Status of the OUAX experiment

#### TWPA tests 00000000





LNF

#### Lab @INFN-LNL

C. Braggio (this presentation), G. Carugno, N. Crescini, R. Di Vora, A. Ortolan, G. Ruoso, A. Lombardi, R. Pengo, L. Taffarello



 $100 \,\mu\text{W}$  at  $100 \,\text{mK}$ 

#### Lab @INFN-LNF

C. Gatti, D. Alesini, D. Babusci, A. D'Elia, D. Di Gioacchino, C.Ligi, G. Maccarrone, A. Rettaroli, S. Tocci



@INFN-Salerno U. Gambardella, G. Iannone, CD. D'Agostino @INFN-Trento P. Falferi, R. Mezzena



- $\rightarrow$  DM axion search (axion-photon coupling) by scanning (8.5 11) GHz frequency range at KSVZ sensitivity
- → LNL and LNF INFN laboratories will work in synergy, operating in different mass ranges and using different low noise amplifiers and single microwave photon detectors.
- $\rightarrow$  EU and US collaborations for the integration of:
  - 1. high-Q cavities (SQMS, Superconducting Quantum Materials and Systems Center, led by Fermilab)
  - 2. state-of-the-art itinerant microwave photon counters (Quantronics group, Saclay)
  - 3. traveling wave PA (N. Roch group, Néel Institute in Grenoble)



SUPERCONDUCTING QUANTUM MATERIALS & SYSTEMS CENTER



Quantronics Group Research Group in Quantum Electronics, CEA-Saclay, France

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LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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HALOSCOPE - resonant search for axion DM in the Galactic halo

- original proposal by P. Sikivie (1983)

- search for axions as cold dark matter constituent: SHM from  $\Lambda_{\text{CDM}}$ , local DM density  $\rho$  $\rightarrow$  signal is a **line** with 10<sup>-6</sup> relative width in the energy( $\rightarrow$  frequency) spectrum

 $\rightarrow$  + sharp (10<sup>-11</sup>) components due to non-thermalized

- an axion may interact with a strong  $\vec{B}$  field to produce a photon of a specific frequency ( $\rightarrow m_a$ )



LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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## QUAX COLLABORATION ROADMAP (2021-2025)



	LNF	LNL
Magnetic field	9 T	14 T
Magnet length	40  cm	$50 \mathrm{~cm}$
Magnet inner diameter	$9~\mathrm{cm}$	$12 \mathrm{~cm}$
Frequency range	$8.5$ - $10~\mathrm{GHz}$	9.5 - 11 GHz
Cavity type	Hybrid SC	Dielectric
Scanning type	Inserted rod	Mobile cylinder
Number of cavities	7	1
Cavity length	0.3 m	0.4 m
Cavity diameter	25.5  mm	58  mm
Cavity mode	TM010	pseudoTM030
Single volume	$1.5 \cdot 10^{-4} \text{ m}^3$	$1.5 \cdot 10^{-4} \text{ m}^3$
Total volume	$7{\otimes}0.15$ liters	0.15 liters
$Q_0$	300 000	1000000
Single scan bandwidth	630 kHz	30 kHz
Axion power	$7\otimes 1.2\cdot 10^{-23}~{\rm W}$	$0.99 \cdot 10^{-22} \ {\rm W}$
Preamplifier	TWJPA/INRIM	DJJAA/Grenoble
Operating temperature	30 mK	30 mK

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LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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#### OUTLINE

- 1. data analysis results of July 2021 run with high-Q dielectric resonator 8 T-field, HEMT readout, 1 MHz-tuning at 10.353 GHz ( $m_a = 42.8 \,\mu\text{eV}$ )
- 2. **TWPA**-based amplification chain characterization generic input test cavity
- 3. July 2022 run with the dielectric resonator ( $\nu_{030} = 10.353$  GHz) 8 T-field, TWPA readout at 10 GHz
- 4. Single Microwave Photon Detectors for "itinerant" photons:
  - preparation of a transmon based-SMPD haloscope readout experiment 3 T-field, 7 GHz NbTi cavity, 10 MHz-tuning
  - preliminary results obtained a underdamped Josephson junction (L. Kuzmin) coupled to a generic input test cavity











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#### DIELECTRIC CAVITY



LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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#### MEASUREMENT SCHEME

- two weeks data taking (June 2021)
- cavity excess-power searched in a small frequency band about 10.353 GHz ( $\sim 42.8 \mu$  eV axion mass)
- three different configurations:
  - 1.  $\beta \simeq 1$ , i.e.  $Q_L > Q_a$ 2.  $\beta \simeq 6$ , i.e.  $Q_L \simeq Q_a$ 3.  $\beta \ge 14$ , i.e.  $O_I < O_a$

• 
$$T_{\rm sys} = 17.3 \pm 1 \, {\rm K}$$
  $T_{\rm sys} =$ 



• a posteriori measured  $T_A = 10 - 12$  K instead of the nominal noise temperature of 4.5 K, crvo HEMT sent for repair



