Matter and the Universe

Cosmic Matter Laboratory FAIR, FZJ, GSI, HIM

Frank Maas – Helmholtz Institute Mainz



Cosmic Matter Laboratory Frank Maas

PAGE 1

Strong force

Bringing cosmic matter into the laboratory

Evolution of the Universe



Grand Challenges

- **Emergence** of matter from the strong interaction
- Role of fundamental **symmetries**
- Generation of the mass of hadrons
- Complexity of hadronic structure
- Properties of matter under extreme conditions

Grand Challenges in Strong QCD

- Quark-gluon dynamics and phases in very dense and/ or very hot nuclear matter ("hot QCD", heavy ion coll.)
- The dynamics, structure and stability of hadrons, the mechanism of hadronization, and strong CP-violation ("cold QCD", hadron spectrum, hadron structure)
- The generation of complex clusters of elementary matter and chemical elements and the limits of stability for exotic nuclei ("neutron rich", "proton rich", "superheavy elements")
- Test of fundamental symmetries, the symmetry between matter and antimatter (EDMs), beyond standard model searches

Impact of Strong Interaction Investigations



Facilities at FZJ and GSI (not complete)



Cosmic Matter Laboratory

Frank Maas

PAGE 7

Additional Infrastructures



Beijing Electron Spectrometer



Antiproton Decelerator CERN



UNILAC SIS18 ESR at GSI



Cooler Synchrotron COSY



Jülich Supercomputing Center



GSI Compute Cluster



HIM Computing



MAMI-C 1.5 GeV electrons



TRIGA Research reactor

Frank Maas



HADES At GSI



ALICE at LHC





Other aspects

- Interplay between theory and experiment
- PhD-programme: former HGS HIRE is continued
- HIC4FAIR, HIM, EMMI,
- Excellence Initiative: Several proposals for excellence clusters in preparation in Mainz, Darmstadt, Frankfurt

Time schedule during POF-IV

- Start of FAIR operation during POF-IV
- GSI research program in POF-IV using existing facilities and FAIR detectors: FAIR PHASE 0
- During Phase 0: Completion of construction of PHASE 1 experiments (Hades/CBM, NUSTAR, PANDA)
- Research using FAIR and FAIR detectors: FAIR PHASE 1
- During Phase 1: Preparation of PHASE 2 experiments









Simulation of nucleosynthesis processes in stellar objects

Almudena Arcones, ERC Starting Grant

Experimental Determination of nuclear properties for stellar nucleosynthesis using storage rings

Yuri Litvinov, ERC Consolidator Grant

Search for transient and oscillating signals from dark matter and dark energy with magnetic resonance and magnetometry (Dark-OST)

Dmitry Budker, ERC Advanced Grant

Search for Electric Dipole Moments using Storage Rings (srEDM)

Hans Ströher, ERC Advanced Grant

COSY – Precision Physics Era



srEDM (Electric Dipole Moments): achievements (JEDI collaboration)



Fig. 2. Deviation of the spin tune v_s , which is defined as the number of spin precessions per turn, as a function of the number of turns in the ring. At t = 38 s (about 28×10^6 turns), the interpolated spin tune amounts to 16097540628.3±9.7 × 10^{-11} , which represents the most precise measurement of this quantity ever performed. The previous best measurement, performed for the muon at the (g–2) experiment, had a precision of 3×10^{-8} per year. The higher precision achieved here is mainly attributed to the much longer measurement time of 100 s compared with 600 µs in the (g–2) experiment.



Fig. 3. One of the longest polarisation lifetimes recorded for the COSY ring. Measurements made at four separate times (to conserve beam) are matched to a depolarisation curve that assumes a Gaussian distribution of transverse oscillation amplitudes. The half-life of the polarisation is 1173 ± 172 s, which is three orders of magnitude longer than previous results using electron beams. δ shows the difference between the model and



First spectroscopic investigation of nobelium (Z=102)



GSİ

hature doi:10.1038/nature19345 Achievements:

- First ever successful laser spectroscopy beyond fermium
- Production rates: ~ 1 atom/s
- Overall efficiency up to 10%
- First ionization potential of nobelium precisely measured
- Nuclear spin and moments extracted for the isotope ²⁵³No

KU LEUVEN





Search for η' mesic nuclei by spectroscopy of ¹²C(p,d) reaction with FRS

d²0/(dΩdE) [µb/(sr MeV)]





◇ FRS used as high-resolution spectrometer
 ◇ extremely good statistics achieved
 ◇ stringent constraints on η'-nucleus potential



Measurement of Excitation Spectra in the ${}^{12}C(p, d)$ Reaction near the η' Emission Threshold

