

MU Programtag 2016

12-13 December 2016
Helmholtz Institute Mainz



Dark Photon Searches at MAMI and MESA

Achim Denig, Mainz



Cluster of Excellence
PRISMA

Precision Physics,
Fundamental Interactions
and Structure of Matter



THE LOW-ENERGY FRONTIER
OF THE STANDARD MODEL



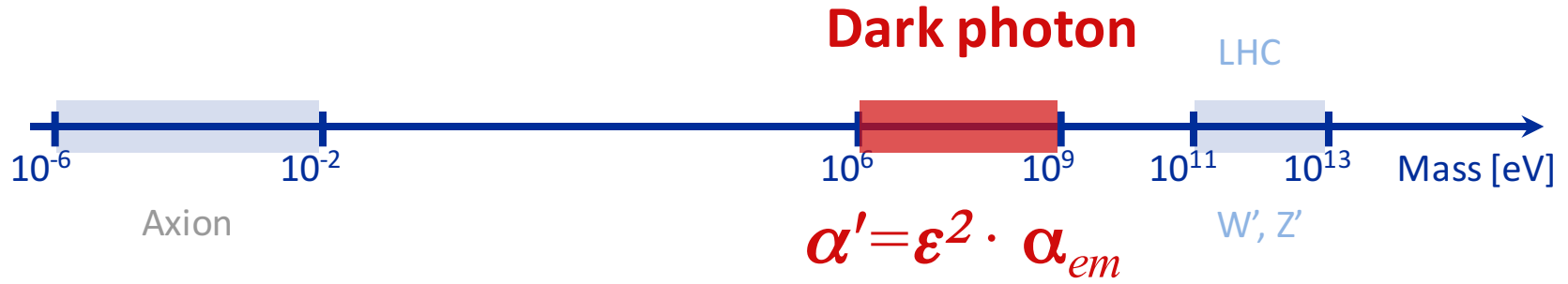
HIM

Helmholtz-Institut Mainz



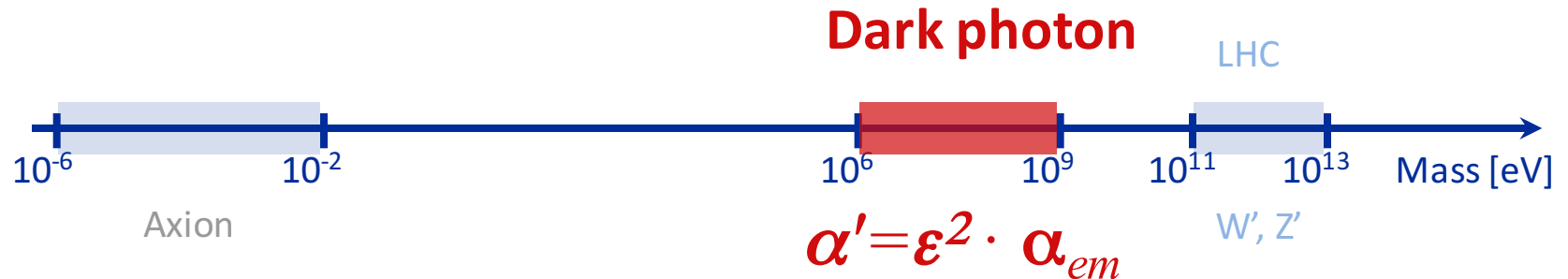
Dark Photon Search

New massive force carrier of extra $U(1)_d$ gauge group;
 predicted in almost all string compactifications



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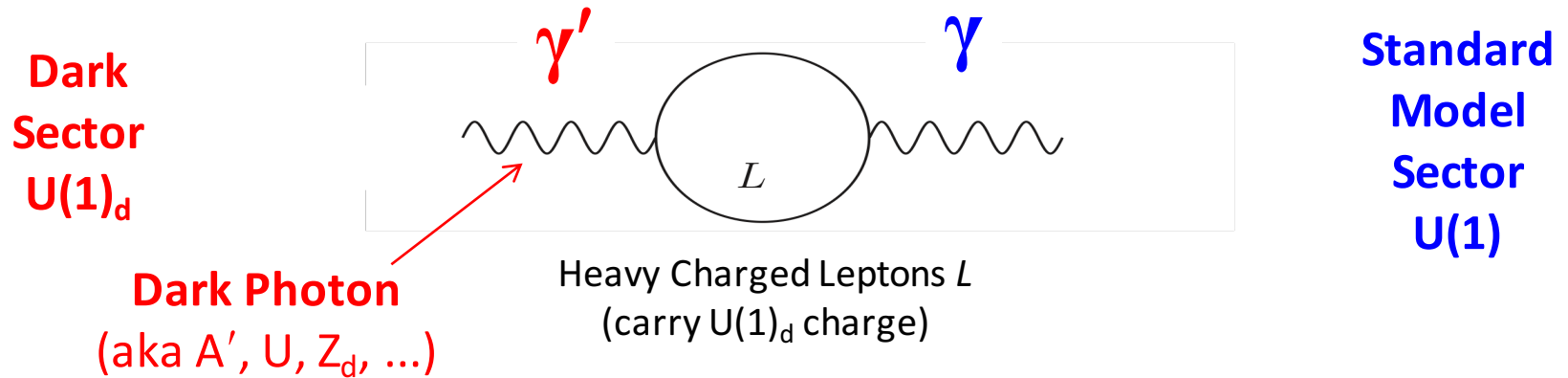
Search for the $O(\text{GeV}/c^2)$ mass scale in a world-wide effort

- Could explain large number of **astrophysical anomalies**
Arkani-Hamed et al. (2009)
Andreas, Ringwald (2010); Andreas, Niebuhr, Ringwald (2012)
- Could explain presently seen **deviation of 3.6σ between $(g-2)_\mu$**
Standard Model prediction and direct $(g-2)_\mu$ measurement
Pospelov(2008)

Kinetic Mixing and Dark Matter

Holdom [1986]

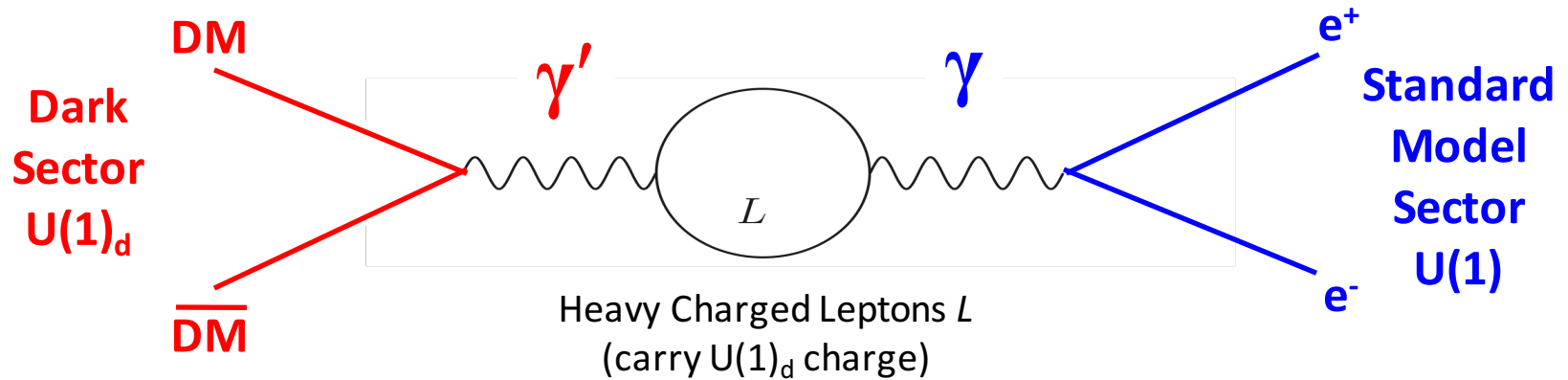
A way to relate the dark sector to the SM (coupling $\sim \epsilon^2$)



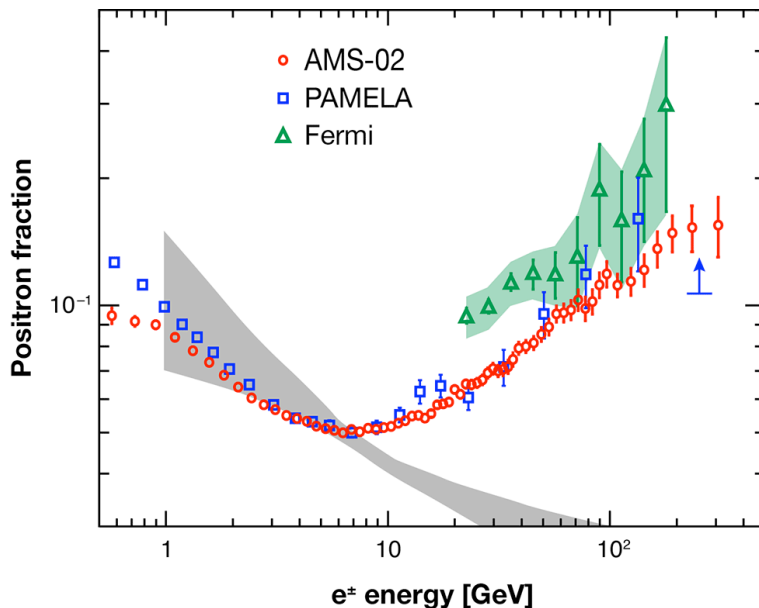
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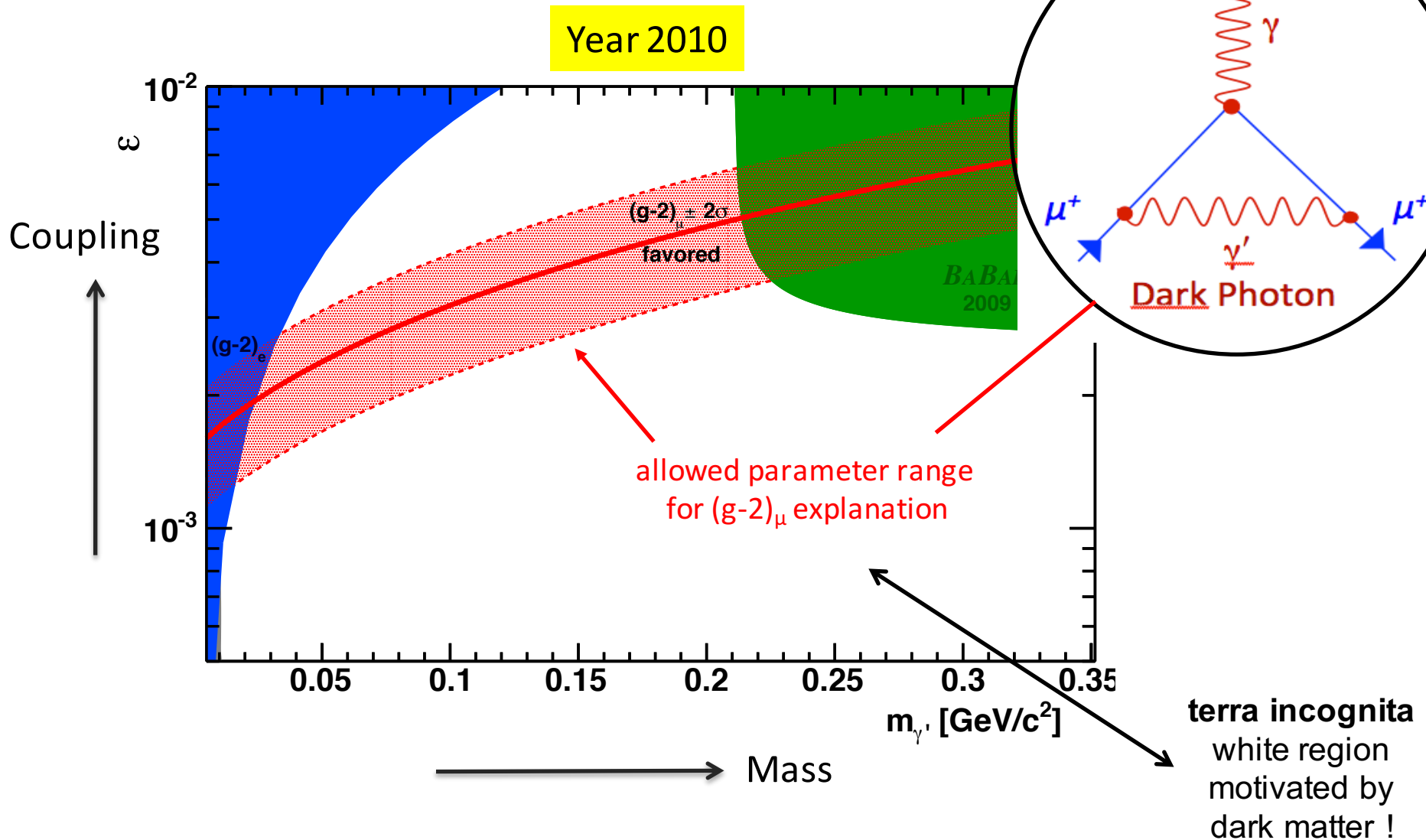


Excess of positrons in cosmic ray spectrum due to Dark Matter annihilation?



