

What Have We Learnt From LHC About Air Showers?

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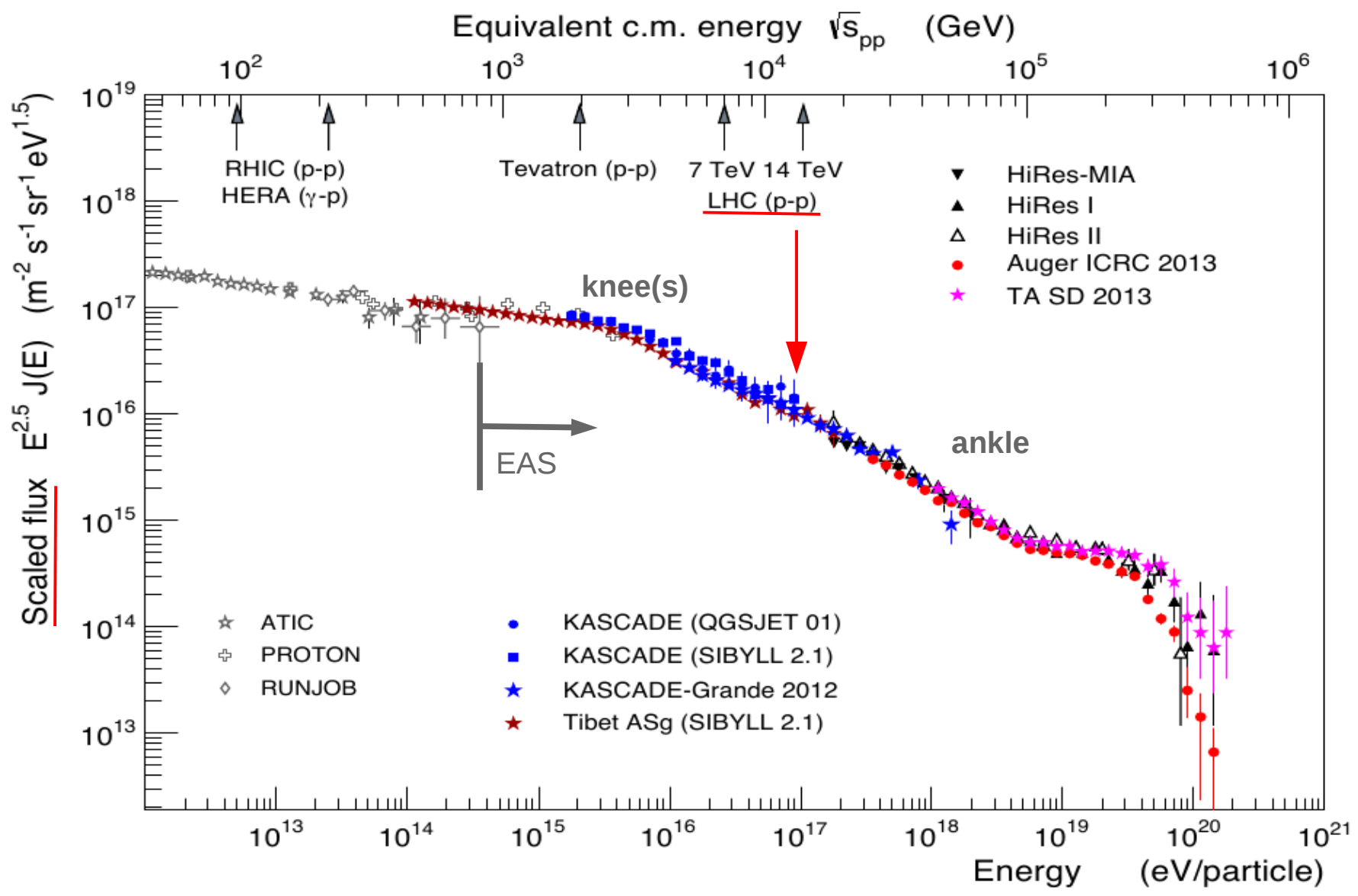
December the 12th 2016

Outline

- Monte-carlo for Cosmic Ray (CR) analysis
- Monte-carlo (MC) comparison to LHC data
- Electromagnetic (EM) signal in extended air showers
 - ➔ X_{\max}
- Muon signal
 - ➔ X_{\max}^{μ}

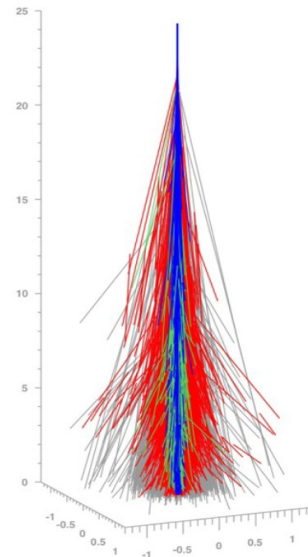
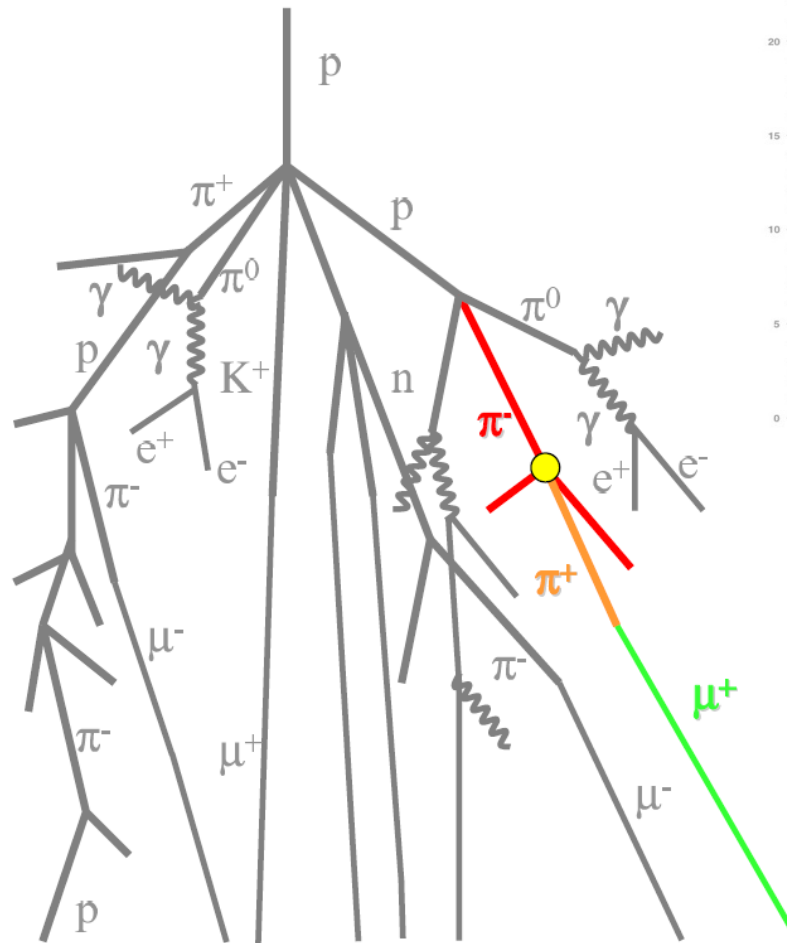
LHC data reduced the model uncertainties and **exclude exotic models** for CR spectrum. **Indirectly** the measurement of diffraction at LHC for protons used in air showers gives information on **pion interactions**.

Cosmic Ray Spectrum



R. Engel (KIT)

Extensive Air Shower



$A + air \rightarrow$ hadrons

$p + air \rightarrow$ hadrons

$\pi + air \rightarrow$ hadrons

initial γ from π^0 decay

$e^\pm \rightarrow e^\pm + \gamma$

$\gamma \rightarrow e^+ + e^-$

$\pi^\pm \rightarrow \mu^\pm + \nu_\mu / \bar{\nu}_\mu$

main source of uncertainties

well known

● Cascade of particle in Earth's atmosphere

Number of particles at maximum

➔ 99,88% of electromagnetic (e/m) particles

➔ 0.1% of muons

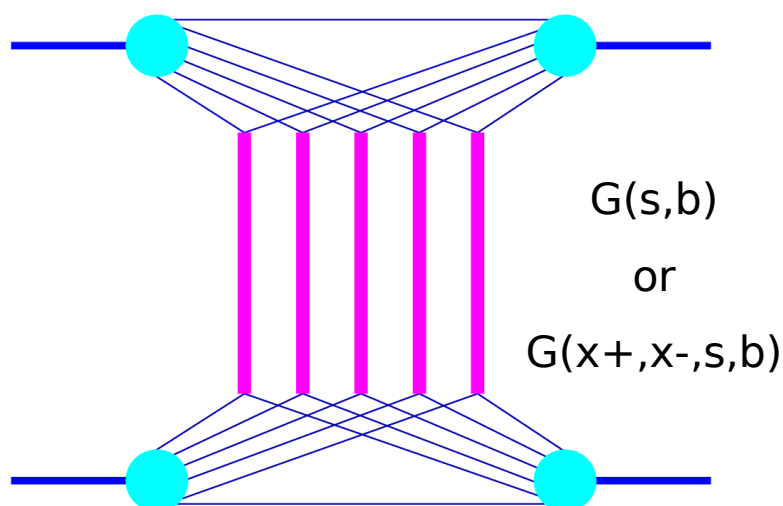
➔ 0.02% hadrons

Energy

➔ from 100% hadronic to 90% in e/m + 10% in muons at ground (vertical)

From R. Ulrich (KIT)

Cross Section and Multiplicity in Models



● Gribov-Regge and optical theorem

- ➔ Basis of all models (multiple scattering) but
 - ◆ Classical approach for QGSJET, SIBYLL and DPMJET (no energy conservation for cross section calculation)
 - ◆ Parton based Gribov-Regge theory for EPOS (**energy conservation at amplitude level**)

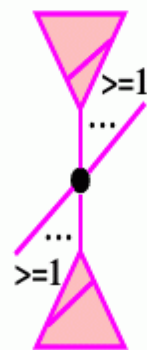
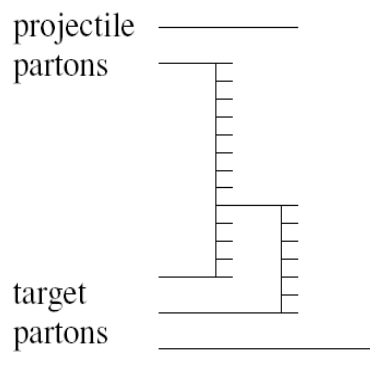
● pQCD

- ➔ Minijets with cutoff in SIBYLL and DPMJET
- ➔ Same hard Pomeron (DGLAP convoluted with soft part : no cutoff) in QGSJET and EPOS but
 - ◆ Generalized enhanced diagram in QGSJET-II
 - ◆ Simplified non linear effect in EPOS

