



Hunting for axions and axion-like particles with the nEDM and n2EDM experiments at PSI

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On behalf of the n2edm collaboration at PSI
ETH Zürich
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Goal

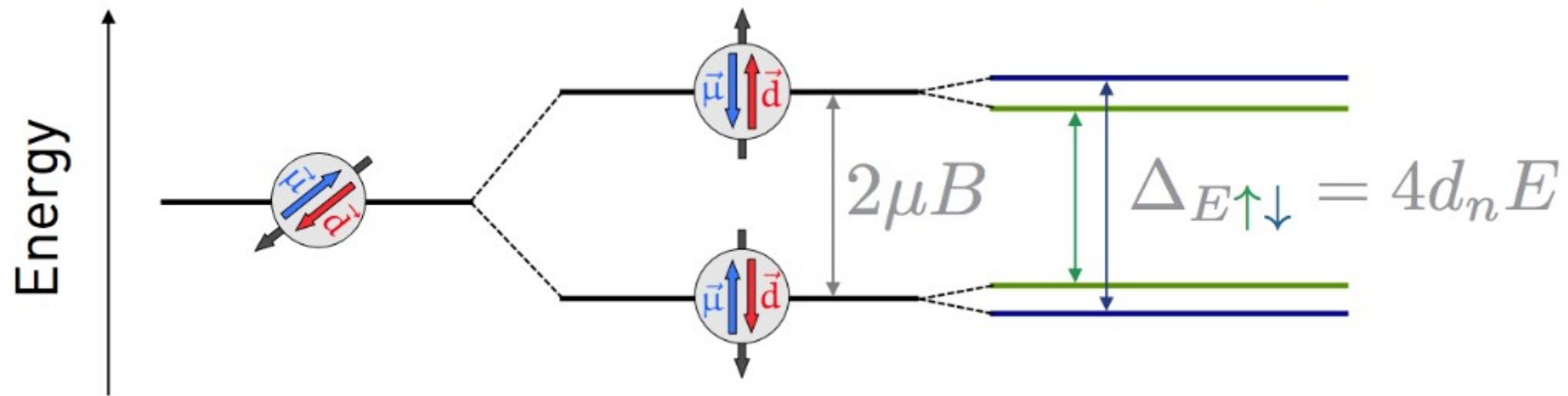
- Ultralight axions with $10^{-22} \text{ eV} < m < 10^{-17} \text{ eV}$
→ DM candidates

$$\mathcal{L}_{int} = \frac{C_g}{f_a} \frac{g_s^2}{32\pi^2} a G_{\mu\nu}^a \tilde{G}^{a\mu\nu} \quad d_n(t) \approx +2.4 \times 10^{-16} \frac{C_g a_0}{f_a} \cos(m_a t) \text{ ecm},$$

Oscillating ultralight dark matter axion field → oscillating neutron EDM

Current limit of $\theta < 10^{-10}$ from nEDM measurements

nEDM Experiment

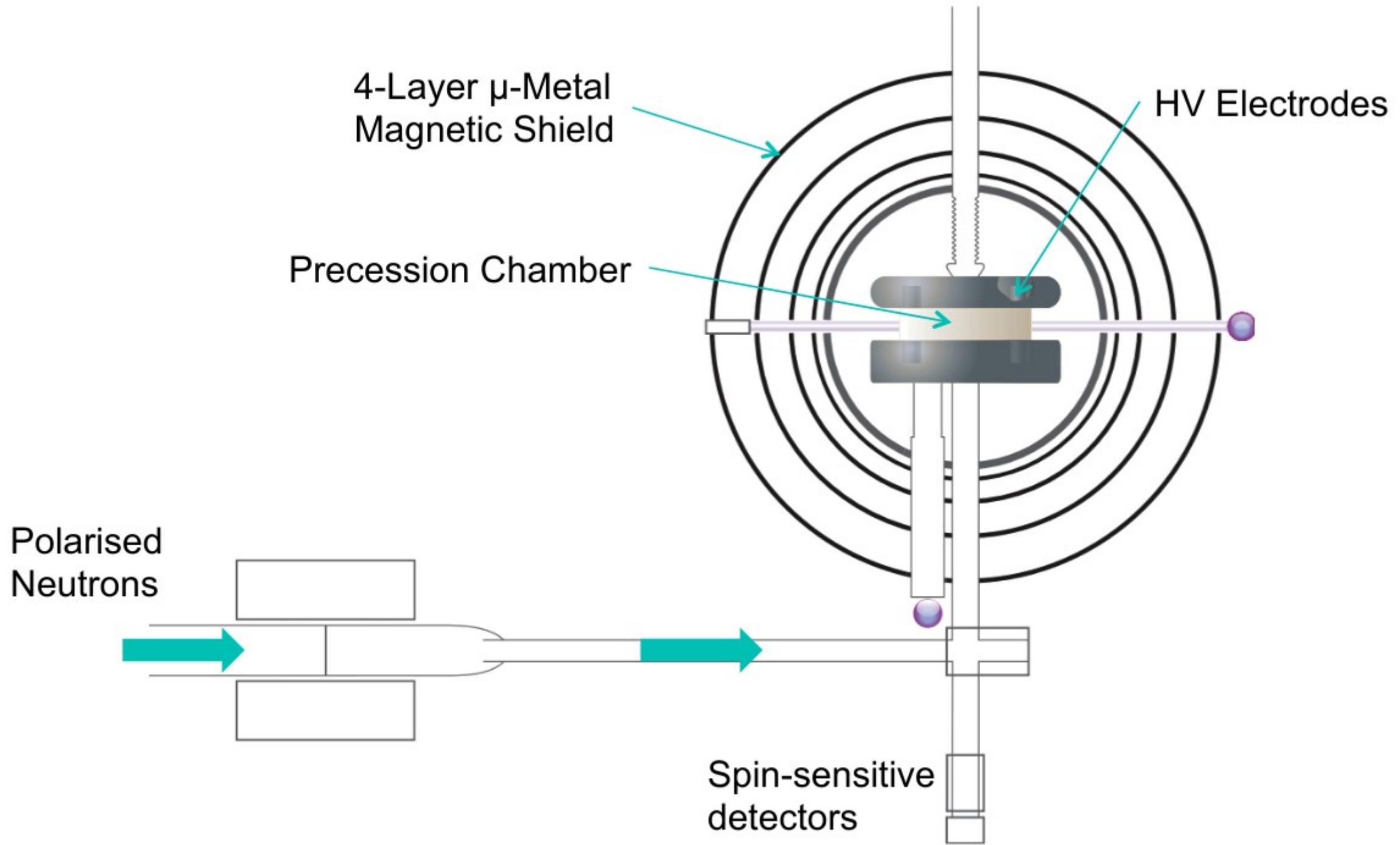


- Energy splitting between EB $\uparrow\uparrow$ and EB $\uparrow\downarrow$
- Measure neutron Larmor frequency
→ Ramsey technique

$$d_n = \frac{\pi\hbar}{2|E|} (f_{n,\uparrow\downarrow} - f_{n,\uparrow\uparrow})$$

