



Axions and WISPs

17th Patras Workshop on Axions, WIMPs, and WISPs
09/08/2022 JGU Mainz, Germany

SungWoo YOUN

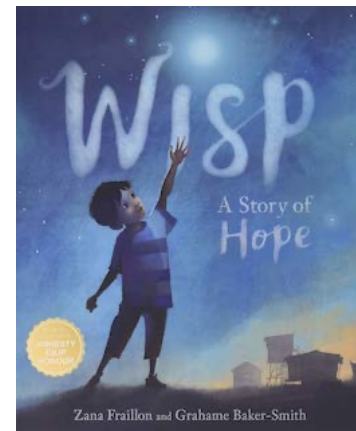
Center for Axion and Precision Physics Research (CAPP)

Institute for Basic Science (IBS)



Outline

- *Background*
 - *Dark matter and WISP*
 - *Search strategies*
- *Axion (ALP) searches*
 - *Photon coupling*
 - *Haloscope / Helioscope / LSW*
 - *Fermion coupling*
 - *nEDM coupling*
- *Other WISP searches*
 - *Hidden photon*
 - *Chameleon*
- *Prospects and summary*

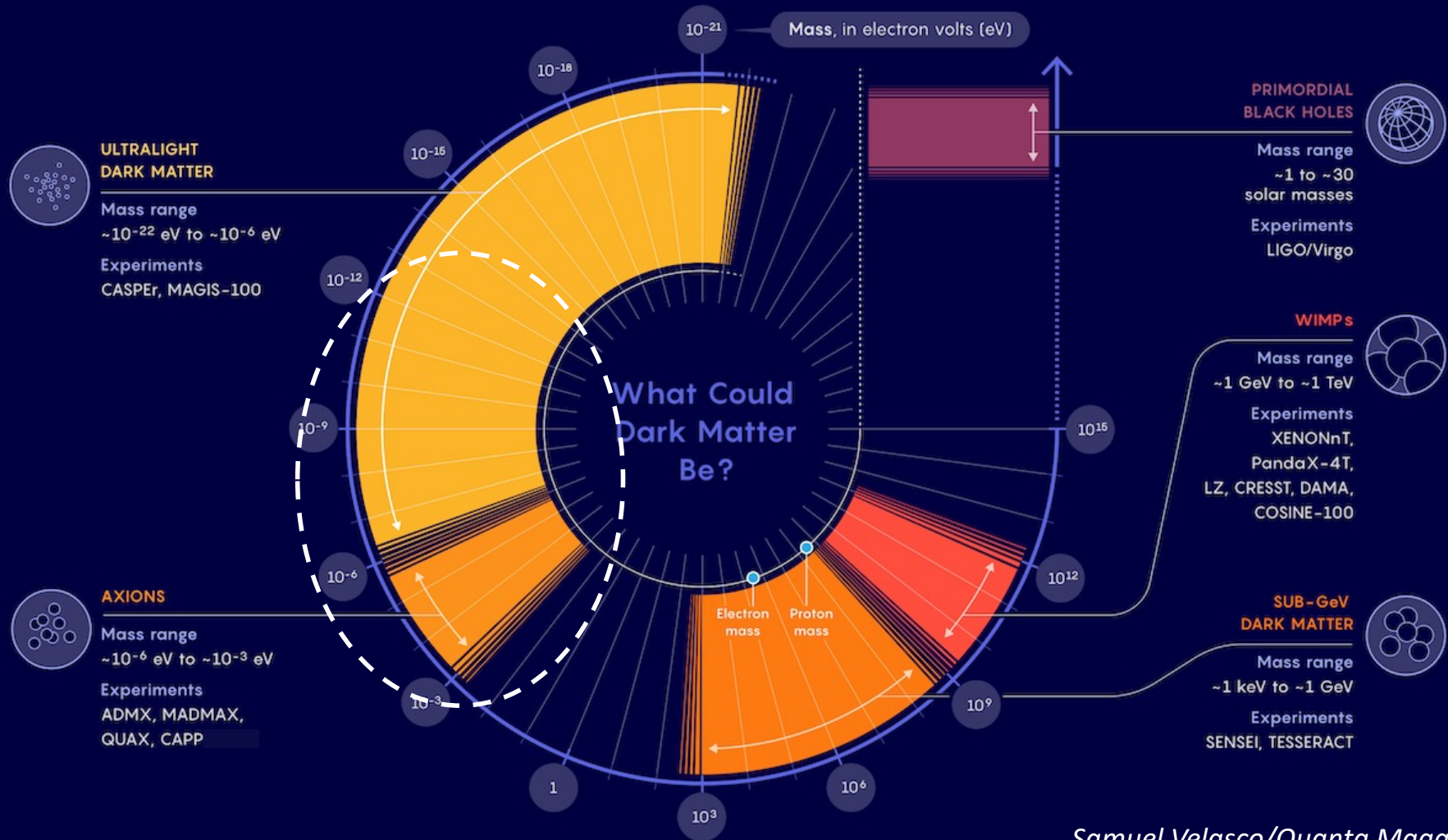


Disclaimer

- *My apologies for not covering*
 - *indirect searches*
 - *all individual efforts*
 - *most recent updates*



Dark matter business expanding



Samuel Velasco/Quanta Magazine



WISP zoo

- **Pseudo-scalar**

- **Axion**

- PQ solution to strong CP problem (1977)
 $m_a f_a \sim \Lambda_{QCD}$
- Invisible axion (1979)
- Dark matter candidate (1983)

- **Axion-Like Particle (ALP)**

- Generic axion w/o solving the strong CP problem
 $m_a f_a \not\sim \Lambda_{QCD}$

- **Scalar**

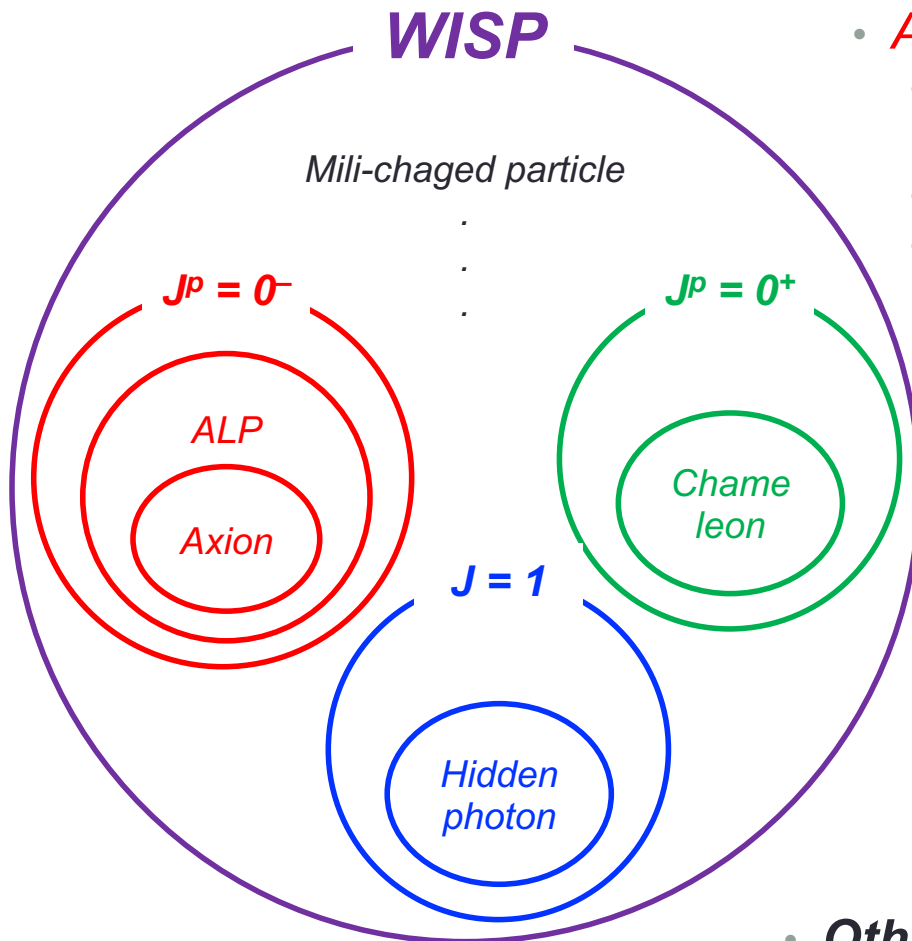
- Chameleon (2003)
- Dark energy candidate

- **Vector**

- Hidden photon
- Gauge field in hidden sector

- **Others**

- Mili-charged particle, ...





Couplings and models

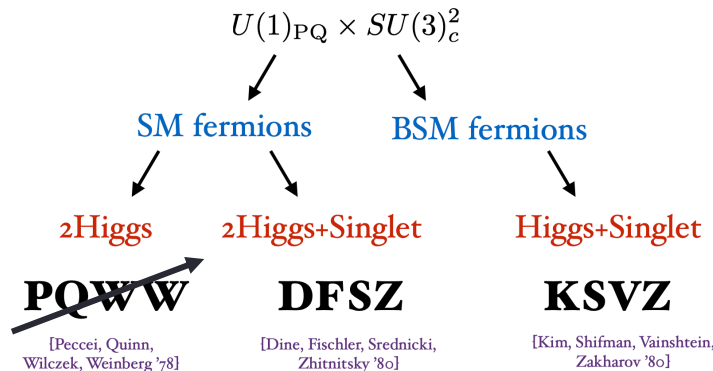
- Axion coupling to SM**

	<i>Photons</i>	<i>Fermions</i>	<i>nEDMs</i>
<i>Lagrangian</i>	$g_{a\gamma\gamma} a \mathbf{E} \cdot \mathbf{B}$	$g_{aff} \nabla a \cdot \hat{\mathbf{S}}$	$g_{EDM} a \hat{\mathbf{S}} \cdot \mathbf{E}$
<i>Observable (measurable)</i>	<i>Photon</i>	<i>Spin precession</i>	<i>Oscillating EDM</i>
<i>Detection</i>	<i>Power spectrum, photon counter, ...</i>	<i>Magnetometer, NMR, ...</i>	<i>NMR, polarimeter, ...</i>

- Axion models**

Specify origin of QCD anomaly (**fermion sector**) and spontaneous PQ breaking (**scalar sector**)

From R. Ziegler

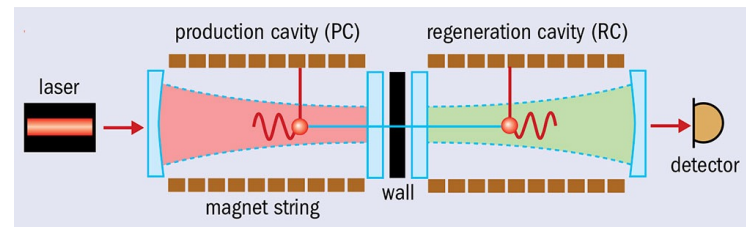
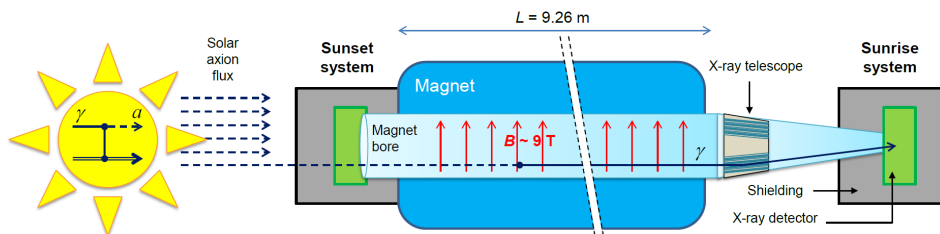
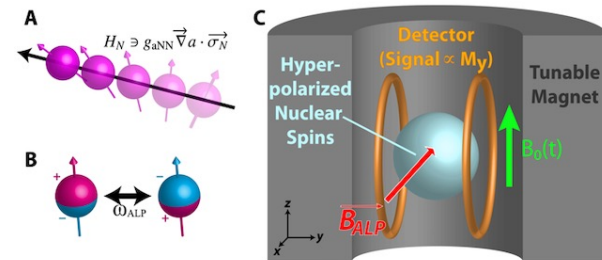
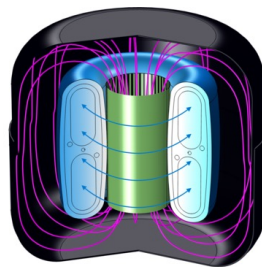
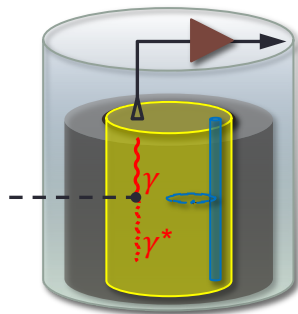