- Phenomenological approach

EPOS

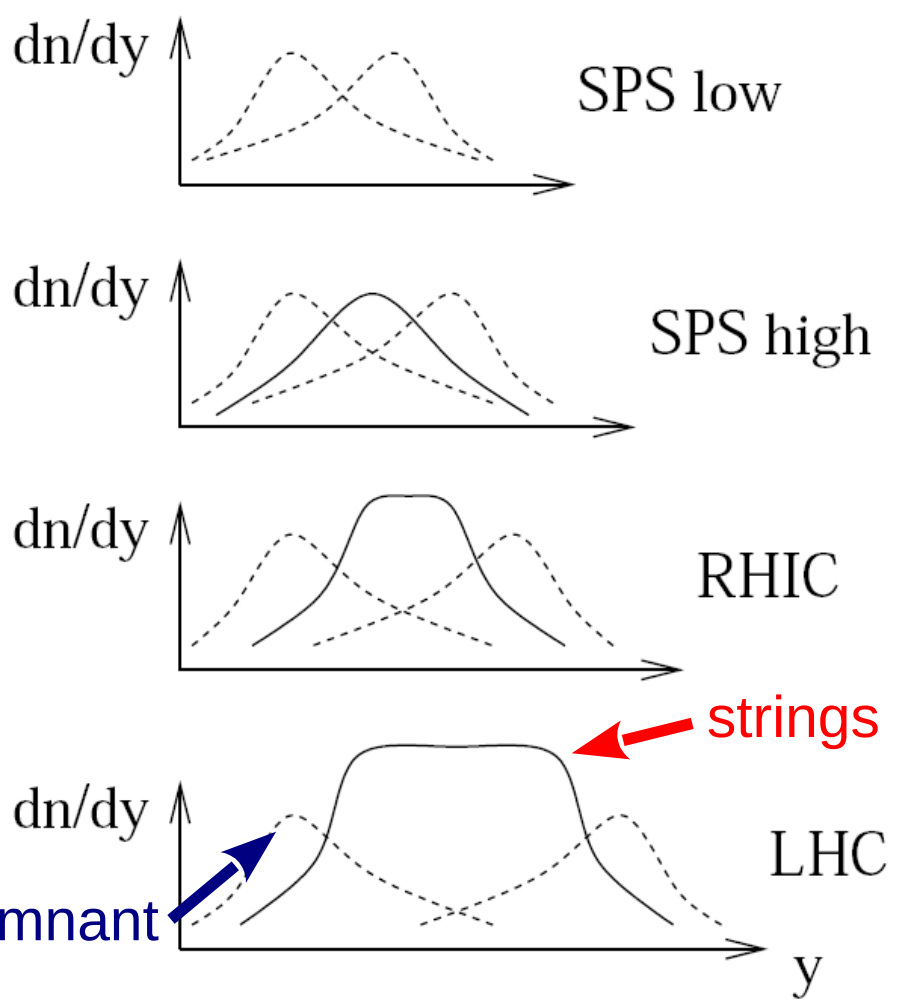
QGSJET II



Remnants

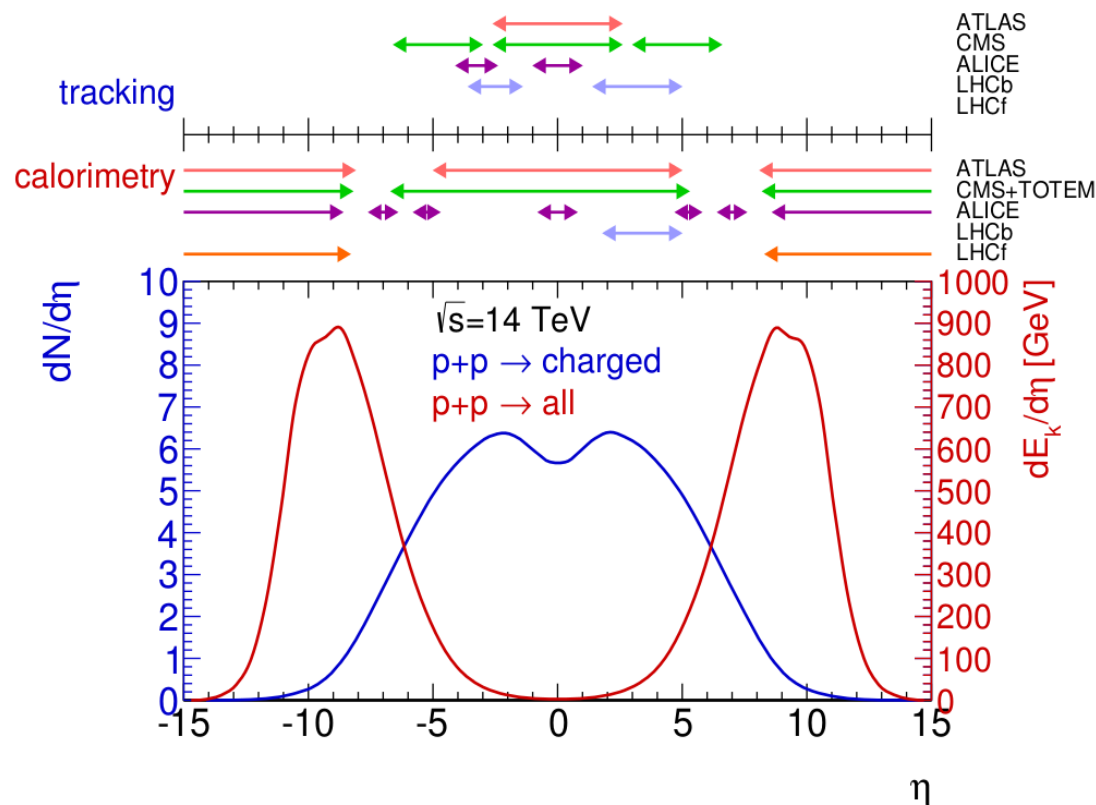
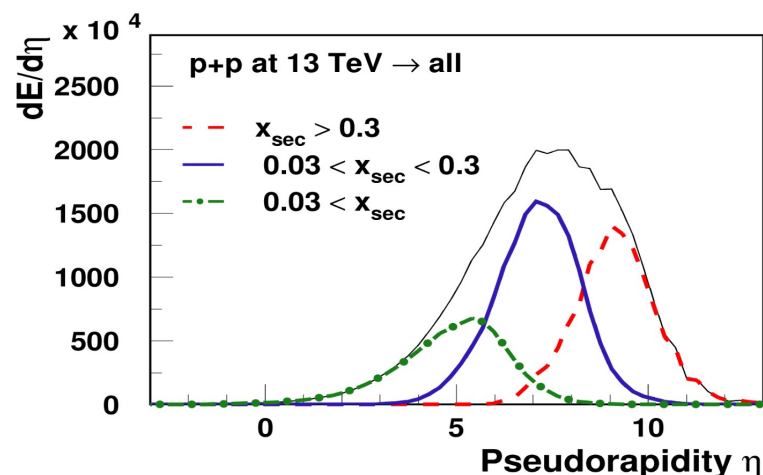
Forward particles mainly from projectile remnant

The (in)elasticity is closely related to diffraction and forward spectra



- ➔ SIBYLL
 - ◆ No remnant except for diffraction
 - Leading particle from string ends
- ➔ QGSJET
 - ◆ Low mass remnants
 - Leading particle similar to proj.
- ➔ EPOS
 - ◆ Low and high mass remnants
 - Any type of leading particle
 - from resonance
 - from string
 - from statistical decay

LHC acceptance

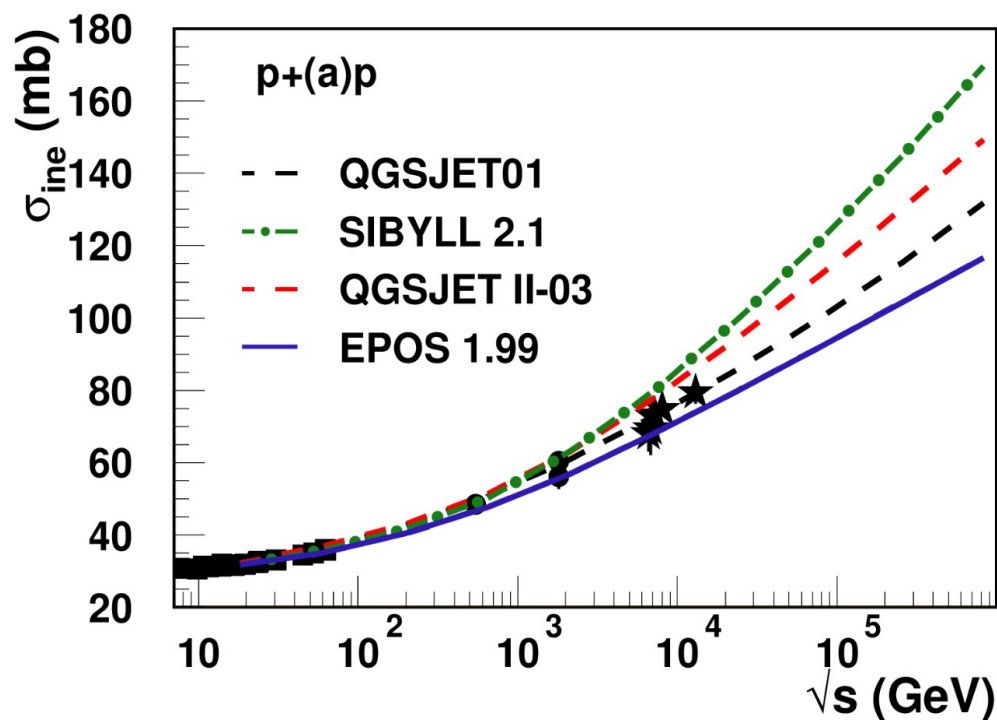


- p-p data of central detectors used to update the models
 - ➔ p-Pb difficult to compare to CR models (only EPOS)
 - ➔ special centrality selection
 - ➔ mostly new tuning of models
- Direct photon energy spectra from LHCf
 - ➔ small phase space but relevant for X_{max}
 - ➔ not yet taken into account into models
- Average elasticity/inelasticity (energy fraction of the leading particle)
 - ➔ all diffraction measurement to be taken into account

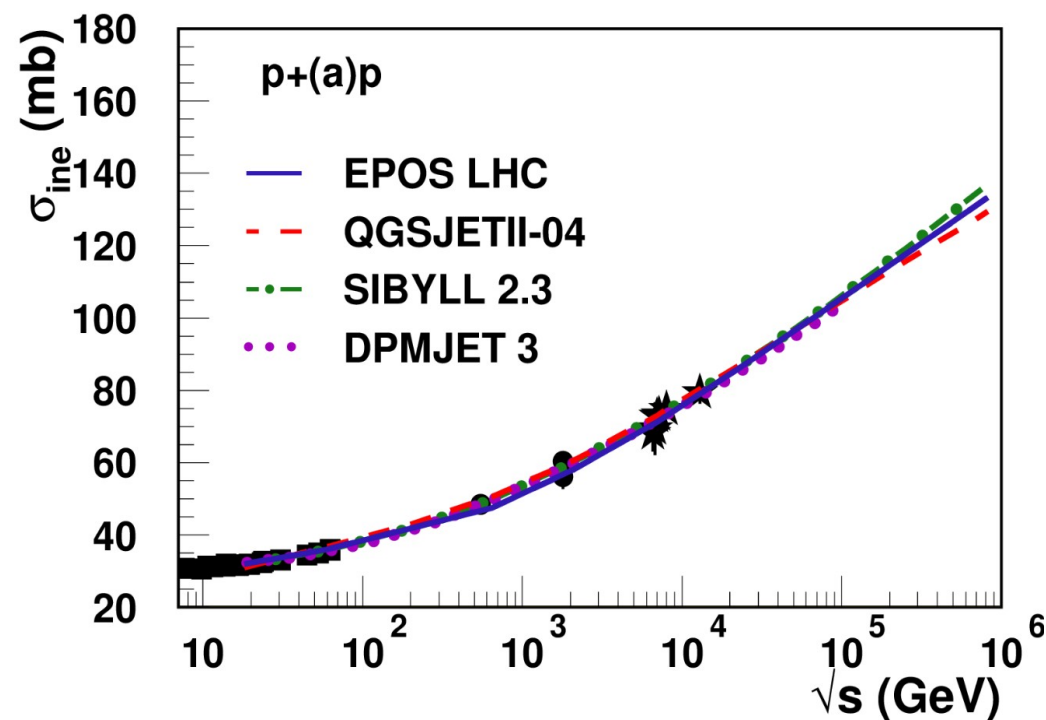
Cross Sections

- ➔ same cross section prediction at pp level and low energy (data for tuning)
- ➔ extrapolation to high energy looks settled
 - ◆ different amplitude and scheme
 - ➔ same extrapolations