Search for Dark Photon at MAMI and MESA

Dark Photon and $(g-2)_\mu$



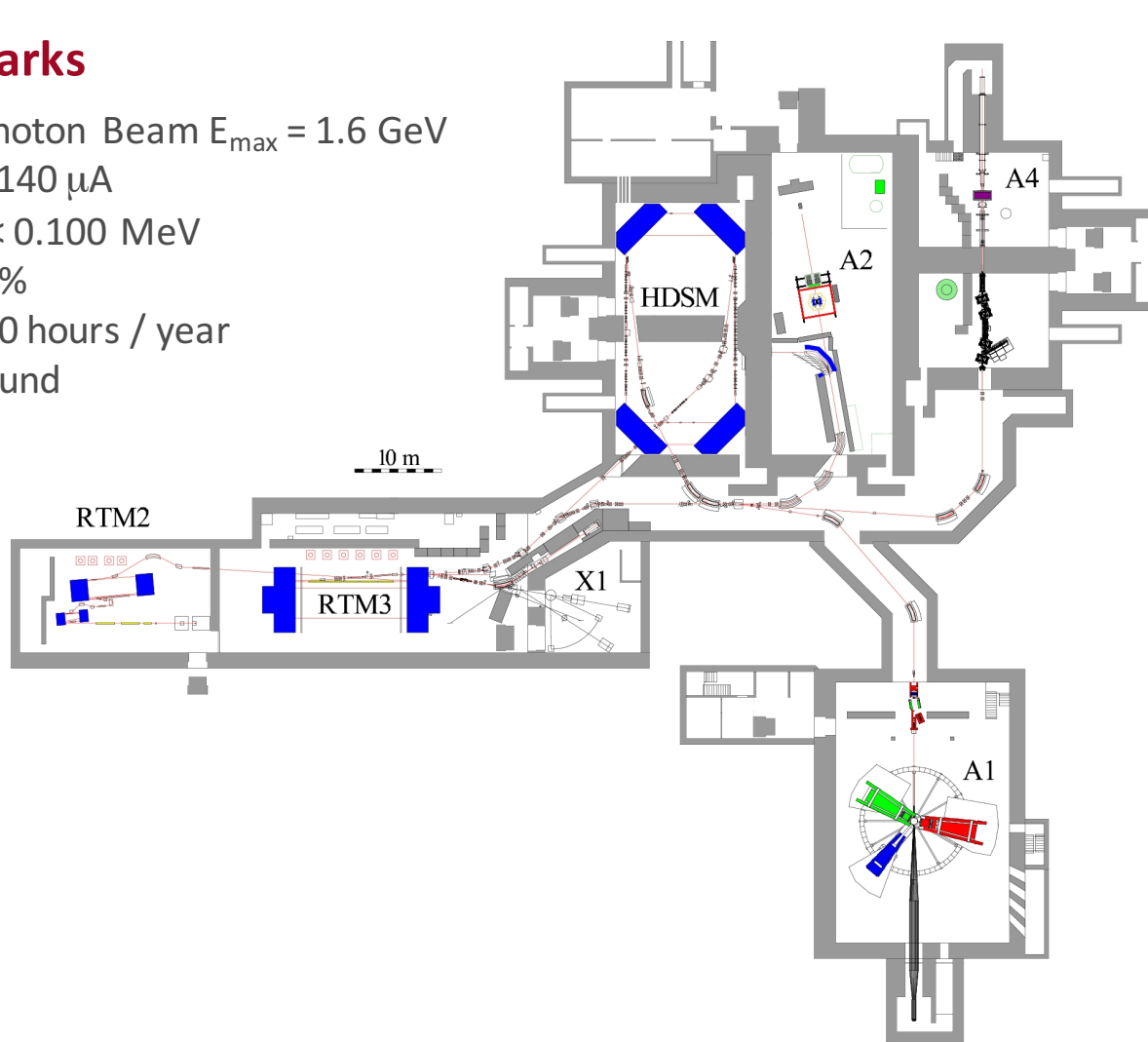


Dark Photon Search at MAMI

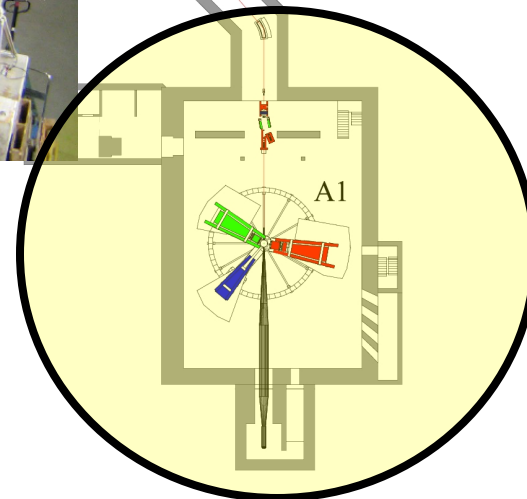
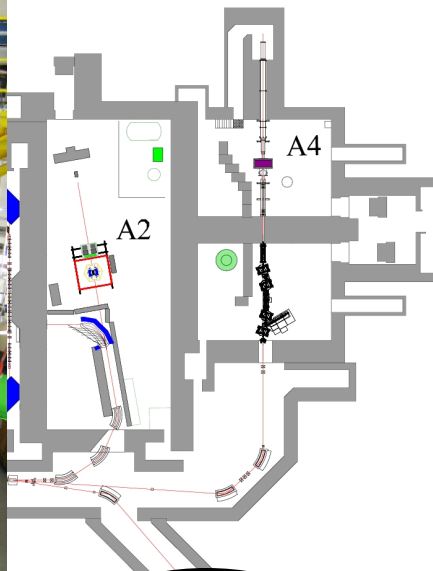
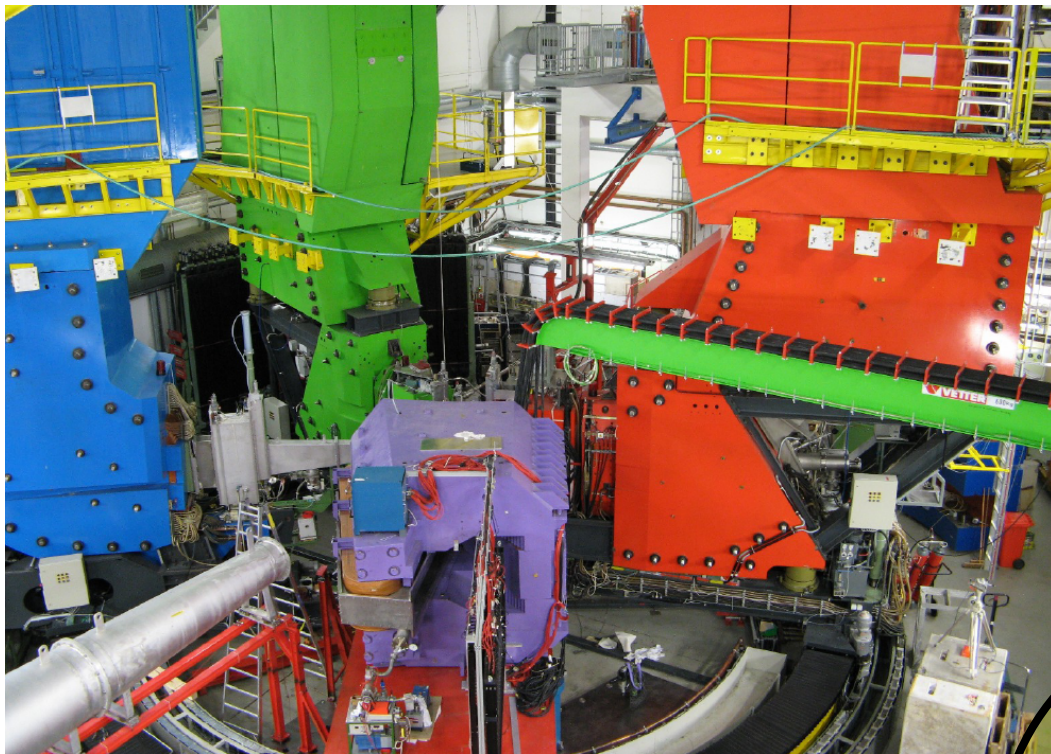
A1: High-Resolution Spectrometers

MAMI Hallmarks

- Electron and Photon Beam $E_{\text{max}} = 1.6 \text{ GeV}$
- Intensity max. $140 \mu\text{A}$
- Resolution $\sigma_E < 0.100 \text{ MeV}$
- Polarization 85%
- Reliability: 7000 hours / year
- 11 m underground



A1: High-Resolution Spectrometers



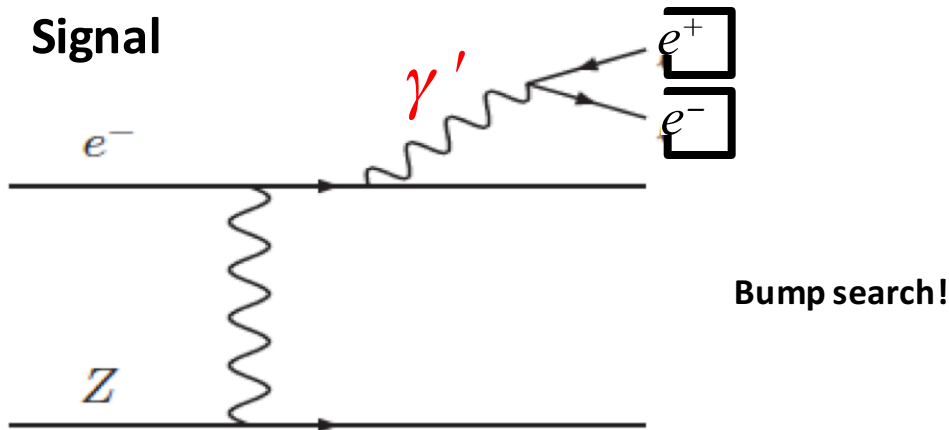
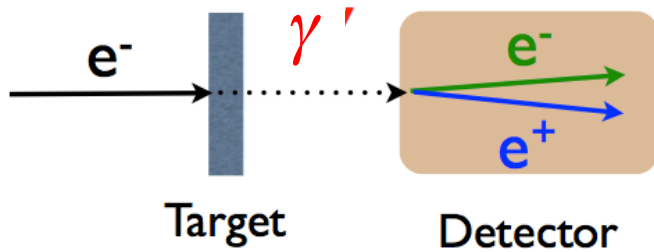
Experiment A1: Electron Scattering

- 4 magnetic focussing spectrometers
- Resolution: $\delta p/p < 10^{-4}$
- Angular acceptance: < 30 mrad

Results from A1 Pilot Run (2011)

Low-Energy Electron Accelerator with high intensity suited for DP search

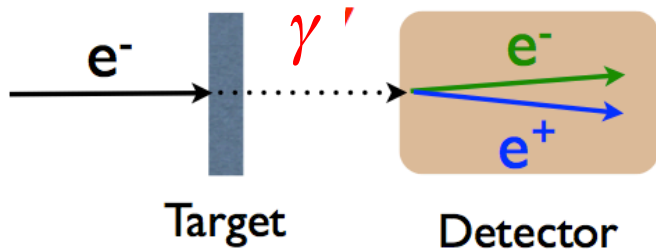
Bjorken, Esssig,
Schuster, Toro (2009)



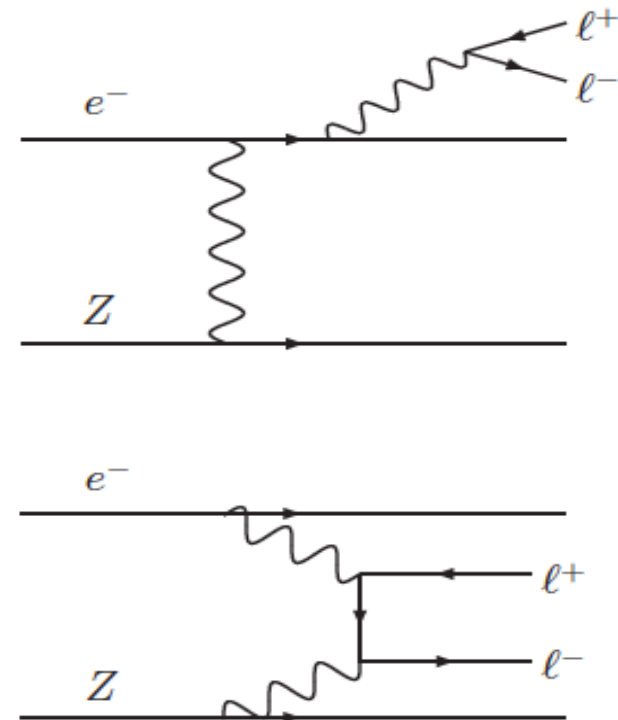
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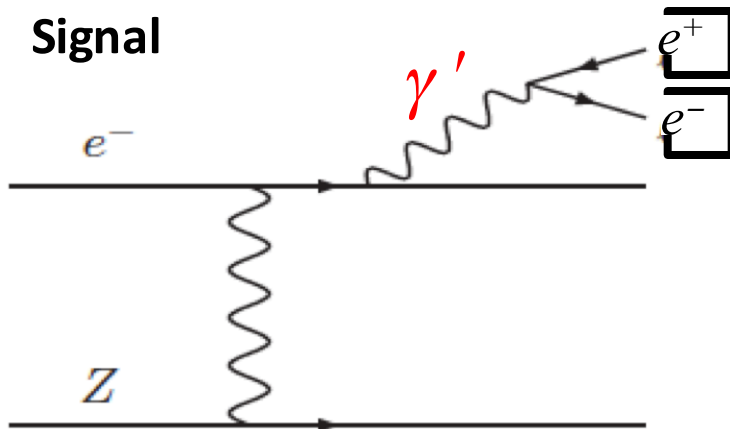
Bjorken, Esssig, Schuster, Toro (2009)



