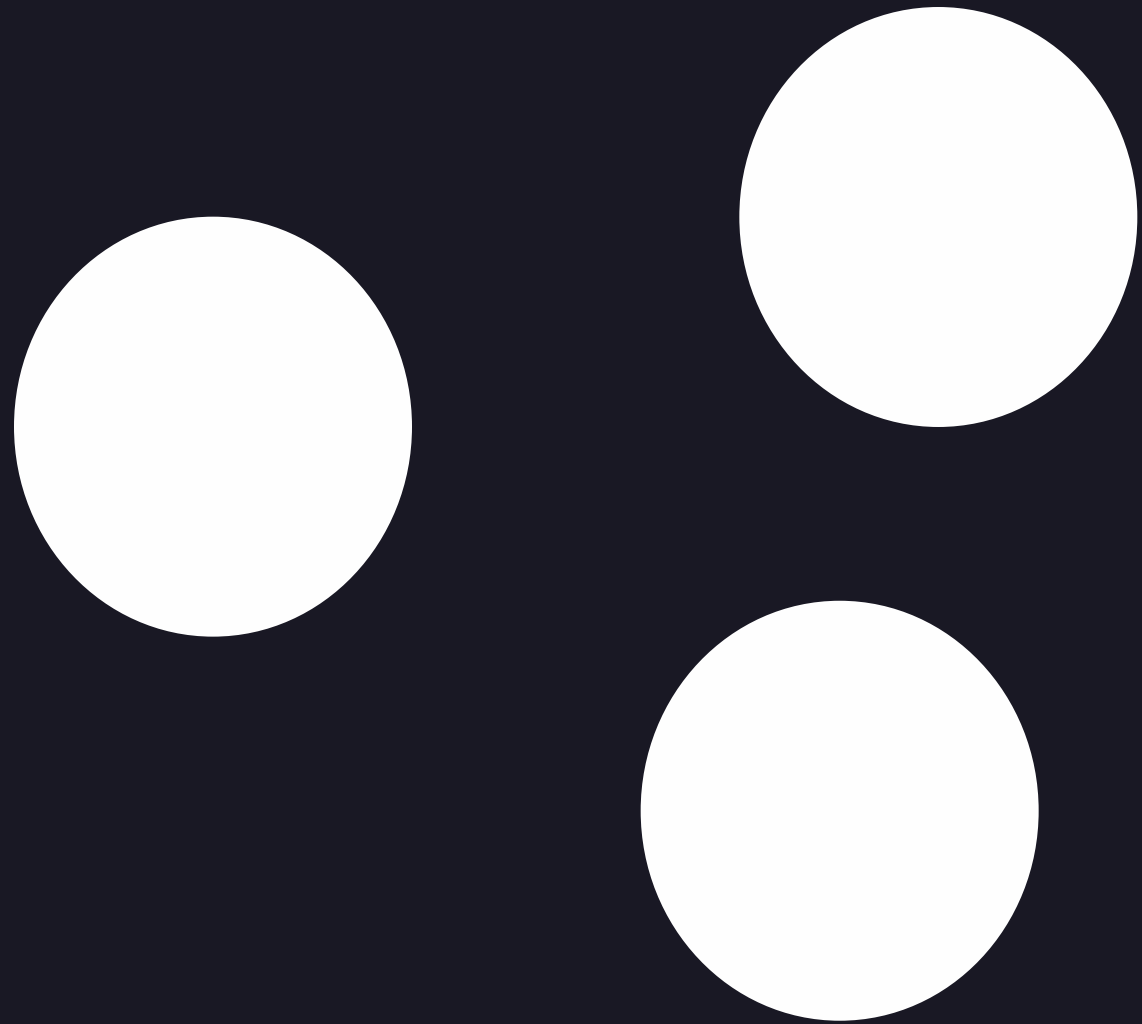


A self-consistent wave description of axion minicluster and their survival in the galaxy

V.Dandoy, T.Schwetz, E.Todarello

ArXiv: [2206.04619]

DM in the galaxy



- In some scenarios, axions are bound in axion miniclusters



Is this still true today?

How to know if they survived?

