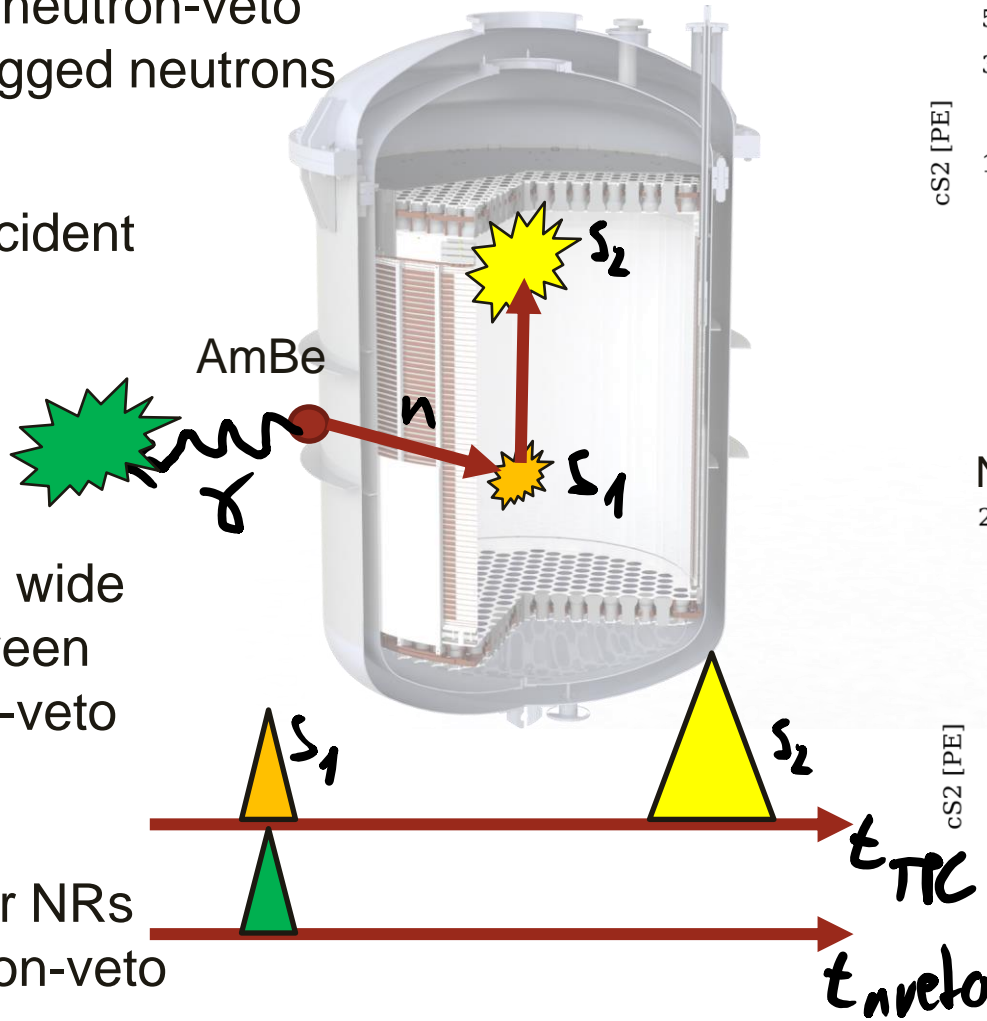
- Calibration of ER response using ^{220}Rn
 - Gives approximately flat spectrum
 - Used to validate cut acceptances and detector threshold
- Detector performance at low energies using ^{37}Ar
 - Mono-energetic line @2.8 keV
 - Allows to study performance with high resolution, due to high statistics
 - Source was produced @TRIGA Mainz
- ER response model based on a combined fit

See also the poster by C. Hils



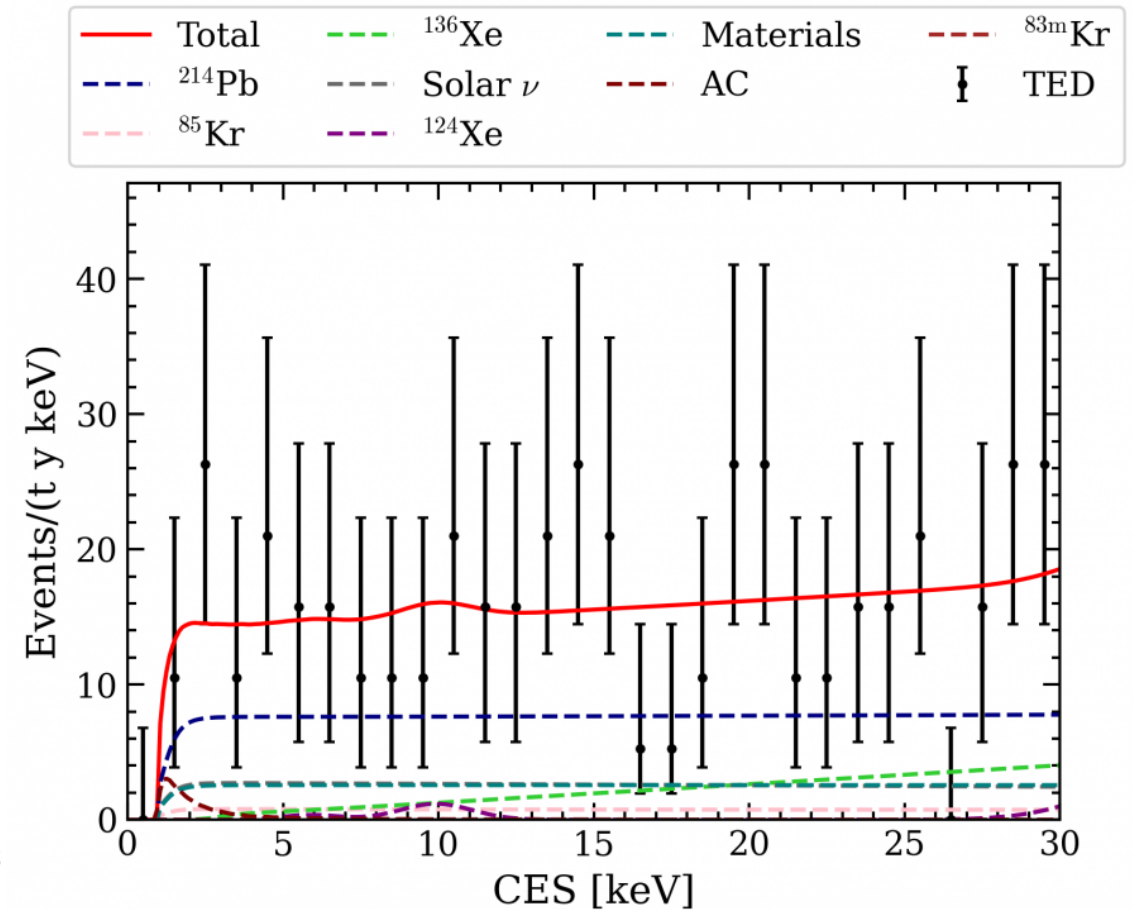
Calibration of NR response:

- Calibration of the neutron-veto and TPC using tagged neutrons
- AmBe emits coincident 4.4 MeV gamma
- Build tight 400 ns wide coincidence between TPC and neutron-veto
- Use single scatter NRs to calibrate neutron-veto tagging efficiency

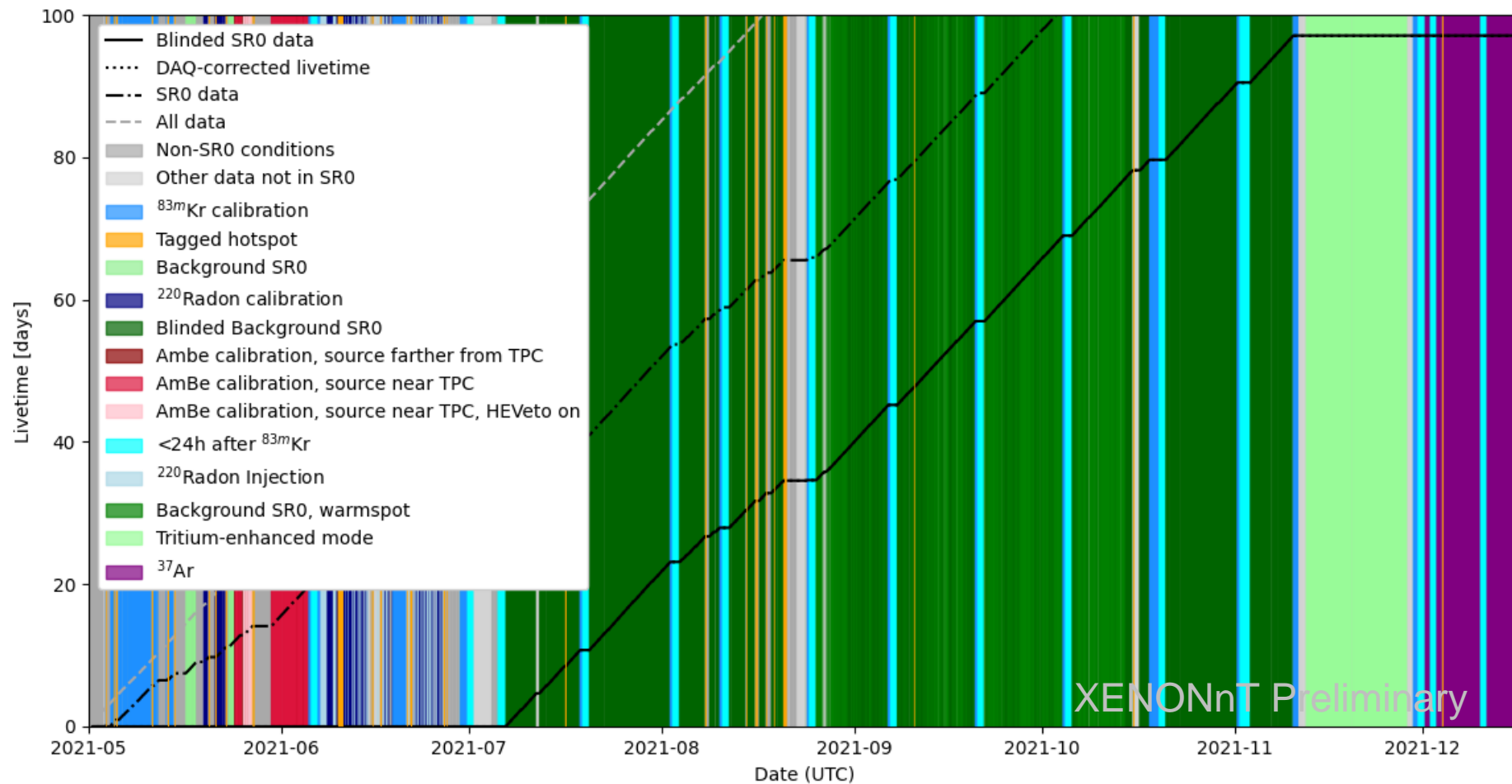


Tritium-enhanced data:

- Tritium as possible explanation for 1T excess.
- Therefore, additional measures in nT:
 - 3 months of outgassing and cleaning procedure
 - All xenon processed through Kr-removal system
 - 3 weeks of GXe circulation
- 14.3 days of operation in tritium-enhanced mode:
 - Bypassing the getter
 - Conservative estimate for tritium enhancement of at least x10 (but could also be a factor x100)
 - No ^3H excess found



SR0 data taking:



97.1 days of exposure between July 6 – Nov 11 2021

Backgrounds:

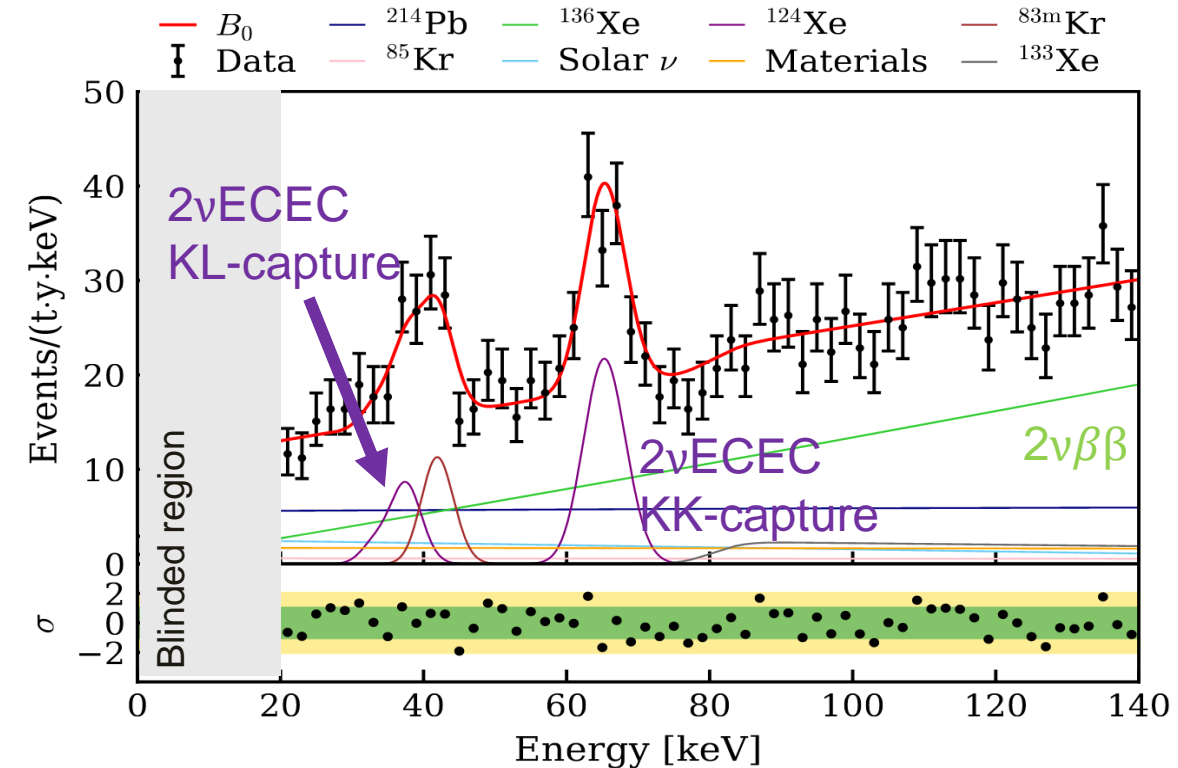
- NR and ER data below 20 keV blinded
- Initial constraints on backgrounds by external measurements and a data-driven coincidence model
- Verification of the background model on side band before unblinding