Background



Signal



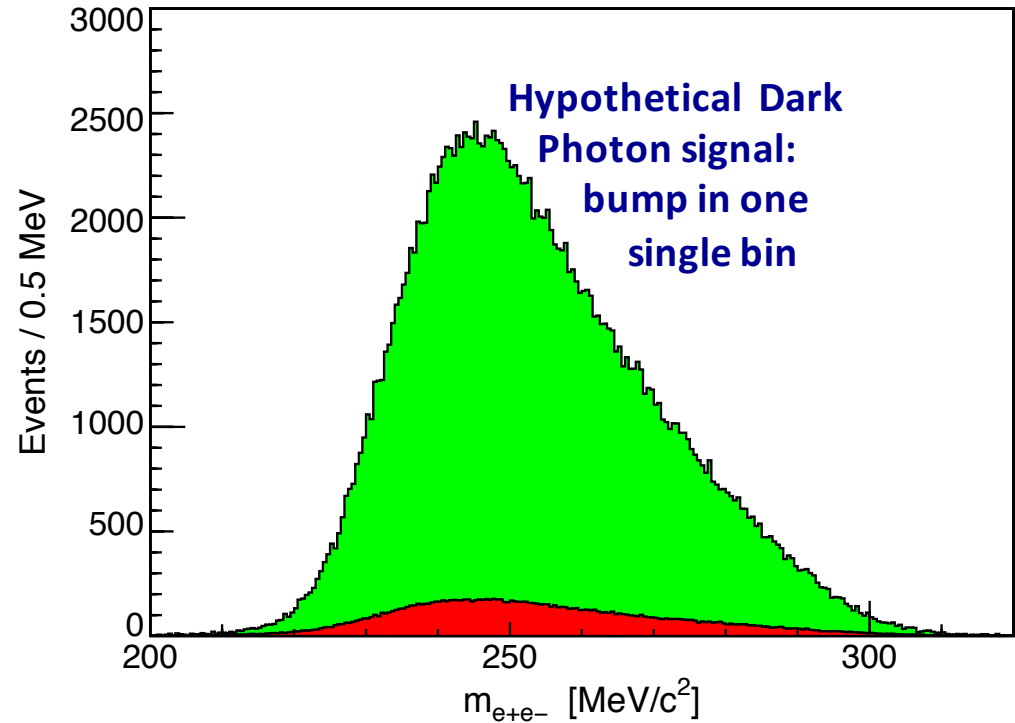
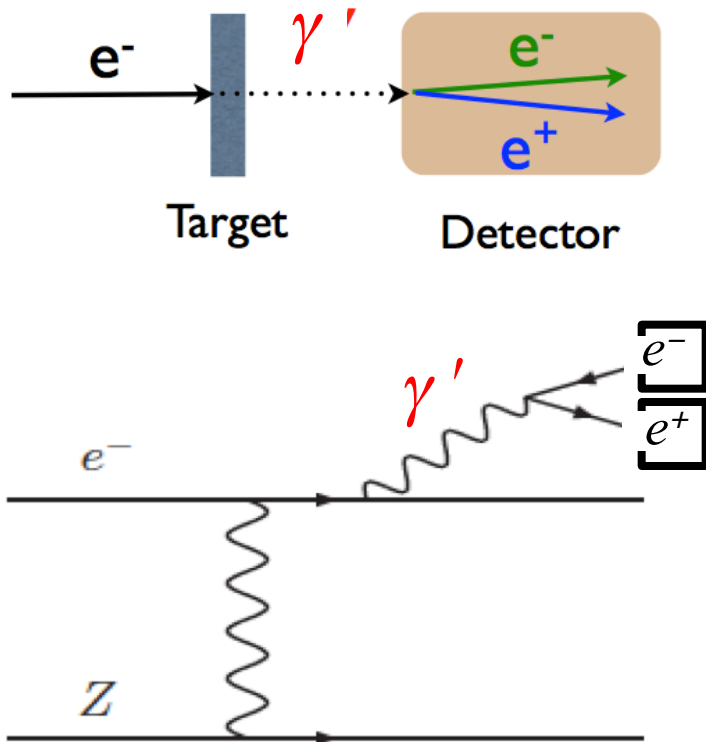
Bump search!

Results from A1/MAMI

Low-Energy Electr. Acceler. with high Intensity suited for DP search

Bjorken, Esssig,
Schuster, Toro (2009)

Signal processes

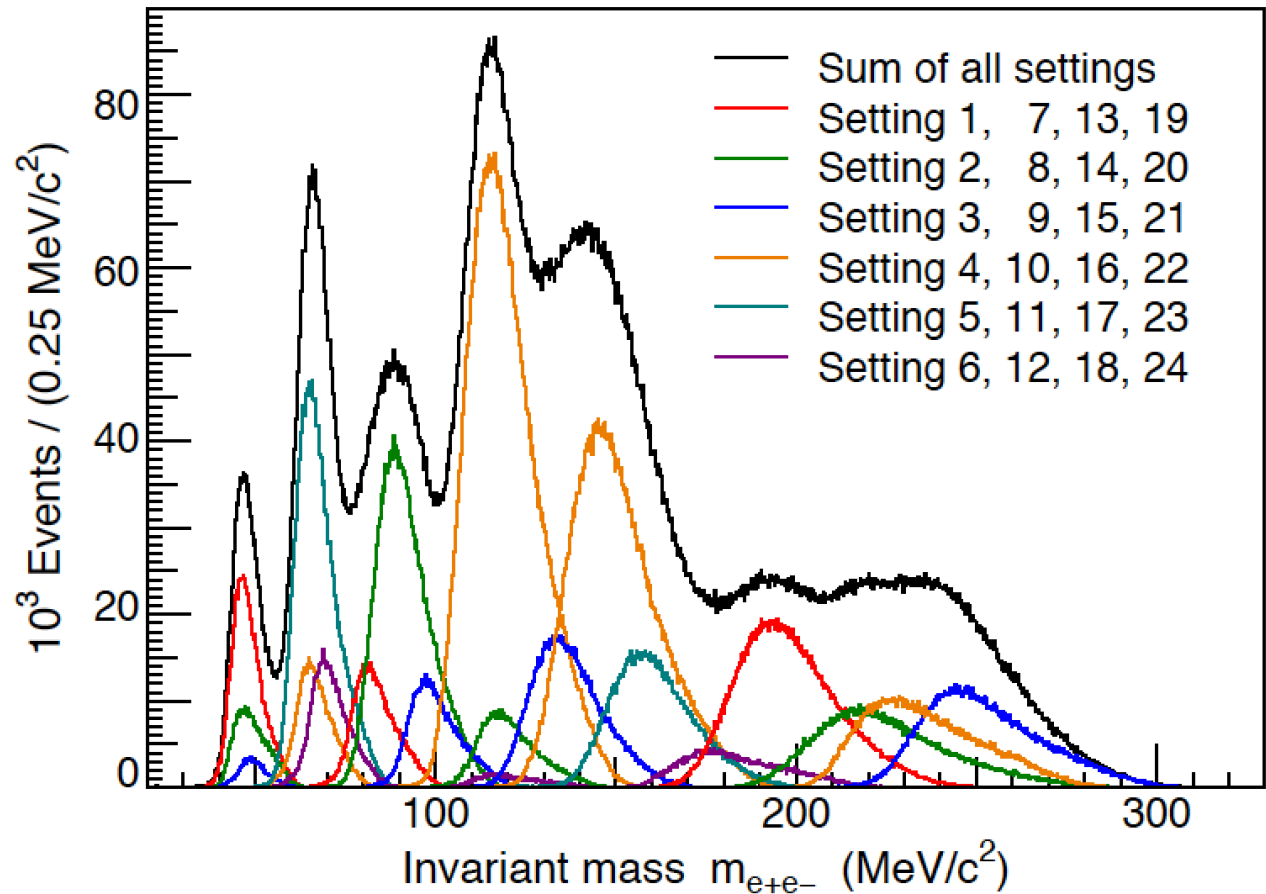


2010 test run
Invariant Mass of e^+e^-

Results from A1

Merkel et al. [A1]
PRL '11
PRL '14

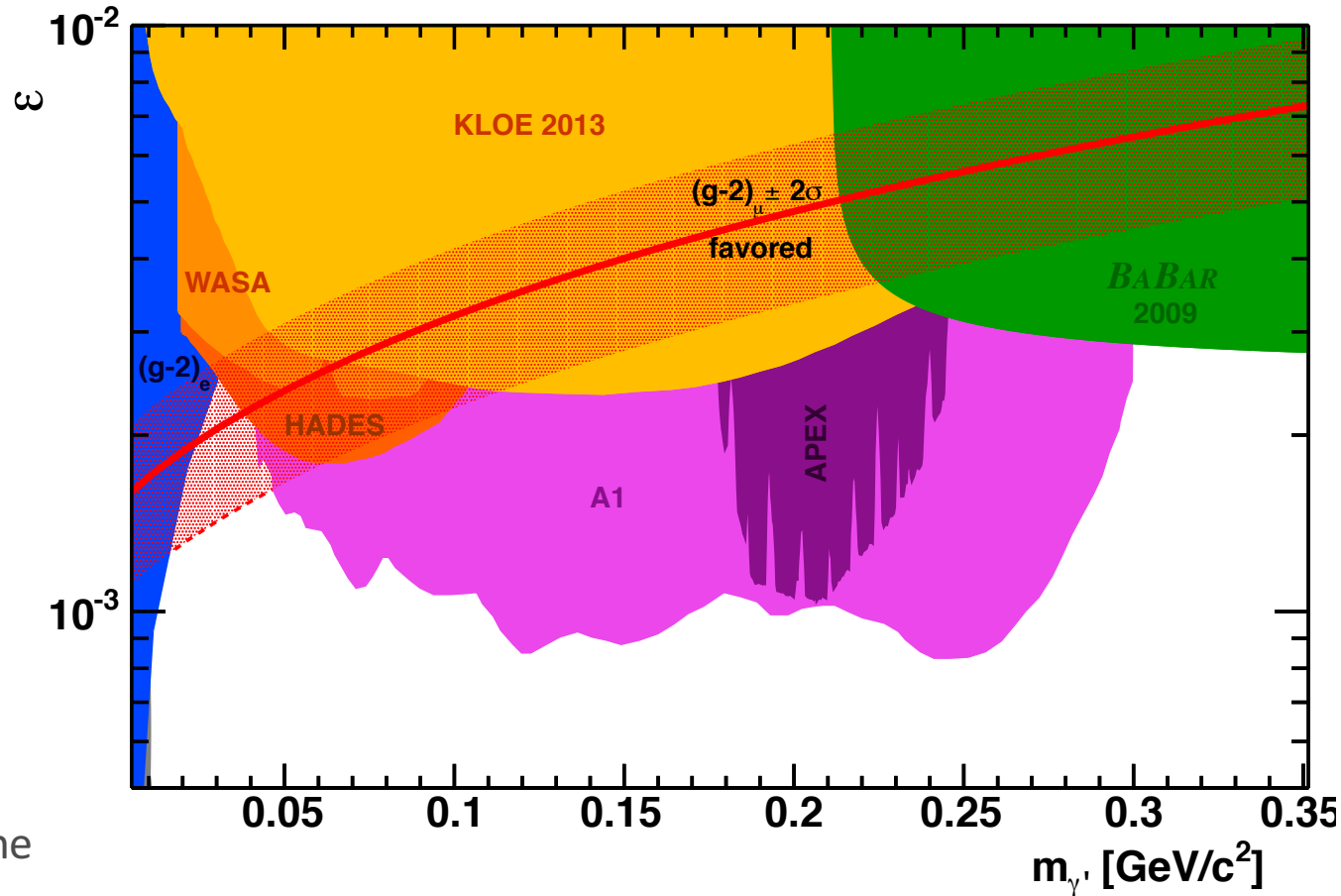
- E_{beam} 180 - 855 MeV
- 100 μA beam current
- Stack of Ta targets
- 22 kinematic settings
- O(1 month) of beam time