$\langle \text{Singlet} \rangle \gg v$:
“invisible” axion models



Experimental searches

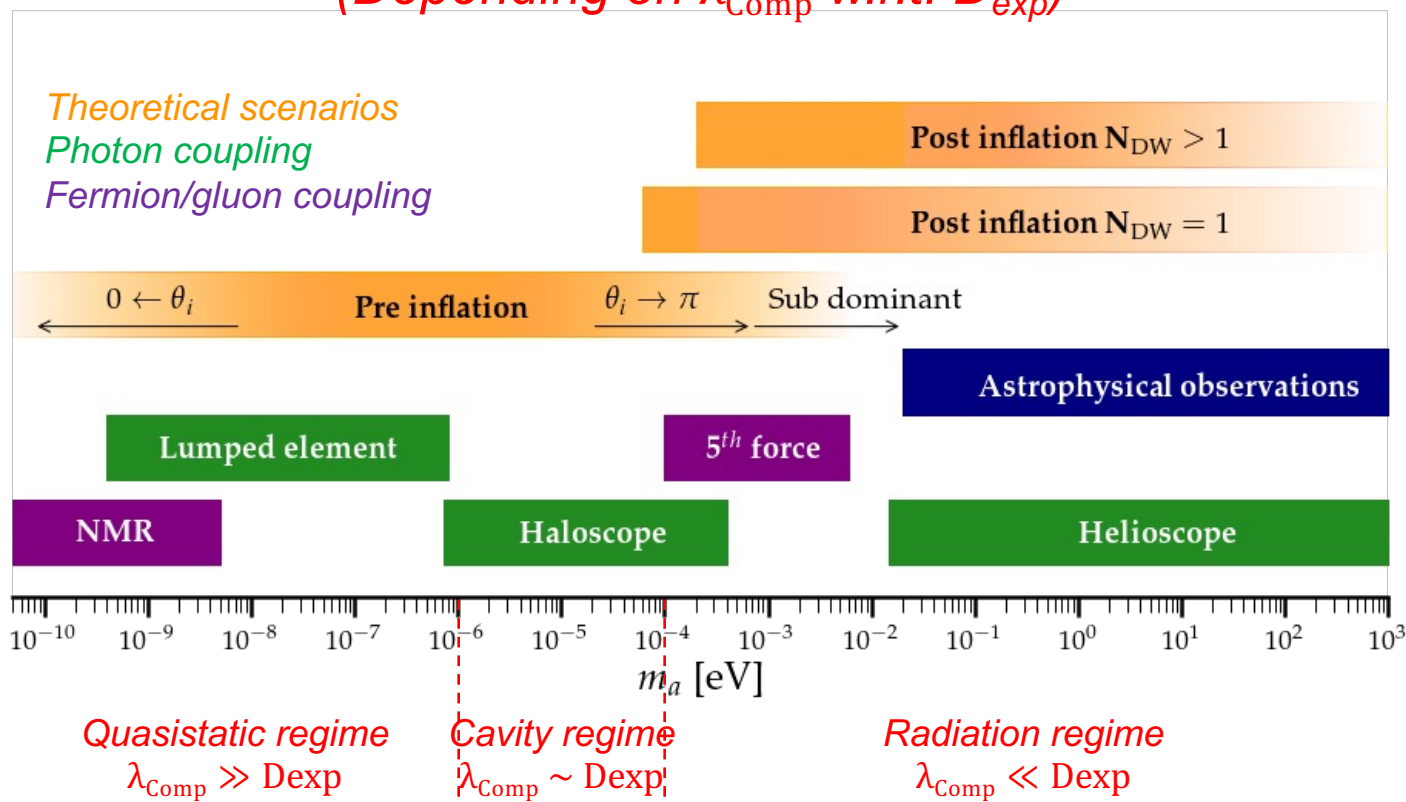
Source \ Coupling	Photons	Fermions	<i>n</i> EDMs
Dark matter	ADMX, CAPP, MADMAX, DM Radio, ...	QUAX-ae, GNOME, CASPER-wind, ...	CASPER-electric, srEDM, ...
Solar	CAST, IAXO		
Laboratory	ALPS (II)	ARIADNE	





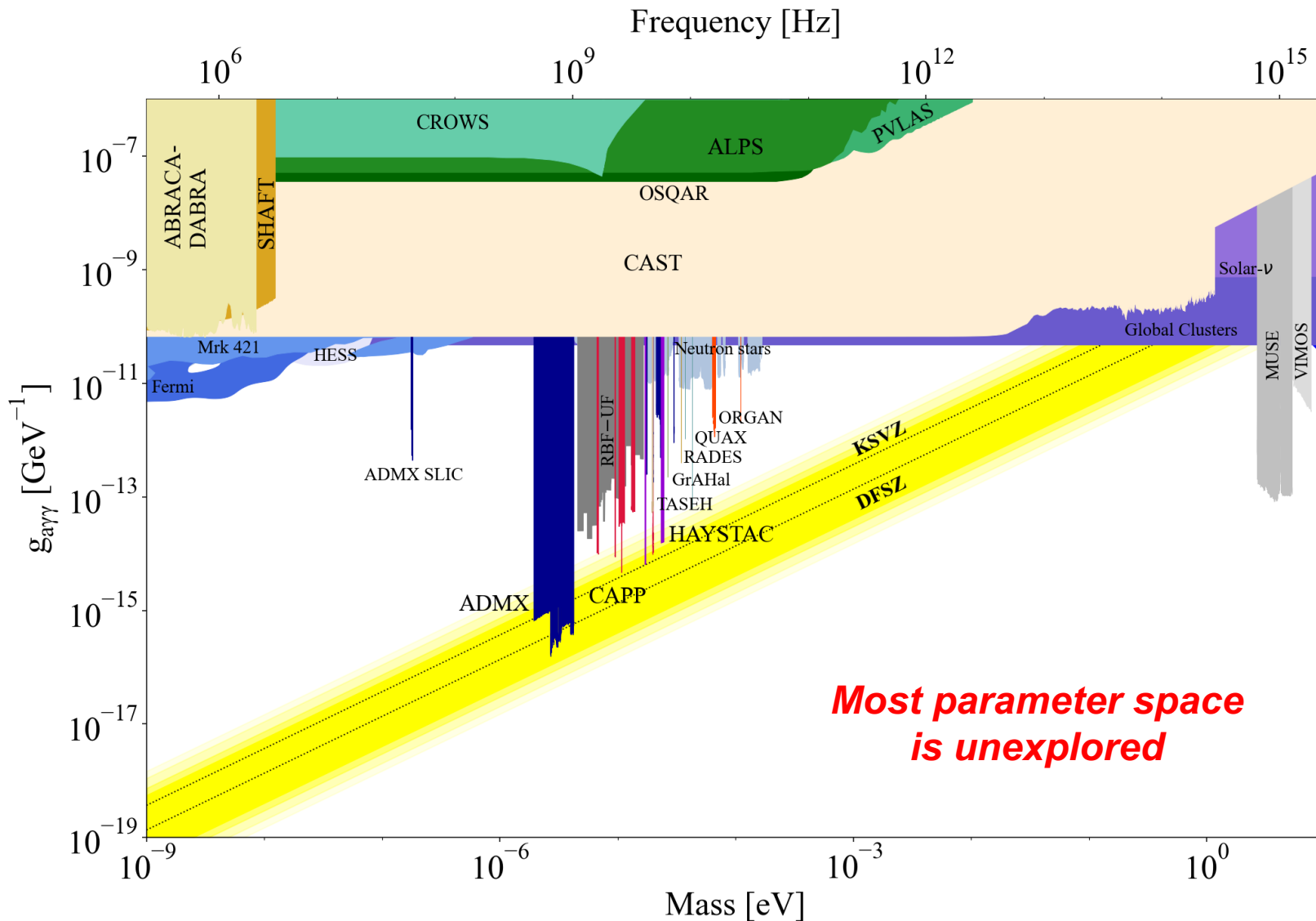
Where are dark matter axions?

- *Different PQ breaking scenarios*
 => *Different mass ranges*
 => *Different search strategies*
 (*Depending on λ_{Comp} w.r.t. D_{exp}*)





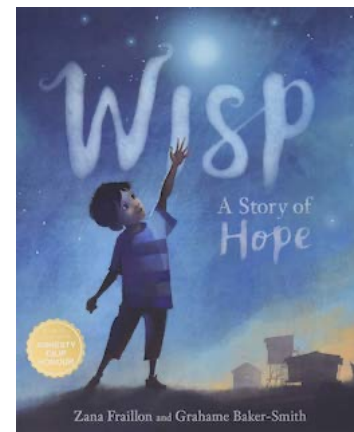
Axion searches – present





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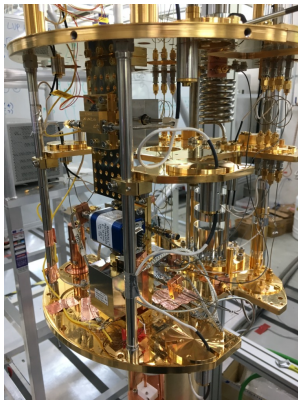


Cavity haloscope – in a nutshell

- *Most sensitive for microwave photons*

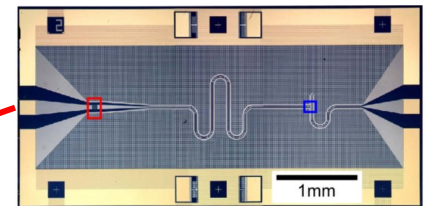
$$\frac{df}{dt} \sim B^4 V^2 C^2 Q_L T_{\text{sys}}^{-2}$$

Cryogenics T

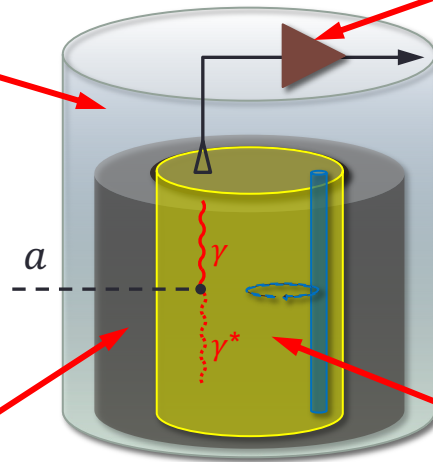


Lowering thermal noise

Quantum noise limited amplifier T

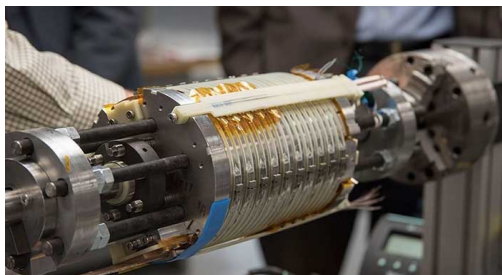


Signal amplification w/ minimal noise added



Tabletop experiments!