Component	Constraint	Fit
^{214}Pb	(584, 1273)	
^{85}Kr	90 ± 59	
Materials	266 ± 51	
^{136}Xe	1537 ± 56	
Solar neutrinos	297 ± 30	
^{124}Xe	—	
AC	0.70 ± 0.04	
^{133}Xe	—	
$^{83\text{m}}\text{Kr}$	—	

Backgrounds:

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Component	Constraint	Fit
^{214}Pb	(584, 1273)	980 ± 120
^{85}Kr	90 ± 59	91 ± 58
Materials	266 ± 51	267 ± 51
^{136}Xe	1537 ± 56	1523 ± 54
Solar neutrinos	297 ± 30	298 ± 29
^{124}Xe	—	256 ± 28
AC	0.70 ± 0.04	0.71 ± 0.03
^{133}Xe	—	163 ± 63
$^{83\text{m}}\text{Kr}$	—	80 ± 16



➔ 2nd order weak processes dominating spectral shape!

$$T_{1/2}^{2\nu ECEC} = (1.15 \pm 0.13_{stat} \pm 0.14_{sys}) \cdot 10^{22} \text{ yr}$$

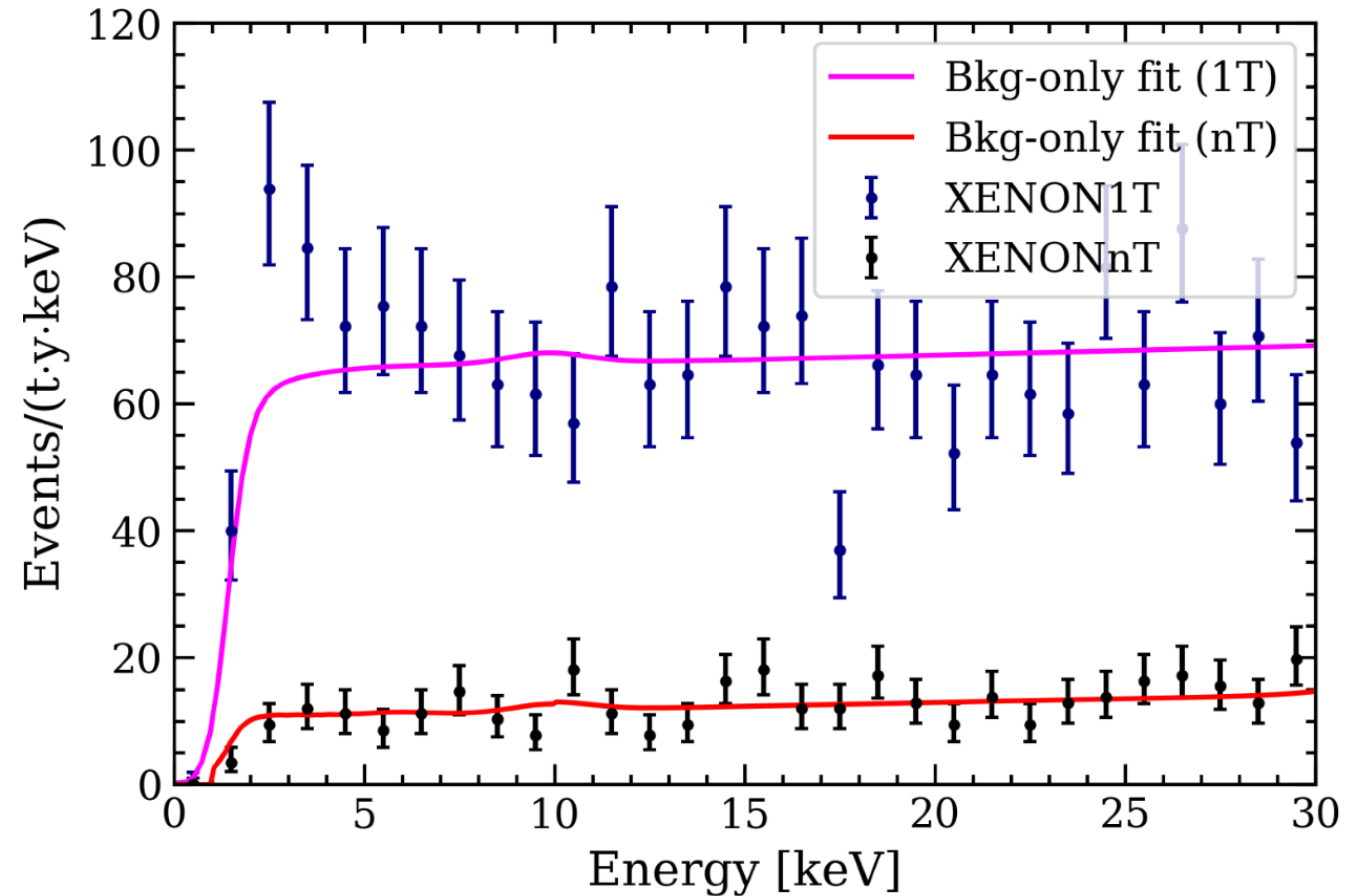
(significance of ~ 10 sigma)

➔ Discovery in XENON1T, now calibration source in nT!