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#### RESULTS

- 1.  $\beta \simeq 1$ , i.e.  $Q_L > Q_a$ 2.  $\beta \simeq 6$ , i.e.  $Q_L \simeq Q_a$
- 3.  $\beta \ge 14$ , i.e.  $Q_L < Q_a$

issues regarding possible systematics for cases 1) and 2)  $\longrightarrow$  we focused on  $Q_L < Q_a$ , with  $Q_L \sim 3 \times 10^5$ 

- $\odot$  Haystac data analysis procedure
- even with a **very bad receiver**, to operate the high-Q cavity allowed for probing realistic QCD axion models, only marginally outside the benchmark QCD axion band
- axion mass not accessible to other running experiments





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- **3.** July 2022 run with the dielectric resonator ( $\nu_{030} = 10.353$  GHz) 8 T-field, TWPA readout at 10 GHz
- **4.** preparation of a single microwave photon counter (SMPD) experiment 3 T-field, 7 GHz NbTi cavity, 10 MHz-tuning





LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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## TRAVELING WAVE PARAMETRIC AMPLIFIERS

In haloscope search amplifiers with **quantum-limited noise performance** and  $\sim$  **GHz amplification bandwidth** are needed.

Standard resonant parametric amplification puts a constraint on the amplification bandwidth. This limitation can be overcome with TWPAs.





LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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Josephson metamaterial and Reversed-Kerr phase matching

- tunable, nonlinear unit element is the "snail"
- phase matching between pump, signal and idler fields is accomplished via **reversed kerr** phase matching



A. Renadive *et al*, Nat. Commun. 13, 1737 (2022) M. Esposito *et al* Appl. Phys. Lett. 119, 120501 (2021) M. Esposito *et al* Phys. Rev. Lett 128, 153603 (2022) **— broadband squeezing!** 



LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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## Measuring TWPA performance in a haloscope setup

 $He^3\text{-}He^4 \ ``wet'' \ dilution \ refrigerator \ (refurbished) \rightarrow recovery \ system + compressor \ at \ LNL$ 



1 mW cooling power at 120 mK  $T_{\it mc}=55$  mK 8T-magnet, charging at 0.07 mA/s ; a 14~T magnet is coming in 2023



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## Measuring TWPA performance in a haloscope setup

• overall detection **gain**, from *G*<sub>xy</sub>s

$$g_4 = \sqrt{\frac{G_{14}G_{34}}{G_{31}}}$$

• system **equivalent noise temperature**, referring noise power at L4 output to the input:

$$P_n^{xy} = g_4 k T_{sys} B + N_{SA} \qquad (xy) = \{14, 34\}$$





- reduced **gain ripples** compared to state-of-the-art TWPAs
- ⊙ in-situ tunability of amplification bandwidth over an unprecedented wide range

#### arXiv:2205.02053

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## Measuring TWPA performance in a haloscope setup



• figures of  $\operatorname{Merit}$  for the complete detection chain K

 $g_4(dB_{0}) = (76.4 \pm 0.01)$   $T_{SUVER} = (2.0 \times 10.06) \text{ K}$ 

• figures of merit for the **TWPA** 

 $G_{\rm TWPA}({\rm dB}) = 24$   $T_{\rm TWPA} \approx 1.8 \,{\rm K}$ 



LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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## SMPDs: MOTIVATION

- QUAX<sub>a−e</sub>, the ferrimagnetic haloscope Phys. Rev. Lett. 124, 171801 (2020)
- $\odot$  QUAX<sub>*a*- $\gamma$ </sub> at 10 GHz
- "Game changer at high frequency and low temperatures": a photon counter measures in the energy eigenbasis beyond SQL

$$- \text{ SNR}_{\text{exc}} = \frac{P_{a \to \gamma}}{kT_{\text{sys}}} \sqrt{\frac{t_m}{\Delta \nu_a}} \quad \text{ SNR}_{\text{SMPD}} = \frac{P_{a \to \gamma}}{h\nu} \sqrt{\frac{t_m}{\Gamma_{dc}}}$$
$$\frac{\text{SNR}_{\text{SMPD}}}{\text{SNR}_{\text{exc}}} > 1 \Longleftrightarrow \frac{\Gamma_{dc}}{\eta} < \frac{\nu_a}{10^6}$$



plot example at 10 GHz given on Tue by SungWoo YOUN

**quantum advantage** can be shown even with relatively high dark count rates  $\Gamma_{dc}$ 

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## SMPDs for itinerant photons

A Single Photon Microwave Counter (SMPD) architecture is significantly different whether it is meant for **cavity photons** or **itinerant (traveling) photons**.

We are interested in the itinerant version due to the intense magnetic fields involved in axion search.