High field Magnet B



Boosting $a \rightarrow \gamma\gamma$ conversion rate

Tunable High-Q resonator $V, Q, C, \Delta f$

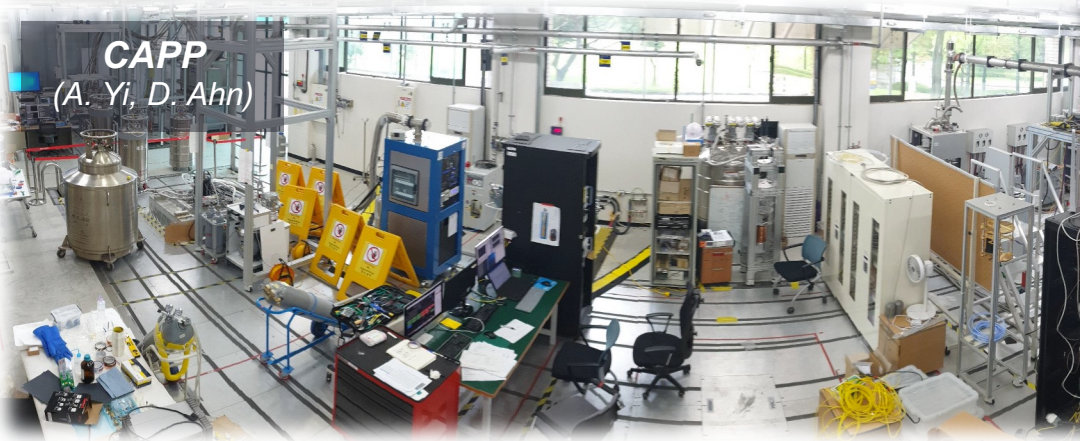
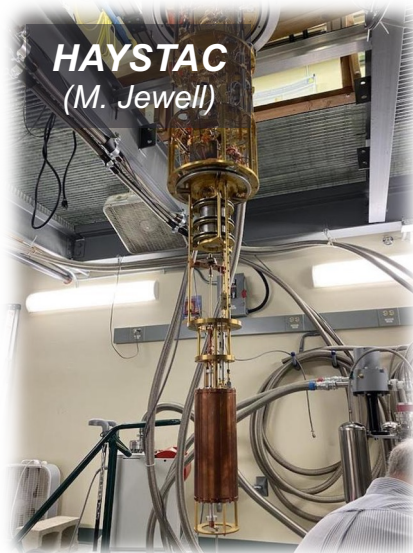
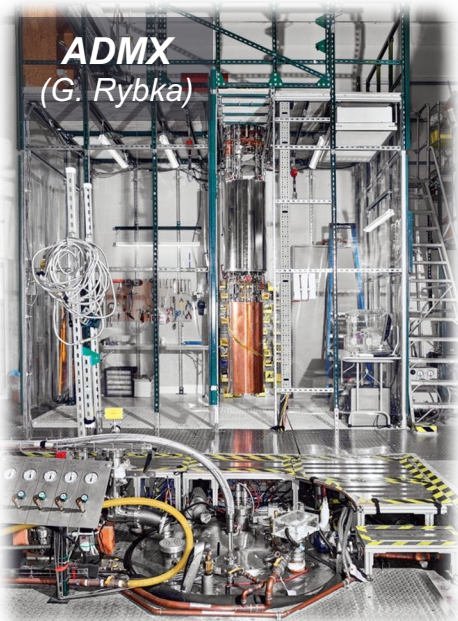


Resonant frequency tuning



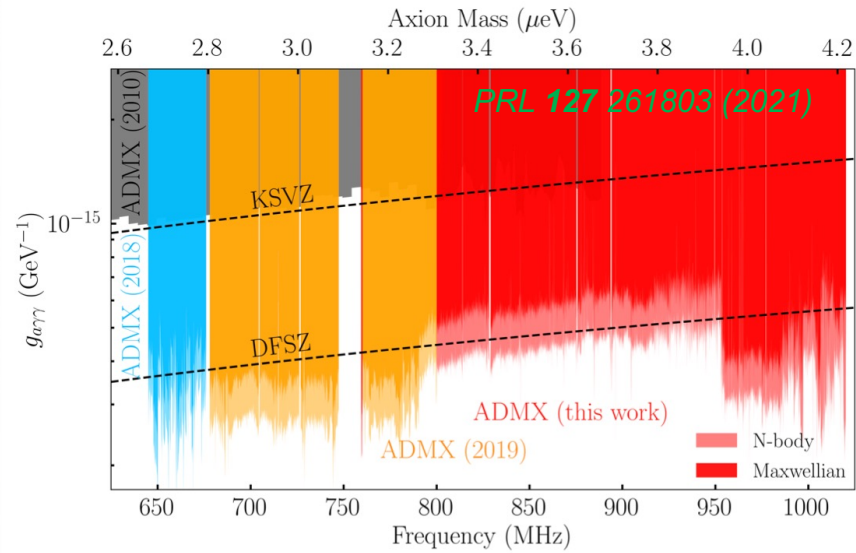
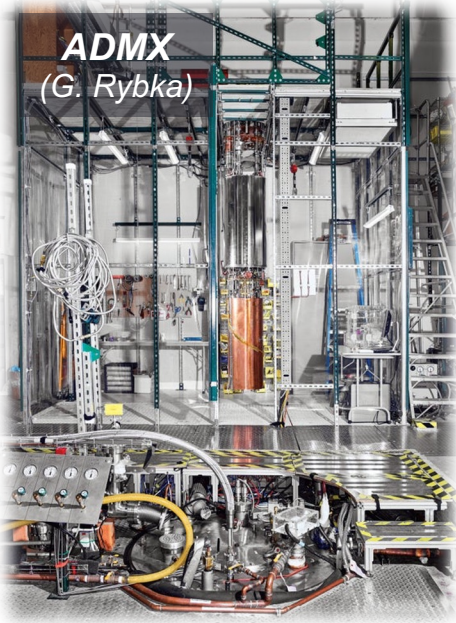
Cavity haloscopes

Cavity regime
 $\lambda_{\text{Comp}} \sim \text{Dexp}$



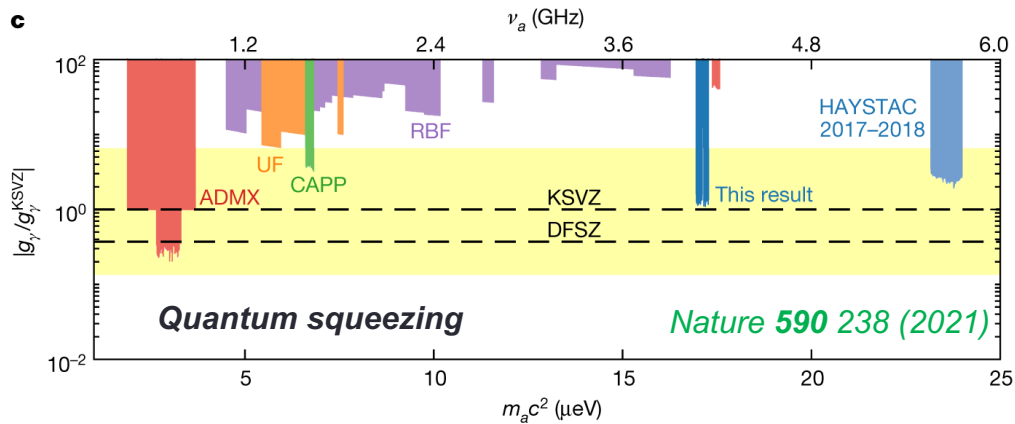
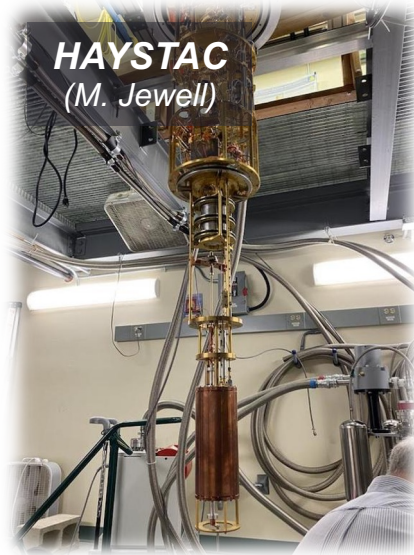


Search highlights



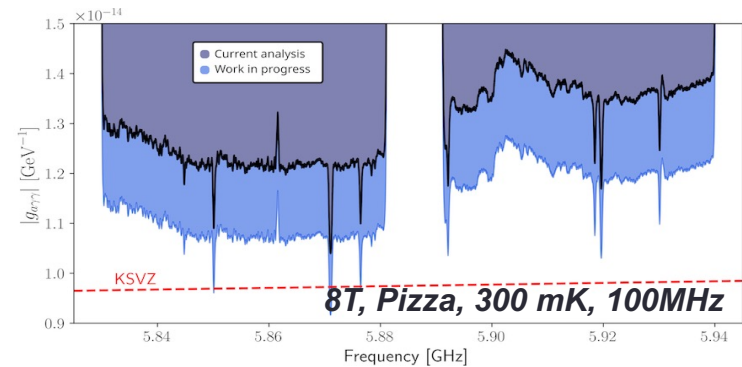
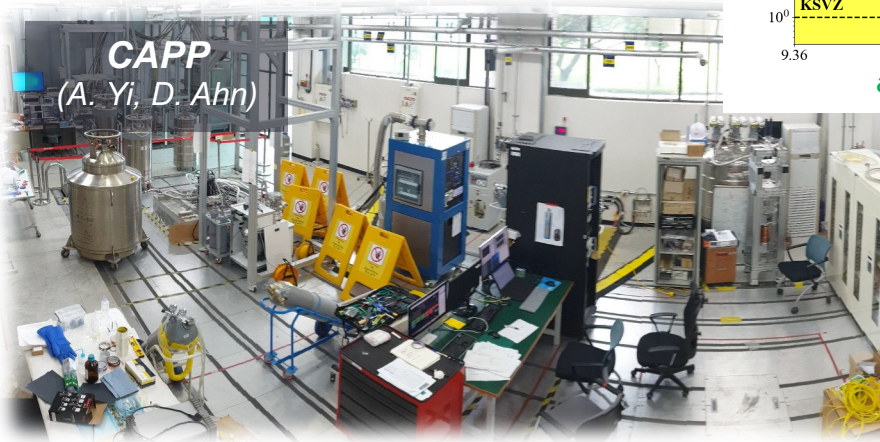
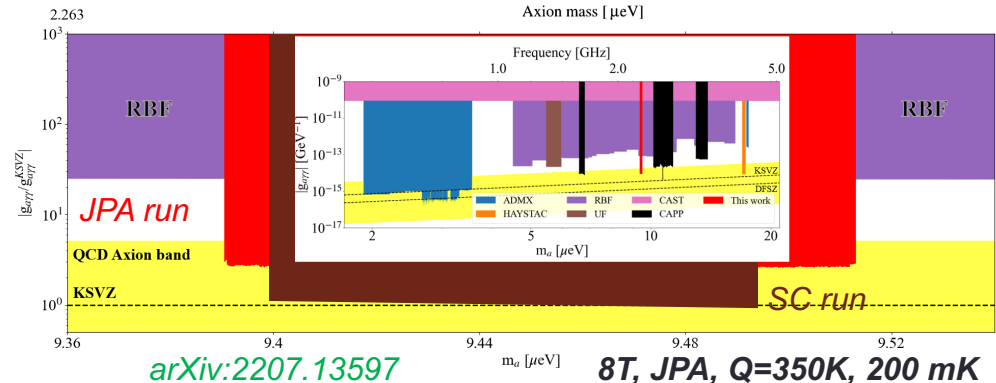
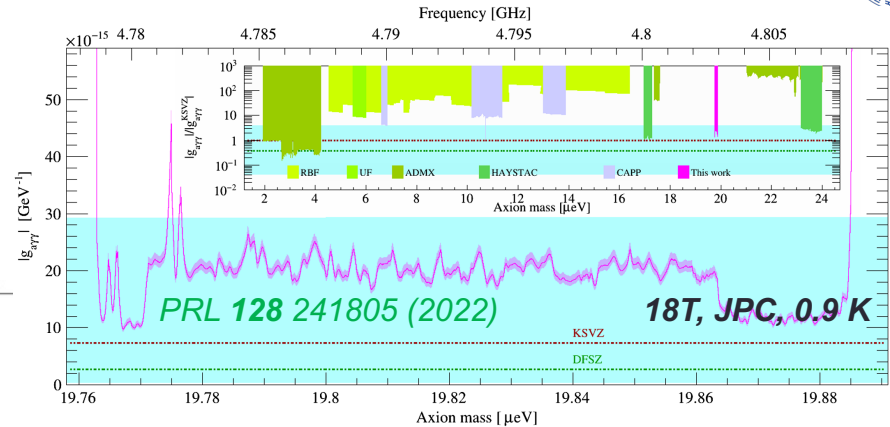
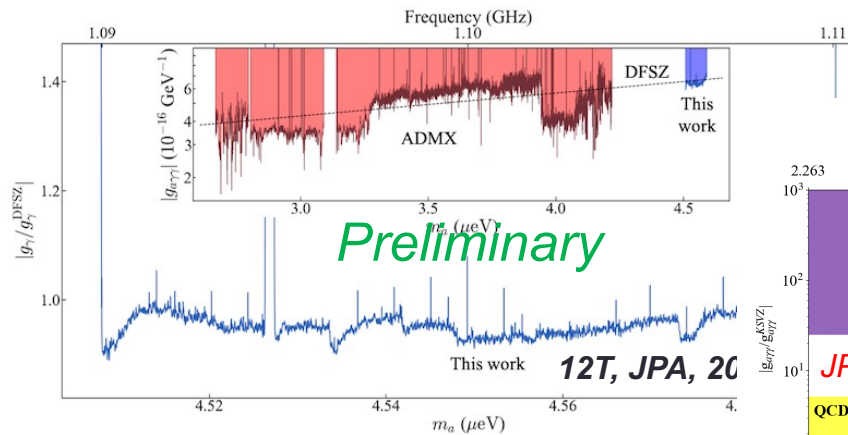


