



Updated Results from HAYSTAC's Quantum-Enhanced Search for Dark Matter Axions

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Axions as Dark Matter

- Solve CP Problem + Dark Matter
 - Axion mass/coupling is unknown
- Post inflation models
 - $m_a > 10 \ \mu eV [1,2]$
- HAYSTAC target
 - 16-20 µeV



Review of Particle Physics 2020, A. Ringwald, L.J. Rosenberg and G. Rybka

[1] E. Berkowitz, M. I. Bucho, and E. Rinaldi. Phys. Rev. D, 92 034507, 2015
[2] S. Borsanyi, Z. Fodor, J. Guenther, et al. Nature, 539 69, 2016

Searching with Haloscopes

- Magnetic Filed (B)
- High Q Resonant Cavity (Q)
- Large Volume (V)
- Tunable Frequency (ν)
- Low Noise Amplifiers (N)
- Cryogenic (N)
- Figure of Merit (Scan Rate):

$$\frac{d\nu}{dt} \propto \frac{\eta Q B^4 V^2 C^2}{N^2}, \frac{d\nu}{dt} \propto \nu^{-\frac{14}{3}}$$



HAYSTAC Detector

- Located at Yale's Wright Lab
- Copper Microwave Cavity
 - Q: ~45k
 - V: 1.5L
 - *v_c*: 4-6GHz
- 8T Superconducting Solenoid
- Dilution Fridge 61mK
- <u>Josephson Parametric Amplifier</u>
 <u>(JPA)</u>



Quantum Limit for Haloscopes

Cavity Hamiltonian:

Vacuum Fluctuations:

Linear Amplifier:

Total SQL:

 $\widehat{H} = \frac{h\nu_c}{2}(\widehat{X}^2 + \widehat{Y}^2) \qquad \left[\widehat{X}, \widehat{Y}\right] = \frac{i}{2}$ $N_v \geq \frac{1}{2}hv_c$ $N_A \ge \frac{1}{2}h\nu_c$

 $N_{total} \ge h\nu_c$

C. M. Caves. Phys. Rev. D, 26 1817-1839, (1982)

H. A. Haus and J. A. Mullen. Phys. Rev., 128 2407-2413, Dec 1962.

AMP

Benefit of JPA

- JPAs achieve near SQL when Phase-Insensitive
- JPAs are Phase-Sensitive Amplifiers
- Each Phase alone is not limited by SQL
- Can produce "Squeezed" States
 - Dump all uncertainty/noise into a single quadrature





Squeezed State Receiver



Bandwidth Enhancement



Bandwidth Enhancement



Bandwidth Enhancement



First Demonstration

- Demonstration by Colorado Group
- SSR speed up for "Fake" Axion signal





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First Results with Squeezing

- Enhancement from Squeezing
 - ~4 dB of Squeezing
 - Speed up of ~2x

- Axion Exclusion
 - Scanned ~73MHz of parameter space
 - Achieved sensitivity to $g_{\gamma} = 1.38 g_{\gamma}^{KSVZ}$



Additional Data Taking

	Days [#]	Spectra [#]	Frequency [MHz]
Phase-IIa	105	861	73
Phase-IIa (rescans)	53	508	"
Phase-IIb	51	791	64
Phase-IIb (rescans)	48	799	"

- Recently completed 2nd data taking campaign
 - Analysis still preliminary
- Main Changes/Improvements
 - New JPAs \rightarrow Higher Frequency
 - Modified DAQ \rightarrow Improved Livetime
 - Realigned Tuning Rod \rightarrow Less Mode Drift



Extending Beyond 4.2GHz

- Previous search limited by JPA Range
 - Max Frequency ~4.2GHz
- New JPAs designed to extend to 4.6-4.7GHz



Improved Livetime

• Ideally always recording cavity field • 1hr of data @ 10 MS/s



Tuning

~12%

- ~100GB per tuning
- >100TB for Phase IIa



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DAQ

Deadtime

~40%

Improving Livetime

- DAQ Deadtime
 - Parallelization of FFT
 - Optimization of Data Transfer
- Tuning Stabilization
 - Reinstalled cavity has less mode drift after tuning
 - Better rod alignment
- Phase IIb achieved 78% average livetime
 - 82% after tuning improvement

	Fractional Time [%]		
	Phase IIa	Phase IIb	
DAQ	40	10	
Tuning	12	8	
Livetime	48	82	



Projected Sensitivity



Conclusion

- HAYSTAC is continuing to search for axions with $m_a > 10 \ \mu eV$
- Second Data taking Phase with the SSR
 - Searched another 62 MHz of parameter space
 - New JPAs \rightarrow Extend search to 4.5GHz
 - Improvement livetime to $\sim 78\% \rightarrow \sim 1.6x$ scan rate enhancement



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Thanks

















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Backups

Limiting Factor

- SSR is limited by loss between two JPAs
 - Loss is dominated by triple junction circulator
 - 4dB of Squeezing is roughly best we can do currently



Installing SSR in HAYSTAC

- Phase-I (2012-2018)
 - Single Phase-Insensitive JPA
 - Scanned 5.6-5.8GHz
 - Phys. Rev. D 97 (2018) 9, 092001
- SSR installed into HAYSTAC
 in 2018







Operating JPAs Near Magnet

- JPAs are extremely sensitive to stray B-Fields
 - << 1 flux quantum (~2G)



Tuning with SSR

- Five parameter optimization
- JPAs tuned to match Cavity Resonance
 - I_{sz} : Squeezer Flux Bias
 - I_{AMP}: Amplifier Flux Bias
- Amplifiers share same Pump Source
 - P_P: Amplifier Gain
 - A: Squeezer Gain
 - θ : Phase difference



What's Next...

- Further enhancement with Squeezing
 - <u>State Swapping (CEASEFIRE):</u> Squeezing before loss (K. Wurtz et al, PRX Quantum 2 (2021) 4, 040350)
- Current setup tops out around ~6GHz
 - Limited by Cavity Range
 - <u>Multi-Rod Cavity:</u> Same cavity but 5-7GHz (*M. Simanovskaia et al, Rev. Sci.Instrum.* 92 (2021) 3, 033305)





Single Photon Detection

- Ultimate goal of "Squeezing" is Single Photon Detection
 - Lose Spectral Information
 - Only shot noise limited
 - Payoff >~10GHz (S. K. Lamoreaux et al., Phys. Rev. D 88, 035020 (2013))



Rydberg Atoms

- Rydberg Atoms have transition frequencies >10GHz
 - First proposed by CARRACK (*M. Tada, Nucl.Phys.B Proc.Suppl.* 72 (1999) 164-168)



Rydberg Atoms at Yale (RAY)

- ³⁹K Rydberg Atoms
 - Demonstrated production of Rydberg states n=60-90
- Transition Frequencies give access to >10GHz



Y. Zhu et al, Phys. Rev. A 105, 042808

Quantum Enhancement



Haloscope At Yale Sensitive To Axion CDM







<u>JPA</u>

JPA Shield



System Diagram



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