Pre - LHC



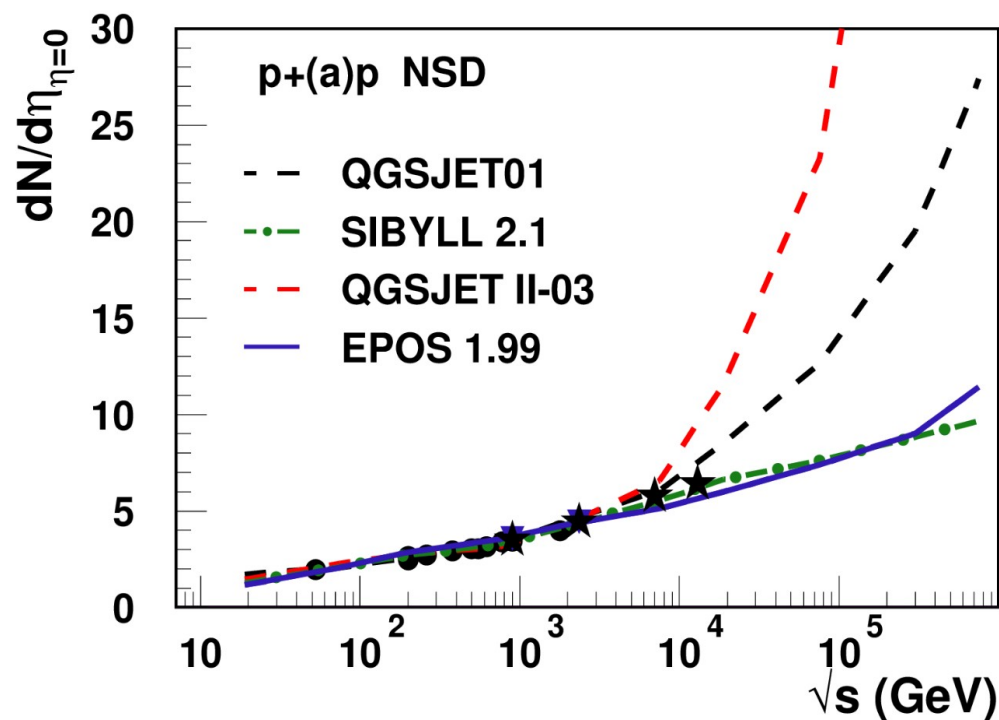
Post - LHC



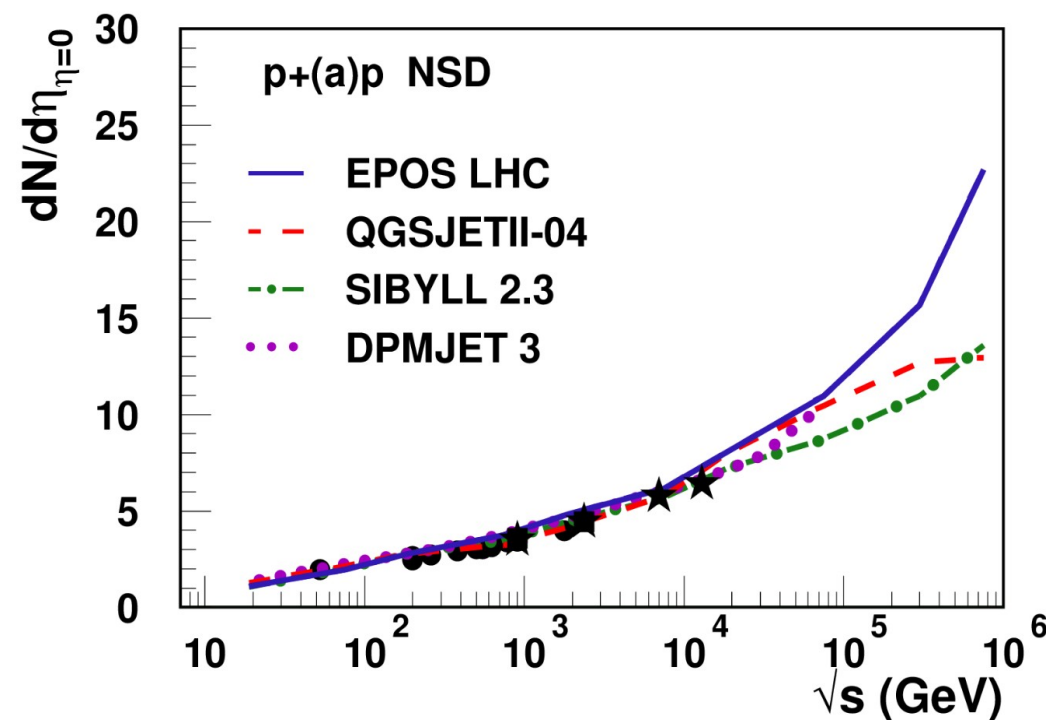
Multiplicity at mid-rapidity

- Multiplicity fixed by data up to 13 TeV
- extrapolation to high energy less model dependent after LHC
- QGSJET01 and QGSJETII-03 extrapolation excluded

Pre - LHC



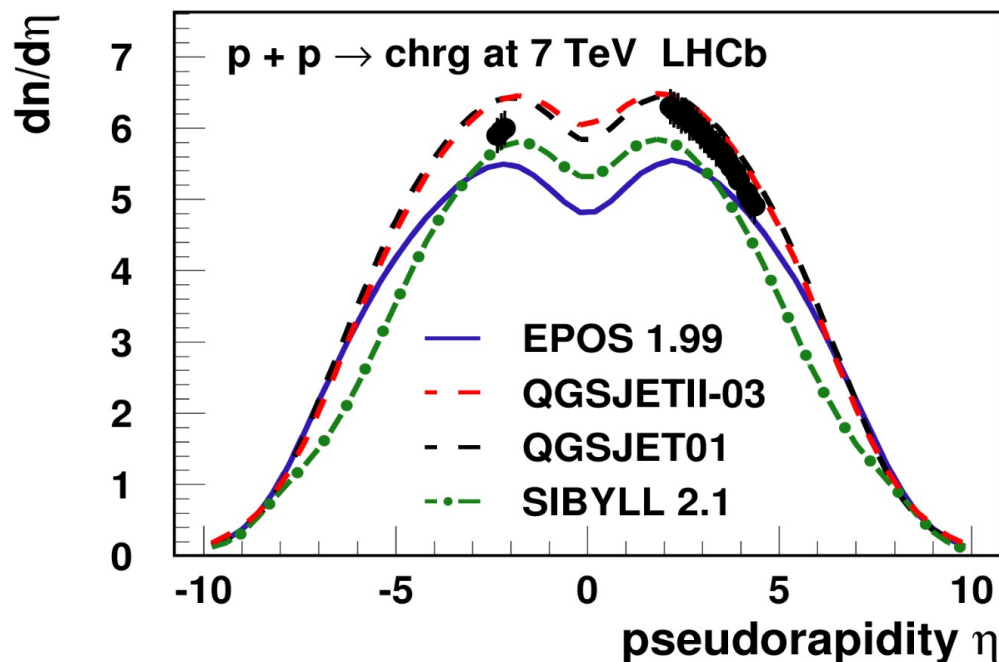
Post - LHC



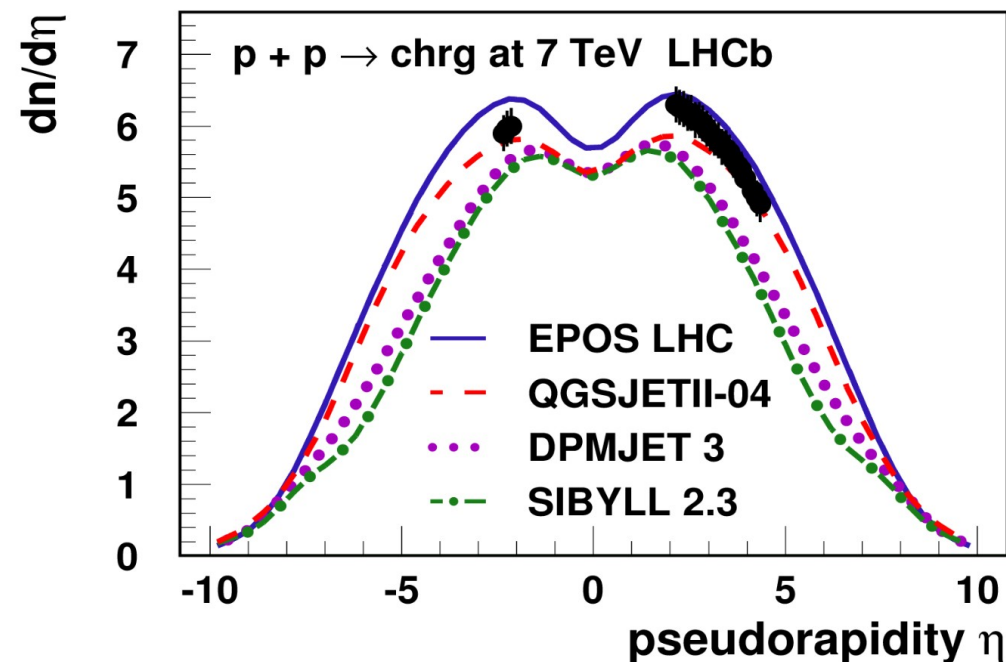
Pseudorapidity

- ➔ The width of the pseudorapidity distributions introduce a difference between mid-rapidity and full multiplicity.
- ➔ From LHC data
 - DPMJET 3 and SIBYLL 2.3 too narrow
 - QGSJETII-04 ~ OK
 - EPOS LHC a bit too large