Search highlights



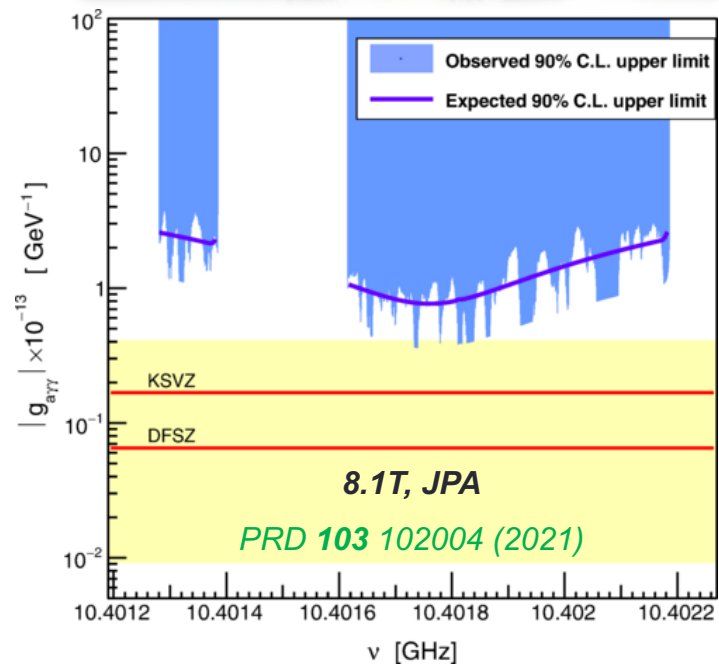


Search highlights



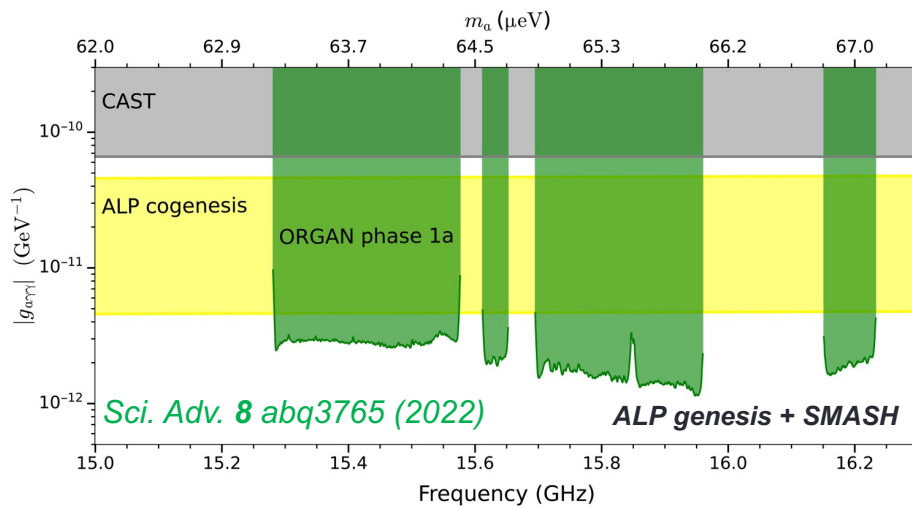


Search highlights





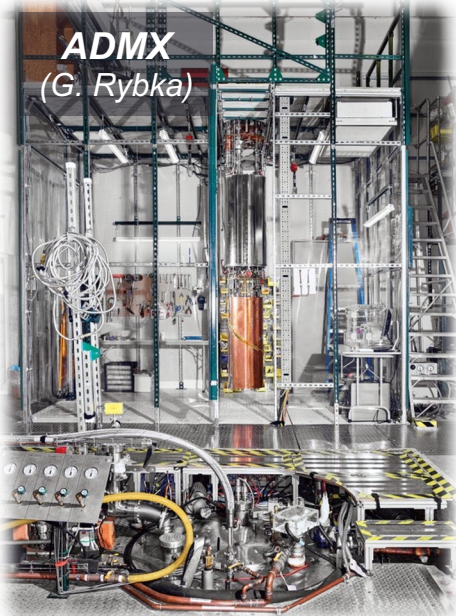
Search highlights





Cavity haloscope

Cavity regime
 $\lambda_{\text{Comp}} \sim \text{Dexp}$



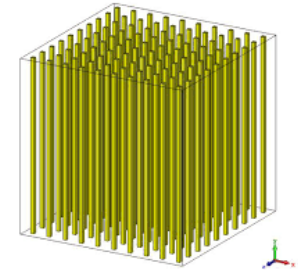
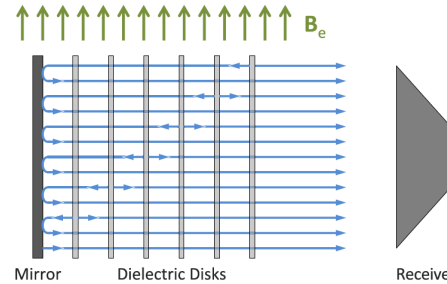


High frequency haloscope

Radiation regime
 $\lambda_{\text{Comp}} \ll \text{Dexp}$

- **Periodic array**

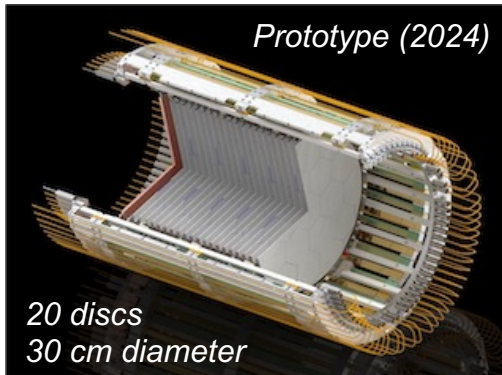
- Search frequency independent of detector size
- Suitable for post-inflation scenario (10–100 GHz)



- **MADMAX**

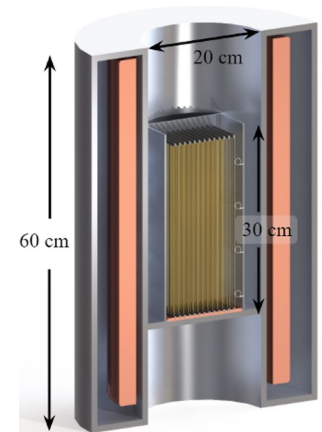
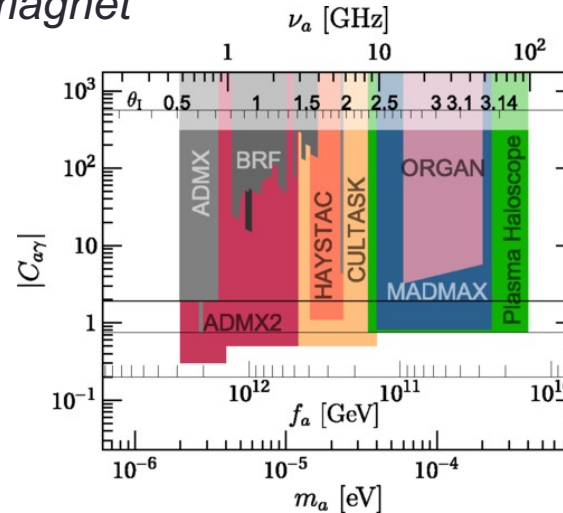
- Power boost w/ dielectric disks
- Full scale experiment
 - 2m-10T dipole / 80 1m² discs ($\epsilon \sim 25$)
- Proof-of-concept and 3D effect
- Prototype using Morpurgo magnet

A. Gardikiotis (Thu)



- **ALPHA**

- Resonance w/ plasma frequency (metamaterial)
- Demonstration of frequency tuning
- Physics data in 2026





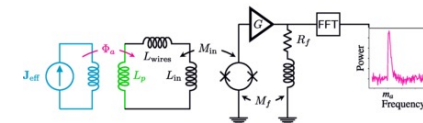
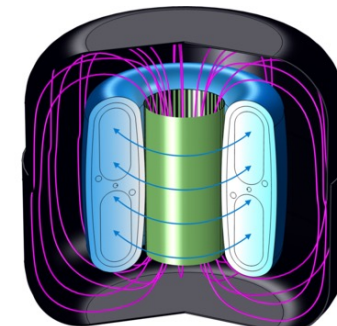
Lumped element haloscope

Quasistatic regime
 $\lambda_{\text{comp}} \gg \text{Dexp}$

- **Broadband low mass search ($< 1 \text{ ueV}$)**
 - *Sensitive to pre-inflation axions*

ABRACADABRA-10cm

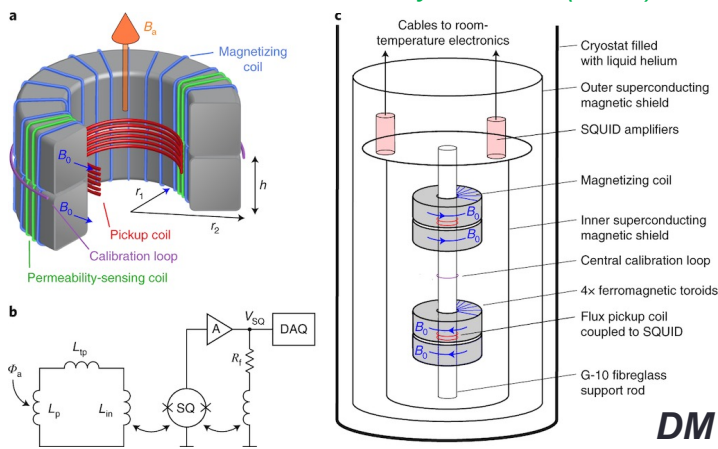
PRL 127 081801 (2021)