Y. K. Tanaka,^{1,*} K. Itahashi,^{2,†} H. Fujioka,^{3,‡} Y. Ayyad,⁴ J. Benlliure,⁵ K.-T. Brinkmann,⁶ S. Friedrich,⁶ H. Geissel,^{6,7}
J. Gellanki,⁸ C. Guo,⁹ E. Gutz,⁶ E. Haettner,⁷ M. N. Harakeh,⁸ R. S. Hayano,¹ Y. Higashi,¹⁰ S. Hirenzaki,¹⁰ C. Hornung,⁶
Y. Igarashi,¹¹ N. Ikeno,¹² M. Iwasaki,² D. Jido,¹³ N. Kalantar-Nayestanaki,⁸ R. Kanungo,¹⁴ R. Knöbel,^{6,7} N. Kurz,⁷
V. Metag,⁶ I. Mukha,⁷ T. Nagae,³ H. Nagahiro,¹⁰ M. Nanova,⁶ T. Nishi,² H. J. Ong,⁴ S. Pietri,⁷ A. Prochazka,⁷ C. Rappold,⁷
M. P. Reiter,⁷ J. L. Rodríguez-Sánchez,⁵ C. Scheidenberger,^{6,7} H. Simon,⁷ B. Sitar,¹⁵ P. Strmen,¹⁵ B. Sun,⁹ K. Suzuki,¹⁶
I. Szarka,¹⁵ M. Takechi,¹⁷ I. Tanihata,^{4,9} S. Terashima,⁹ Y. N. Watanabe,¹ H. Weick,⁷ E. Widmann,¹⁶ J. S. Winfield,⁷
X. Xu,^{6,7} H. Yamakami,³ and J. Zhao⁹

(η-PRiME/Super-FRS Collaboration)



Frank Maas Cosmic Matter Laboratory

HADES Virtual photon radiation off Au+Au collisions





- Inclusive excess mass spectrum
 - o all known sources subtracted
 - o fully corrected for acceptance
- Almost exponential spectrum up to vector meson region.
- Fit to dN / dM \propto M^{3/2} × e^{-M/T} \rightarrow T_{emitting Source} = 95 ± 5 MeV

- Observed radiation of virtual photons much more intense as expected from assuming incoherent superposition of photons emitted in NN collisions
 - \rightarrow Regeneration of baryonic resonances
 - \rightarrow Strong modification of spectral functions
- Excess yield driven by temperature and size/ lifetime (four-volume integral)





ALICE: LHC run 2, Pb-Pb at $\sqrt{s_{NN}} = 5.02 \text{ TeV}$



Characteristics of quark-gluon plasma produced at the highest energy ever.

Nuclear modification factor of inclusive charged particles.

Substantial reduction of systematic uncertainties wrt previous analyses!

Transverse momentum spectrum harder than at lower energy: similar

R_{AA} means stronger energy loss in denser medium.





HELMHOLTZ Institut Mainz

Highlights from Lattice QCD: (g-2)



- Pseudoscalar meson exchange expected to dominate hadronic light-by-light scattering contribution to the muon g-2
- Compute transition form factor between π^0 and two off-shell photons





 $\epsilon_{\mu\nu\alpha\beta} q_1^{\alpha} q_2^{\beta} \mathcal{F}_{\pi^0\gamma^*\gamma^*}(m_{\pi}^2; q_1^2, q_2^2) \equiv M_{\mu\nu}$

Fit form factor data to lowest meson dominance model:

 $(a_{\mu}^{\text{hlbl}})_{\pi^{0}} = \begin{cases} (68.2 \pm 7.4) \cdot 10^{-11} & (\text{LMD}) \\ (65.0 \pm 8.3) \cdot 10^{-11} & (\text{LMD+V}) \end{cases}$

- Pseudoscalar meson exchange contributes about 50% of the estimated total light-by-light scattering part
- Agrees well with phenomenological studies



Cosmic Matter Laboratory

Highlights from HIM BES-III Analysis: New Resonance discovered



Precise cross section measurement of $e^+e^- \rightarrow \pi^+\pi^- J/\psi$ from 3.77 to 4.60 GeV at BESIII



> The e⁺e⁻ $\rightarrow \pi^+\pi^-$ J/ ψ cross section was measured with unprecedented precision with BESIII data.

The mass & width precision of the Y(4260) resonance was improved by a factor of ~3; a new resonance near 4.32 GeV/c² was observed for the first time with 7.6o significance.

arXiv:1611.01317, BES-III collaboration, corresponding author: Z. Liu, HIM

12.12.16



HELMHOLTZ

Highlights from studies for PANDA



EPJ A Highlight COVER EPJA October issue (2016) "Feasibility studies of time-like proton electromagnetic form factors at PANDA at FAIR" Corresponding author:

D. Khaneft (PhD-Student HIM)

A. Dbeyssi (Helmholtz-Postdoc HIM)



Frank Maas



Time schedule during POF-IV

- Start of FAIR operation during POF-IV
- GSI research program in POF-IV using existing facilities and FAIR detectors: FAIR PHASE 0
- During Phase 0: Completion of construction of PHASE 1 experiments (Hades/CBM, NUSTAR, PANDA)
- Research using FAIR and FAIR detectors: FAIR PHASE 1
- During Phase 1: Preparation of PHASE 2 experiments