Pre - LHC



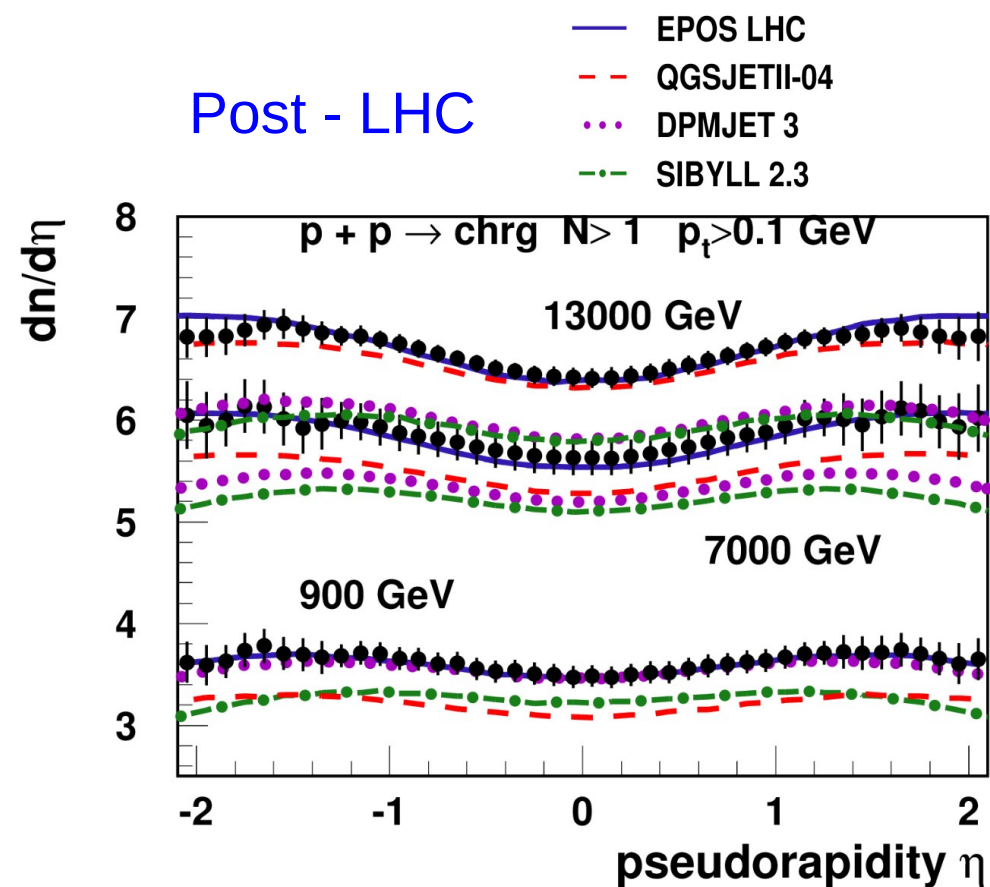
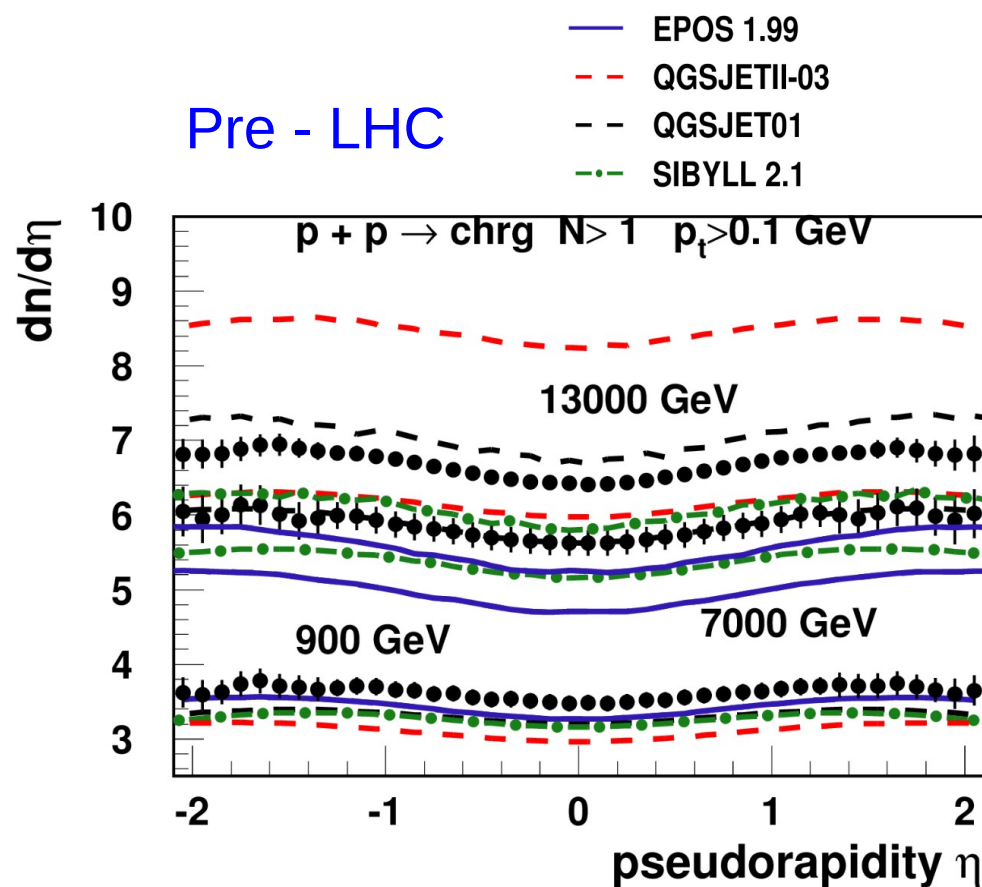
Post - LHC



Test of Models vs Accelerator Data

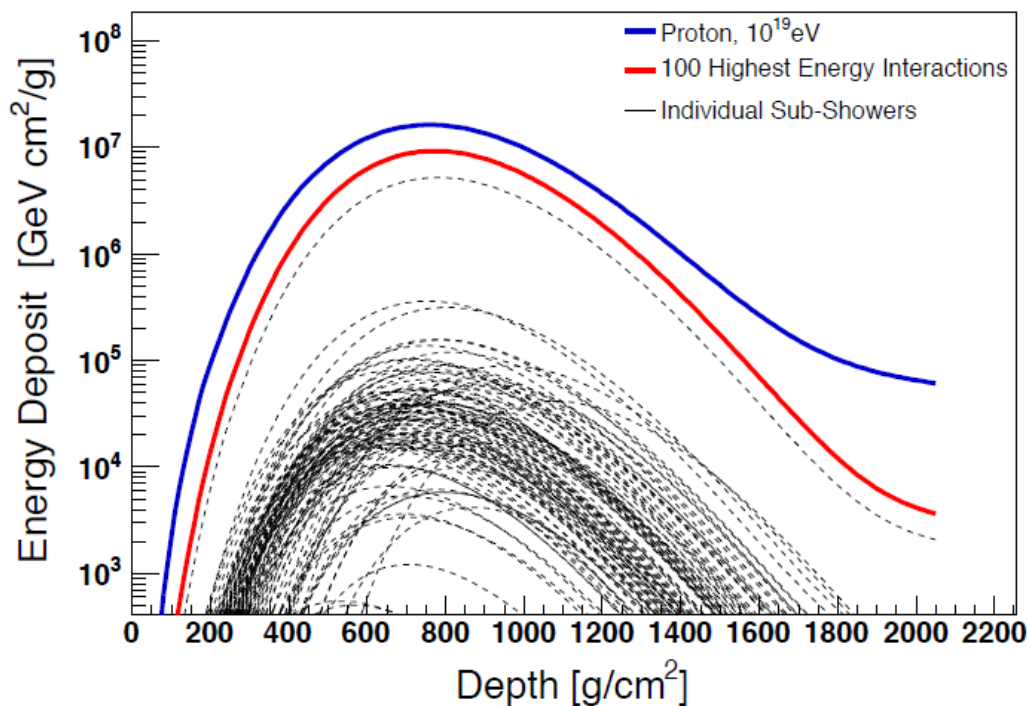
→ From LHC data

- All pre-LHC models extrapolation excluded
- DPMJET 3 and SIBYLL 2.3 underestimate multiplicity
- QGSJETII-04 and EPOS LHC ~ OK (and similar to Pythia 8)



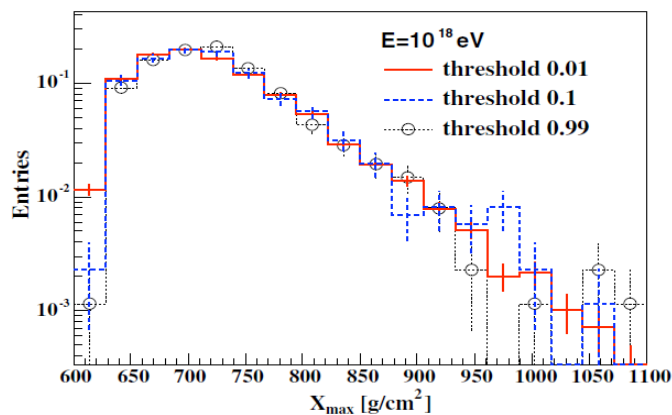
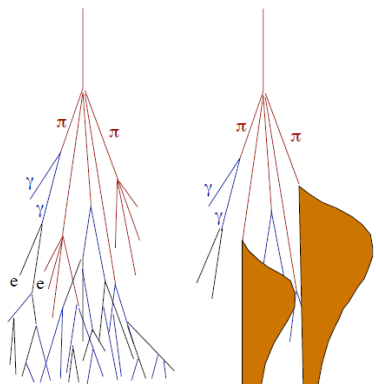
Hadronic Interactions for X_{\max}

Electrons



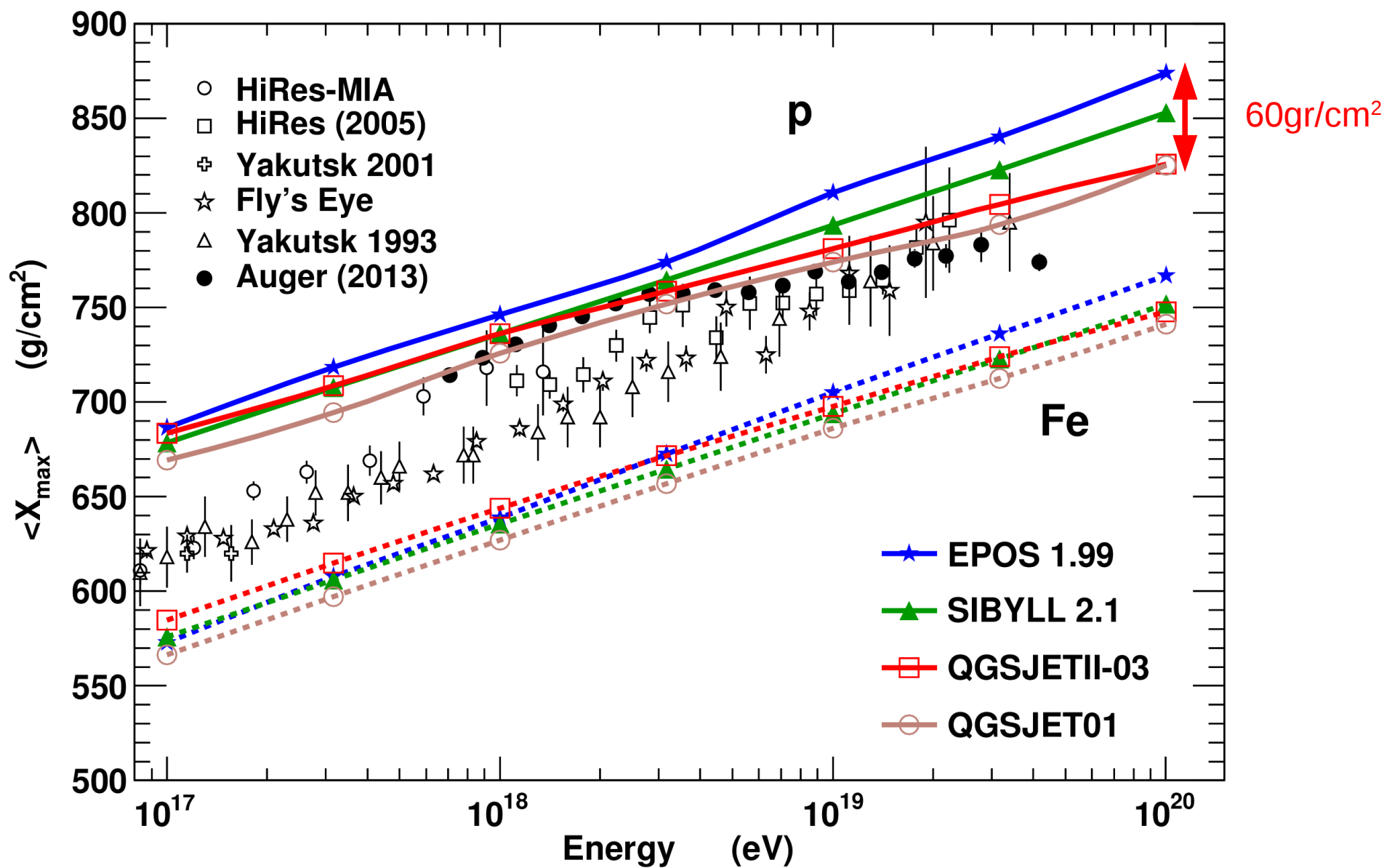
Shower particles produced in 100 interactions of highest energy

X_{\max} dominated by first (high energy) interaction(s) : proton (nucleus) - Air