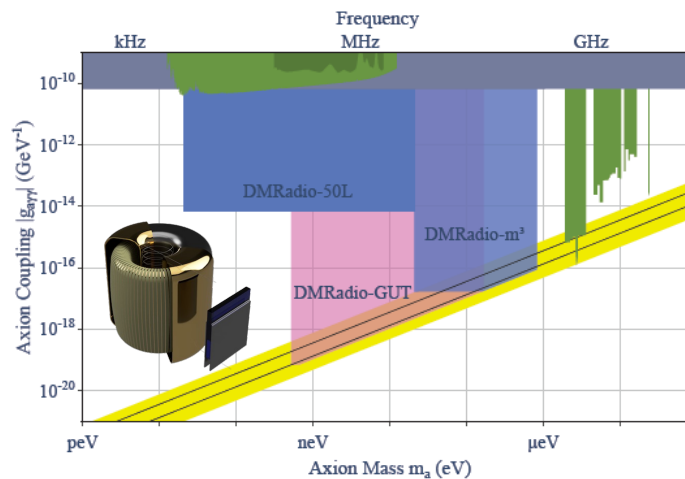
M. Smanovskaia (Thu)

SHAFT

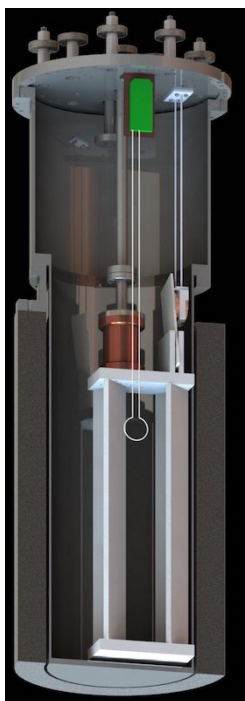
Nature Phys. 17, 79 (2021)



DM Radio



ADMX SLIC



PRL 124, 241101 (2020)



Other experiments

• Running

- CAST dipole (9T)
 - **CAST-CAPP**: phase-matched cavities, ~ 20 μeV
 - **RADES**: microwave fiber, ~ 34 μeV

M. Maroudas (Tue)

JHEP 2021 75 (2021)

• New

- **Grenoble Axion Haloscope**
 - 14T/52mm magnet, ~ 26 μeV
- **Taiwan Axion Search Experiment with Haloscope**
 - 4.7 GHz, $11 \times g_{\text{arr}}^{\text{KSVZ}}$
- **Superconducting axion search**
 - SC cavity, 14T, 8.4 GHz (under construction)

ArXiv:2110.14406

arXiv:2205.05574

K. Schmieden (Mon)

• Proposed

- **Broadband Reflector Experiment for Axion Detection**
 - Parabolic reflector, THz region
- **Canfranc Axion Detection Experiment**
 - 90 GHz (W-band), Kinetic Induction Detectors

PRL 128 131801 (2022)

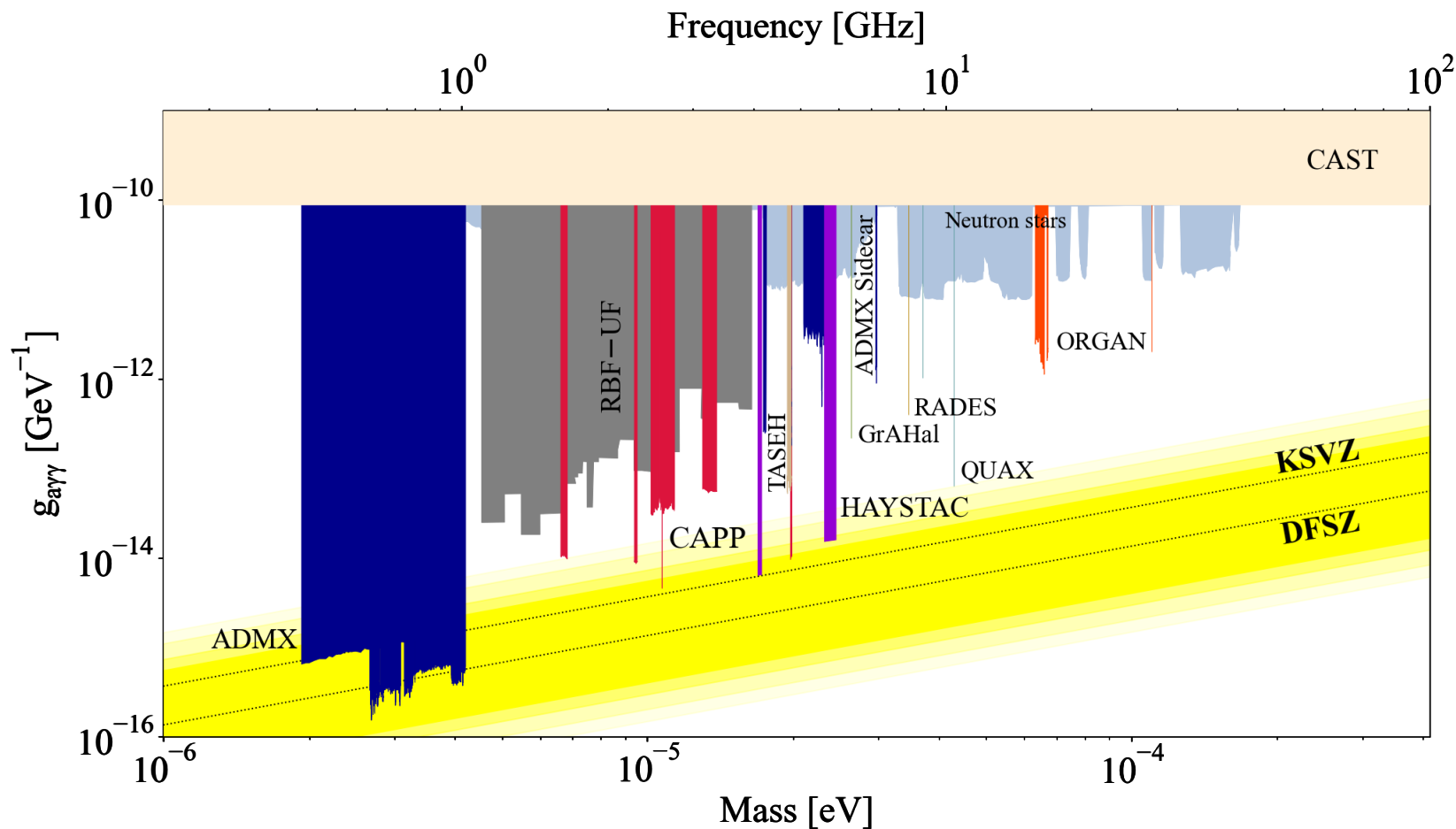
S. Knirck (Wed)

arXiv:2206.02980



Haloscope searches

Present (~2022)





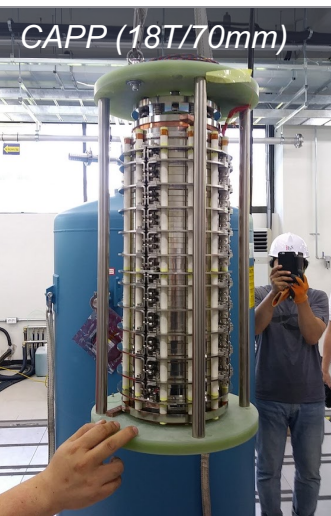
Magnet



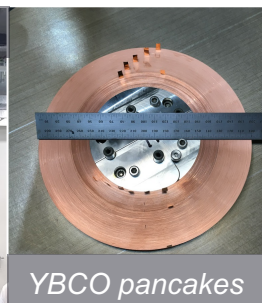
ADMX (8T/600mm)



CAPP (12T/320mm)



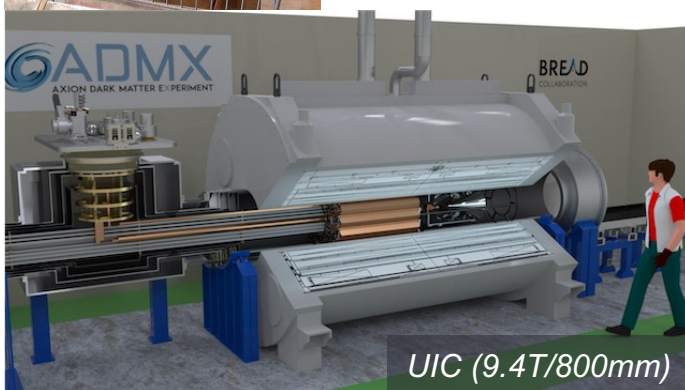
CAPP (18T/70mm)



YBCO pancakes



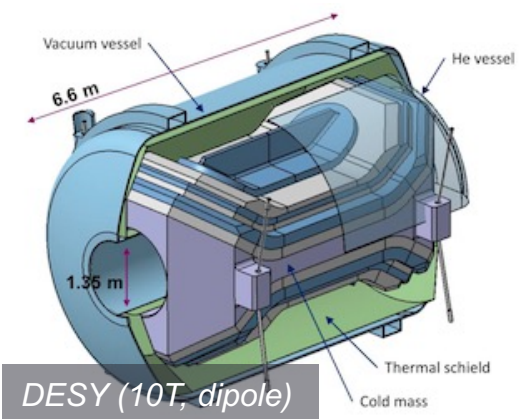
Grenoble (43T/34mm)



UIC (9.4T/800mm)



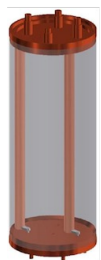
CERN (1.6T)



DESY (10T, dipole)



Cavity

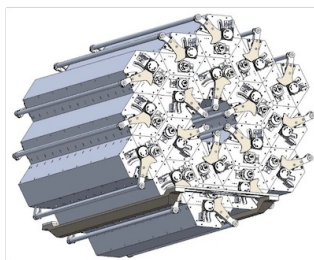


Run 1A-C

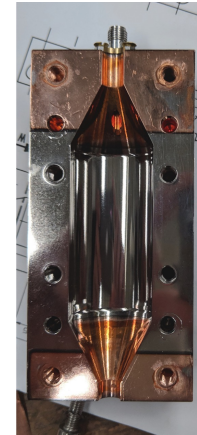


Run 2

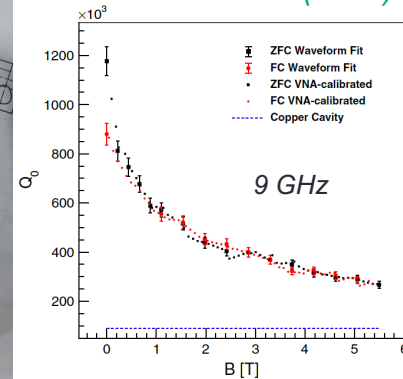
arXiv: 2203.14923



EFR

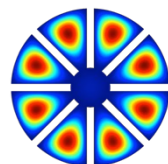
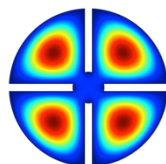
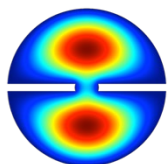


PRD 99 101101 (2019)