→ at time of publication most stringent
limit ruling out major part of the parameter range motivated by $(g-2)_\mu$

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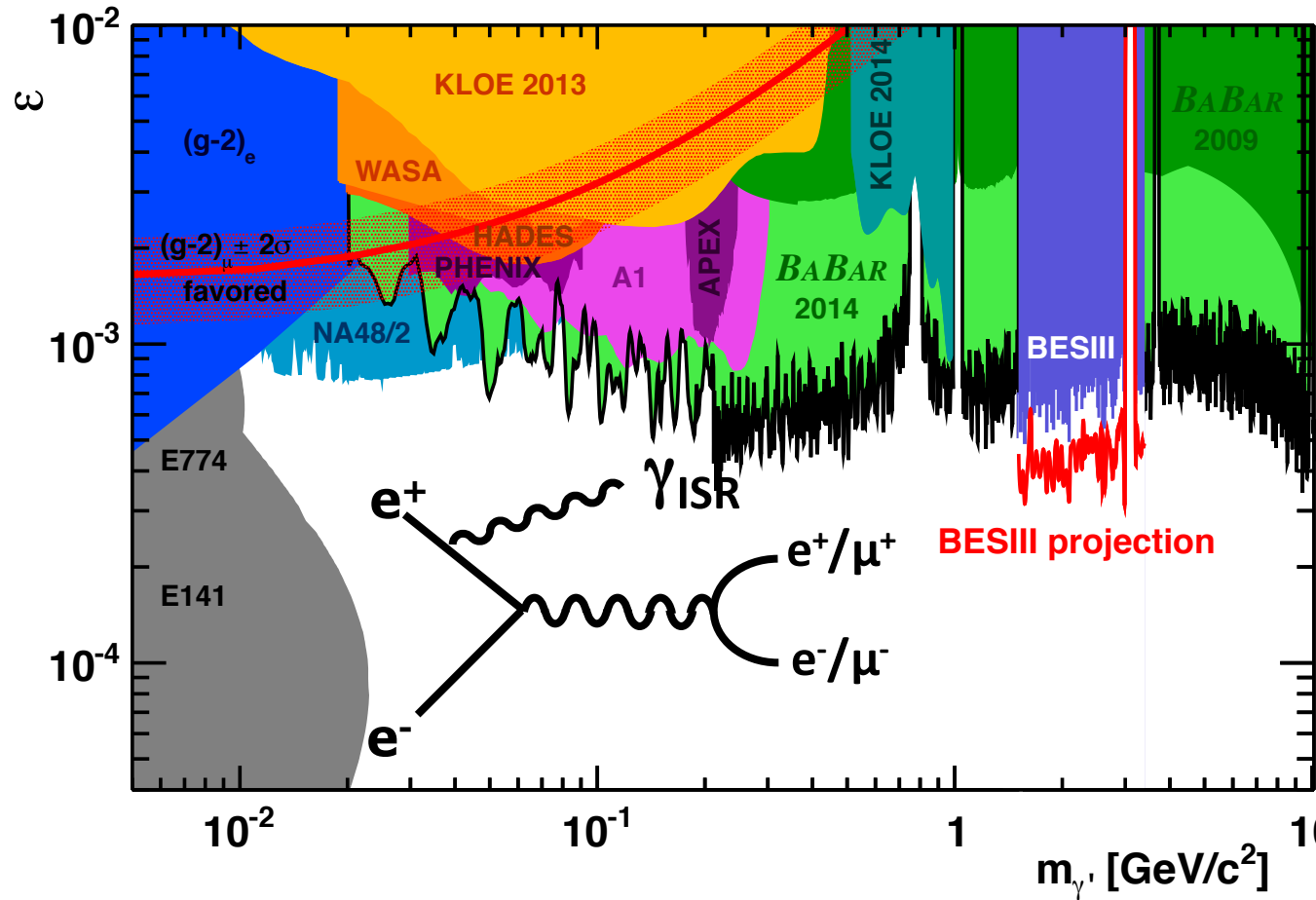


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Dark Photon Search at MESA

New Tool for Low-Precision Physics: MESA

Mainz Energy-Recovering Superconducting Accelerator

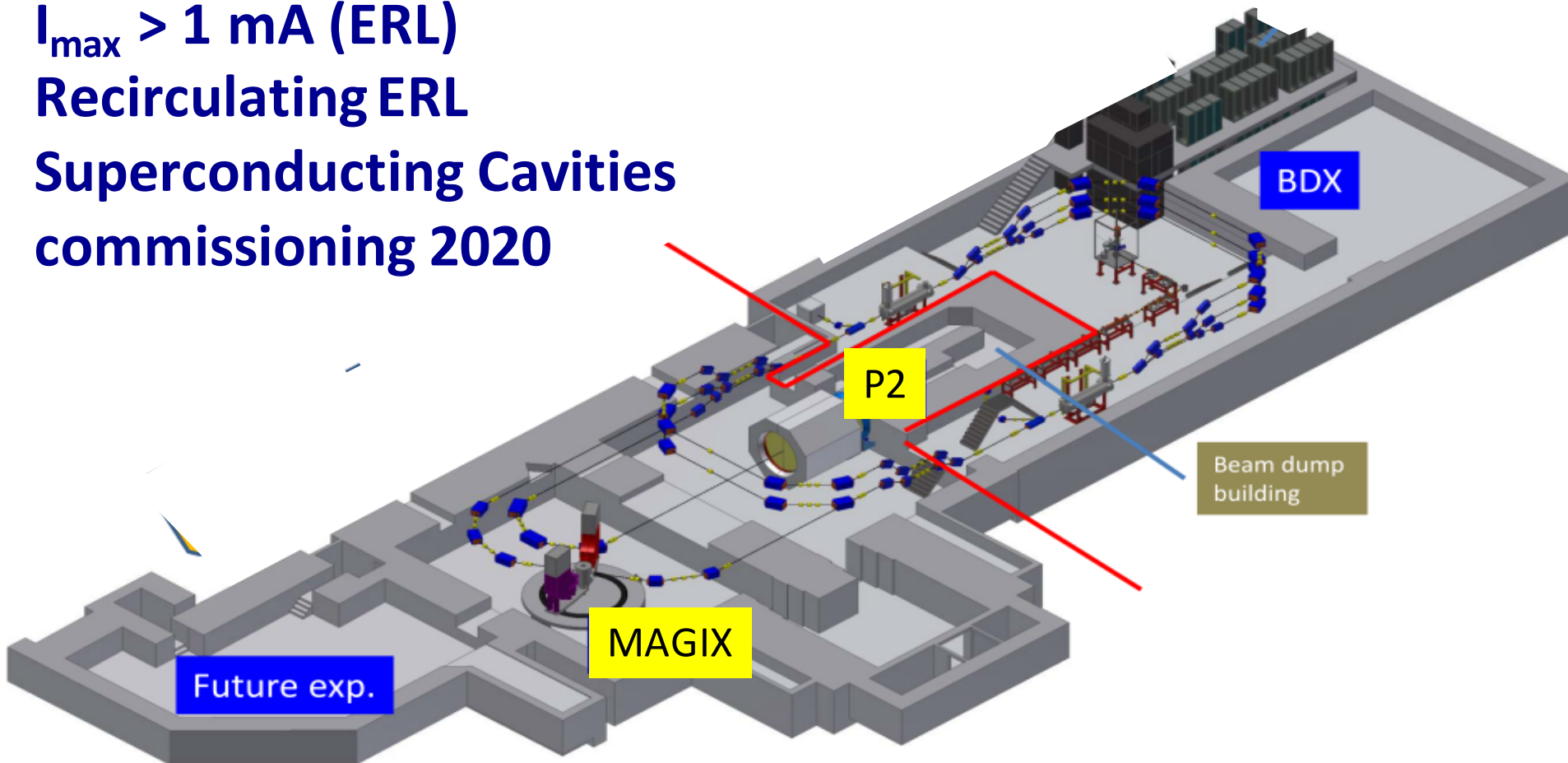
$E_{\max} = 155 \text{ MeV}$

$I_{\max} > 1 \text{ mA (ERL)}$

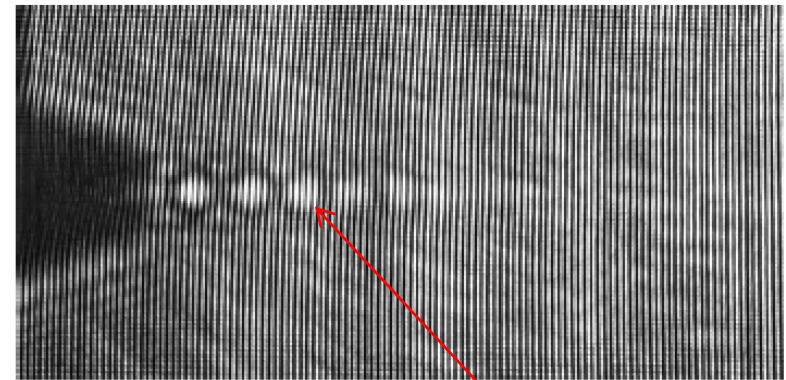
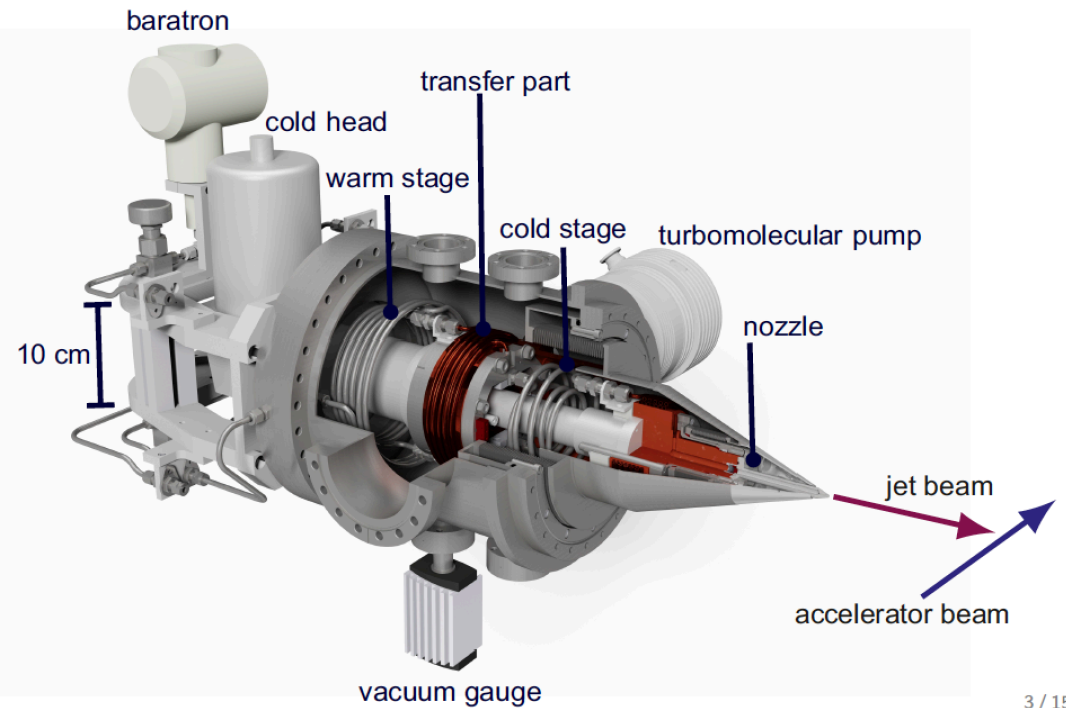
Recirculating ERL

Superconducting Cavities
commissioning 2020

7



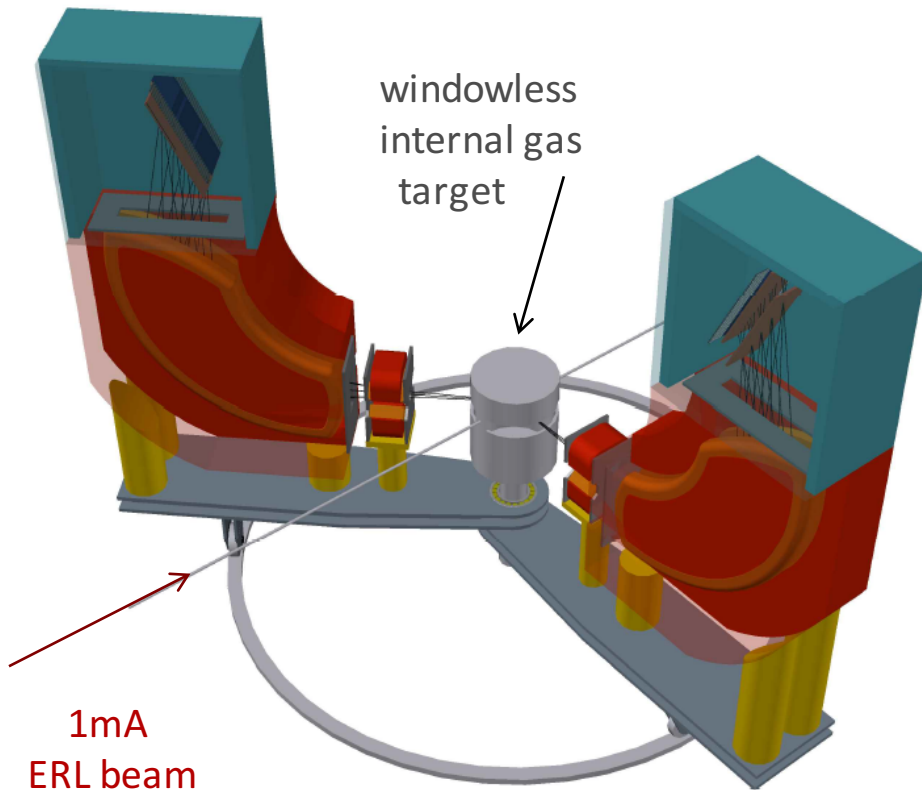
Internal Gas Jet Target



Characteristic profile of a supersonic gas jet

- Supersonic gas /cluster jet
- Higher gas density ($10^{19}/\text{cm}^2$)
- O(mm) target length
- Estimated luminosity $O(10^{35} \text{ cm}^{-2} \text{ s}^{-1}) @ 10^{19}/\text{cm}^2$
- Windowless !
- Ready in 2016 !