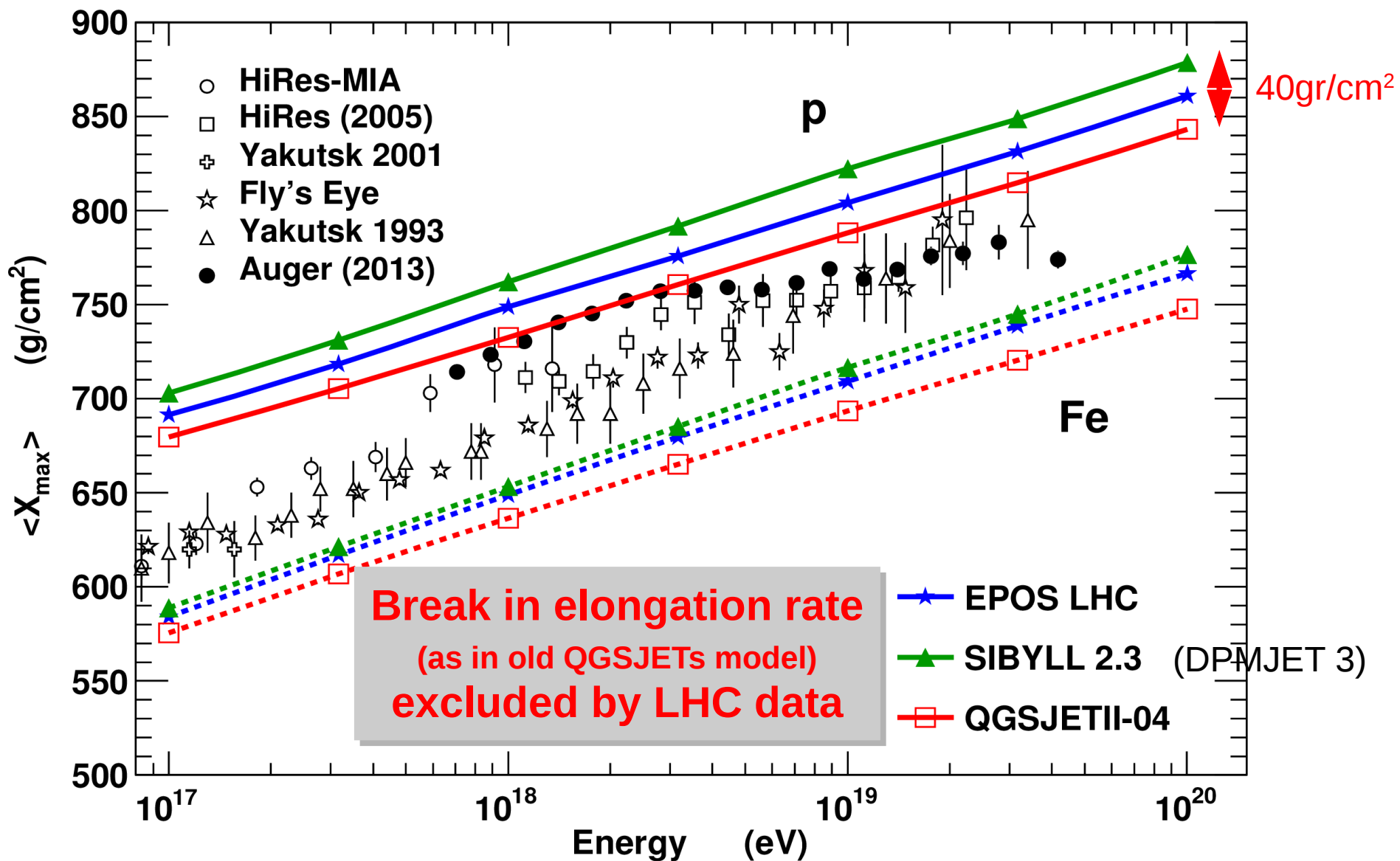


Fluctuations mainly coming from the first hadronic interaction.

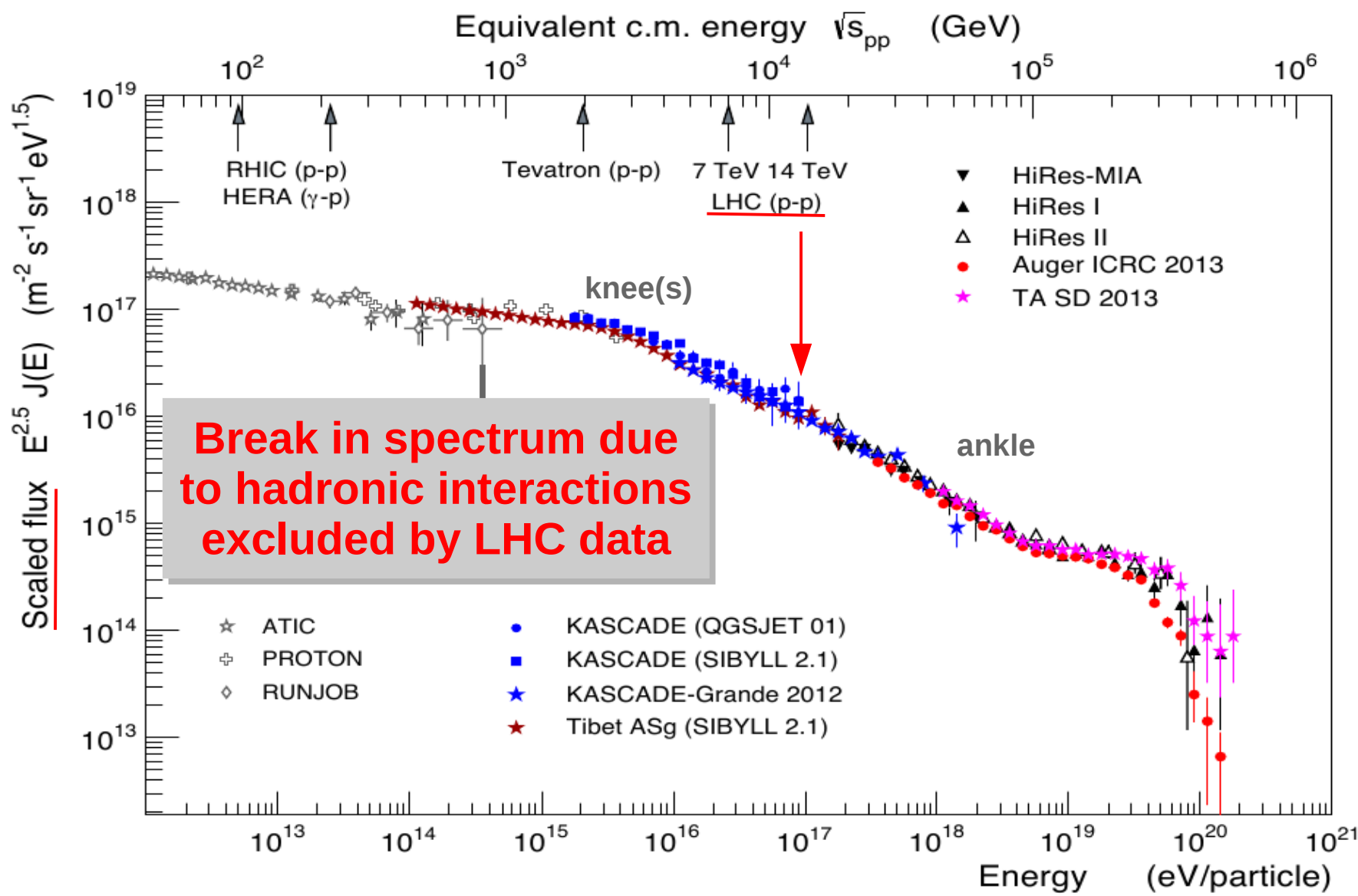
EAS with Old CR Models : X_{\max}



EAS with Re-tuned CR Models : X_{\max}

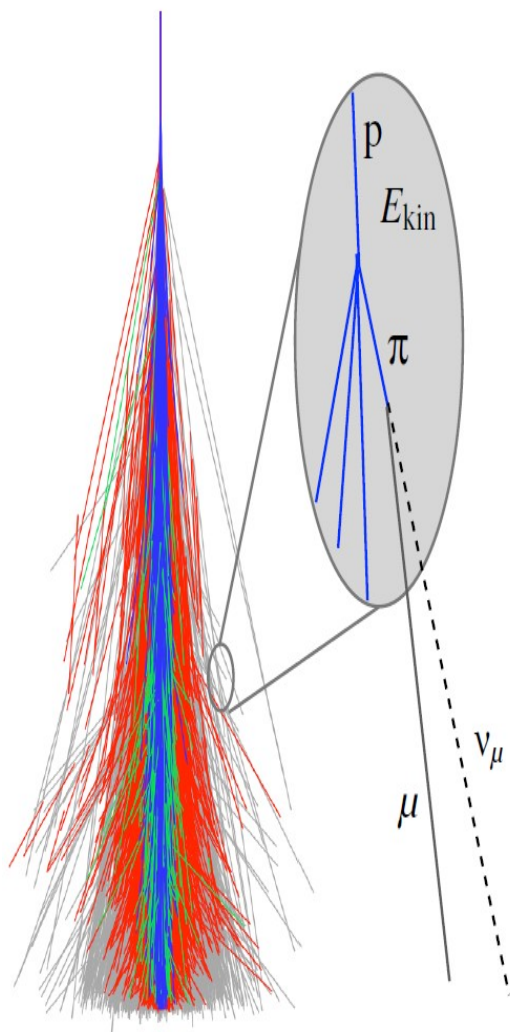


Cosmic Ray Spectrum

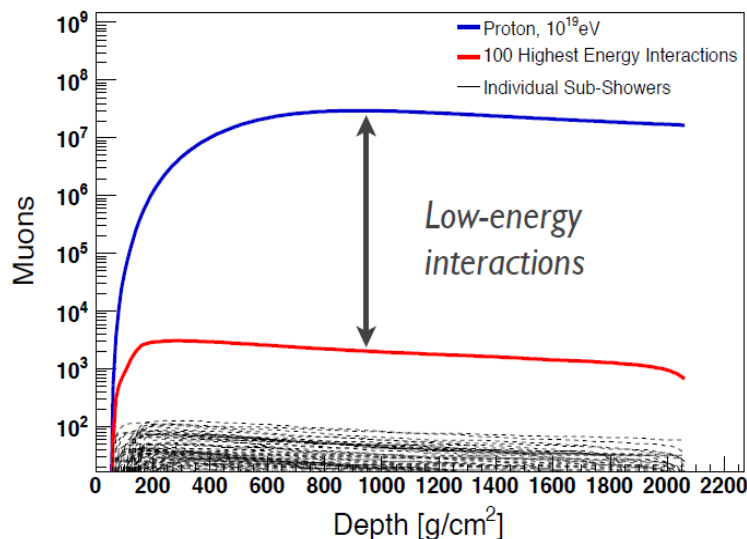


R. Engel (KIT)

Muon production by low energy interactions



Muons

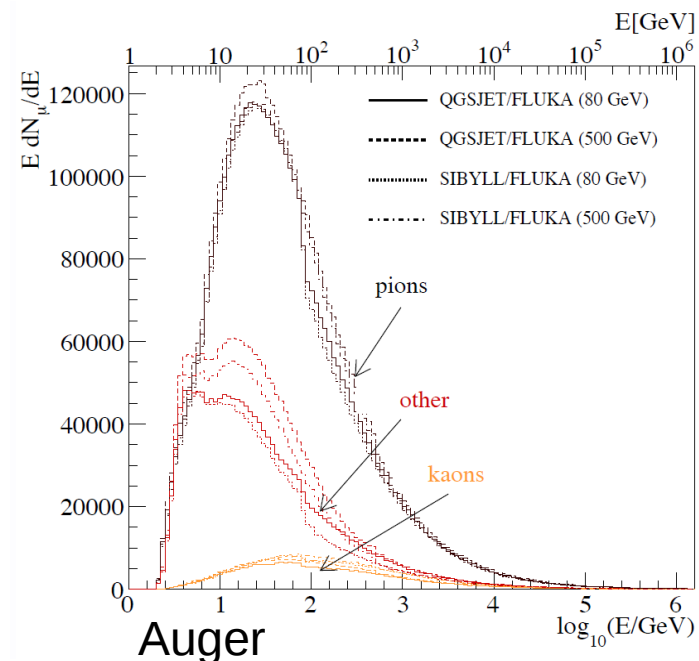
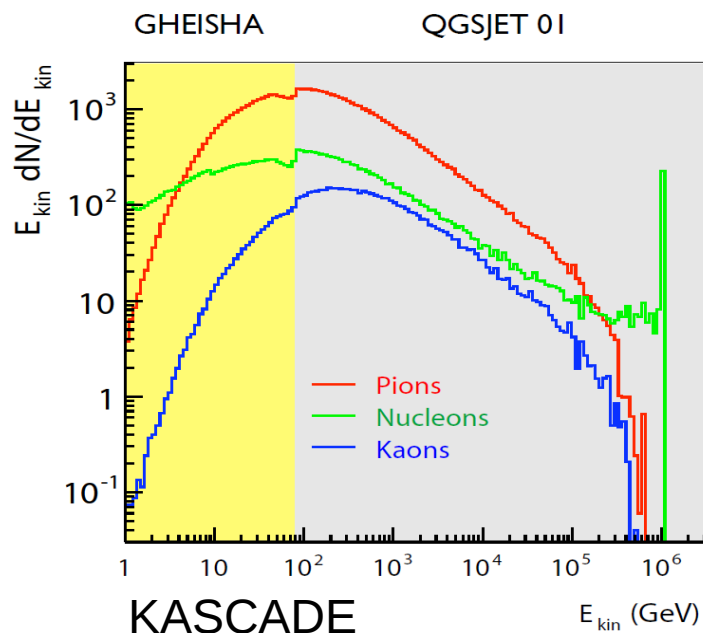