- $-\,$  detection of individual microwave photons is a challenging task because of their  $low~energy\sim 10^{-5}\,\rm eV$
- a solution: use "artificial atoms" introduced in circuit QED, their transition frequencies lie in the ~GHz range
- or: rely on a single current-biased Josephson junction (L. Kuzmin)

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#### ARTIFICIAL ATOMS: the TRANSMON QUBIT





 $E_{01} = E_1 - E_0 = \hbar \omega_{01} \neq E_{02} = E_2 - E_1 = \hbar \omega_{21}$  $\rightarrow$  good **two-level atom** approximation

control internal state by shining laser tuned at the transition frequency:

$$H = -\vec{d} \cdot \vec{E}(t)$$
, with  $E(t) = E_0 \cos \omega_{01} t$ 

toolkit: capacitor, inductor, wire (all SC) 
$$\begin{split} \omega_{01} &= 1/\sqrt{LC} \sim 10\,\mathrm{GHz} \sim 0.5\,\mathrm{K} \\ \rightarrow \mathrm{simple}\,\mathrm{LC}\,\mathrm{circuit}\,\mathrm{is}\,\mathrm{not}\,\mathrm{a}\,\mathrm{good}\,\mathrm{two-level}\,\mathrm{atom} \\ \mathrm{approximation} \end{split}$$

$$\begin{split} I_{J} &= I_{c} \sin \phi \qquad V = \frac{\phi_{0}}{2\pi} \frac{\partial \phi}{\partial t} \\ V &= \frac{\phi_{0}}{2\pi} \frac{1}{I_{c} \cos \phi} \frac{\partial I_{J}}{\partial t} = L_{J} \frac{\partial I_{J}}{\partial t} \\ L_{J} &= \frac{\phi_{0}}{2\pi} \frac{1}{I_{c} \cos \phi} \qquad \text{NL Josephson inductance} \end{split}$$

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## transmon-based SMPD

In the Quantronics group (CEA, Saclay) a transmon-based counter has been developed and used to make spin fluorescence measurements, paving the way to **single spin flip detection** with SMPDs.





 $\omega_a + \omega_p = \omega_q + \omega_b$ 

R. Lescanne *et al*, Phys. Rev. X 10, 021038 (2020) E. Albertinale *et al*, Nature 600, 434 (2021)



#### Quantronics Group

Research Group in Quantum Electronics, CEA-Saclay, France

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#### Quantronics Group Research Group in Quantum

Electronics, CEA-Saclay, France



- a three-step process repeated several times \_
- qubit reset (R) performed by turning on the pump pulse + a weak resonant coherent pulse to the waste port
- detection (D) step with the pump pulse on \_
- measurement (M) step probes the dispersive shift of the \_ buffer resonator to infer the gubit state

LNL and LNF haloscopes	July 2021 run	TWPA tests	SMPDs	LNF
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#### QUANTUM SENSING

"Quantum sensing" describes the use of a quantum system, quantum properties or quantum phenomena to perform a measurement of a physical quantity Rev. Mod. Phys. 89, 035002 (2017)

- 1. Use of a **quantum object** to measure a physical quantity (classical or quantum). The quantum object is characterized by quantized energy levels, i.e. electronic, magnetic or vibrational states of superconducting or spin qubits, neutral atoms, or trapped ions.
- 2. Use of **quantum coherence** (i.e., wave-like spatial or temporal superposition states) to measure a physical quantity
- 3. Use of **quantum entanglement** to improve the sensitivity or precision of a measurement, beyond what is possible classically.



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### BASIC PROTOCOL

quantum sensing experiments typically follow a generic sequence of processes known as:

- 1. sensor initialization into a known basis state
- 2. interaction with the signal
- 3. sensor readout
- 4. signal estimation



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## PILOT SMPD-HALOSCOPE EXPERIMENT

- copper cavity sputtered with NbTi magnetron sputtering in INFN-LNL
- $\odot~$  right cylinder resonator, TM\_{010} mode  $\nu_c \sim 7.3~{\rm GHz}$  to match the new generation SMPD bandwidth (7.280 7.380) GHz
- $\odot$  system of sapphire triplets to tune the cavity frequency  $\sim 10$  MHz tuning without impacting *Q*
- Attocube nanopositioner to change the sapphire rods position



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- $\odot$  system of sapphire triplets to tune the cavity frequency  $\sim 10$  MHz tuning without impacting *Q*
- Attocube nanopositioner to change the sapphire rods position
- developed and tested a **3 T magnet** (U. Gambardella, INFN Salerno)





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More cooling power coming soon...



Leiden Cryogenics commissioning in October 2022



"wet" delfridge from PTB (ongoing refurbishing)

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## Not that **practical** to use ... but that's it!



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## Current-biased Josephson Junction

working principle:

- voltage switching of an underdamped JJ
- ► phase diffusion regime



L.S. Kuzmin *et al* IEEE Trans Appl Supercond 28, 2400505 (2018); A. L. Pankratov *et al* npj Quantum Information 8:61 (2022)

"washboard potential"



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# QUAX Haloscope at LNF



Leiden CF-CS-110-1000 dilution refrigerator with 8 mK base temperature

Probe KSVZ axions in 1 GHz band at 9 GHz

- Multi cavity for fast scanning rate
- Wide band TWJPA quantum amplifier
- Superconducting cavities





9 T magnet from AMI



First run with single 8.5 GHz OFHC Cu cavity 1