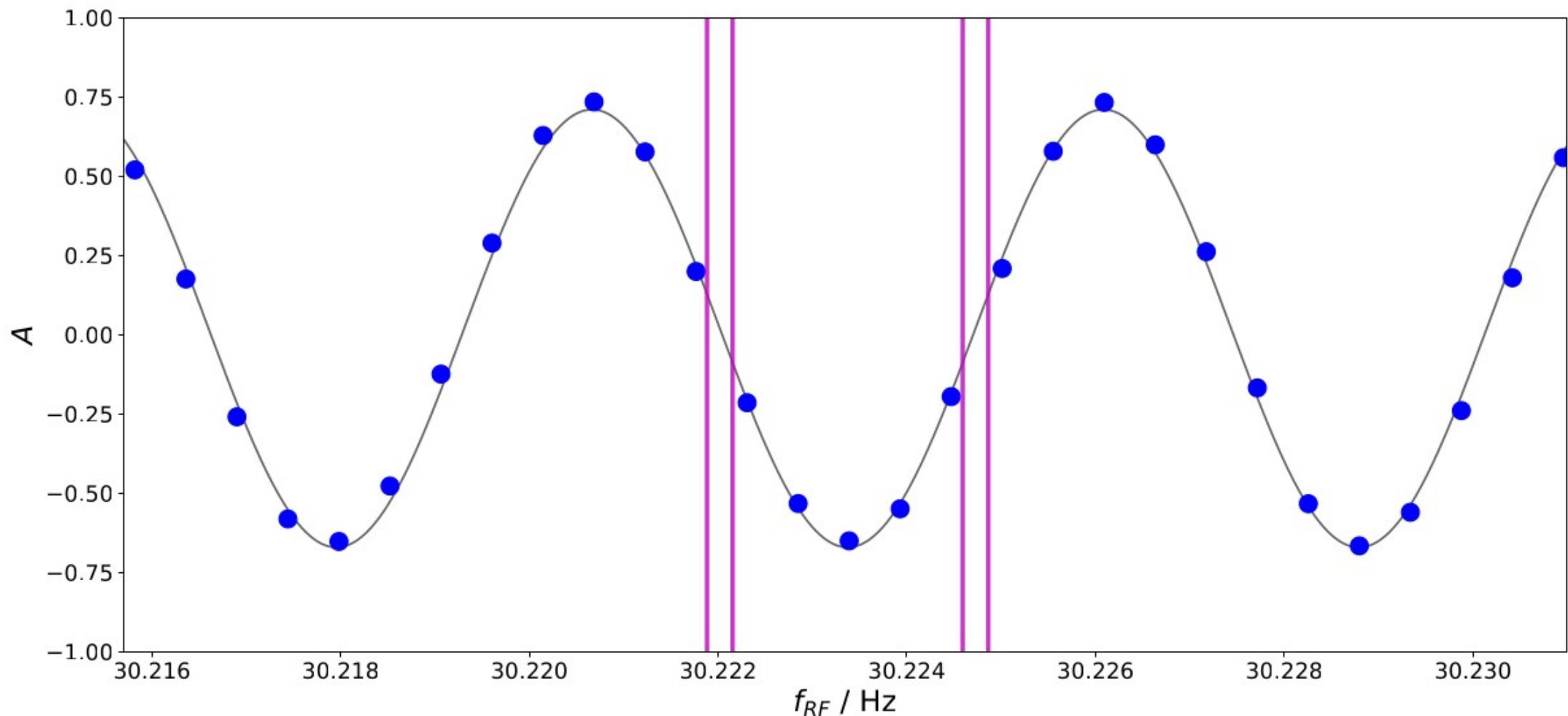
Ultracold neutrons (UCN)

- Free neutrons with low kinetic energy < 250-300 neV
→ gravity becomes important (100 neV → 1 m height)
- Undergo total reflection at any angle
→ Can be stored in relatively ordinary, macroscopic containers



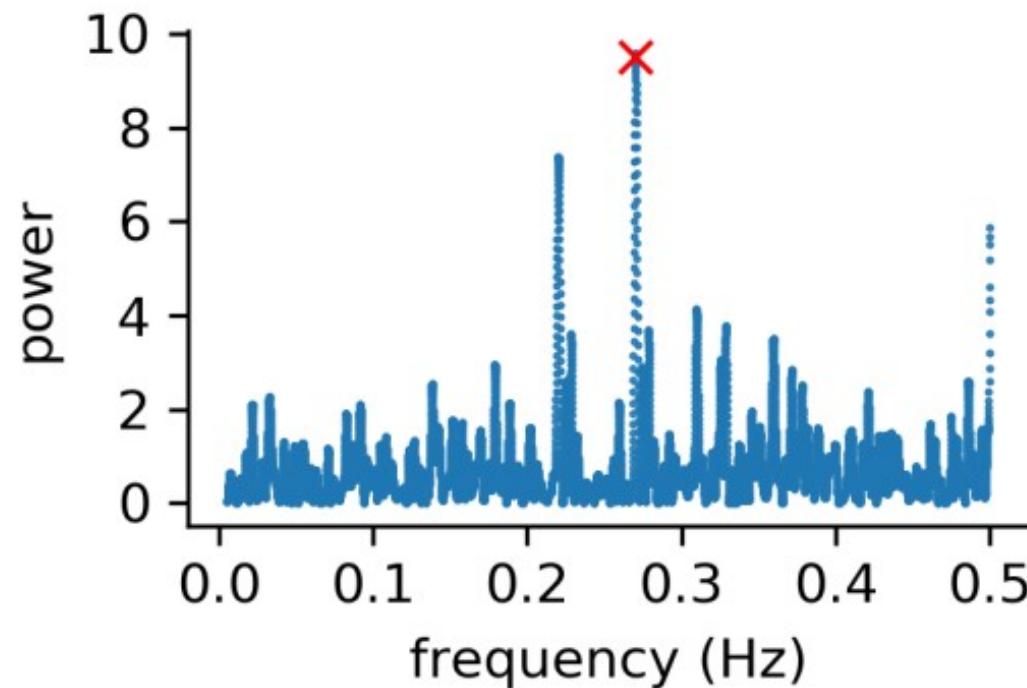
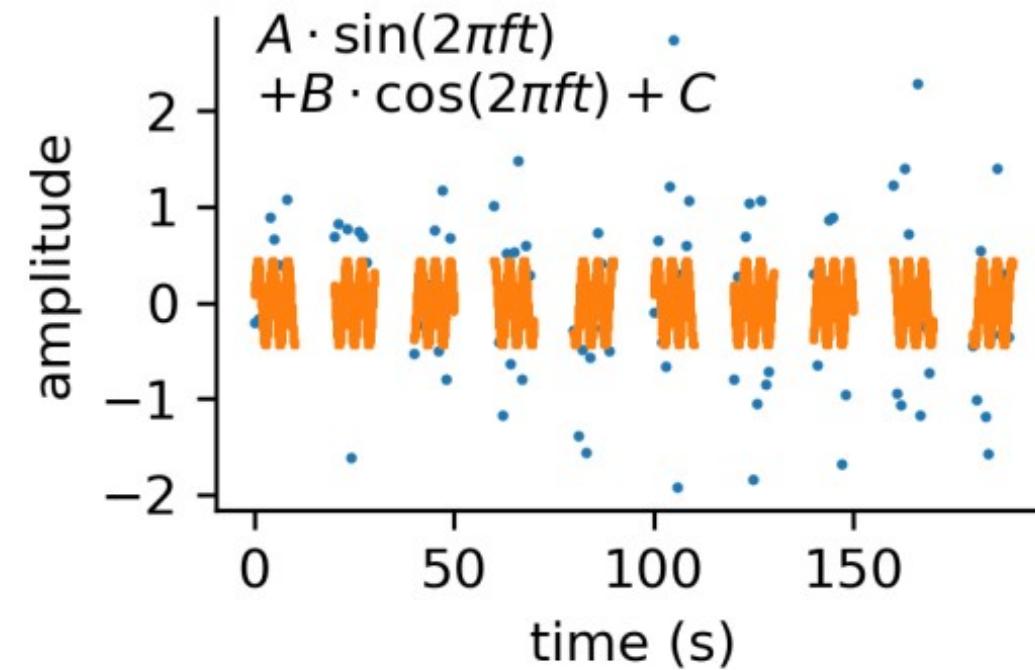
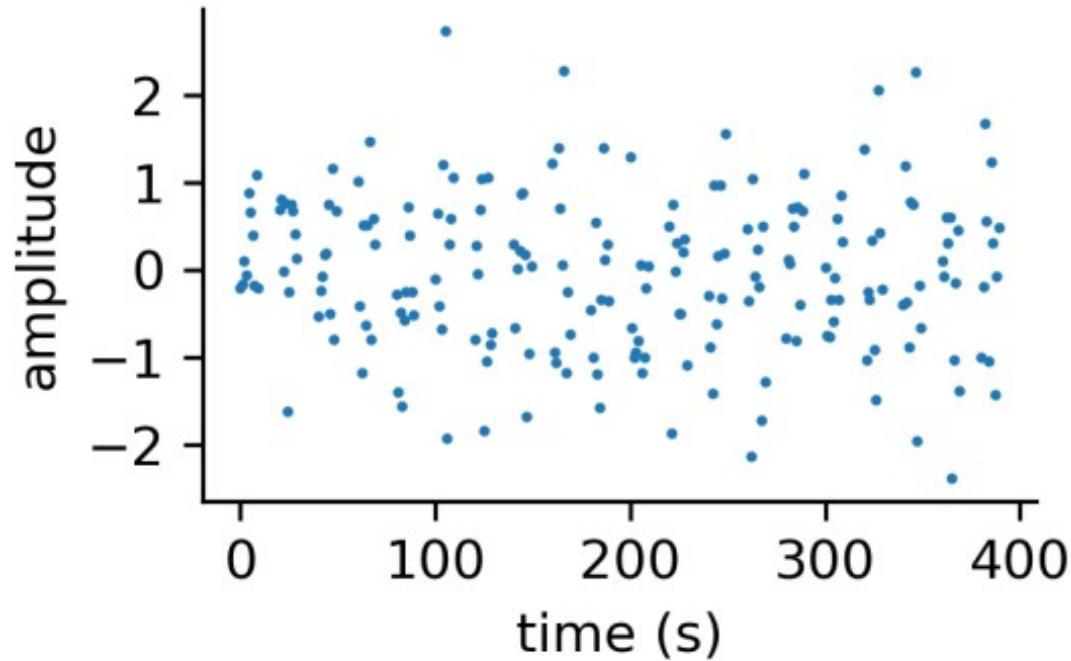
nEDM measurement