Energy of Last generation:

~ 100 GeV for KASCADE

~ 30 GeV for Auger

N_{μ} generated by all (low energy) interactions : pion-Air

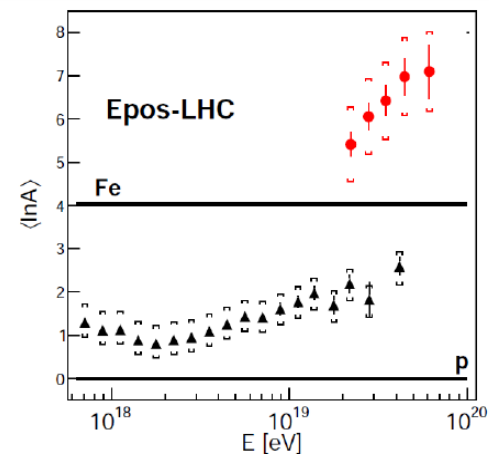
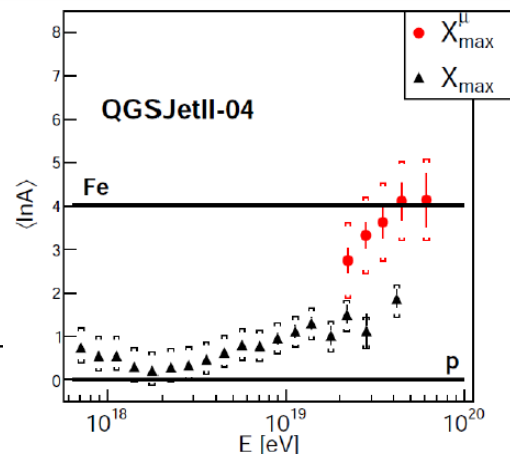
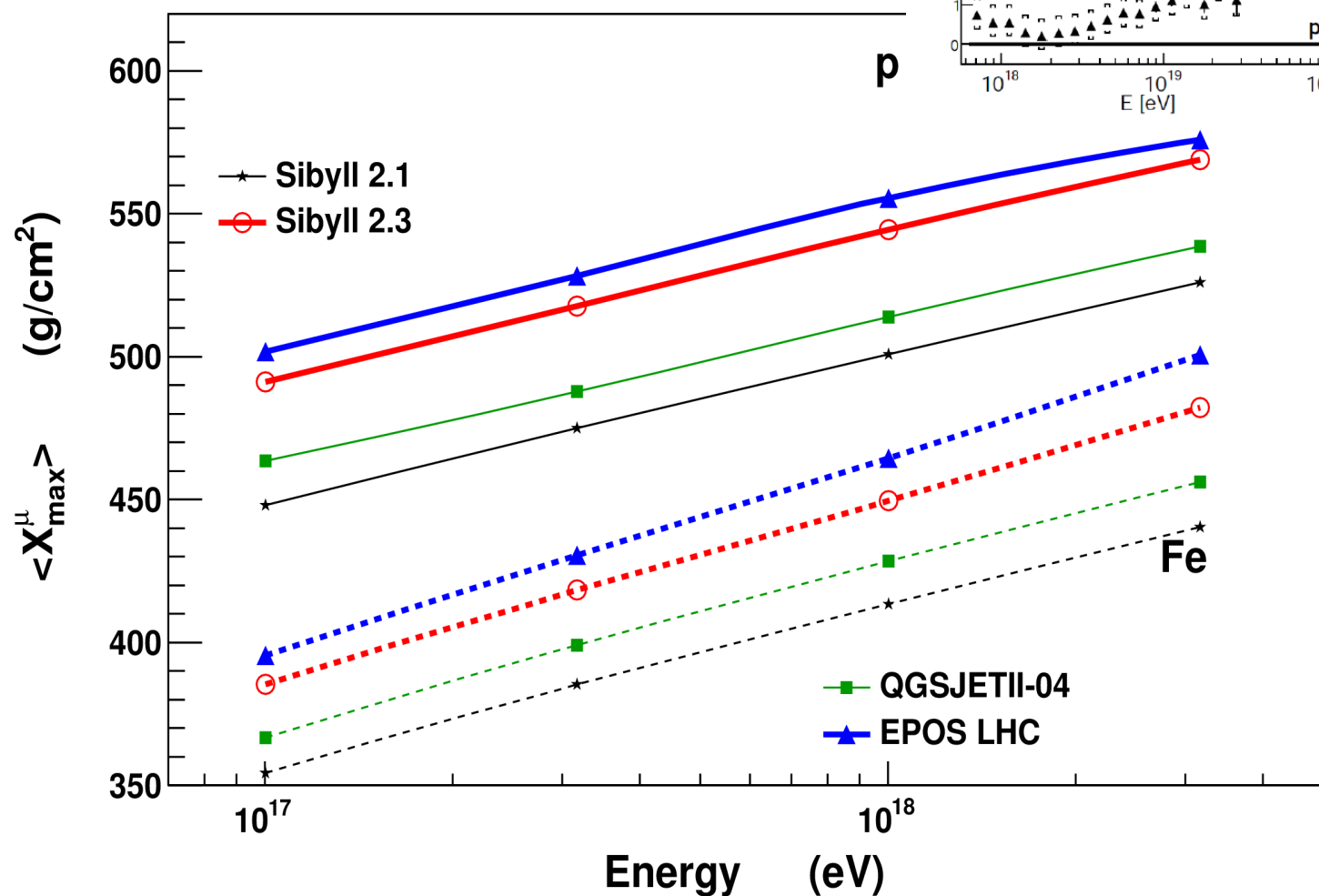


Muon Production Depth

Same for EPOS LHC and SIBYLL 2.3

➔ deeper X_{\max}^{μ}

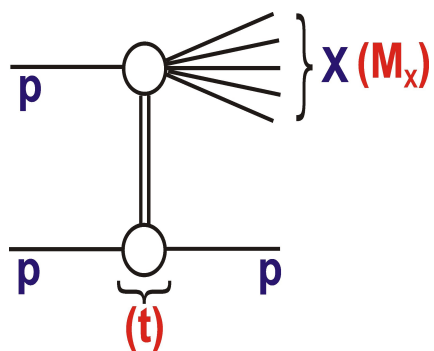
➔ link to rapidity gap measurement



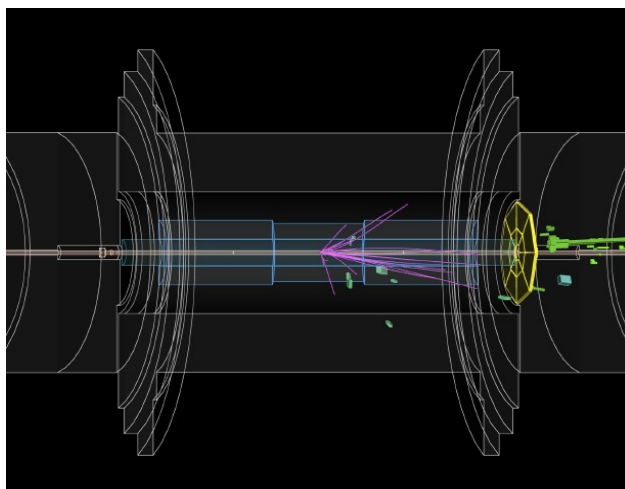
MPD changed in EPOS and SIBYLL after LHC data

Rapidity Gap

diffractive process

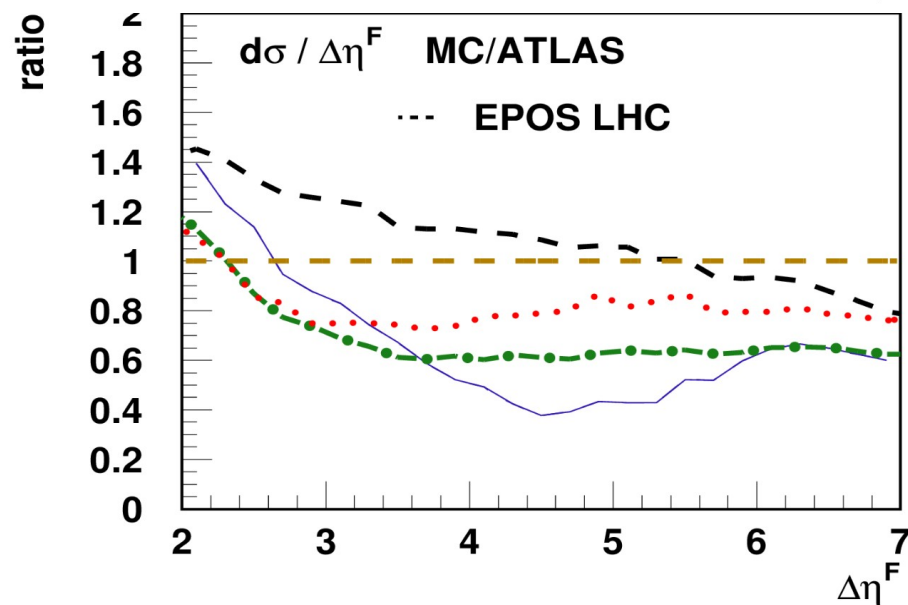
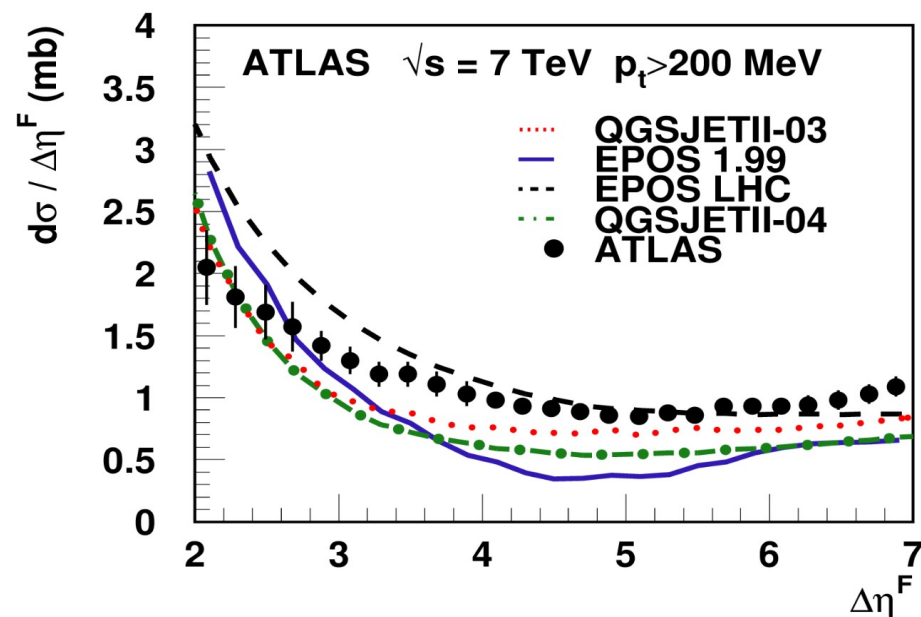