ER results XENONnT:

- Factor x5 improved background compared to XENON1T
- No excess below 5 keV found
 - 8.6σ exclusion of XENON1T-size peak

E. Aprile *et al*
**Search for New Physics in Electronic Recoil
Data from XENONnT**
[arXiv:2207.11330](https://arxiv.org/abs/2207.11330)

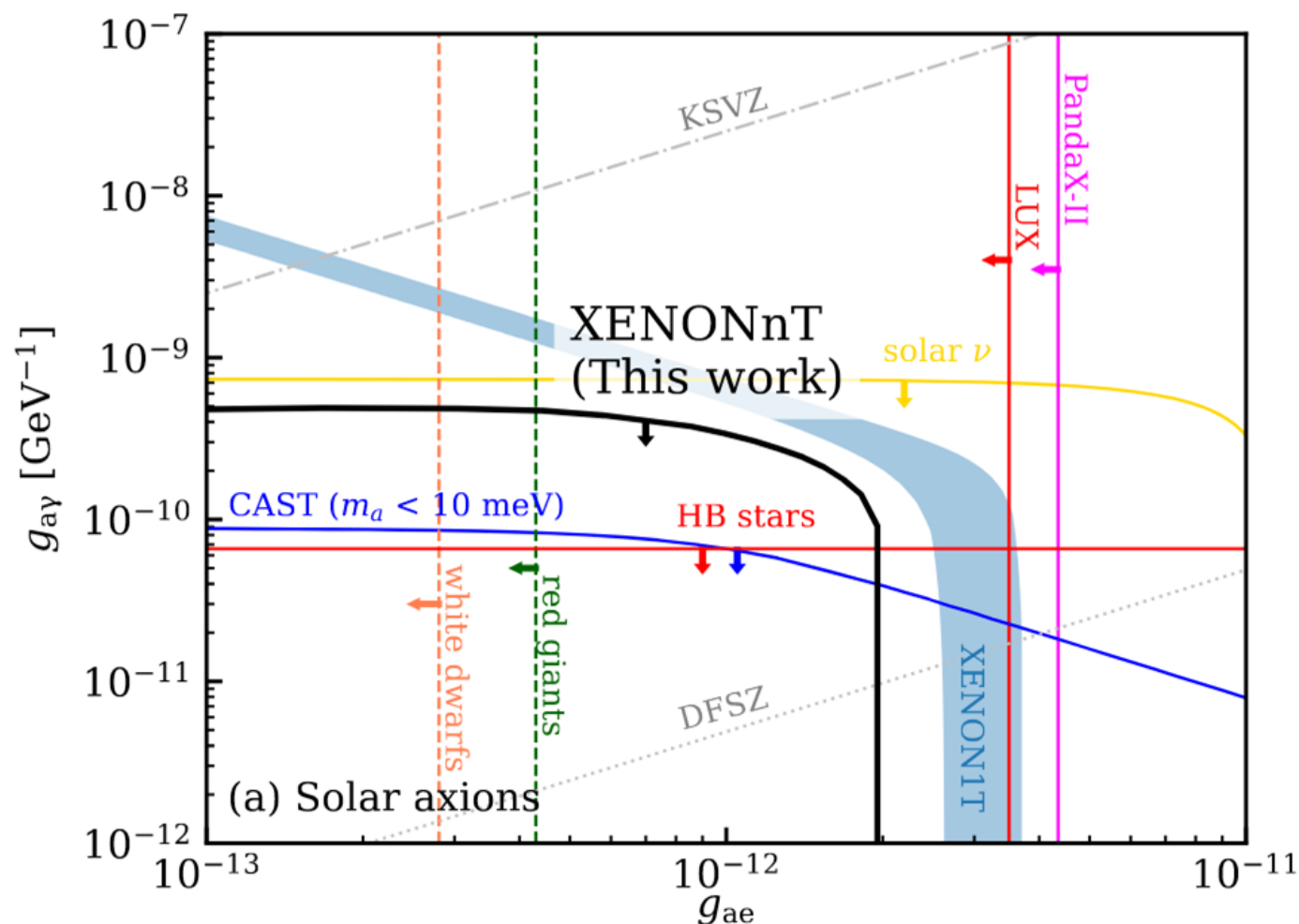


ER results XENONnT:

- Axion signal includes axio-electric and reverse Primakoff effect
- Improved constraints on the axion-gamma, axion-electron and axion-nucleon couplings
- Upper limit on the ^{57}Fe solar axion rate