1. How to characterize the minicluster?



Axions are described by a classical field



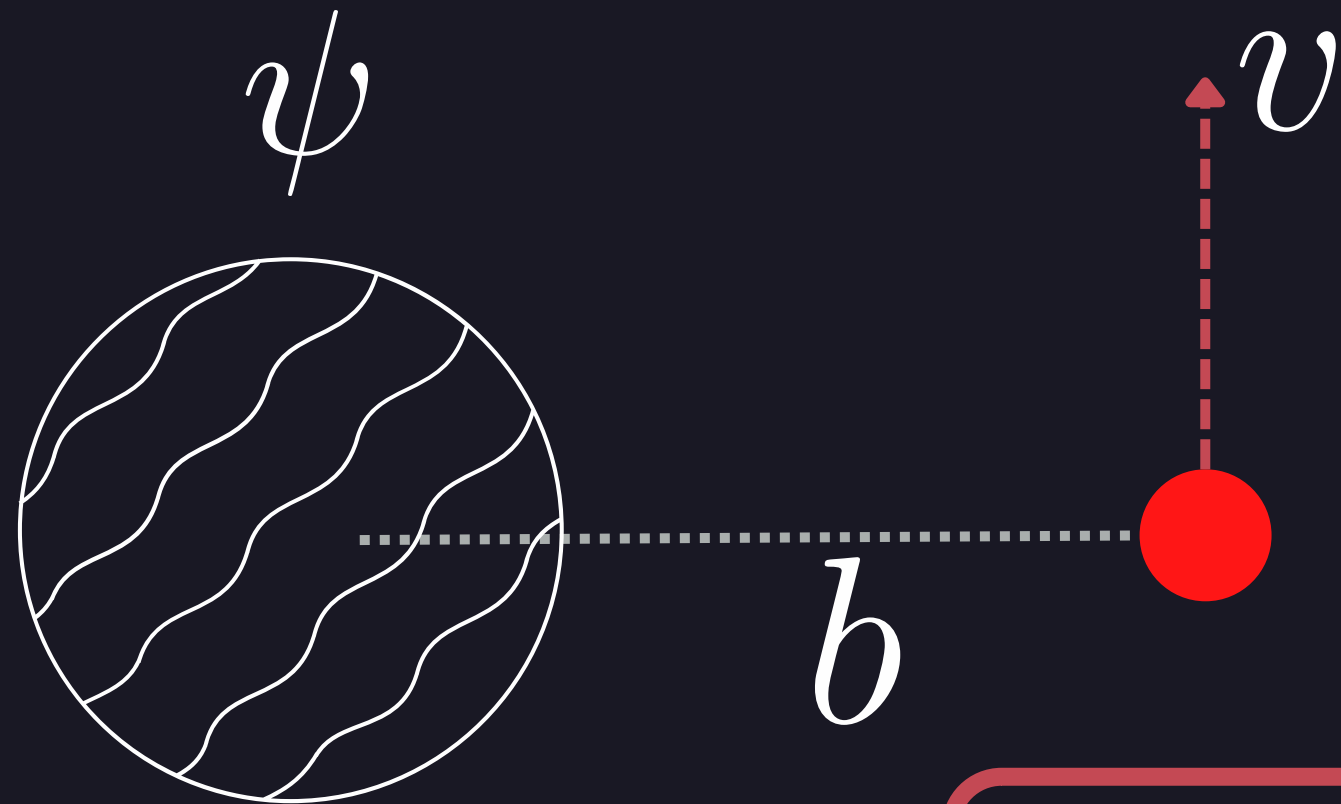
Wave function for a given density profile?