ATLAS detector



ATLAS Collaboration

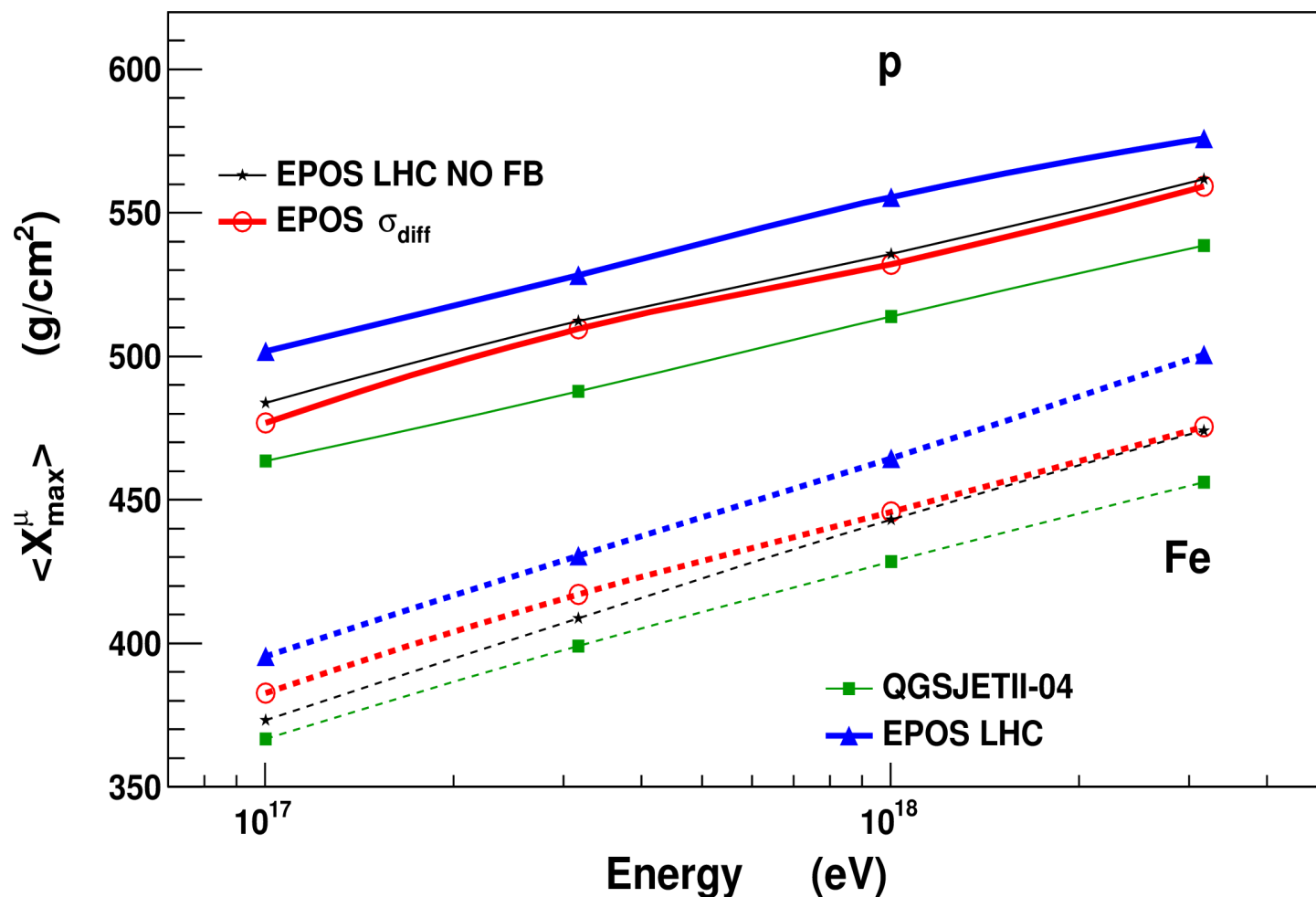
- **Rapidity gap closely related to diffraction**
 - ➔ diffractive cross-section
 - ➔ AND diffractive mass distribution
- **Hard constraint for CR**
 - ➔ high elasticity for proton seen at LHC not compatible with low value needed in air showers : pion inter.



$\langle X_{\max}^{\mu} \rangle$ with modified EPOS LHC

EPOS LHC without forward baryons or more inelastic pion int.

- softer meson spectra (lower elasticity) : smaller X_{\max}^{μ}
- less forward baryons: smaller X_{\max}^{μ}



-25 g/cm² for diff

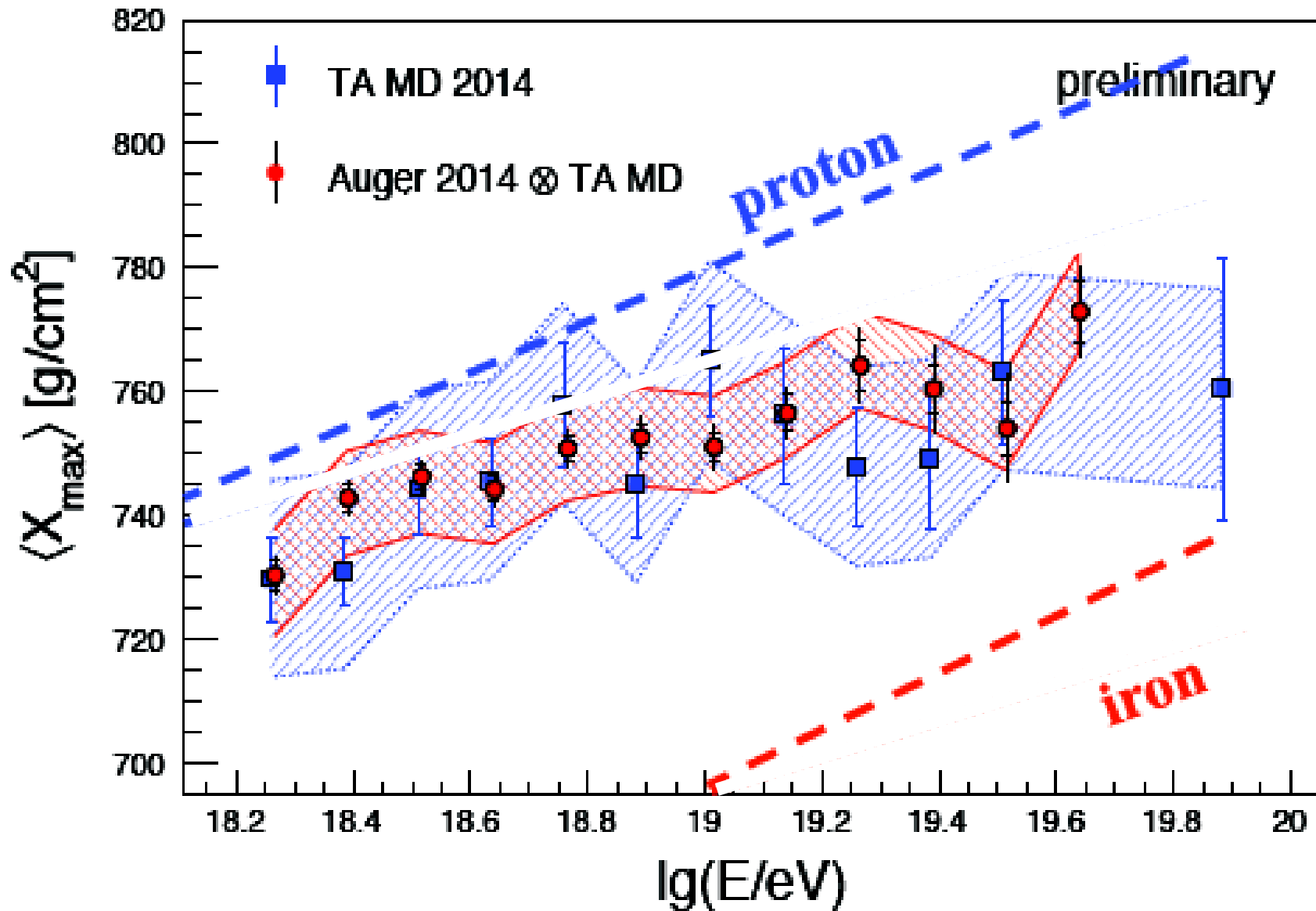
-20 g/cm² for baryons

MPD sensitive to pion interactions which should have a lower elasticity than in proton inter.

Summary

- **LHC data show no special feature and soft increase**
 - ➔ data bracketted by pre-LHC models : **break in CR not due to hadronic inter.**
 - ➔ post LHC models EPOS LHC and QGSJETII-04 predicted new 13 TeV correctly
- **Central particle production at LHC reduced model uncertainties in slope of X_{\max}**
 - ➔ same energy evolution in models important for mass of primary cosmic rays
 - ➔ **all pre-LHC models in contradiction with LHC data** (central and forward prod.)
 - ➔ using latest model version reduce uncertainties and avoid unphysical behavior
- **Improvements to come (EPOS 3 for ICRC 2017, others ?)**
 - ➔ **forward physics:** photon and neutron spectra and diffraction measured at LHC, and baryon stopping and resonance production at SPS
 - ➔ **effect of extrapolation to p-Air interaction:** p-Pb measurements can be used to constrain nuclear effects (p-O would be the best check).
 - ➔ **effect of (very) low energy:** extension to very low energy (few GeV) to have a better control on the muon production.

TA ...

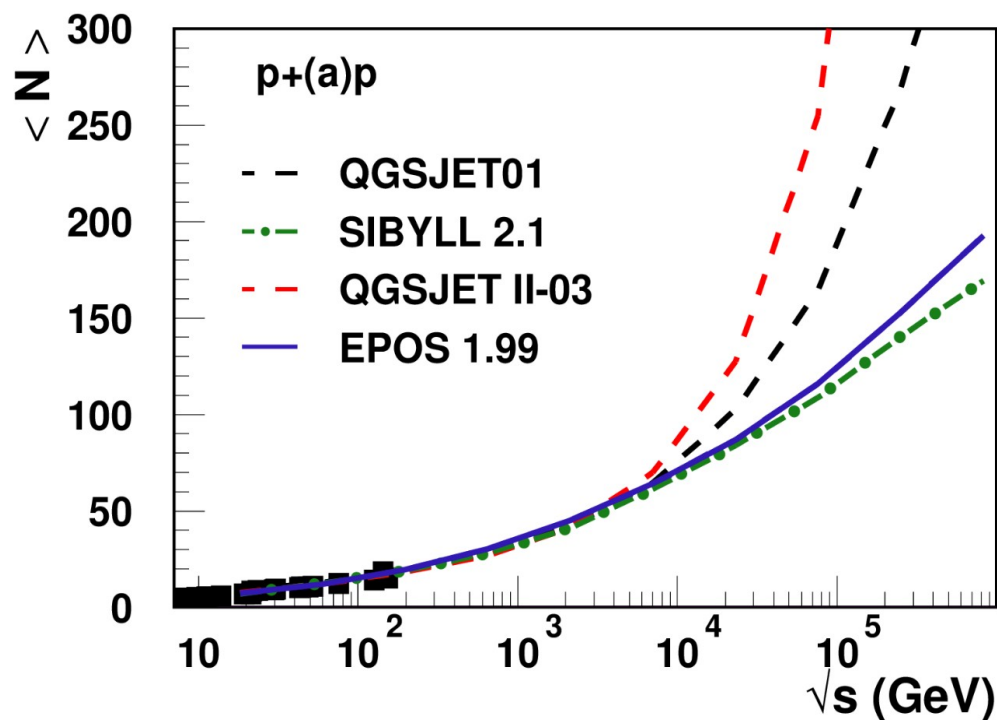


From Roberto Aloiso talk (2015 working group)