0. Polarised neutrons
1. Apply EM pulse with $f_{RF} \approx |\gamma_n|/2\pi B_0$ to flip the spins
2. Let them precess for $T = 180$ s
3. Apply second pulse, flipping them back onto z-axis
4. Measure asymmetry $A = (N_\uparrow - N_\downarrow)/(N_\uparrow + N_\downarrow)$

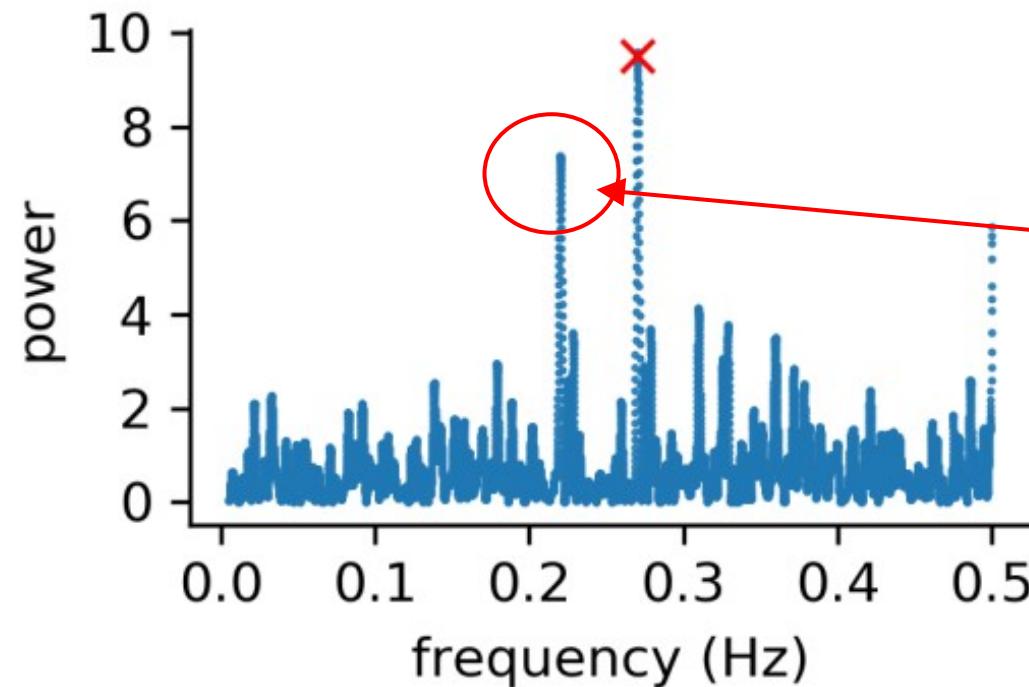
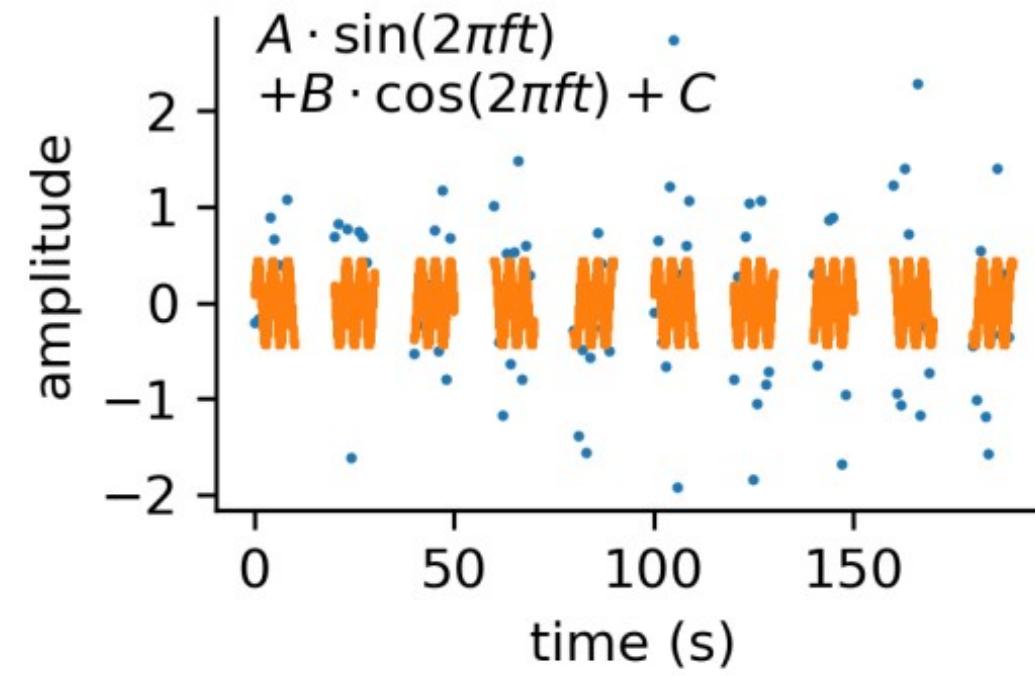
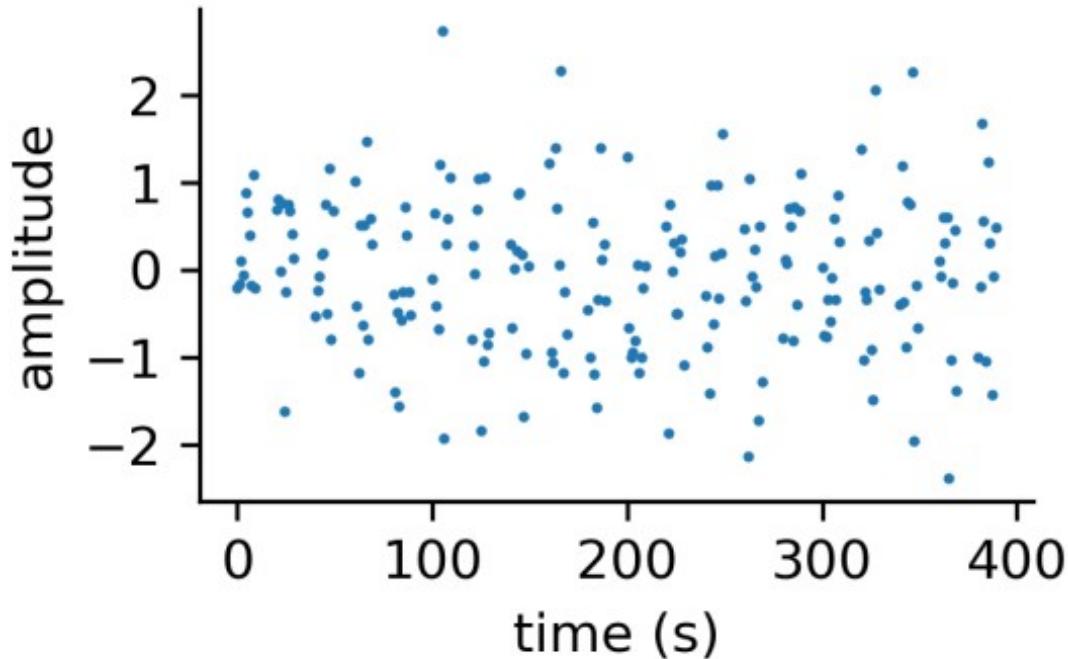