$$i\partial_t\psi = \left(-\frac{\nabla^2}{2m_a} + m_a\phi(r) \right) \psi$$

$$\nabla^2\phi = 4\pi Gm_a|\psi|^2$$

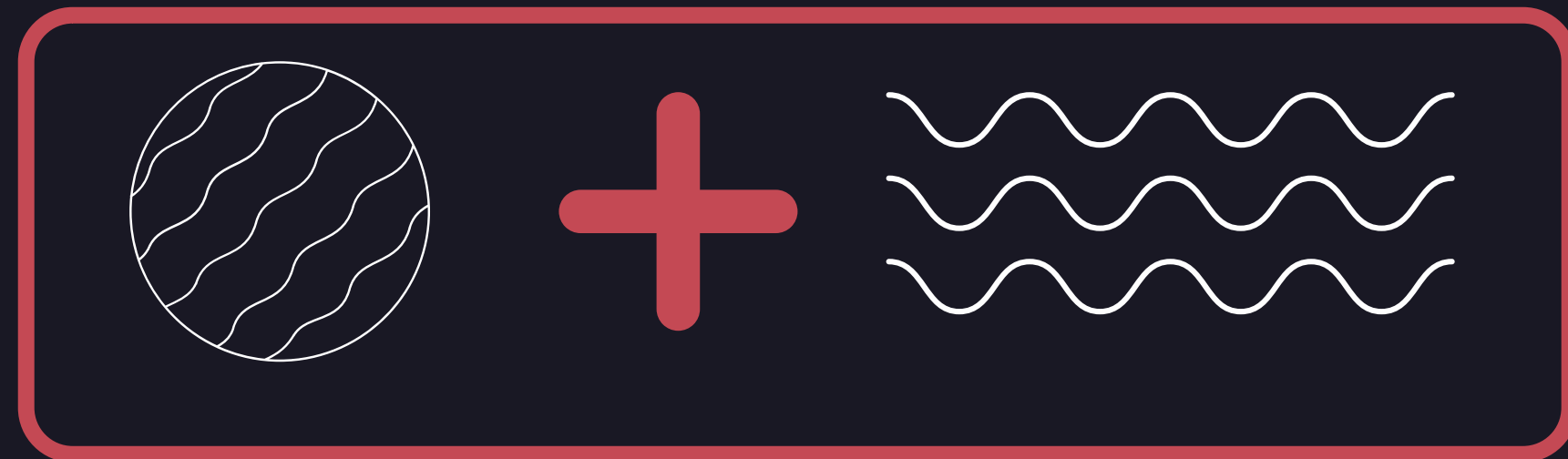
How to know if they survived?

2. How does it interact with a star?



The star creates a time dependent perturbation

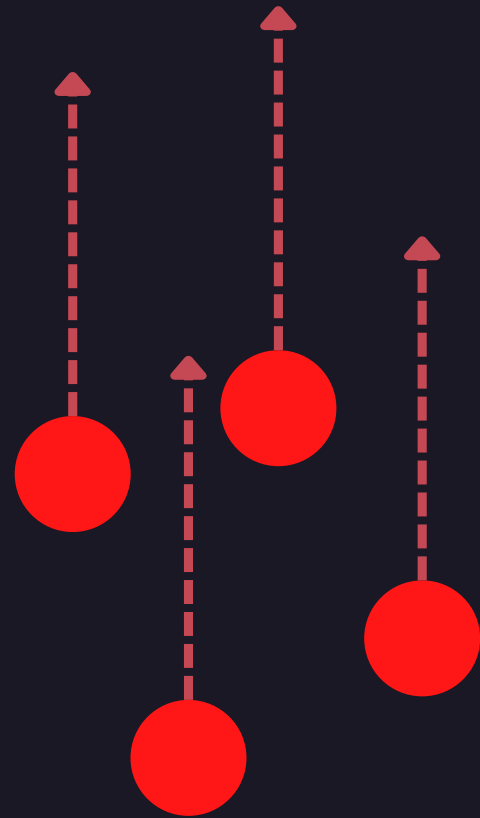
$$H_1(t) = -\frac{GM_* m_a r^2}{(b^2 + v^2 t^2)^{3/2}} P_2(\cos \gamma(t))$$