**Operation of a high-intensity (polarized) ERL beam
in conjunction with light internal target
→ a novel technique in nuclear and particle physics**



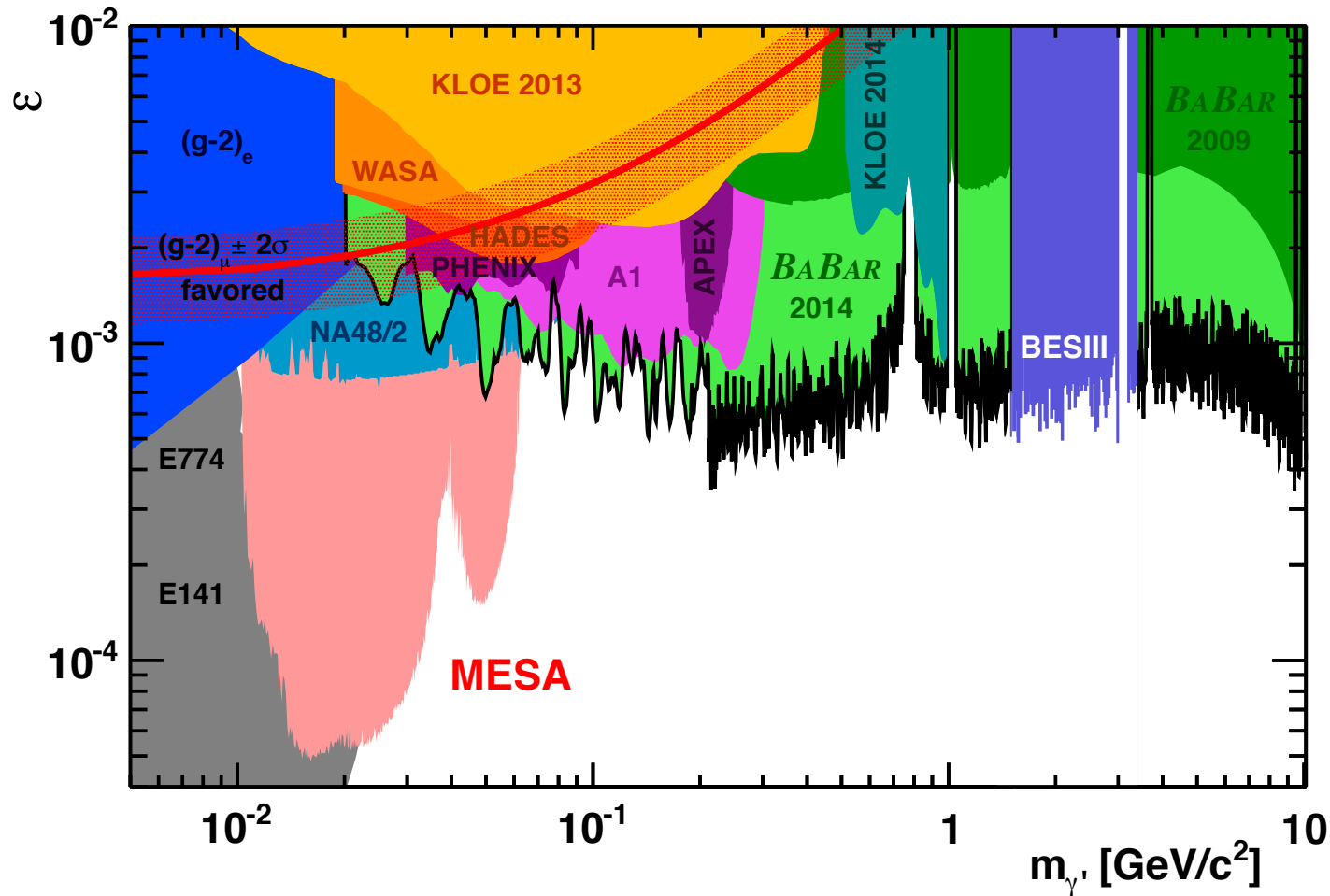
High resolution spectrometers MAGIX:

- double arm
- compact design
- momentum resolution: $\Delta p/p < 10^{-4}$
- acceptance: ± 50 mrad
- GEM-based focal plane detectors
- Gas Jet or polarized T-shaped target

Dark Photon at MAGIX/MESA

Features:

- Xe gas target
- Luminosity $10^{35} \text{ cm}^{-2}\text{s}^{-1}$
- 6 month of data taking





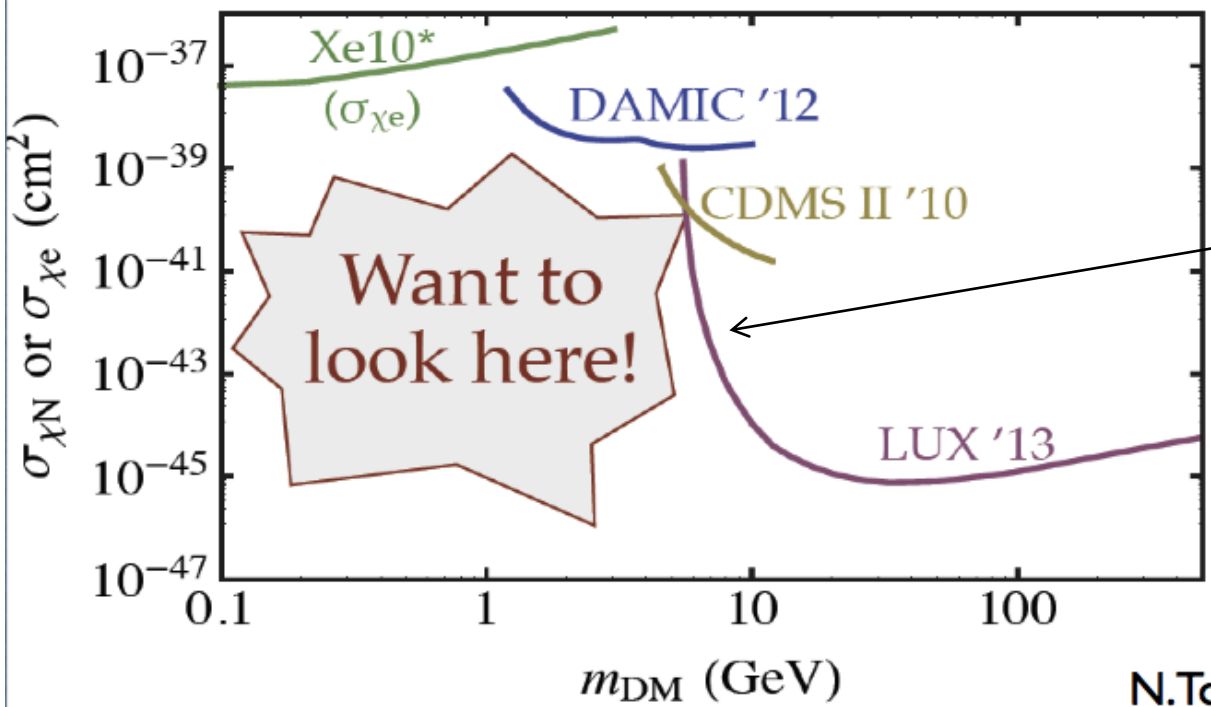
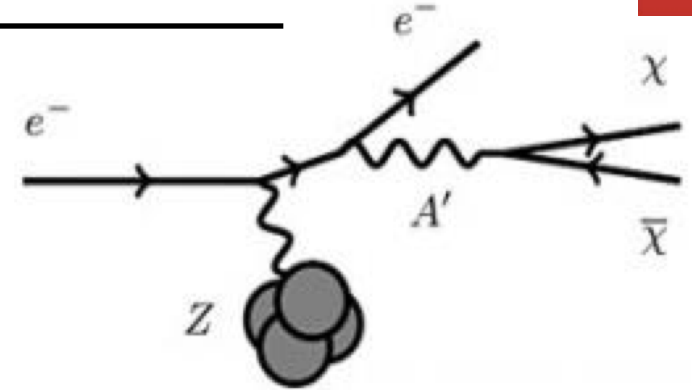
Invisible Dark Photon Decays

Invisible Decay of Dark Photon

So far Kinetic Mixing: $M(\gamma') < 2M(\chi)$

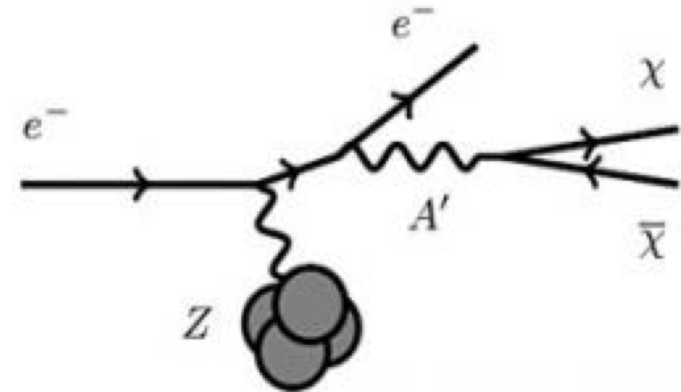
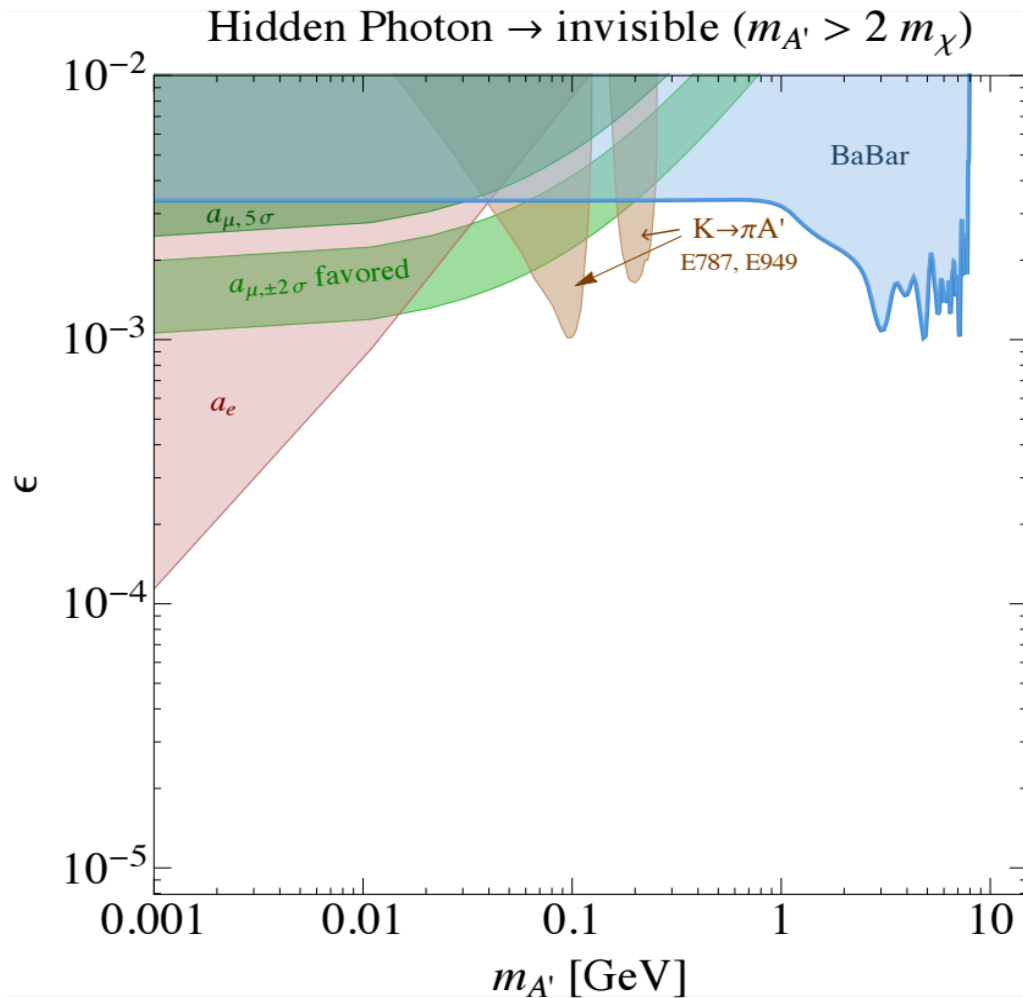
Consider now: $M(\chi) < 2M(\gamma')$