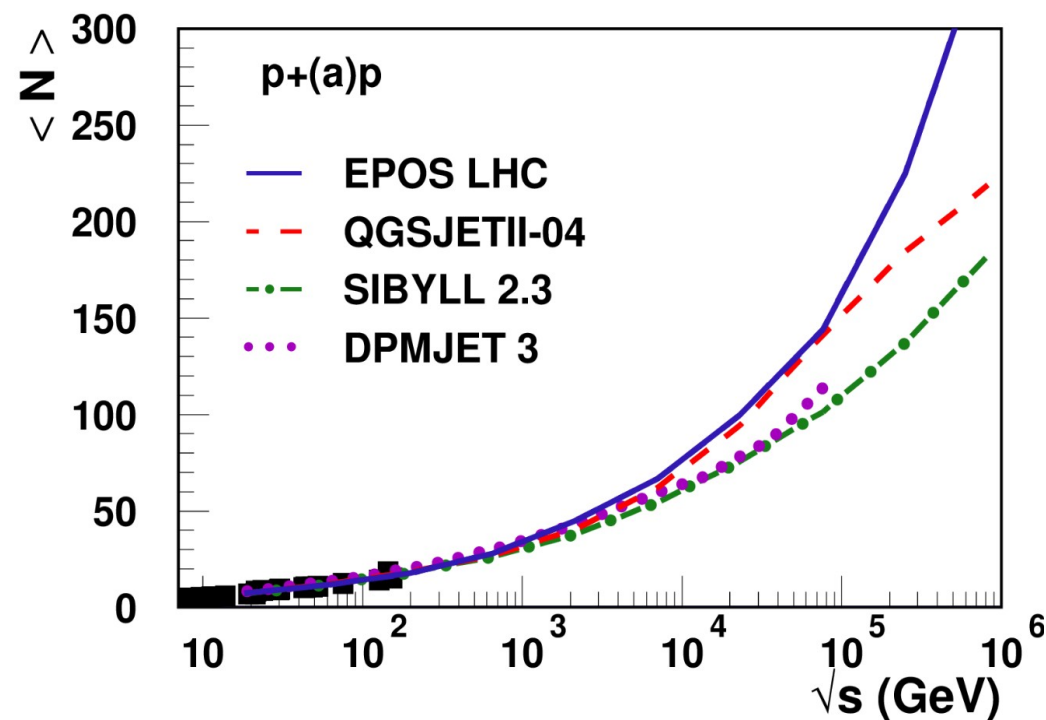
Multiplicity

- Multiplicity fixed by data up to 900 GeV
- extrapolation to high energy is still model dependent ?

Pre - LHC



Post - LHC



Hadronic Interaction Models in CORSIKA

(HDPM)

Old generation : (SIBYLL 2.1 QGSJET01 DPMJET 2.55 VENUS) (<2001)

All Glauber based
But differences in hard, remnants, diffraction ...

New (!) generation :

LHC tuned :

LHC inspired : **SIBYLL 2.3**

Motivation :

- update with latest LHC results in simple model

Engel et al.

semi-hard

(QGSJET II-03)

QGSJET II-04

Ostapchenko

QGSJET III (?)

Motivation :

- Hard Pomeron-Pomeron connexion

NEXUS 3.97

soft

Attempt to get everything described in a consistent way (energy sharing)

(EPOS 1.99) (2005-2012)

EPOS LHC (2013-)

Pierog & Werner

EPOS 3 (2017-)

Motivation :

- binary scaling in hard probes
- better diffraction

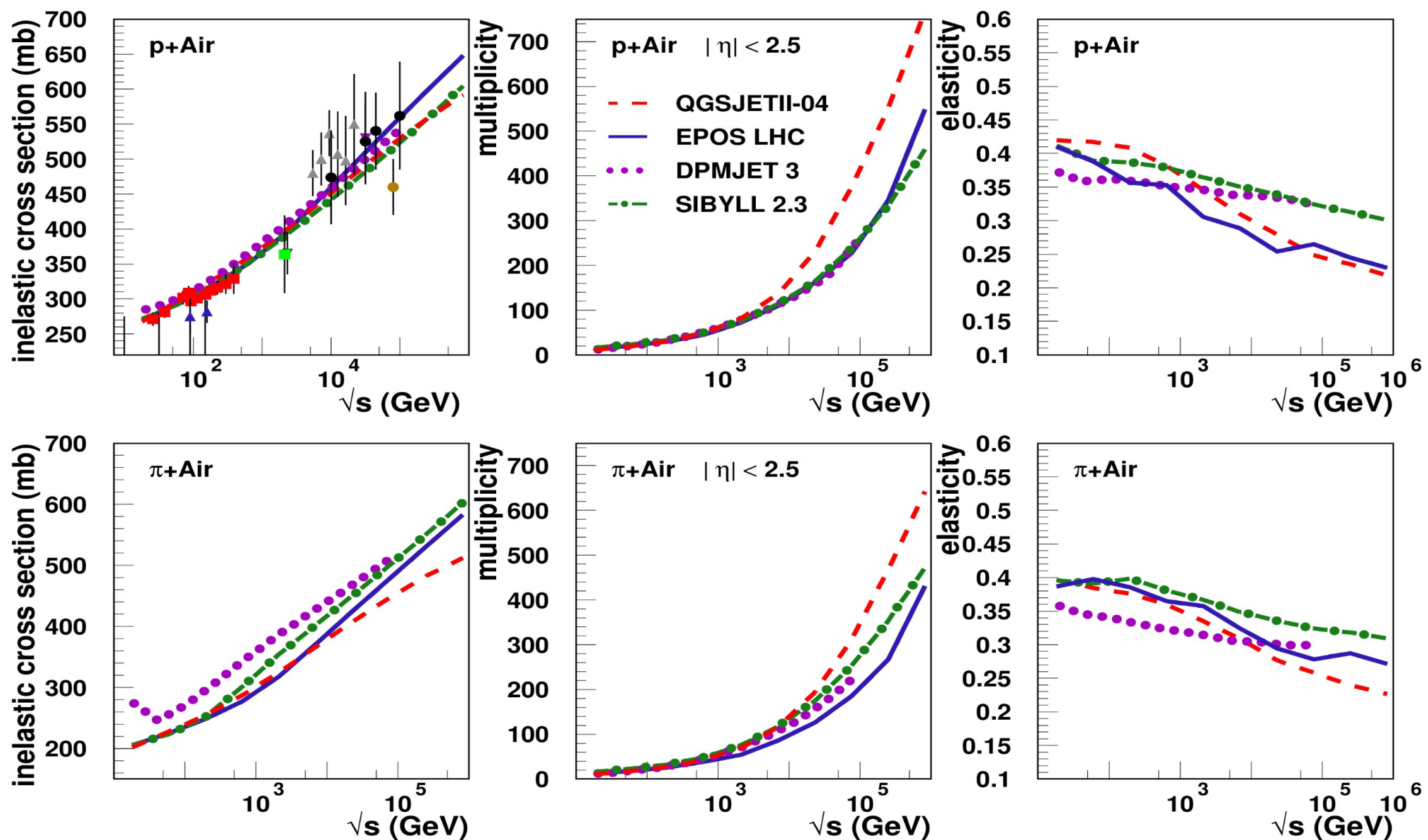
Fedinitch & Engel

DPMJET III

Motivation :

- update with LHC results
- fix high energy

Ultra-High Energy Hadronic Model Predictions π -Air

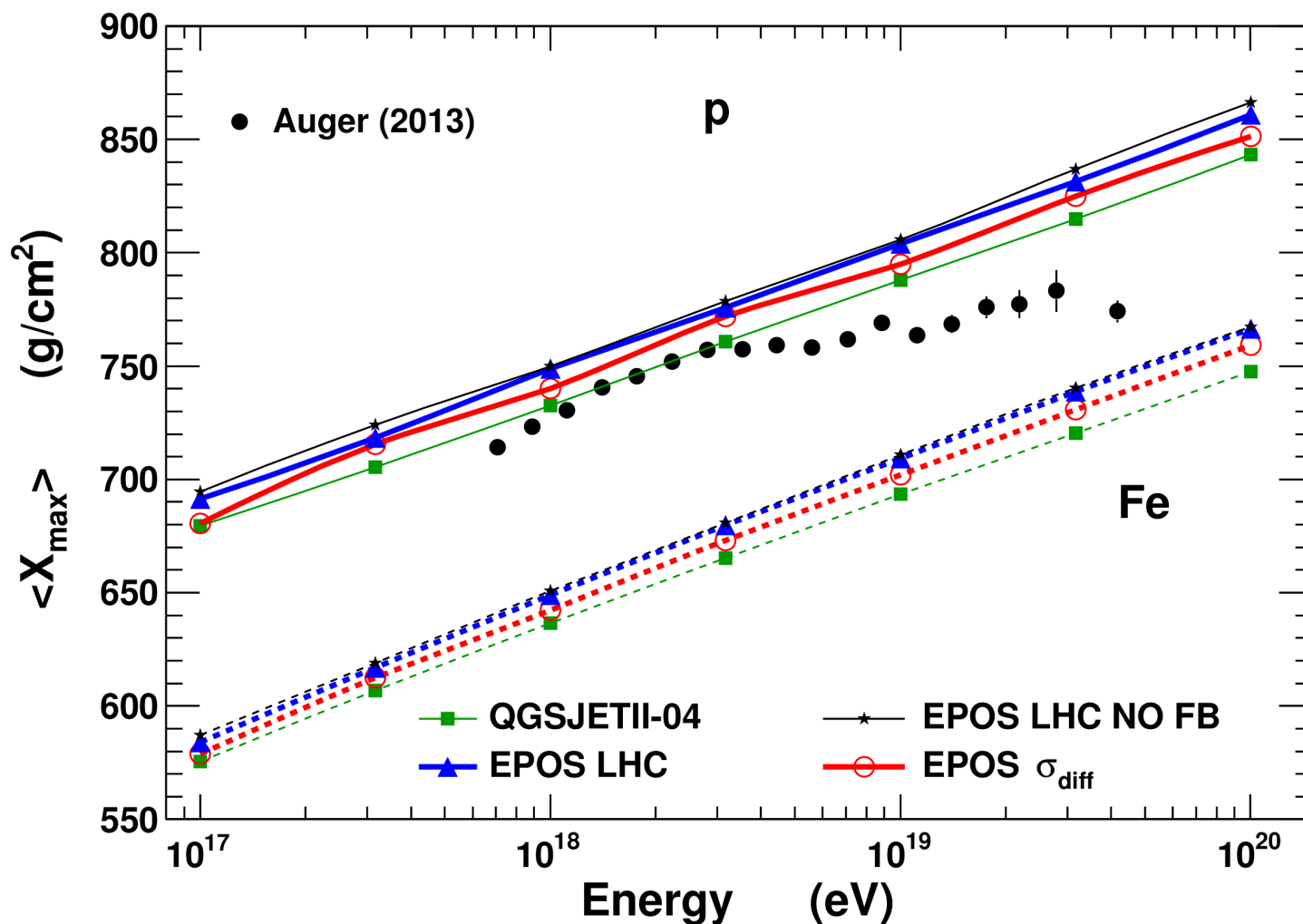


$\langle X_{\max} \rangle$ with Modified EPOS LHC

EPOS LHC without forward baryons or more inelastic pion int.

→ softer meson spectra: smaller X_{\max}

→ forward baryons: negligible effect



-10 g/cm² for diff
 ~0 g/cm² for baryons

X_{\max} less sensitive to baryon spectra than to pion spectra in pion interactions