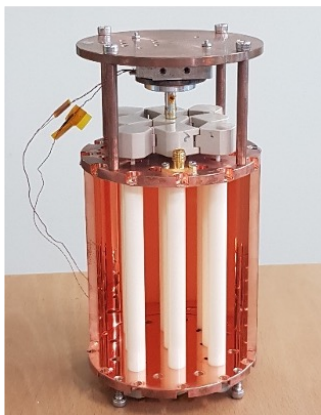


D. Ahn (Tue)

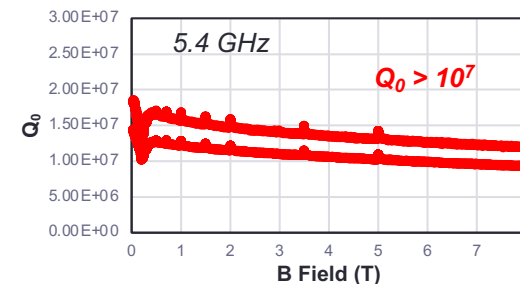
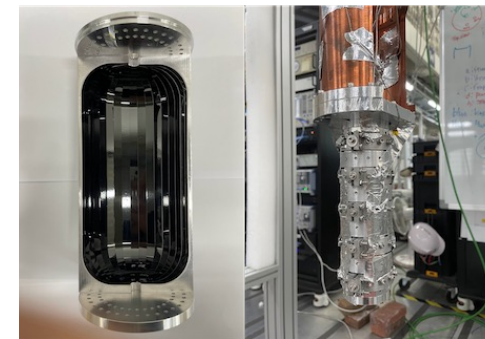
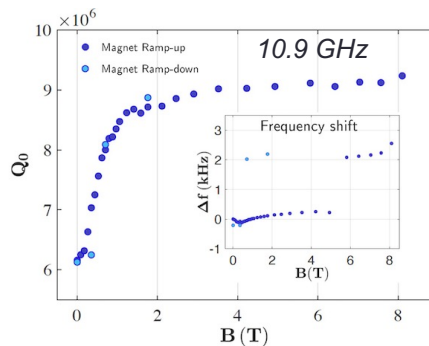
PRL 125 221302 (2020)



arXiv: 2205.08885



NIM 985 164641 (2021)

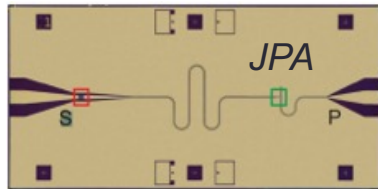




Microwave photon detector

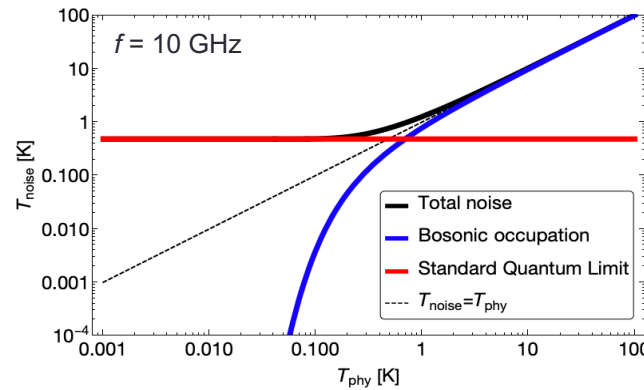


Transistor-based
($T_N \sim K$)

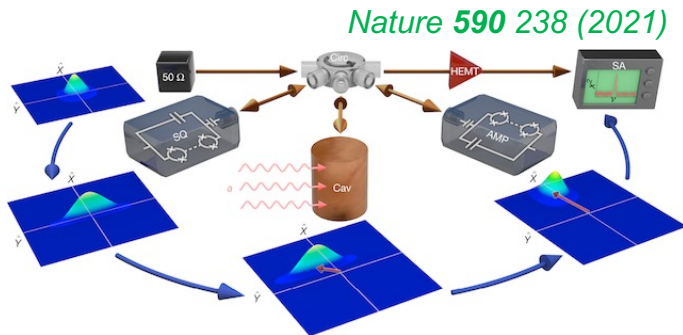
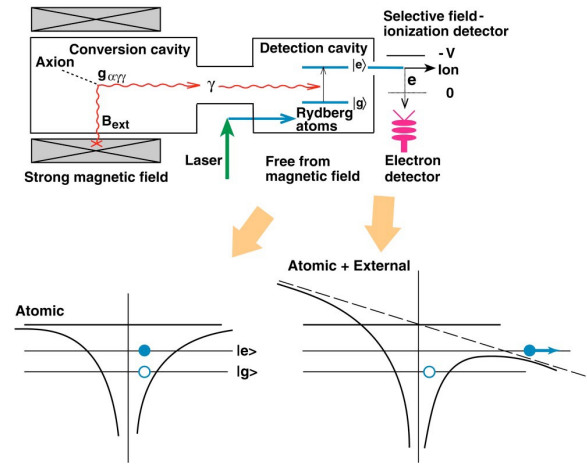


Quantum limited
($T_{SQL} \sim 50 \text{ mK} \times f \text{ [GHz]}$)

Power detection vs. photon counting (w/ amplifiers) (w/ single photon detector)



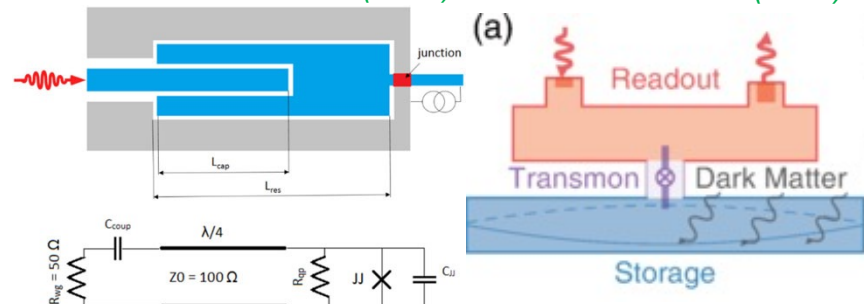
**Game changer
at high freq. and low temp.**



Quantum squeezing ($T_N < T_{SQL}$)

IEEE TASC 2850019 (2018)

PRL 126 141302 (2021)



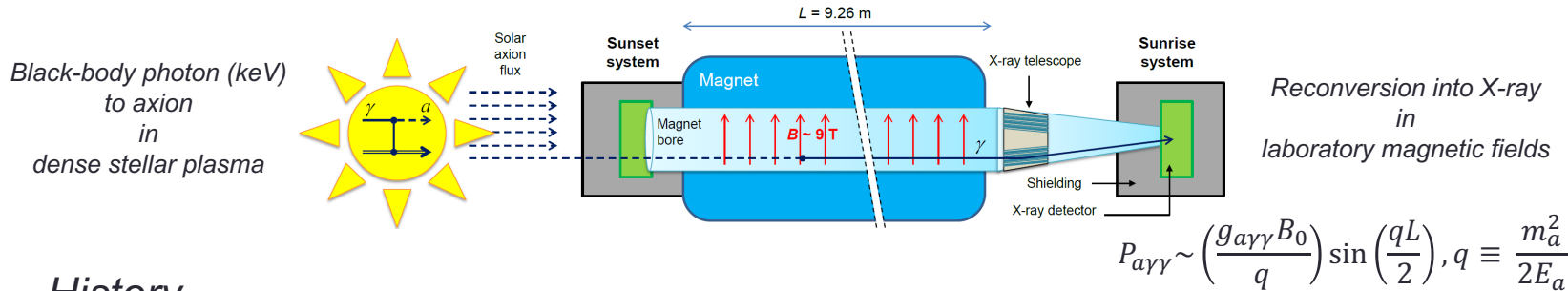
Single photon counting ($T_N \ll T_{SQL}$)



Helioscope



Solar axion telescope



History

- BNL => SUMICO => CAST => IAXO

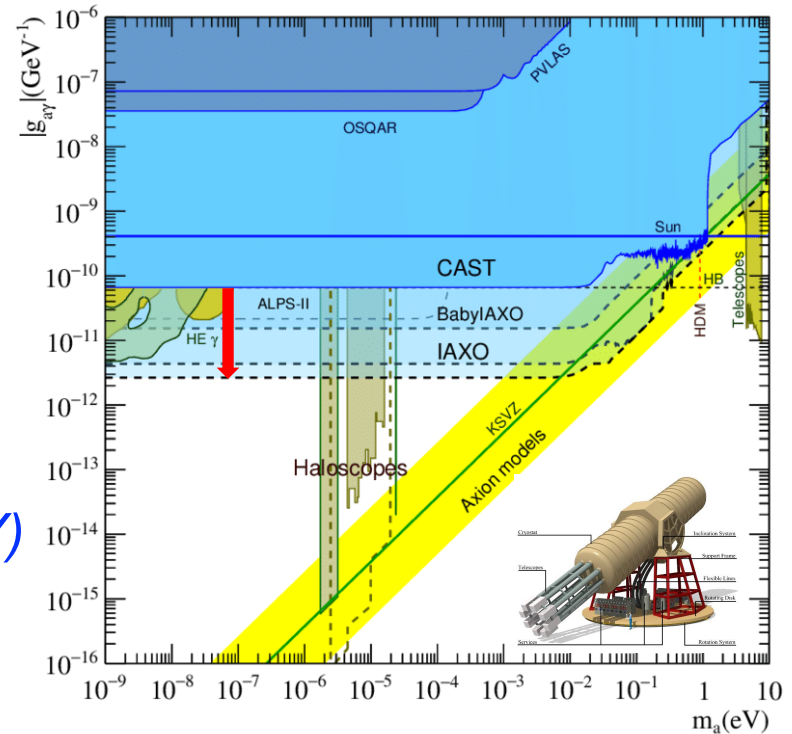
International AXion Observatory

- 8 dipoles (5.4 T, 20 m, 600 mm)
- Goal: $g_{a\gamma} \sim 10^{-12} \text{ GeV}^{-1}$
- Diverse physics over wide range
 - QCD axions
 - ALP miracle (DM & inflation)
 - Astrophysical hints

T. Schiffer (Tue)

Baby-IAXO under construction (DESY)

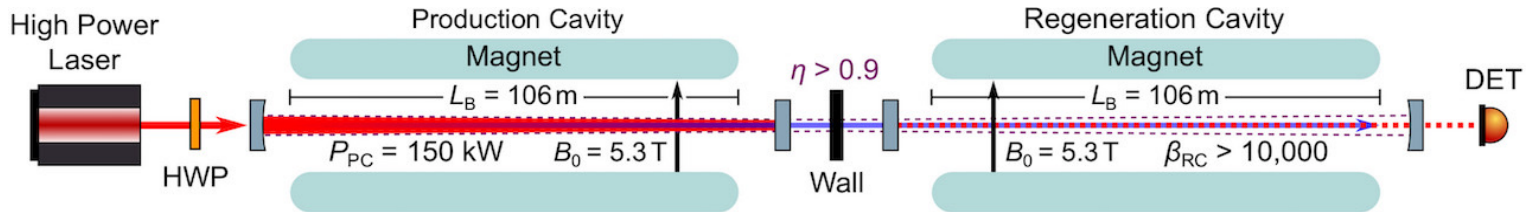
- First step towards full IAXO
- $4 \text{ T} / 10 \text{ m} \Rightarrow 10 \times \text{MFOM}_{\text{CAST}}$