Dark Photon decays dominantly into
Light Dark Matter (LDM), which is not yet constraint experimentally!



exclusion limits from classical Dark Matter experiments (nuclear recoil) not sensitive to LDM

Invisible Decay of Dark Photon



- Dark Matter particle not seen
- Few constraints
- Could again explain $(g-2)_\mu$

- \rightarrow Missing energy / mass
- \rightarrow Search for Dark Matter particle directly using dedicated low-background detectors

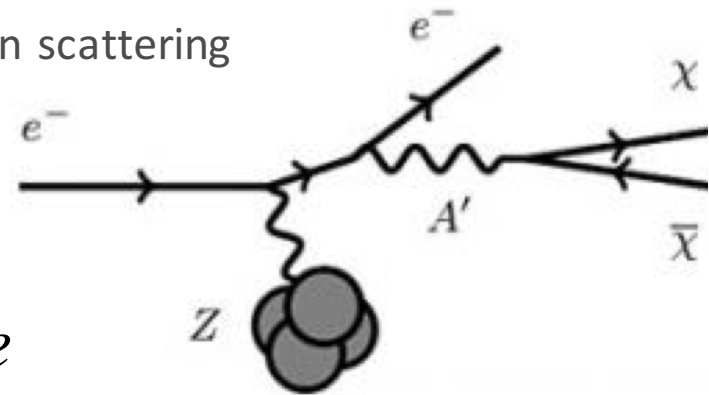
Dark Sector Searches at MAGIX

Model 2: Dark Photon coupling to Dark Matter

- could still explain $(g-2)_\mu$ discrepancy
- exploit excellent momentum resolution of MAGIX (proton recoil!)
- Main background: Virtual Compton scattering

$$e + p \rightarrow e' + p + X$$

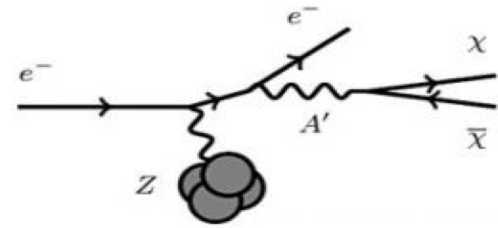
\hookrightarrow invisible



$$m_{\gamma'}^2 = (e + p - e' - p')^2$$

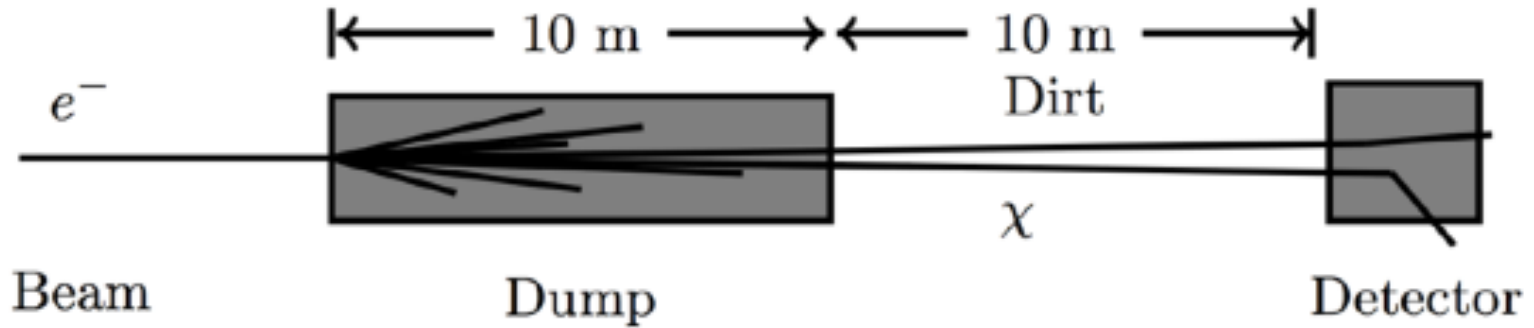
Sensitivity at MAGIX currently calculated within a bachelor thesis
(use of thin HVMAPS detectors for proton recoil under study)

Beam Dump Experiment (BDX)

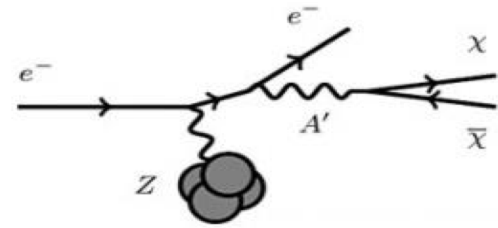


Electron Scattering on Beam Dump

→ Collimated and boosted (!) pair of Dark Matter particles !

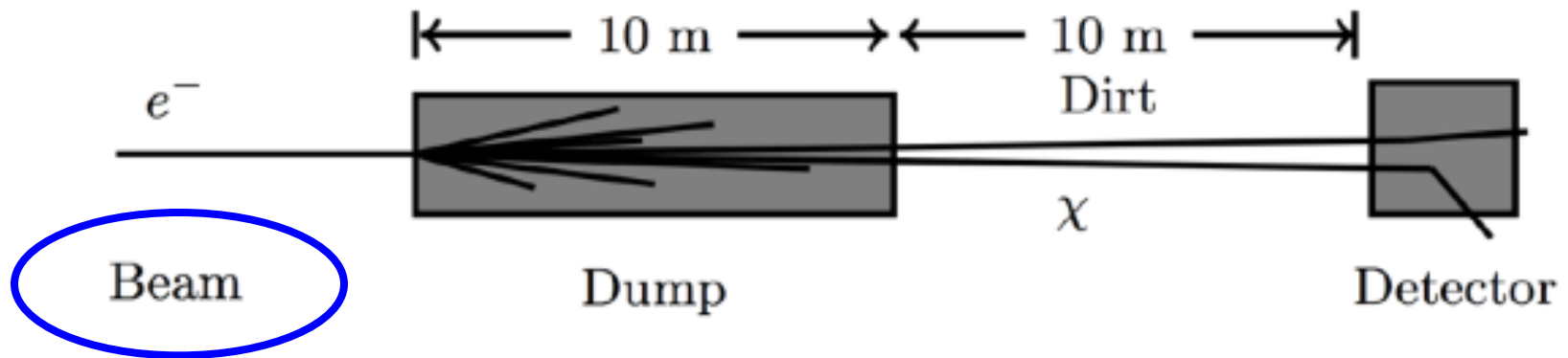


Beam Dump Experiment (BDX)



Electron Scattering on Beam Dump

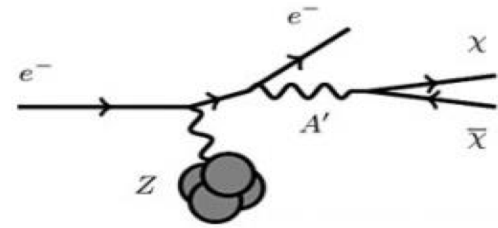
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Beam:

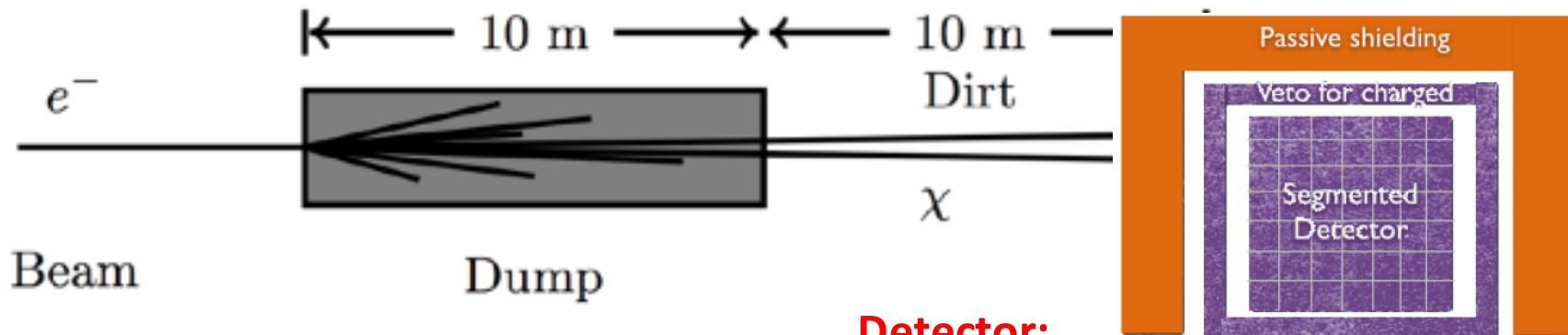
- High intensity electron beam $O(10^{20} - 10^{22})$ EOT/year
- \sim GeV energy scale
- Passive shielding and active vetos for background subtraction (μ , n , ν)
- Time structure for TOF measurement?

Beam Dump Experiment (BDX)



Electron Scattering on Beam Dump

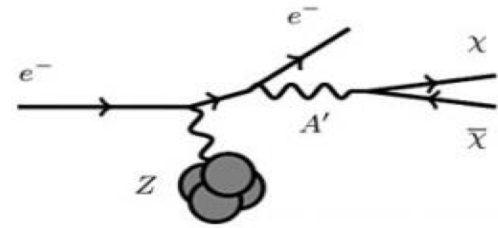
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Detector:

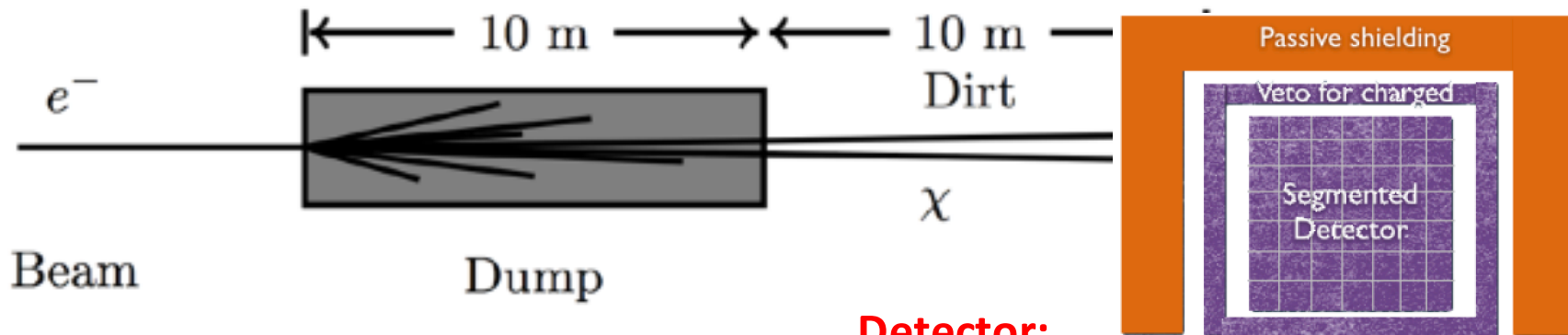
- $\sim 1 \text{ m}^3$ size detector (1-5 tons)
- High density to maximize event yield
- Low threshold for nucleon recoil detection (MeV) + EM shower detection
- CsI (TI) crystals from BaBar EMC (?)
High-tech for DM research considered

Beam Dump Experiment (BDX)