Spectral Analysis

- Normal FT not possible → unevenly spaced datapoints
- Use **LSSA** instead (least-squares spectral analysis) / (Lomb-Scargle periodogram)
 - fit a trigonometric function to the data: $A \sin \omega t + B \cos \omega t + C$
 - obtain power spectrum from fit parameters
 - repeat for all ω



Scan ~43'000 frequencies from 10^{-9} Hz to 10^{-3} Hz



The highest peak is not necessarily a signal!

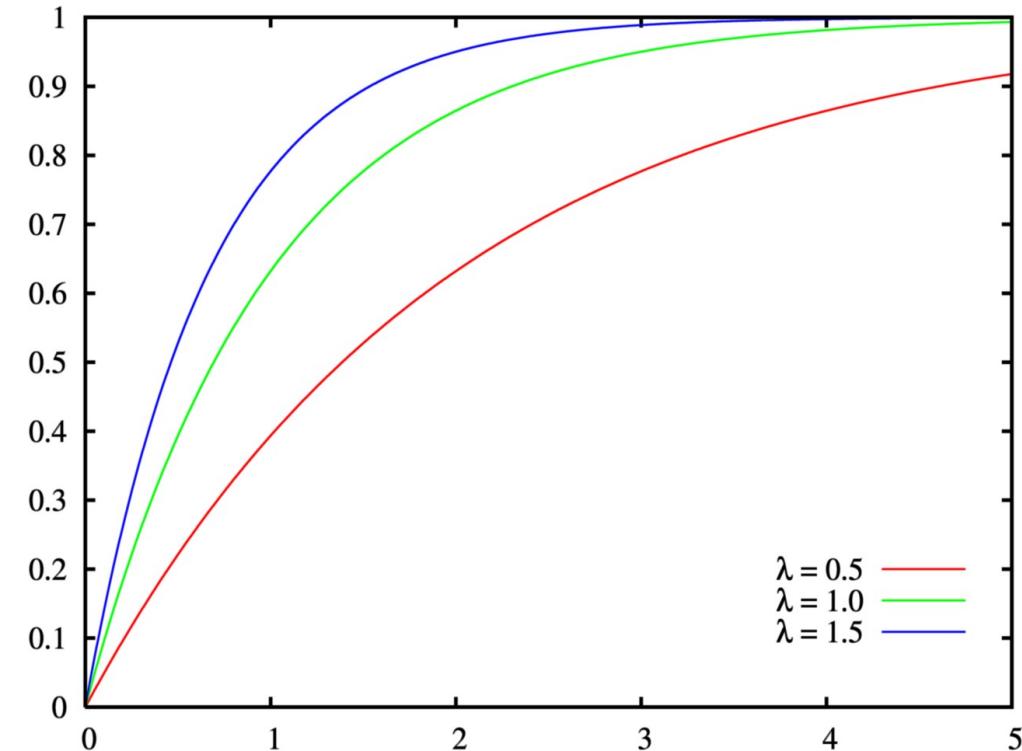
Null Hypothesis

What is the periodogram of nEDM data with no signal?

- MC simulations
- Fit CDF to the simulations
- Obtain probability for peak being just random fluctuations

CDF: cumulative distribution function
Probability, that a random variable X at point x is smaller than or equal to x.

e.g. $e^{\lambda x}$



Null Hypothesis

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False alarm probability:

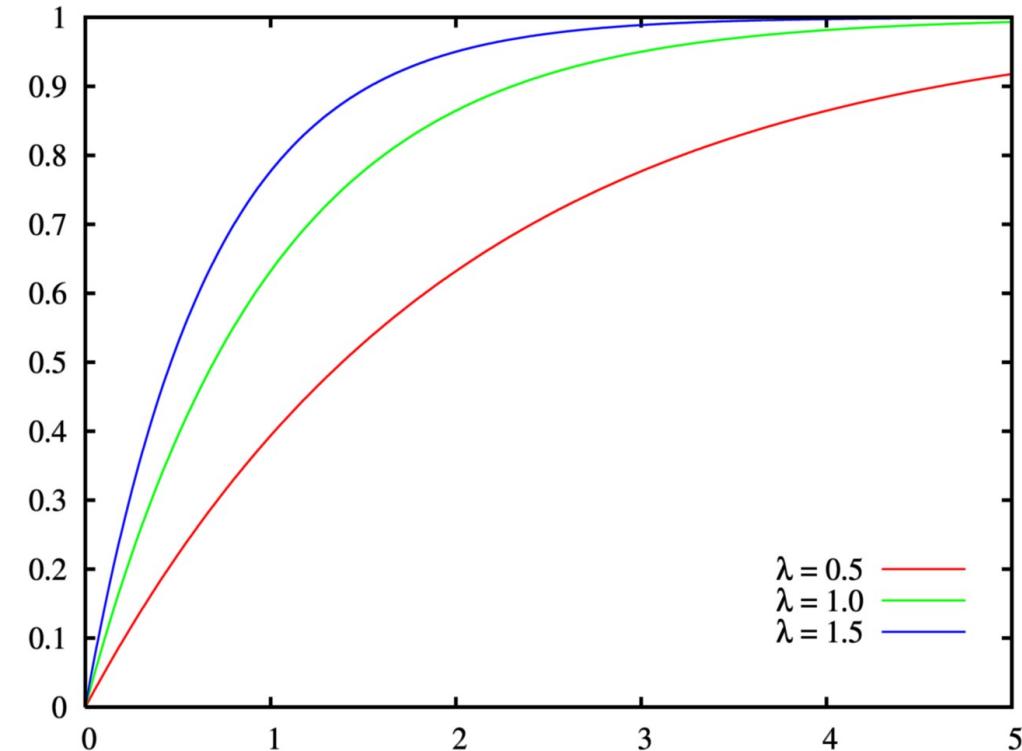
Lots of data → very rare events **will** happen.