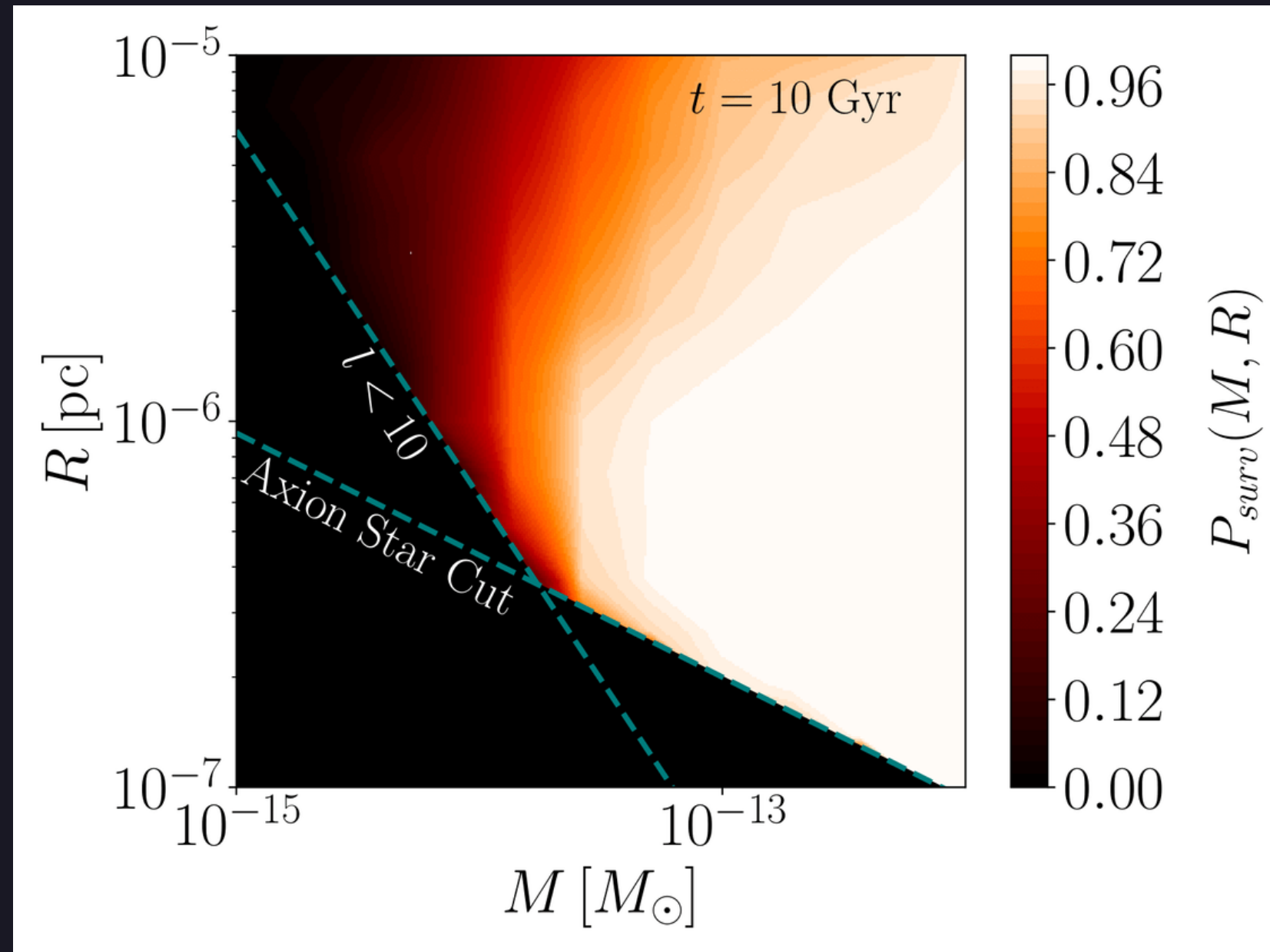
Final state

How to know if they survived?

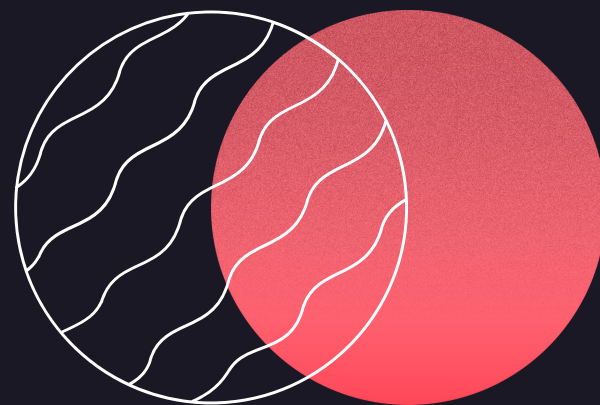
3. Simulate their evolution in the galaxy



Simulate all their lifetime interactions and extract the survival



Thank you!



This project has received funding /support from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska -Curie grant agreement No 860881-HIDDeN