Electron Scattering on Beam Dump

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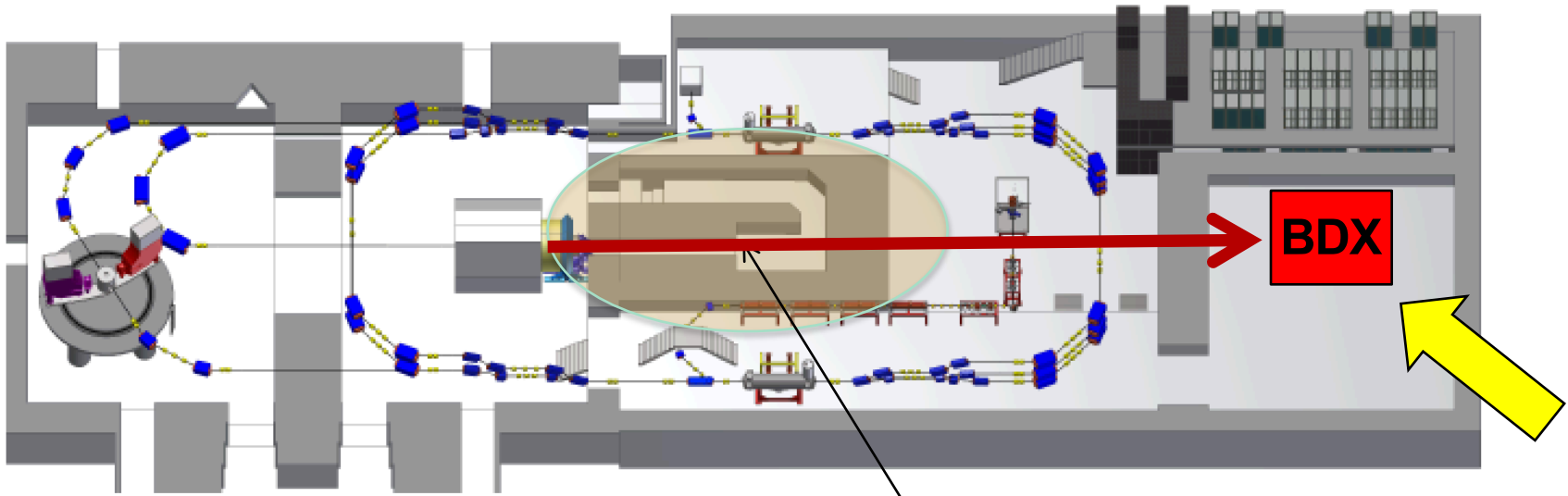
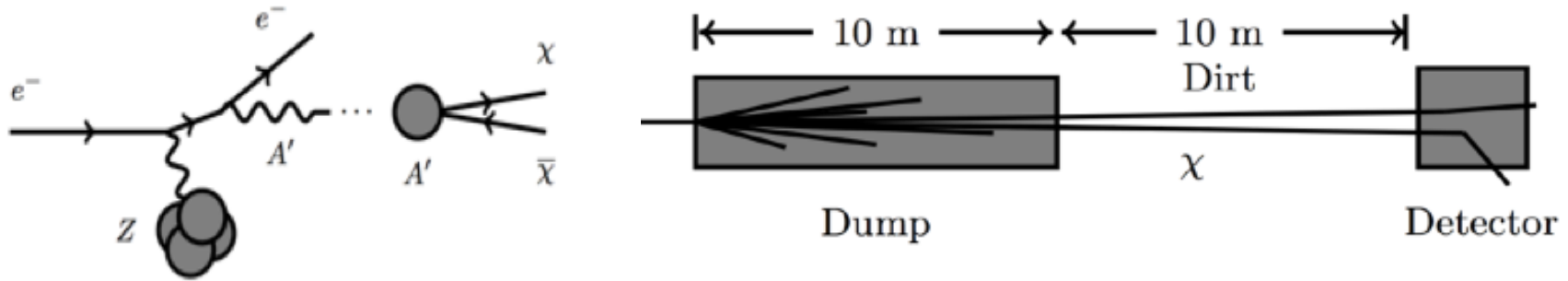


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Beam Dump Experiment (BDX) @ MESA

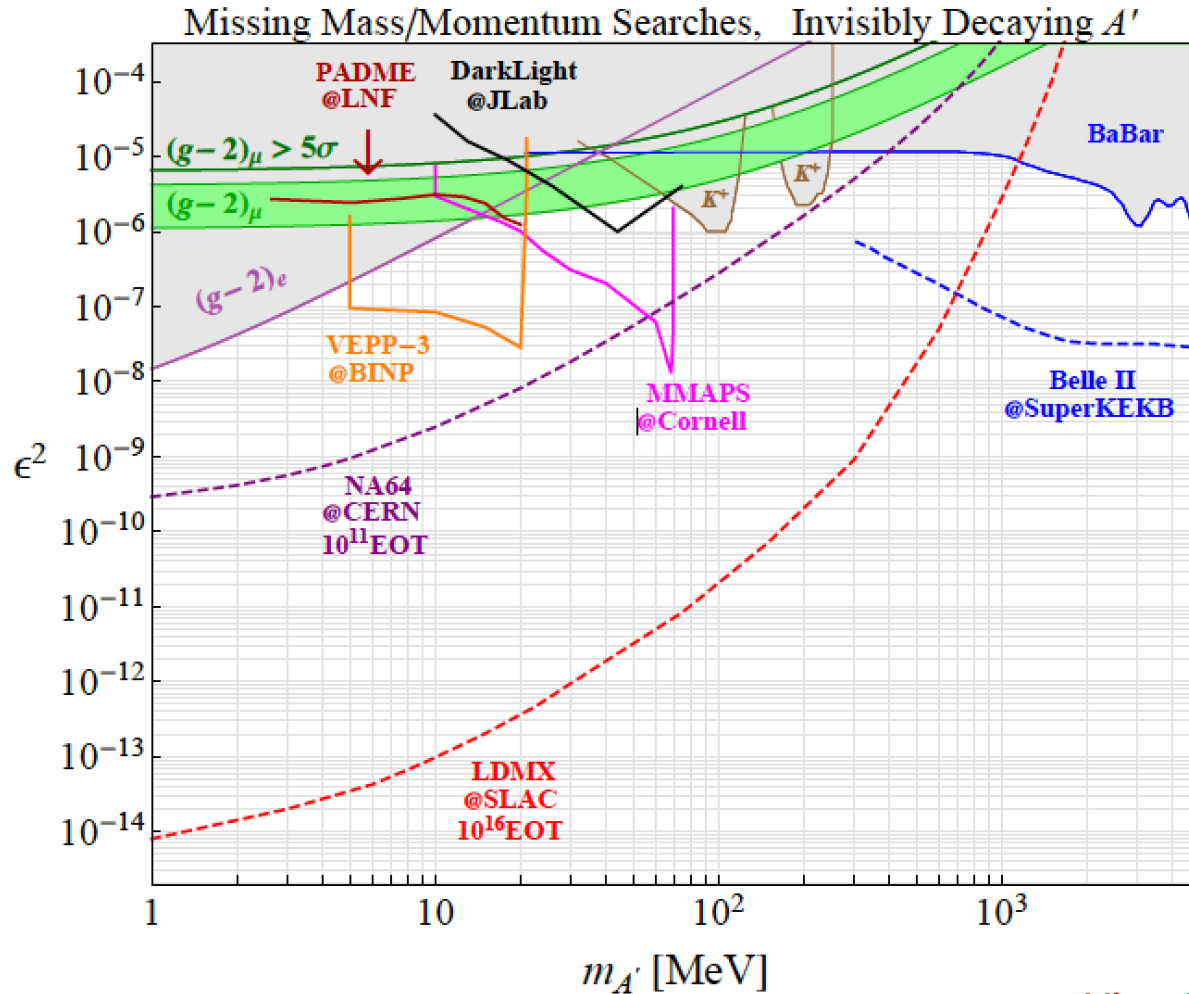
Electron Scattering on Beam Dump → Collimated pair of Dark Matter particles !



This existing beam dump is going to be the P2 beam dump

10,000 hours @ 150 μ A
→ 10²³ electrons on target (EOT)

Future of Invisible Dark Photon Searches

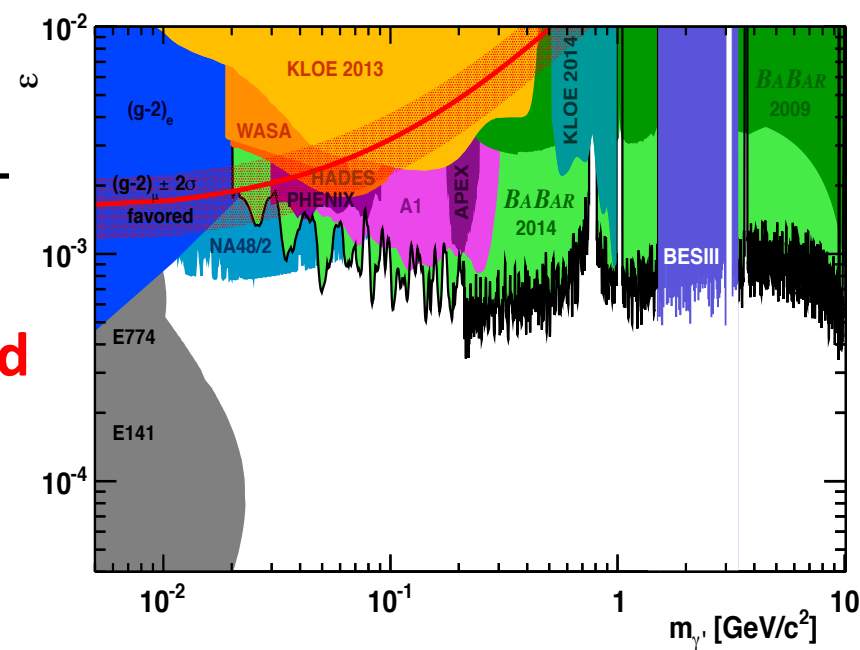


arXiv: 1608.08632



Conclusions

Take Home Message



- Searches for Dark Photons motivated by Dark Matter and by $(g-2)_\mu$ puzzle
- Tremendous exper. progress seen
- Explanation for $(g-2)_\mu$ ruled out in simplest kinetic mixing model
- Dark Photon searches now making more and more direct connection to Dark Matter candidates itself (invisible decays)
- Thriving field and more to come in next 5 – 10 years
- Anomalous behaviour in ${}^8\text{Be}^* \rightarrow {}^8\text{Be} e^+e^-$ nuclear transition ...
Claim for a 17 MeV Dark Photon with ~ 7 sigma significance

Thank you for your attention



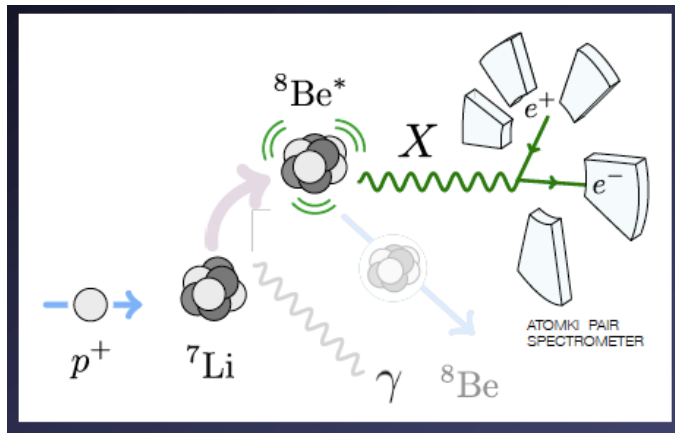
Legenda:

- Publishing
- Approved
- Proposals

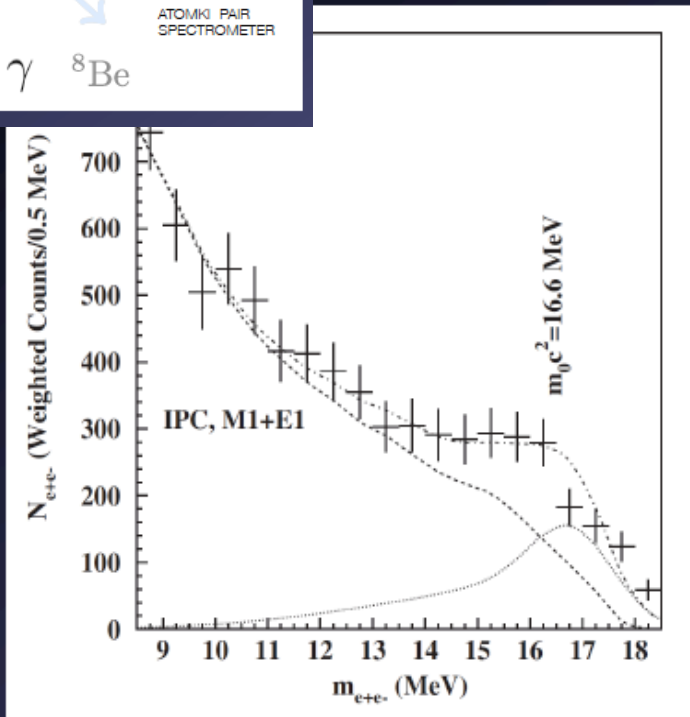
courtesy: Mauro Raggi

BACKUP

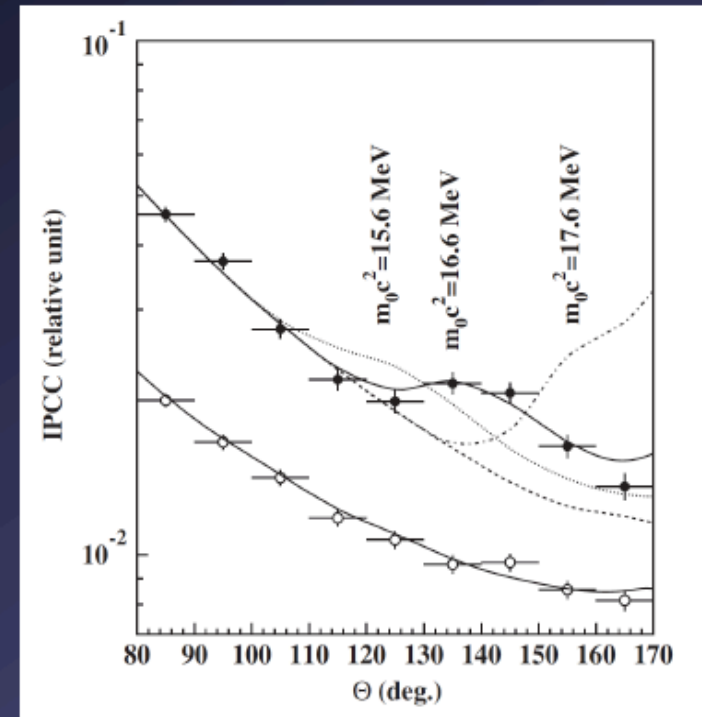
17 MeV Dark Photon ???