$$p_{\text{global}} = 1 - F_Z(z)^N$$

N number of investigated events.

CDF: cumulative distribution function
Probability, that a random variable X at point x is smaller than or equal to x.

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Signal Hypothesis

No statistically significant signal consistent with the hypothesis was observed.

Determine Exclusion:

1. MC simulations with axion signals with parameters (A, ω)

2. Calculate the CDF

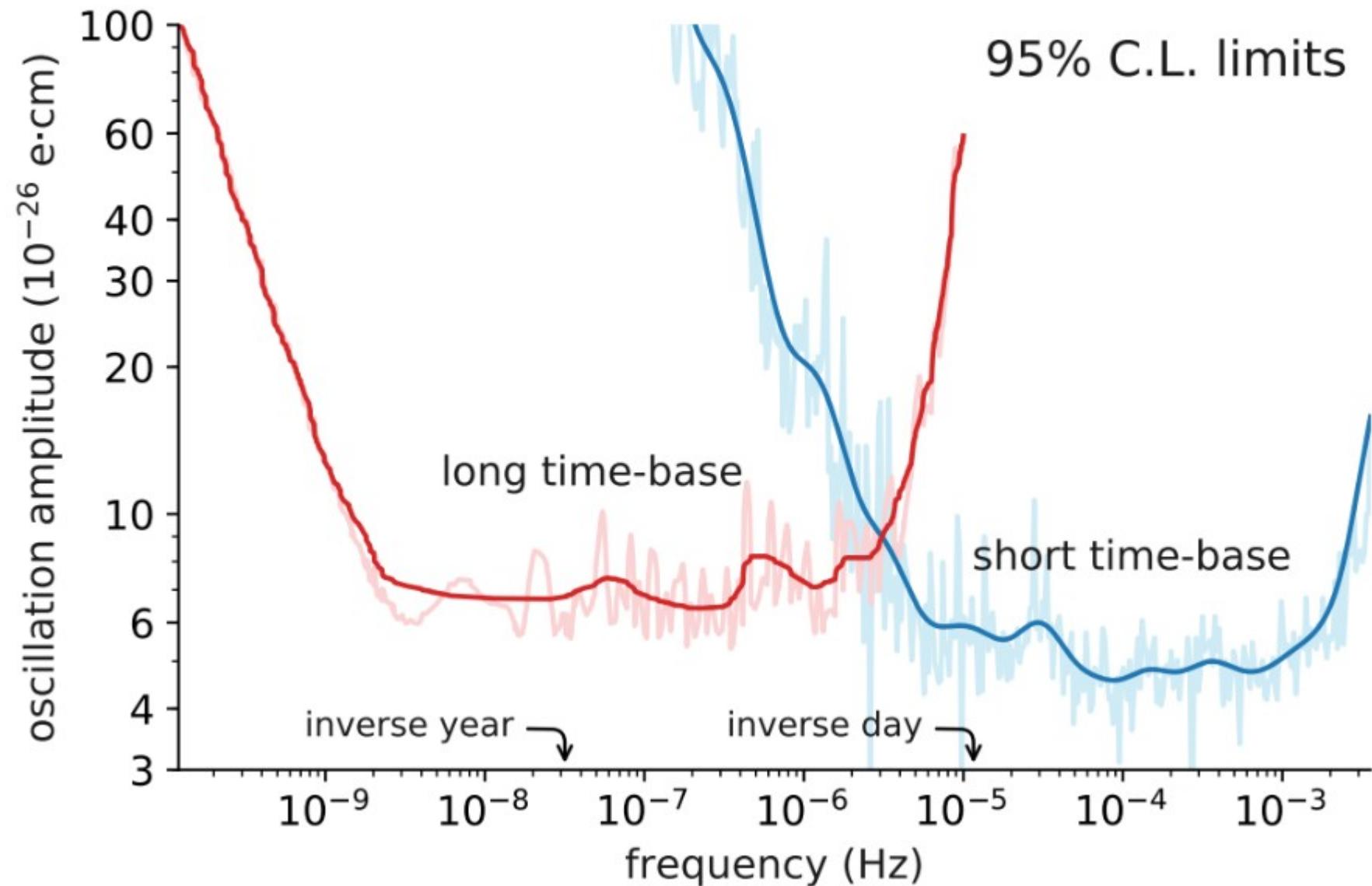
3. Obtain 2D exclusion zone