Light shining through a wall

Axion-production/photon-regeneration at the lab

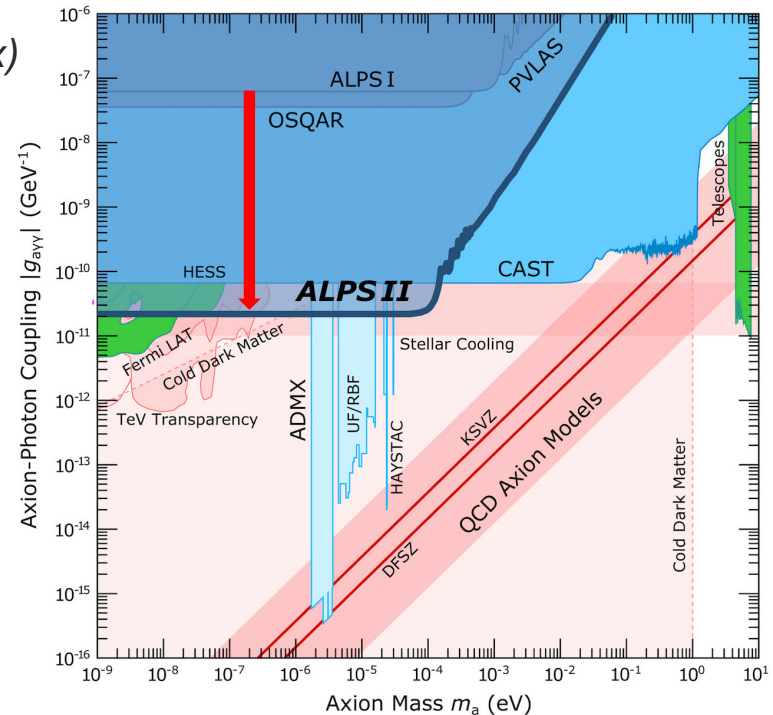


- Model independent search
 - No need of cosmo./astrophys. source
 - cf. helo(helio)scope search needs density(flux)
- History
 - BFRT => OSQAR => ALPS => ALPS II

G. Othman (Tue)

Any Light Particle Search II

- 2 x 12 Hera dipoles (8.8 m & 5.3 T)
- Dual high-finesse optical cavities
- Alignment sensing: PDH/DWS
- Photon detector: HET/TES
- Goal: $g_{a\gamma\gamma} \sim 10^{-11} \text{ GeV}^{-1}$ below 0.1 eV
 - Three orders of magnitude of ALPS





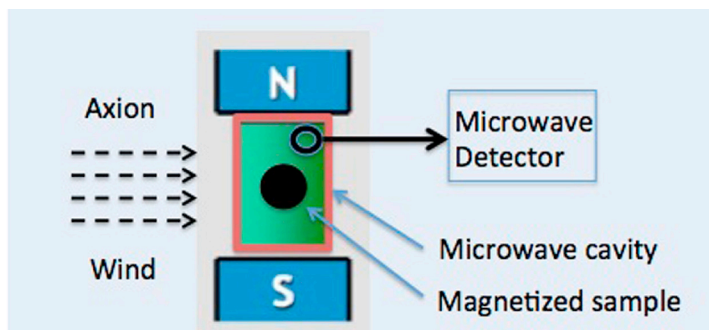
Fermion coupling



• QUAX-ae

- Ferromagnetic haloscope
 - *Axion-electron spin*
- Photon-magnon system
 - Series of YIG spheres
 - TM_{110} of a cylindrical cavity
- Upgraded with JPA
- Best limit near $m_a \sim 43 \mu\text{eV}$

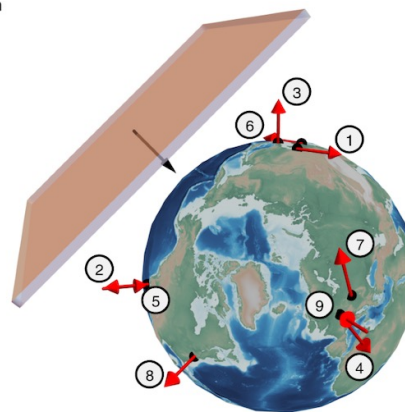
PRL 124 171801 (2020)



• GNOME

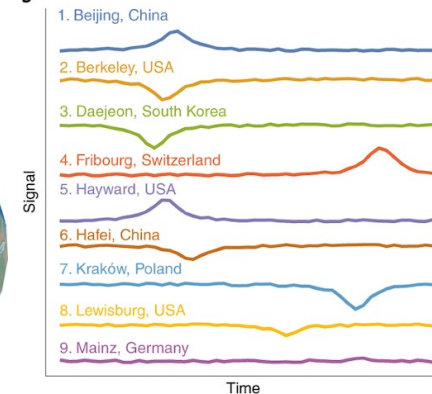
- Topological defect dark matter
 - *Axion-nucleon spin*
- Global network of optical magnetometers
 - Correlated signal

a



Nature Phys. 17 1396 (2021)

b



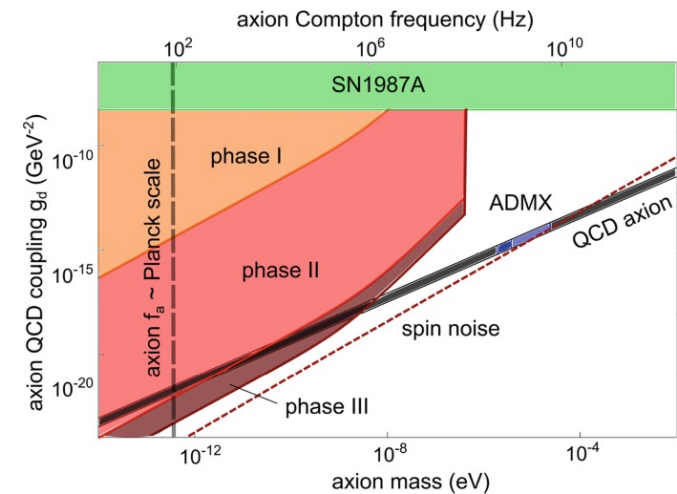
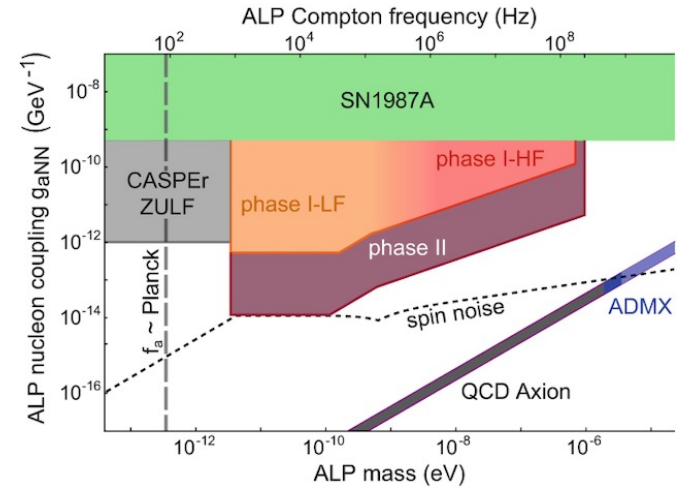
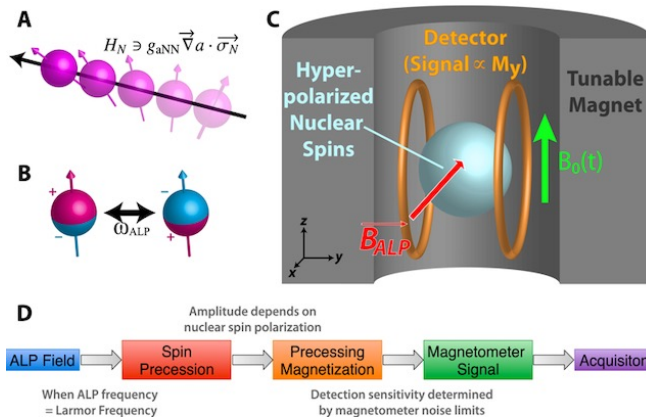


Fermion/EDM coupling

• Cosmic Axion Spin Precession Experiment

H. Bekker, Thu

- **CASPER-wind(gradient)**
 - Axion-nuclear spin (axion wind)
- **CASPER-electric**
 - Axion-nEDM (electric field)
 - Probe for $aG\tilde{G}$ (QCD axion?)
- Nuclear magnetic resonance
 - $\omega_L(\propto B_0) = m_a$
 - Highly sensitive to $m_a < 10^{-8}$ eV





CPV and EDM



CPV in QCD

- $\theta_{eff} \equiv \theta_{QCD} - \frac{a}{f_a} = 0$
- QCD axion $\Rightarrow EDM = 0$

vs.

Additional CPV in nature

- $\theta_{eff} \equiv \theta_{QCD} - \frac{a}{f_a} + \dots \neq 0$
- $EDM \neq 0 \Rightarrow$ non-QCD axion field?

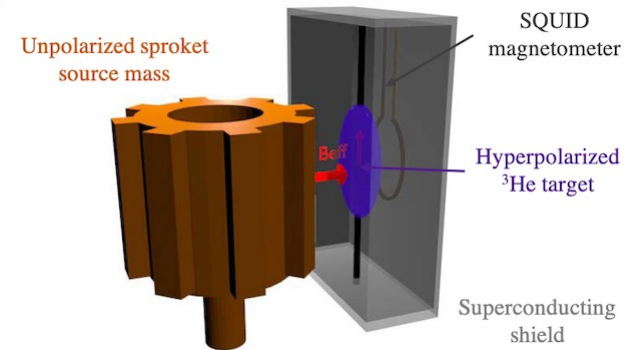
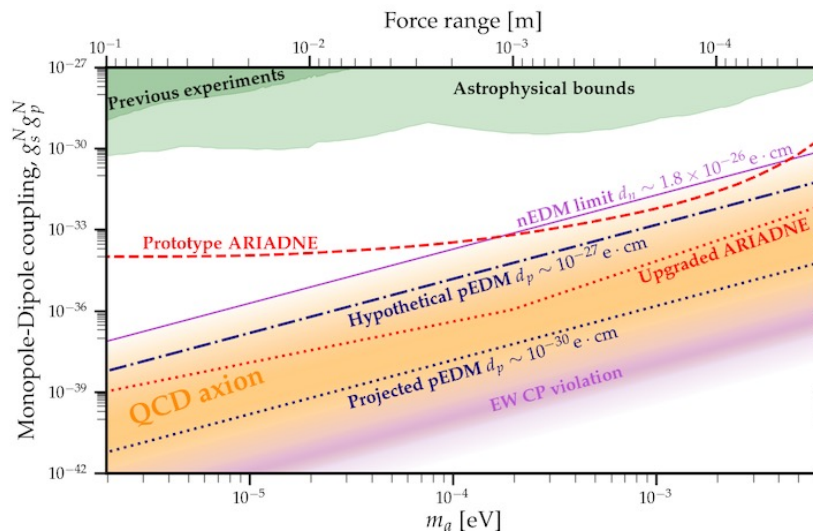
Storage ring proton beam

- Direct probe for pEDM
- Sensitivity improvement
 - $pEDM \sim 10^{-29} e \cdot cm$

ARIADNE

A. Geraci (Thu)

- Nucleon-nucleon interactions
 - 5th force mediator
- No cosmological assumptions



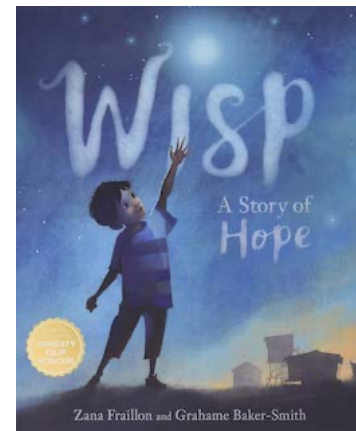
Decisive exclusion

- Negative results from the two independent experiments
- $0.1 \text{ meV} < m_a < 10 \text{ meV}$



Outline

- *Background*
 - *Dark matter and WISP*
 - *Search strategies*
- *Axion (ALP) searches*
 - *Photon coupling*
 - *Haloscope / Helioscope / LSW*
 - *Fermion coupling*
 - *nEDM coupling*
- *Other WISP searches*
 - *Hidden photon*
 - *Chameleon*
- *Prospects and summary*