distribution of e^+e^-



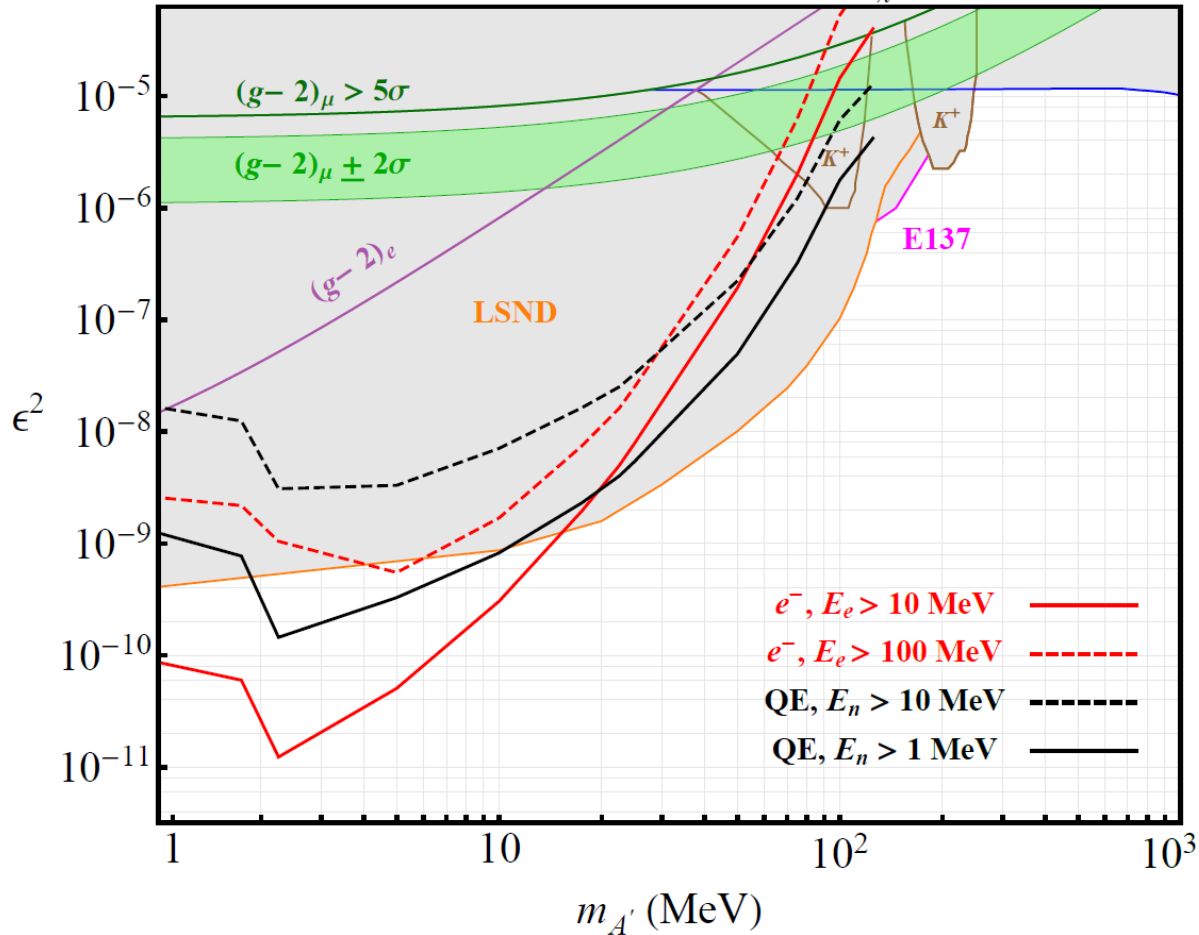
Angular correlation of e^+e^- pairs



Krasznahorkay et al., PRL 116 (2016) 042501

Testing competitive parameter range

BDX@MESA, 10^{22} EOT, $E_e = 150$ MeV, $d = 3$ m, $m_\chi = 1$ MeV, $\alpha_D = 0.1$



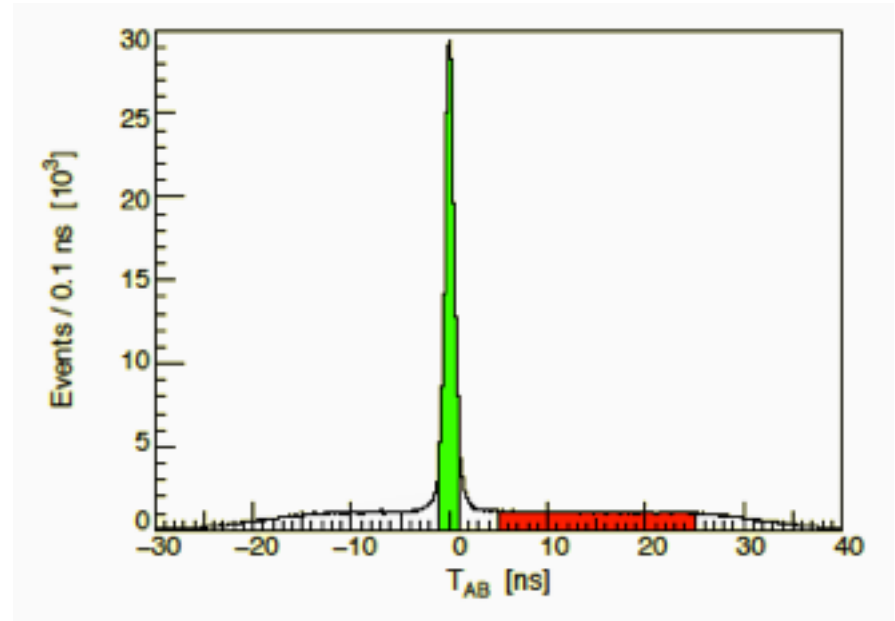
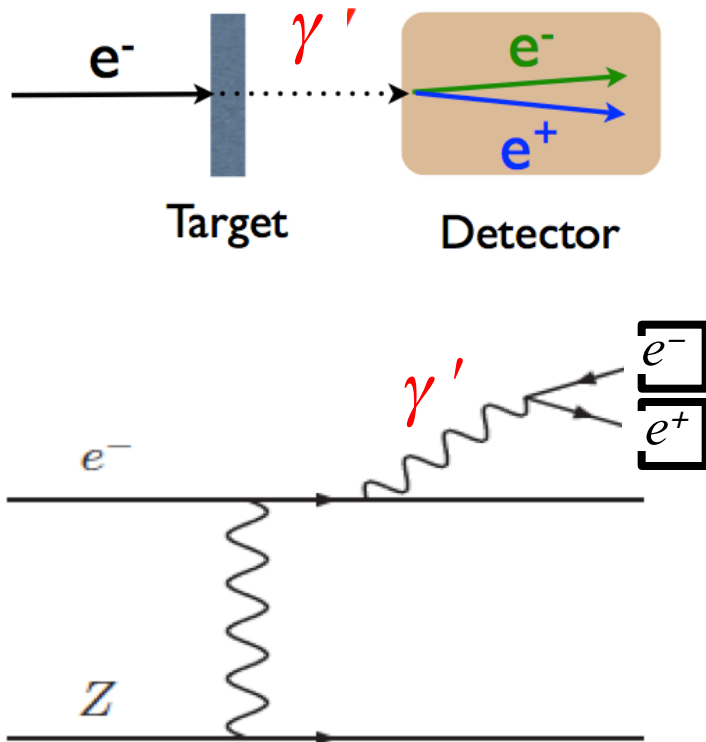
G. Krnjaic +
E. Izaguirre
Perimeter Inst.

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Low-Energy Electron Acceler. with high intensity suited for DP search

Bjorken, Esssig,
Schuster, Toro (2009)

Signal processes

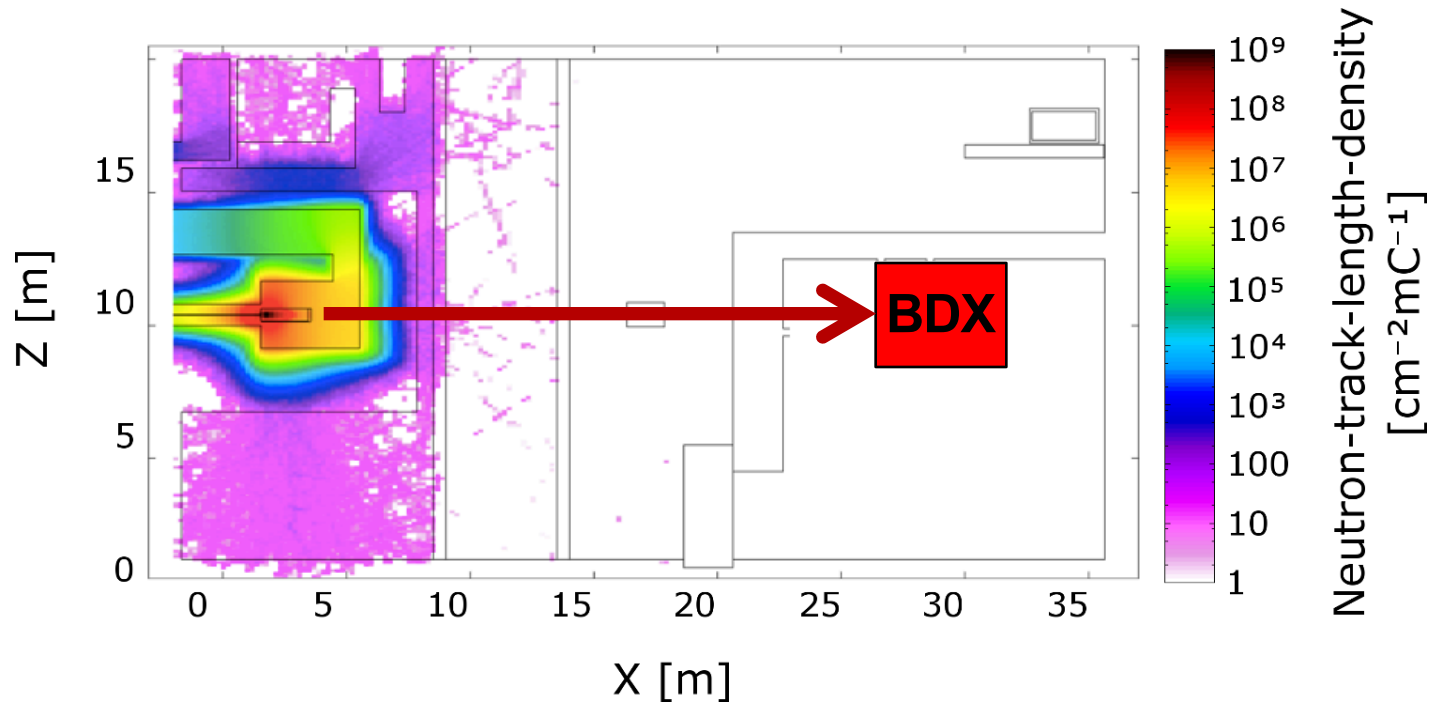


Time difference T_{AB} between both spectrometers

MESA Physics Program

- Parity Violation: $\sin^2\Theta_W$ Measurement
Neutron Skin of Nuclei
- Electromagnetic Form Factors of the
Nucleons (proton radius puzzle)
- Few Body Physics
- Nucl. Physics of astrophysical relevance
- Searches for Particles of the Dark Sector

©Klaus Hansen



Background situation

- FLUKA simulation of neutron background promising ($\sim 10^{11}$ EOT)
- MESA running below pion production threshold \rightarrow no neutrinos!

The Mainz Microtron MAMI

MAMI Hallmarks

- Electron and Photon Beam $E_{\max} = 1.6 \text{ GeV}$
- Intensity max. $140 \mu\text{A}$
- Resolution $\sigma_E < 0.100 \text{ MeV}$
- Polarization 85%
- Reliability: 7000 hours / year
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