4. Use CLs method: $CL_s = \frac{\Pr(P^{H(\omega,A)}(\omega) < P^D(\omega))}{\Pr(P^{H_0}(\omega) < P^D(\omega))}$

→ correct exclusion in parameter space where the experiment is not sensitive

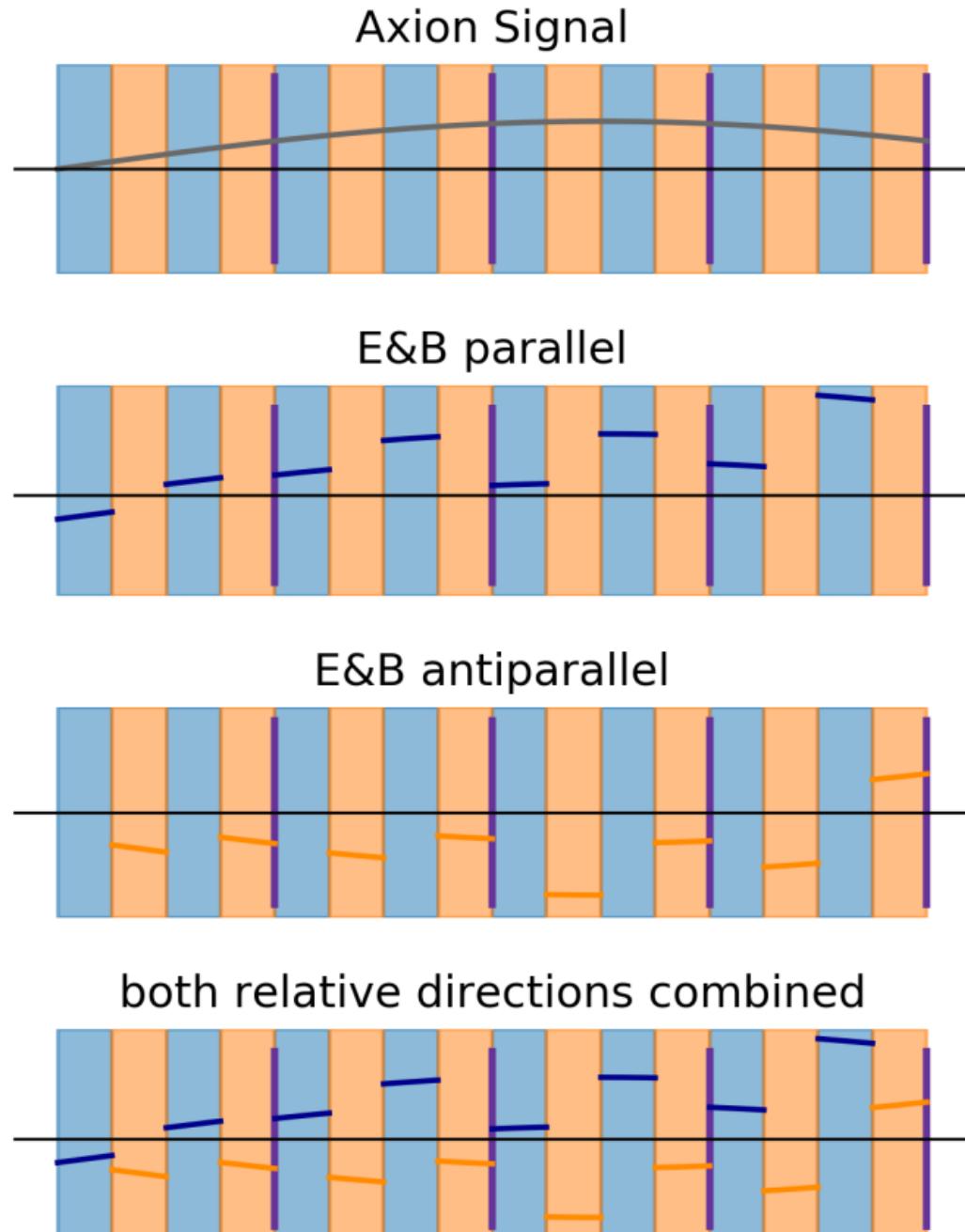
Results

Red: Sussex-RAL-ILL (2017)
Blue: PSI (2017)



Improvements: Better fitting

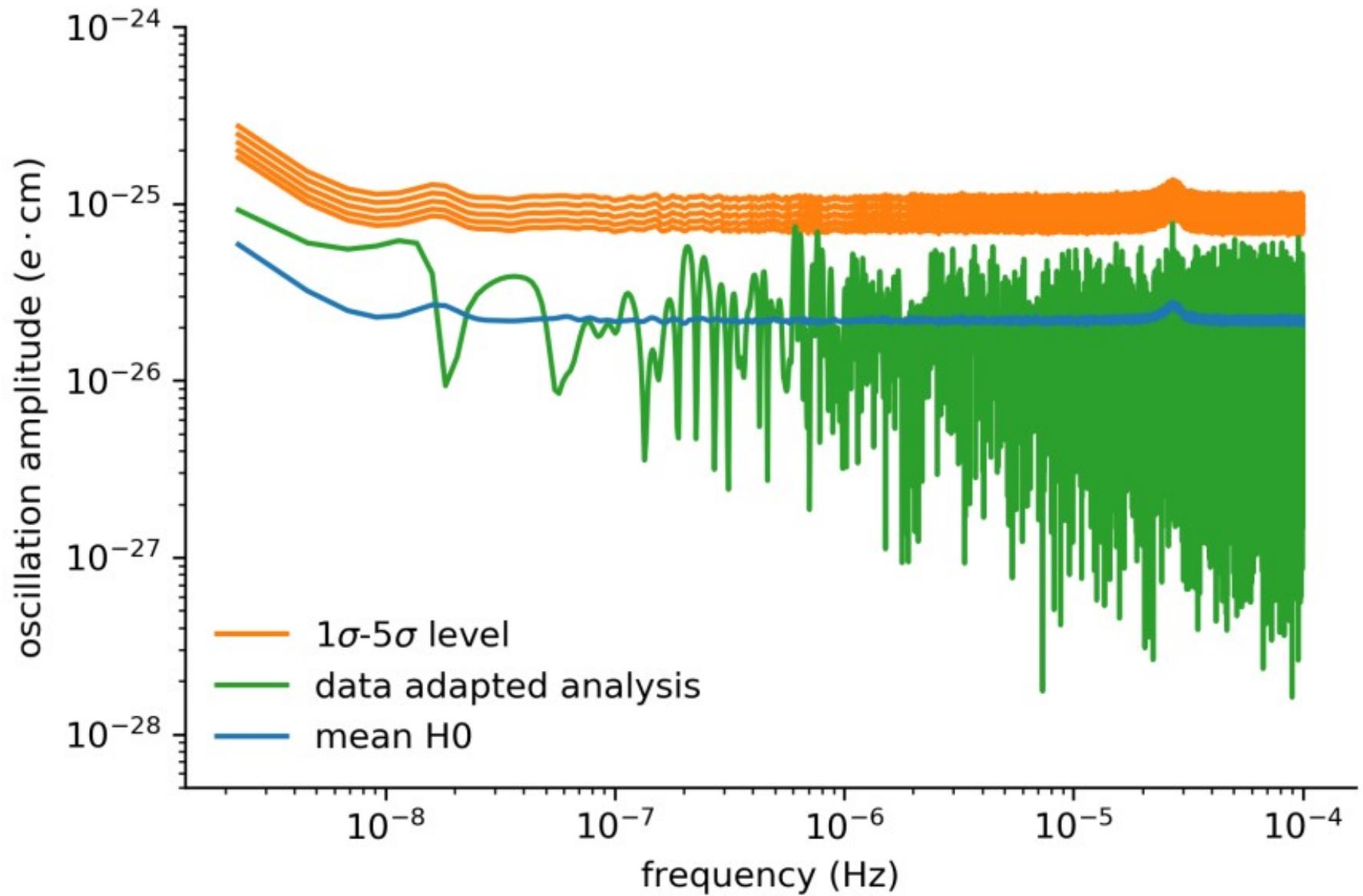
- PSI 2017: $EB^{\uparrow\uparrow}$ and $EB^{\uparrow\downarrow}$ fitted separately
- Better: Fitting both at the same time
 - determine offset without loss of sensitivity
 - better exclusion at low frequencies
 - separation reveals axion signal