Disclaimer

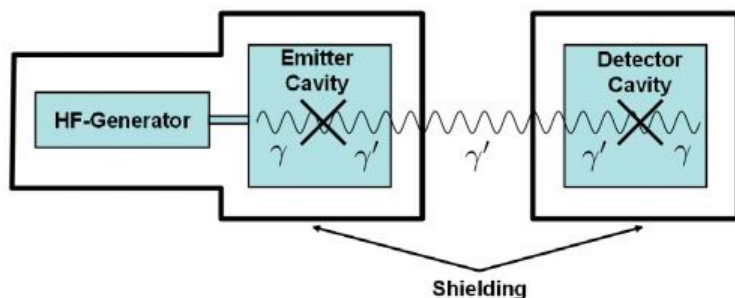
- *My apologies for not covering*
 - *indirect searches*
 - *all individual efforts*
 - *most recent updates*



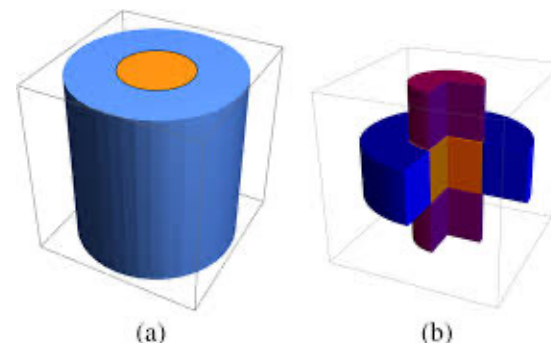
Hidden photon



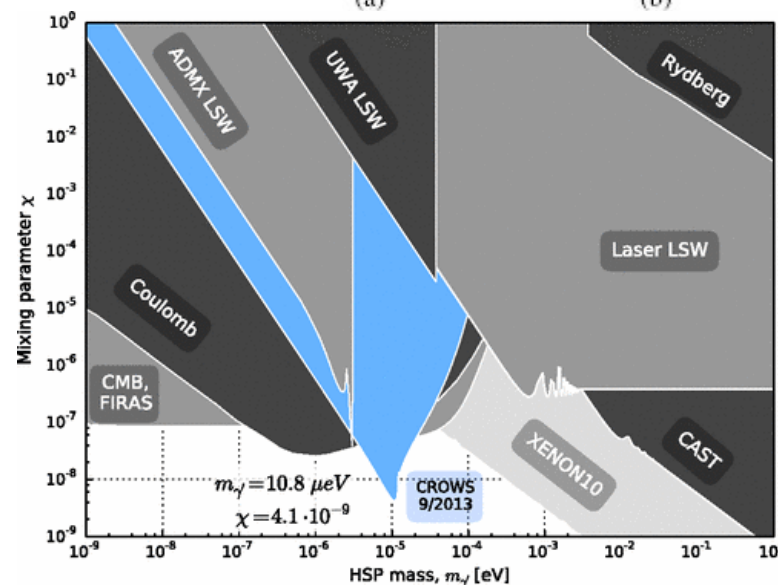
- **Vector field**
 - Gauge boson in hidden sector + dark matter
 - Kinetic mixing with SM gauge fields
- **LSW-type w/ no magnets** *PLB 659 509 (2008)*



PRD 103 055004 (2021)



- **Cavity searches**
 - ADMX *PRL 105 171801 (2010)*
 - UWA *PRD 82 052003 (2010)*
 - CROW *PRD 88 075014 (2013)*
 - SC cavities and ...
 - New ideas *R. Cervantes, O. Tajima (Thu)*





Chameleon



- **Scalar field**

- Dark energy candidate *PRD 69 044026 (2004)*

- **Effective potential dependent on ambient matter / EM field**

$$V_{eff}(\phi, \vec{r}) = \Lambda^4 e^{\frac{\Lambda^n}{\phi^n}} + e^{\frac{\beta_m}{M_{Pl}}\phi} \rho_m(\vec{r}) + e^{\frac{\beta_\gamma}{M_{Pl}}\phi} \rho_\gamma(\vec{r})$$

Matter coupling

Photon coupling

Phys. Dark Univ. 26 100367 (2019)

- **Searches**

- **Afterglow effect**

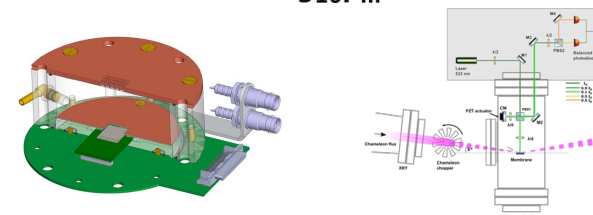
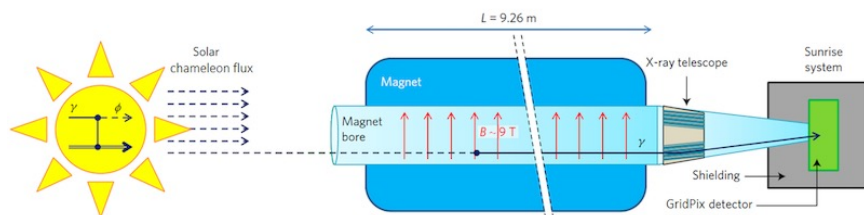
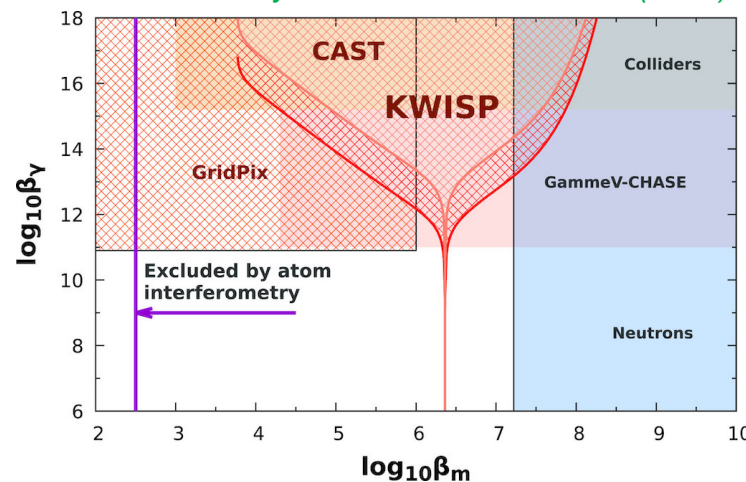
- GammeV, CHASE, ADMX, ...

- **Solar chameleon**

- Similar to solar axion search

- CAST ($\phi - \gamma$) and KWISP ($\phi - m$)

- $P_{\phi \rightarrow \gamma, m} \propto \beta_{\gamma, m}^2 B^2 L^2$





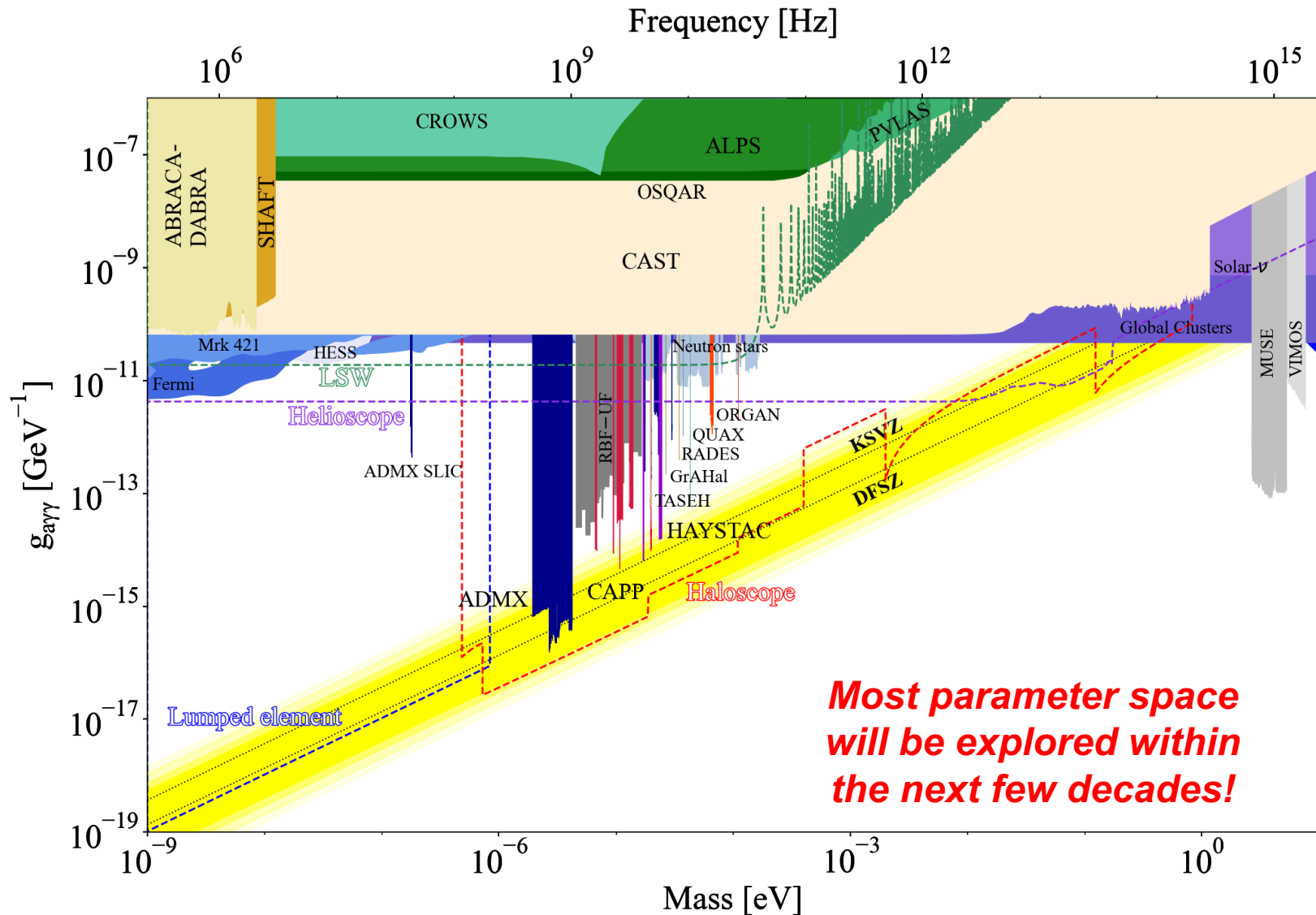
New ideas



- *The piezoaxionic effect* (A. Madden, Mon)
- *WISPF1: Searching for ALPs-photon conversion on fiber interferometer* (J. M. B. Berenguer, Tue)
- *T-RAX: Transvers Resonant Axion eXperiment* (C. Lee, Thu)
- *New axion dark matter search technologies* (M. Tobar, Thu)
- *Searching for wavelike dark matter with SRF cavities* (R. Cervantes, Thu)
- *Radiometric broadband searches for light dark matter with BRASS-p* (L. H. Nguyen, Fri)
- *WISPLC: Search for Dark Matter with LC Circuit* (Z. Zhang et al. Mon)



Axion searches – future





Conclusion



- *Axions and WISPs are leading candidates for dark matter*
 - *QCD axions solves the strong CP problem*
- *Theoretically well motivated but experimentally challenging*
 - *Weak coupling and unknown mass*
- *Tremendous search efforts*
 - *Different technologies targeting at different mass ranges*
- *Axion community is getting larger*
 - *New results, new groups and new ideas (during the workshop)*
- *Next a few decades must be critical/exciting*
 - *Covering a substantial portion of the parameter space*
 - *Uncovering the nature of dark matter*