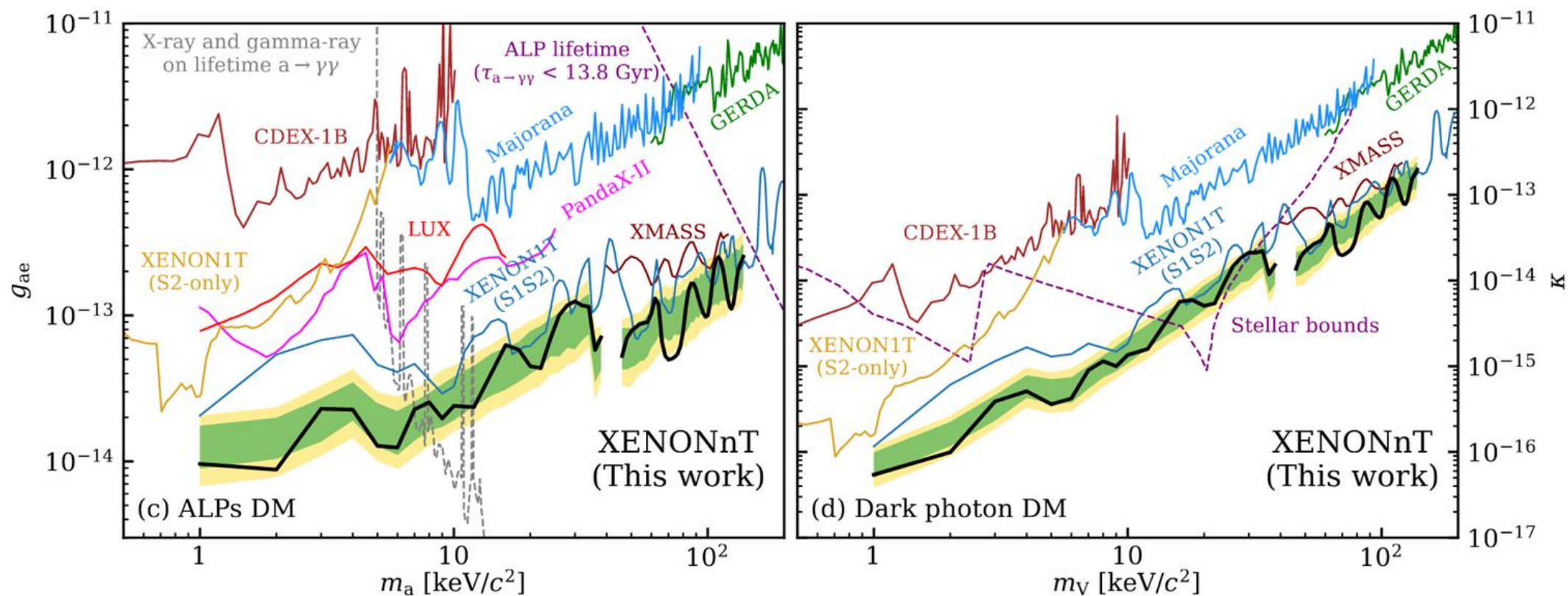


$< 20.4 \text{ ev}/(\text{t}\cdot\text{y})$ (90% C.L.)



ER results XENONnT:

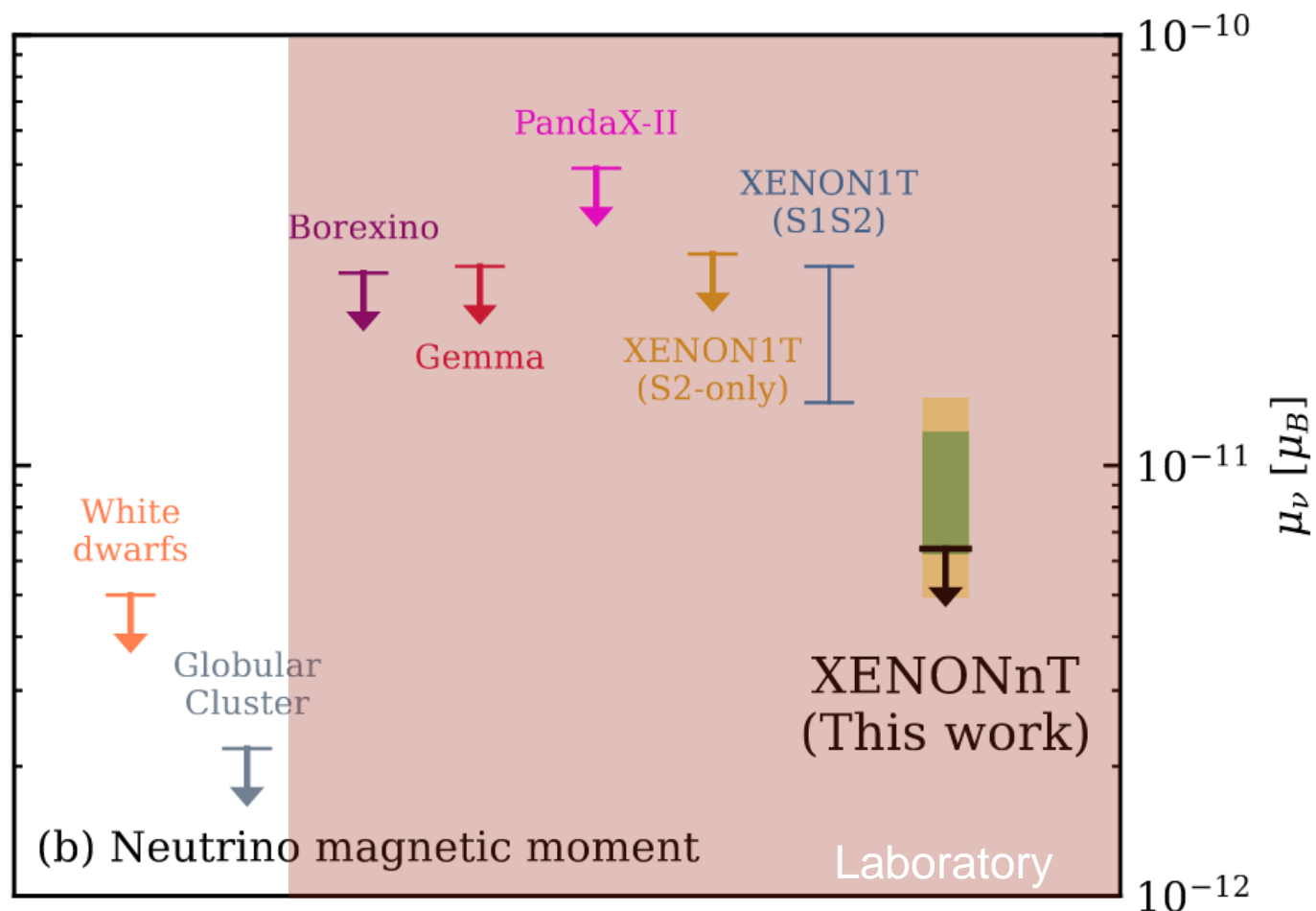
- Put constraints on new physics
- No peak-like signals as expected from axion like particles or dark photons



ER results XENONnT:

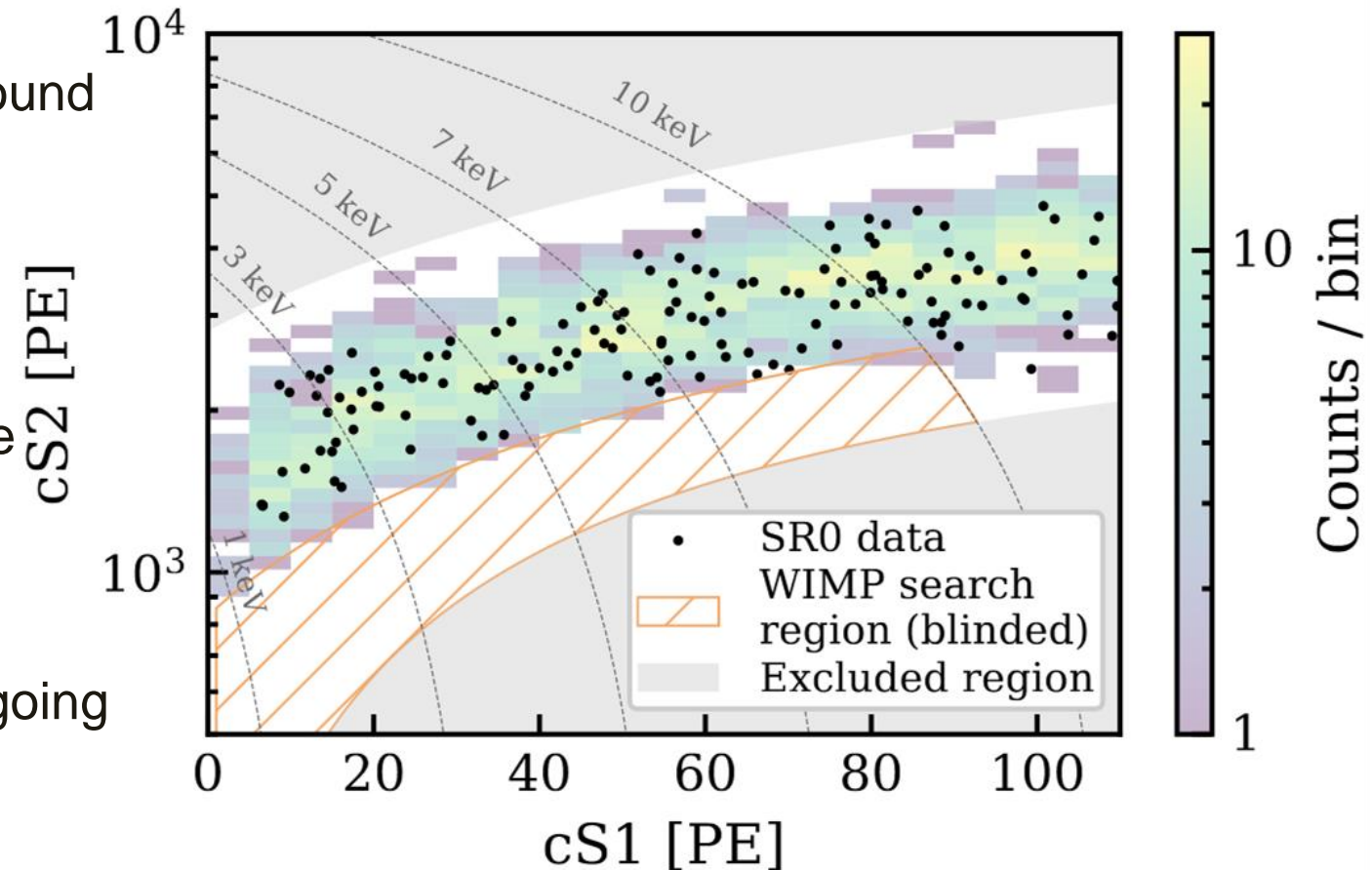
- New constraints on neutrino-magnetic moment:

➔ $\mu_\nu < 6.3 \times 10^{-12} \mu_B$



Outlook and conclusion:

- XENONnT achieves excellent background levels and xenon purity!
- New subsystems work as expected
- ER analysis disfavors a XENON1T like excess, but puts new constraints on different dark matter models
- NR unblinding and WIMP analysis ongoing expect also here new results soon!



Is it a Bird?... Is it a Plane? No it's

- Joining effort and expertise between XENON, LZ and DARWIN
- See **xlzd.org**



XLZD meeting at Karlsruhe

WIMP Dark Matter

- Spin-independent
- Spin-dependent
- Sub-GeV
- Inelastic

Extended Dark Matter

- Dark photons
- Axion-like particles
- Planck mass

Sun

- pp neutrinos
- Solar metallicity
- ^7Be , ^8B , hep

Neutrino Nature

- Neutrinoless double beta decay
- Double electron capture
- Magnetic moment

Supernova

- Early alert
- Supernova neutrinos
- Multi-messenger astrophysics

LZ **DARWIN**



J. Aalbers et al
 XLZD white paper
<https://arxiv.org/abs/2203.02309>

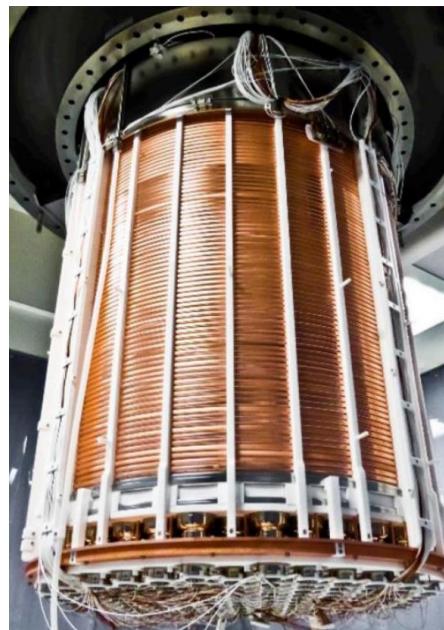
Theresa Fruth
 LZ status Wednesday 11:10

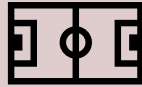



Adam Brown
 DARWIN status Friday 11:10



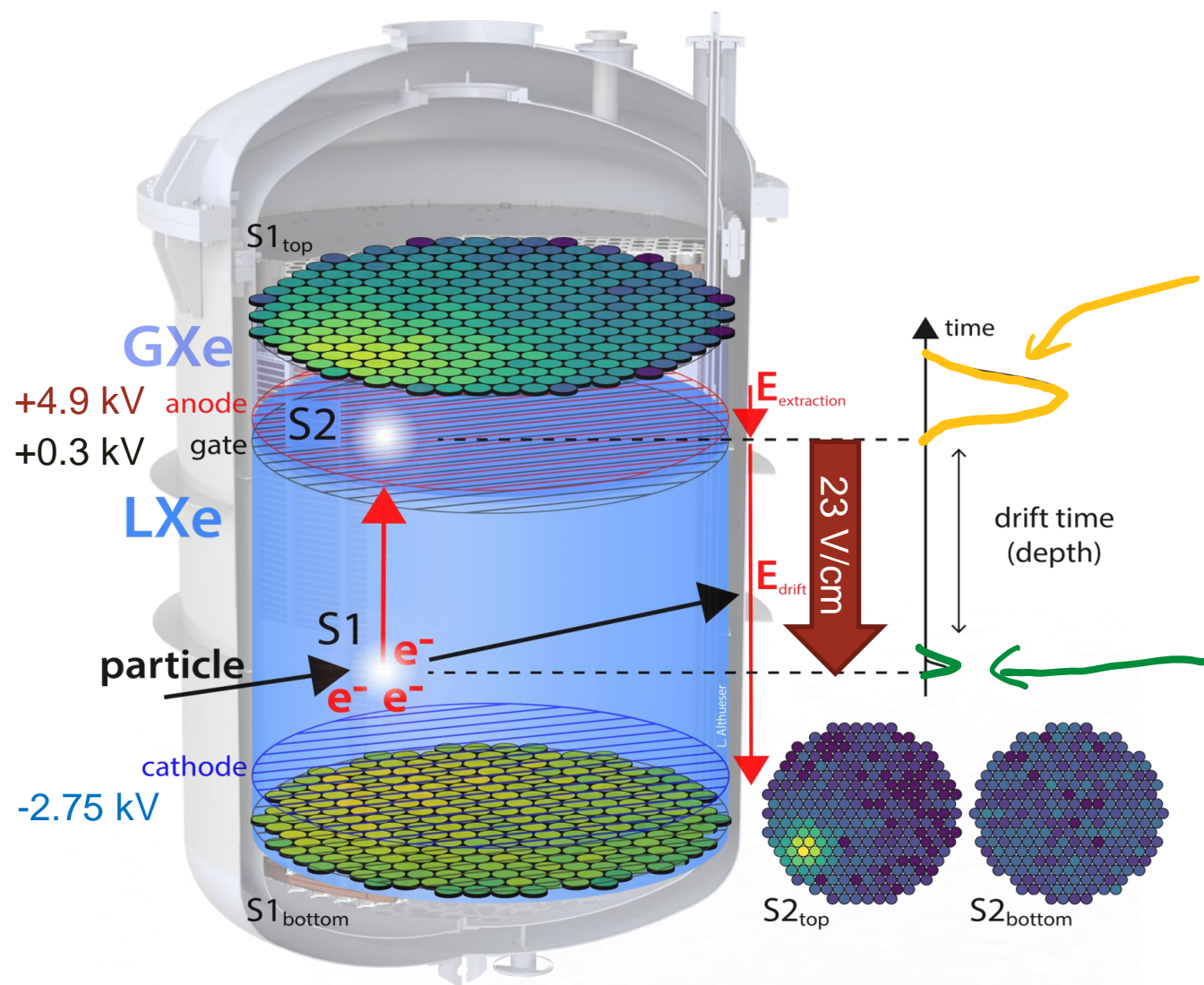
Back-up:

The XENON evolution:



XENON10	XENON100	XENON1T	XENONnT
2005-2007	2008-2016	2012-2019	2020-2026
14 kg Xe target	62 kg Xe target	2 t Xe target	~6 t Xe target, 8.6 t total mass
~ 10^{-43} cm ² 	~ 10^{-45} cm ² 	$4 \cdot 10^{-47}$ cm ² 	$1.4 \cdot 10^{-48}$ cm ² (projected for 20 t·y exposure)  (Flea)
~2M background ER / (keV· t·y)	1800 background ER / (keV· t·y)	82 background ER / (keV· t·y)	16.1 background ER / (keV· t·y)

Liquid noble gas time projection chamber:



Scintillation #2 signal (S2)

- Electrons from ionized Xe^+ drift upwards between anode and gate
- Extraction from LXe into GXe by higher field
- Electroluminescence yields S2 which is prop. to number of e^-
- Signals $O(100 \text{ pe} - 100000 \text{ pe})$

Scintillation #1 signal (S1)

- Excited Xe atoms form excimers Xe_2^*
- Excimers deexcite via emission of VUV-photons (178 nm)
- Signals $O(3 \text{ pe} - 1000 \text{ pe})$

Liquid noble gas time projection chamber:



Electronic Recoils:

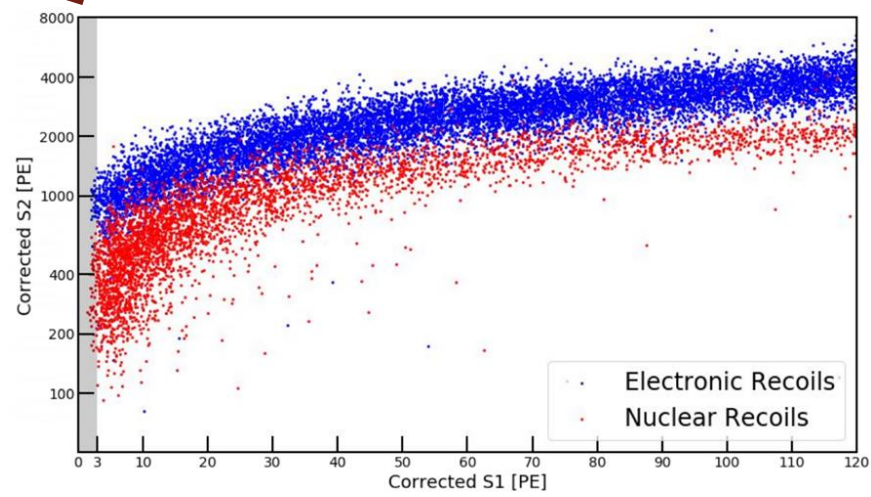


Nuclear Recoils:



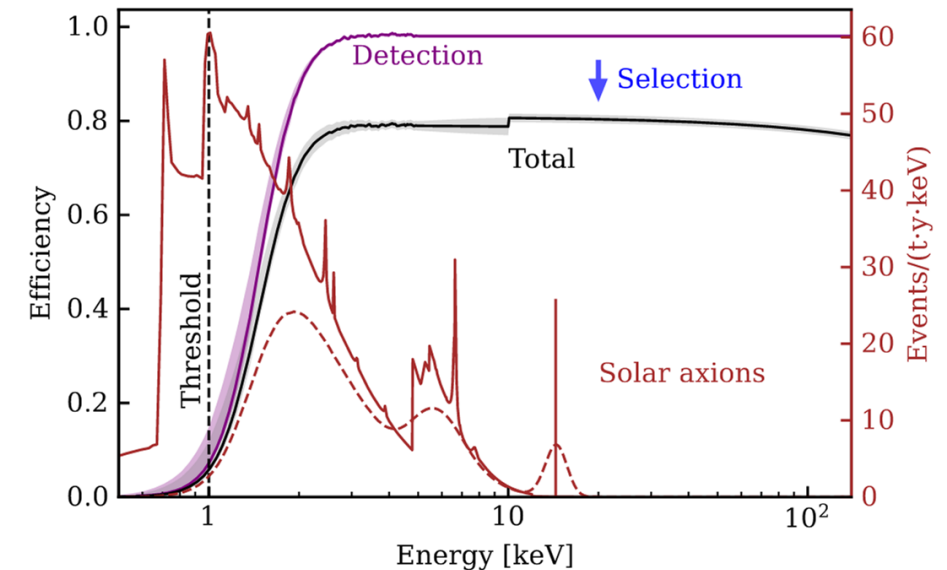
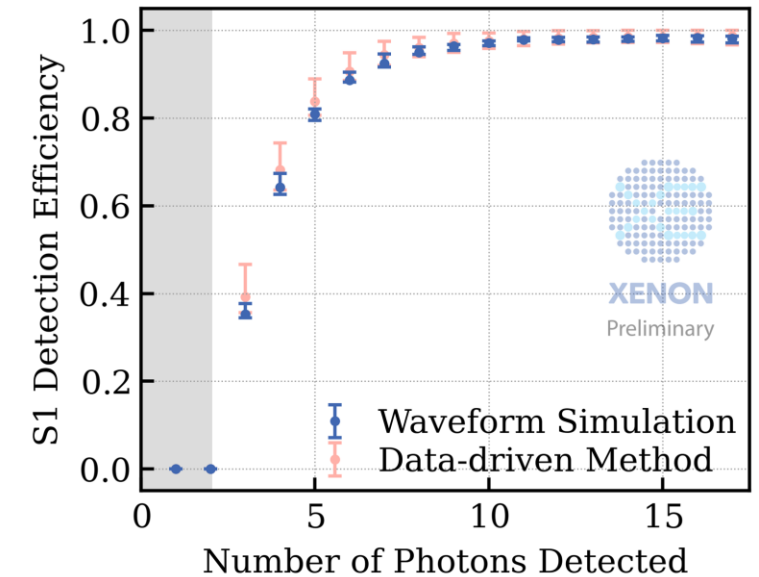
Ratio: S_2/S_1 (ER) > S_2/S_1 (NR)