Improved Results

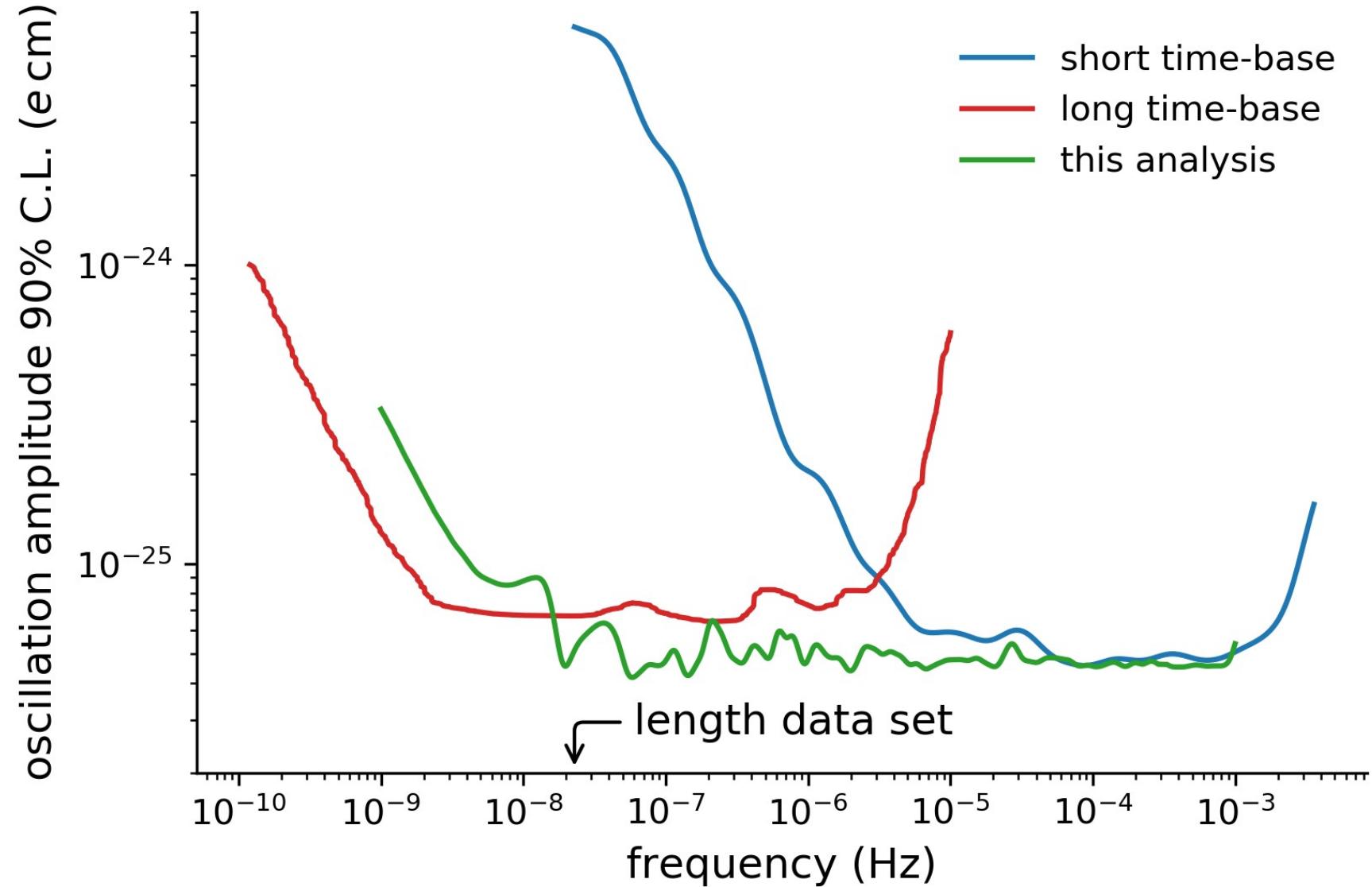
No peak above detection limit

Bump due to 112 cycle periodicity

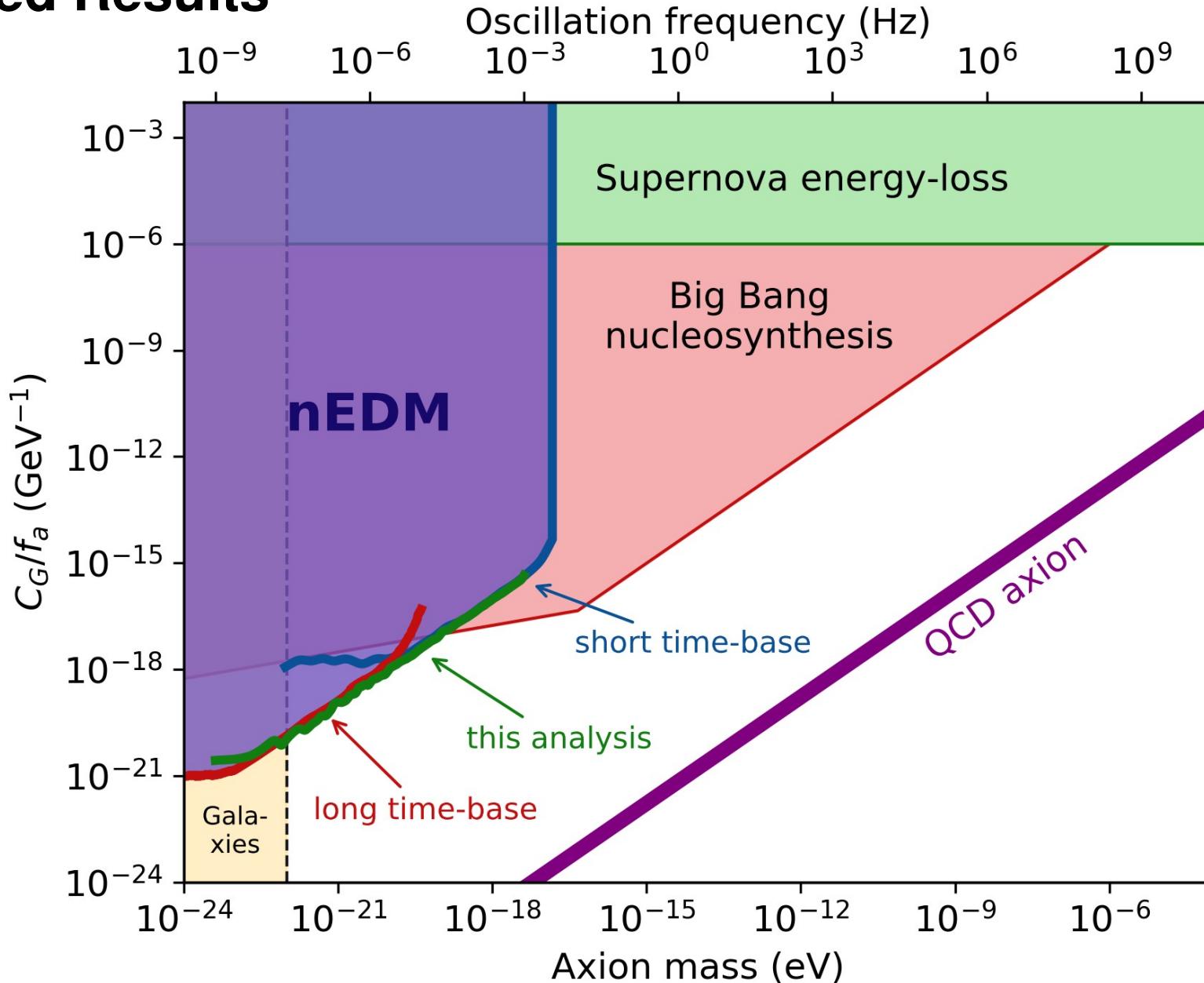


Improved Results

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Blue: PSI (2017)
Green: PSI



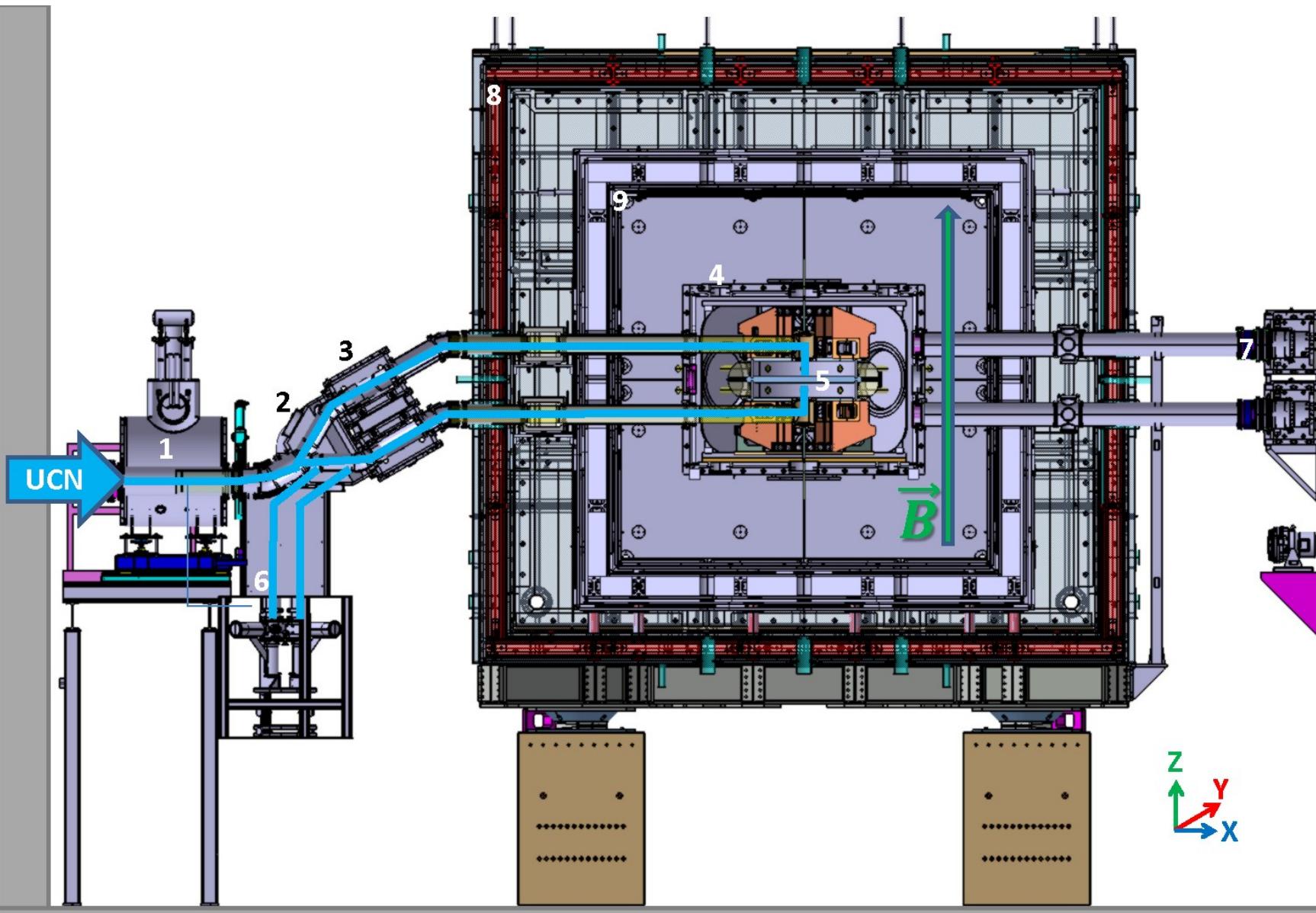
Improved Results



Outlook n2EDM

Apparatus improvements:

- 2 Chambers
→ EB $\uparrow\uparrow$ and EB $\uparrow\downarrow$ together
