

The VMB@CERN experiment

Giuseppe Messineo

INFN Ferrara

messineo@fe.infn.it

Poster #34

on behalf of the VMB@CERN Collaboration



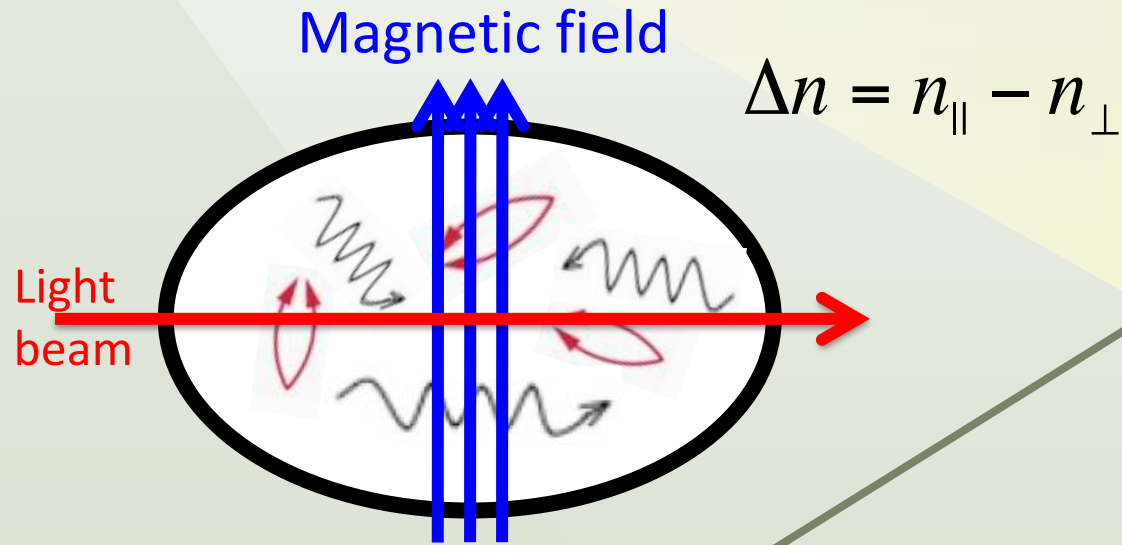
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OPTICAL PROPERTIES OF QUANTUM VACUUM

Vacuum is structured and has properties that can be studied experimentally.



$$\Delta \tilde{n} = \Delta n_B + i \Delta \kappa_B$$

BIREFRINGENCE DICHOISM

QED + ALPs

$$\Delta n_{QED} = 3A_e B^2 = 4 \times 10^{-24} \quad @ 1 \text{ T magnetic field}$$

VACUUM MAGNETIC BIREFRINGENCE

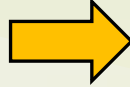
ALPs

Model independent search for 'light' DM

Plot showing constraints on Axion Mass (eV) vs $|g_{A\gamma\gamma}|$ (GeV⁻¹). Key features include: SN1987A gammas, Chandra, Fermi, Haloscopes (ADMX and Others), KSVZ, DFSZ, LSW (OSQAR and Others), PVLAS, Haloscopes (CACT), and Telescopes.

Experimental method:

- Perturb with an external B field
- Probe with a (polarised) light beam
- Detect changes in the polarisation state



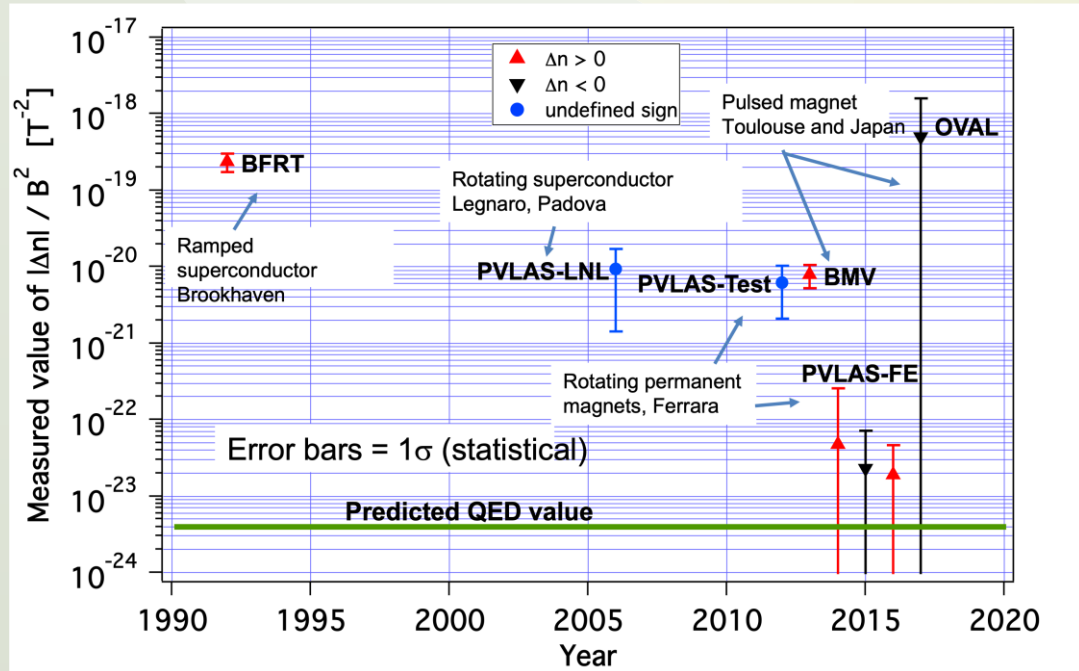
Key ingredients:

- High magnetic field
- Long optical path
- High sensitivity polarimeter

Effect $\propto B^2$

Optical cavity

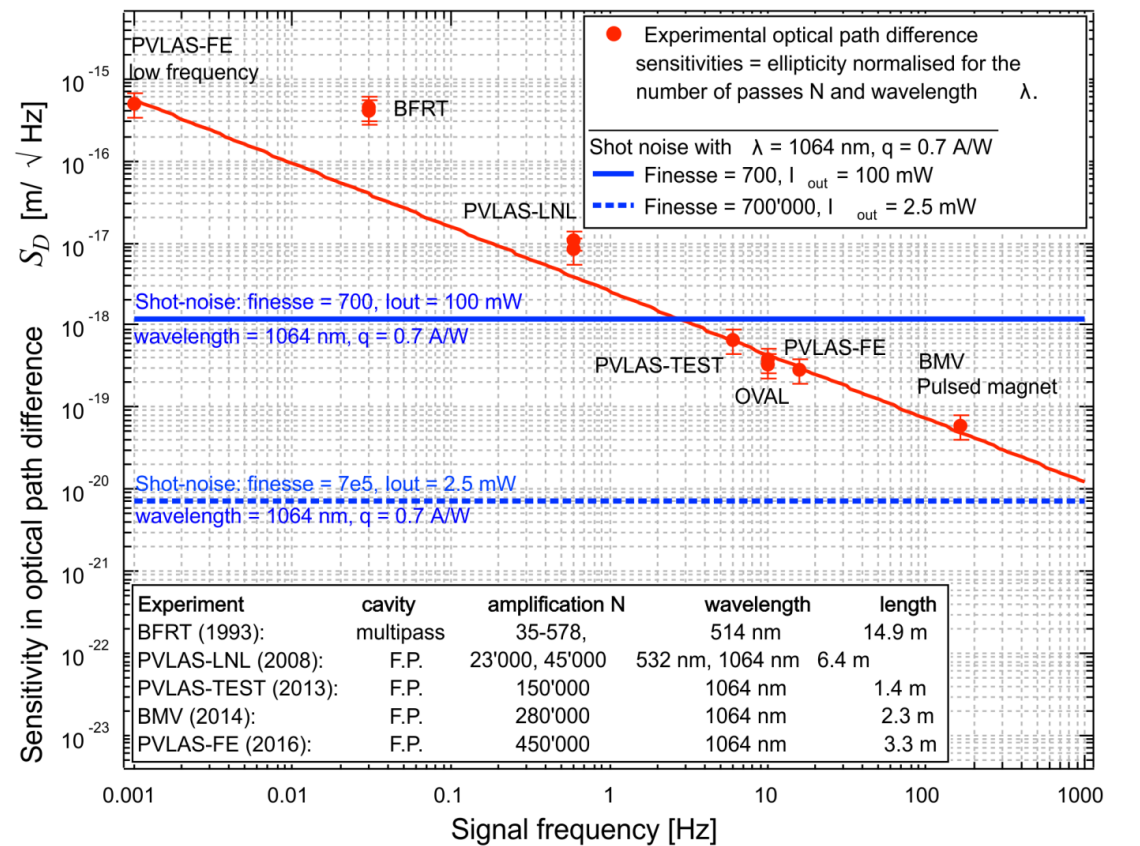
Modulation of the signal



Need to increase the signal!

LHC dipole magnet $B^2 L \approx 1200 \text{ T}^2 \text{ m}$

PVLAS permanent magnets $B^2 L \approx 10 \text{ T}^2 \text{ m}$

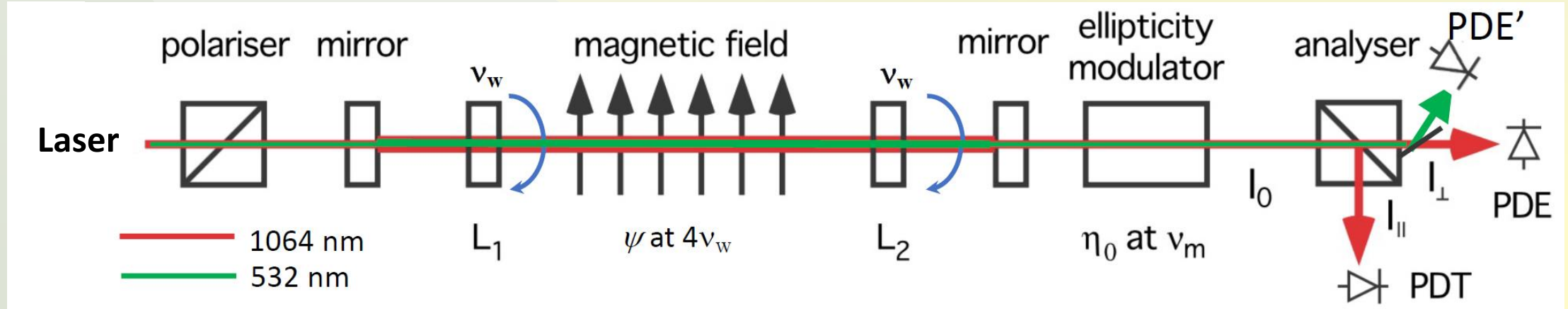


VMB@CERN INITIATIVE

Magnetic field in superconducting magnets cannot be modulated (fast enough).

Rotate polarization instead of magnet!

Lol submitted to CERN: CERN-SPSC-2018-036/SPSC-I-249



We use two co-rotating half waveplates inside the optical cavity:

- Polarisation rotates inside the magnetic field but is fixed on the mirrors (no mirror birefringence signal).
- Maximum finesse $\approx 800 - 3000$ (depending on the losses of the waveplates).
- Auxiliary green laser beam @ 532 nm allows real-time control of systematics of the individual waveplates.