recombination leads to S2 quenching for NR



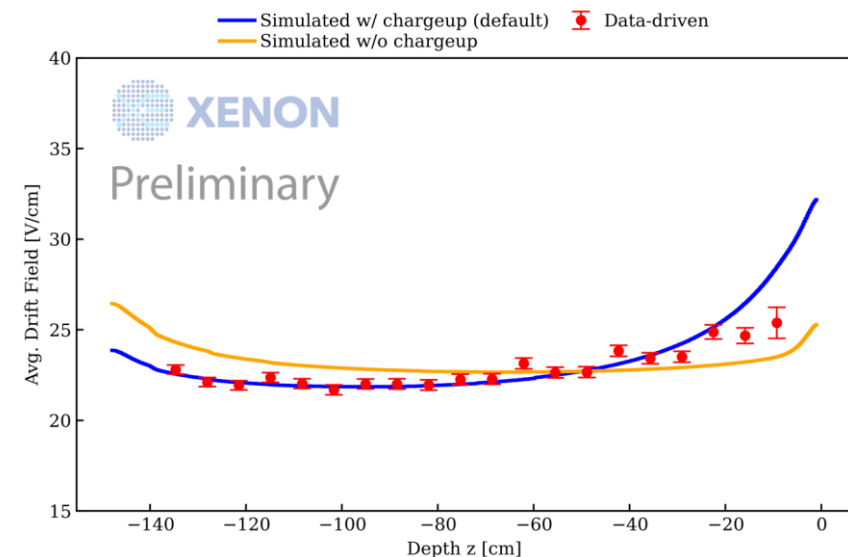
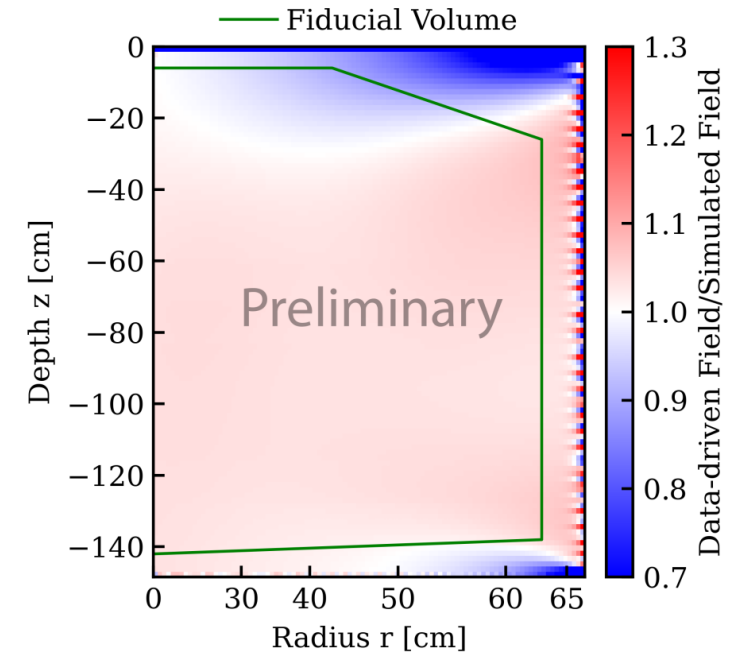
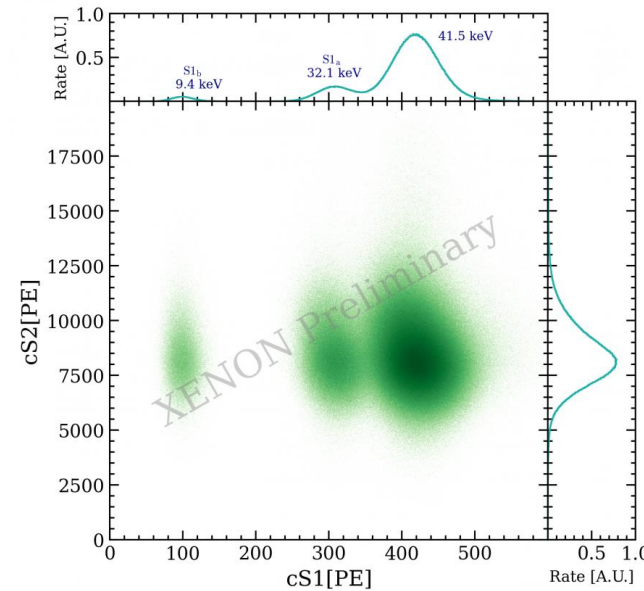
Calibration of detector threshold and acceptance:

- Detector threshold estimate using a data and a simulation driven method
 - Threshold driven by a 3-fold PMT coincidence requirement for S1 signals
 - Simulation of full waveforms which are analyzed with the same processing framework
 - Data-driven uses higher energetic peaks to sample toy peaks.
- Average data-quality cut acceptance ~86 %



Detector uniformity and electric field validation:

- ^{83m}Kr , decays slow enough to be diffused uniformly over the LXe volume.
 - Used to calibrate position dependent light collection efficiencies for S1/S2
 - To calibrate for field distortions in the position reconstruction
 - Validate COMSOL field simulation by comparing its 32.1 keV and 9.4 keV signal



Detector stability:

- Light and charge yield monitoring using background source and bi-weekly $^{83\text{m}}\text{Kr}$ calibrations
 - Light yield stability $\sim 1\%$
 - Charge yield stability $\sim 1.9\%$