- Improved magnetometry
→ better gradient field measurement



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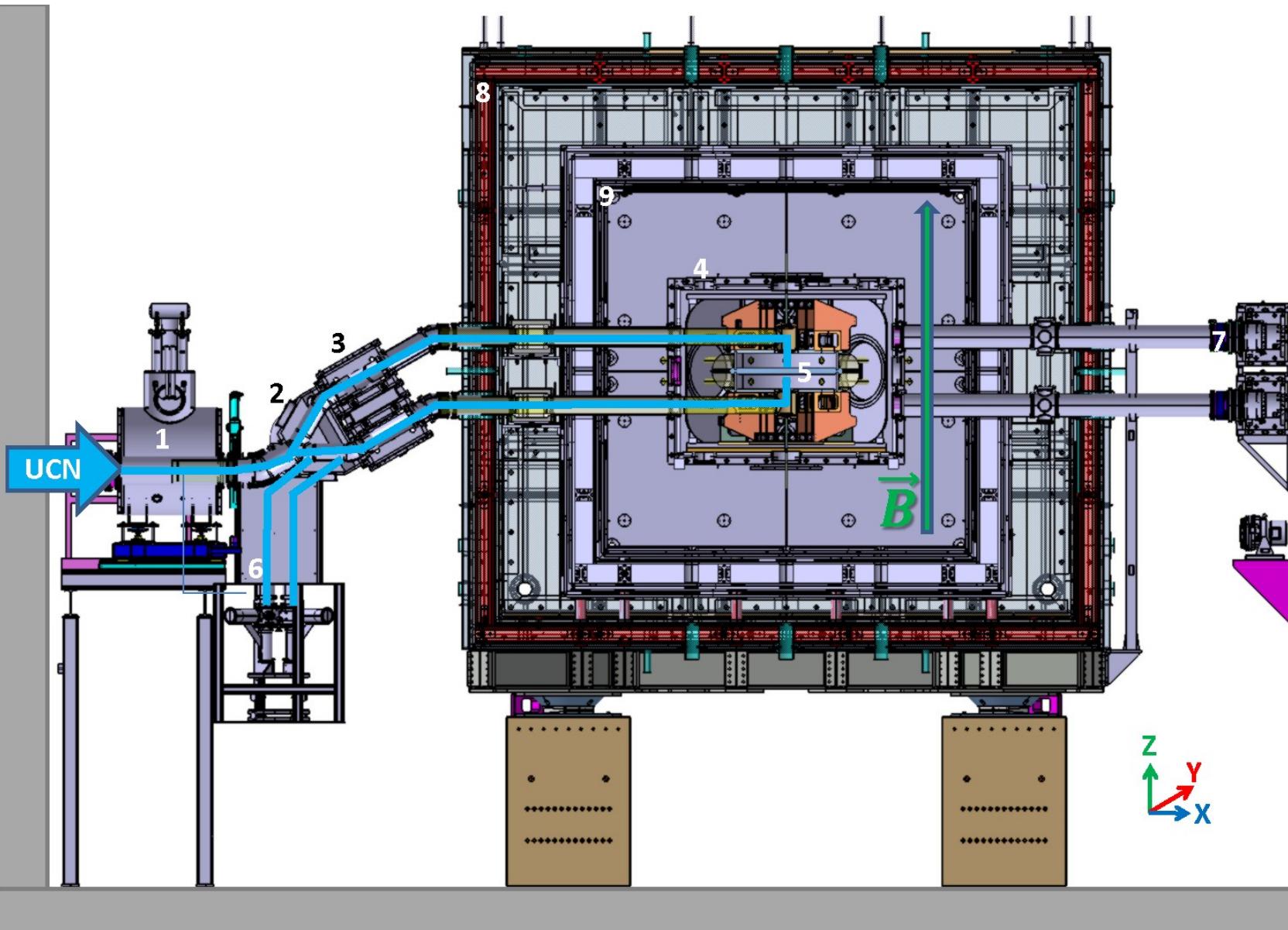
Statistical improvement:

- Bigger chambers
- Dynamic magnetic shielding
- ...

Better overall nEDM sensitivity

=

Better axion sensitivity!



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Expect factor 10 overall improvement

